

From Predictive Calibration to Forward Analysis

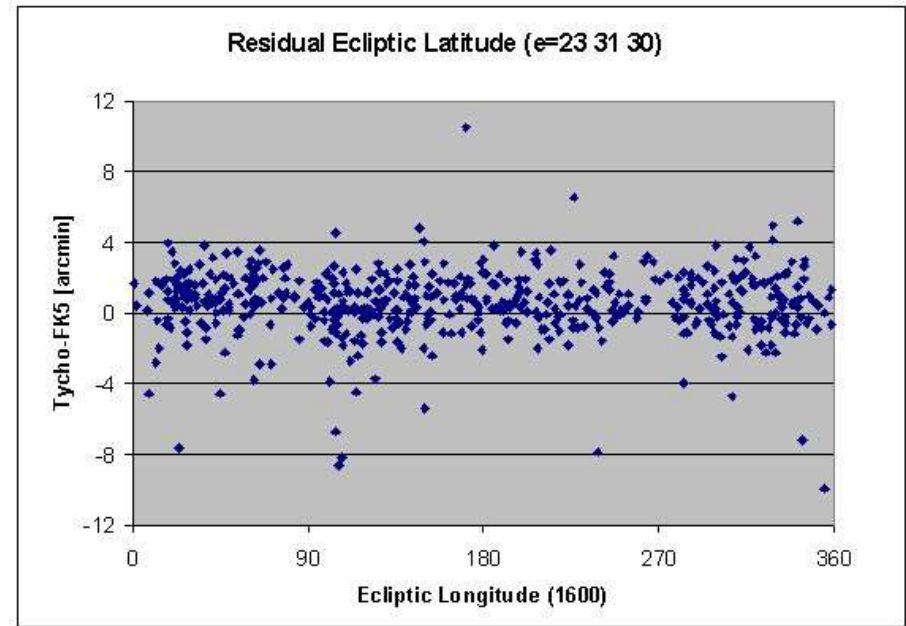
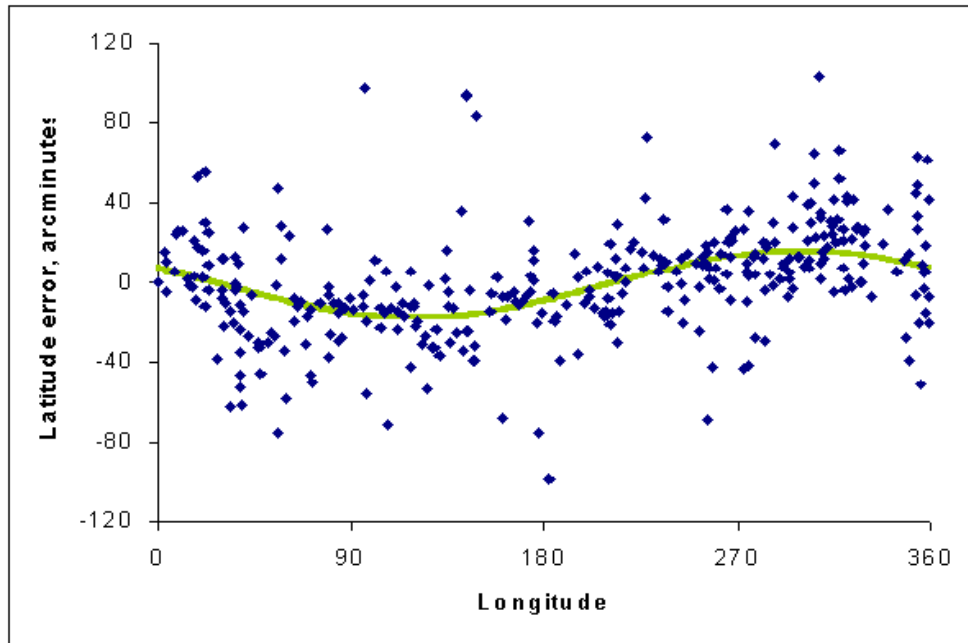
Preparing for the ELT era

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ESA/ESO

Hipparchos → Ptolemaios → Tycho



Tycho's Strategy

- **Be Better / Advanced in ALL areas of concern**
- **Suite of much better Instruments** VLT/ELT
- **Rigorous Calibration Plan**
 - Nightly cross-calibration / Frequent base checks CalPlan, Std. Progr.
- **Data Quality Check + Pipeline** DQ + DF
 - Usually at night spheric. trigon. results within 1 hr to check
 - “Pipeline” to solve ~ 50 000 spherical triangles
 - Logarithms not yet invented !! **Need sub - ” accuracy**

Agenda

- **Key Points**

- Keeping the pace between upgrades of Scientific Aims (Ambitions) , Instrumentation and Methods
- Consolidating ground conquered
- Preparing for greater challenges

- **Key Phrases Predictive Calibration & Forward Analysis**

- Predictive: utilizing physical (first) principles “a priory knowledge”
- Forward: do justice to the (precious raw) “observables” by enabling to map into and compare theoretical models of targets in the raw data domain

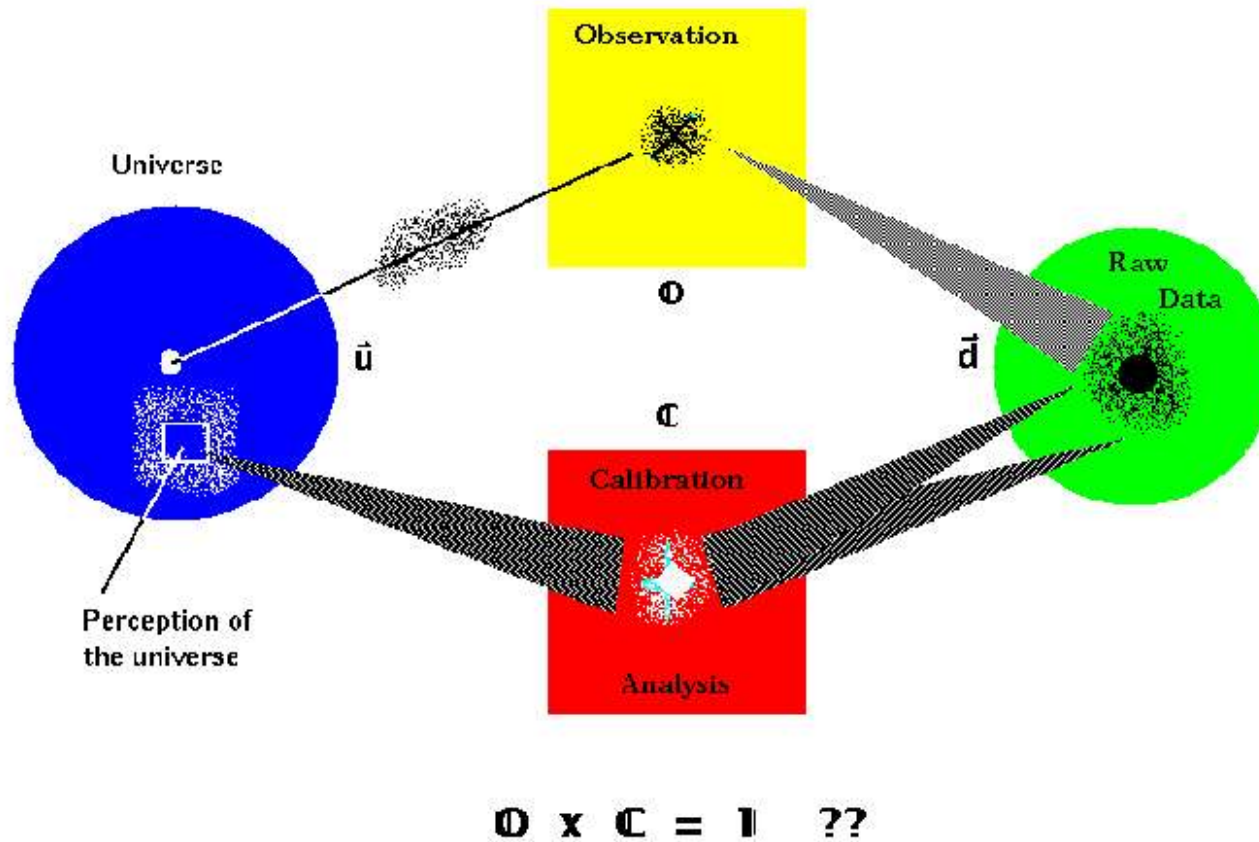
Heritage 1

- **1995 ST-ECF / ESO Calibration WS**
 - “Predictive Calibration based on Physical Instrument Models”
- **In parallel (→ 1998) ESO formulates**
 - VLT Operations Plan → Requirements for “Data Quality & Flow”
- **1997 - 1999 implementations of Physical Models**
 - CASPEC + UVES (Ballester & Rosa 1977 theory paper)
 - UVES W-Calib Bootstrap + more in ETCs (Ballester + team)
 - HST FOS (initially Rosa & Kerber)
- **1999 ESA → Instrument Physical Modeling Group**
 - thanks to former DG Riccardo Giacconi (see AnnRevAstrAstroph 2005)

Heritage 2

- **2000 - 2005 ST-ECF Team on FOS & STIS**
 - Alexov, Bristow, Fiorentino, Kerber, Rosa + contr. Modigliani (DMD)
 - FOS Post Operational Archive - based on FOS model
 - STIS Model + SimulatedAnnealing - demonstrated factor 10 +
 - Verified on superior entirely new Pt-Ne/Cr line catalogue (NIST collab.)
- **2005/6 CRIRES / Xshooter model**
 - Bristow + Kerber integrated into to ESO-INS
 - thanks to DG Catherine Cesarsky (bringing them back from ESA)
- **2007 → Physical Models established Instr. Support**
 - Spectrograph kernel + Simulated Annealing ready to ...
 - ... support many more spectrographic instruments

The Observational Information Loop

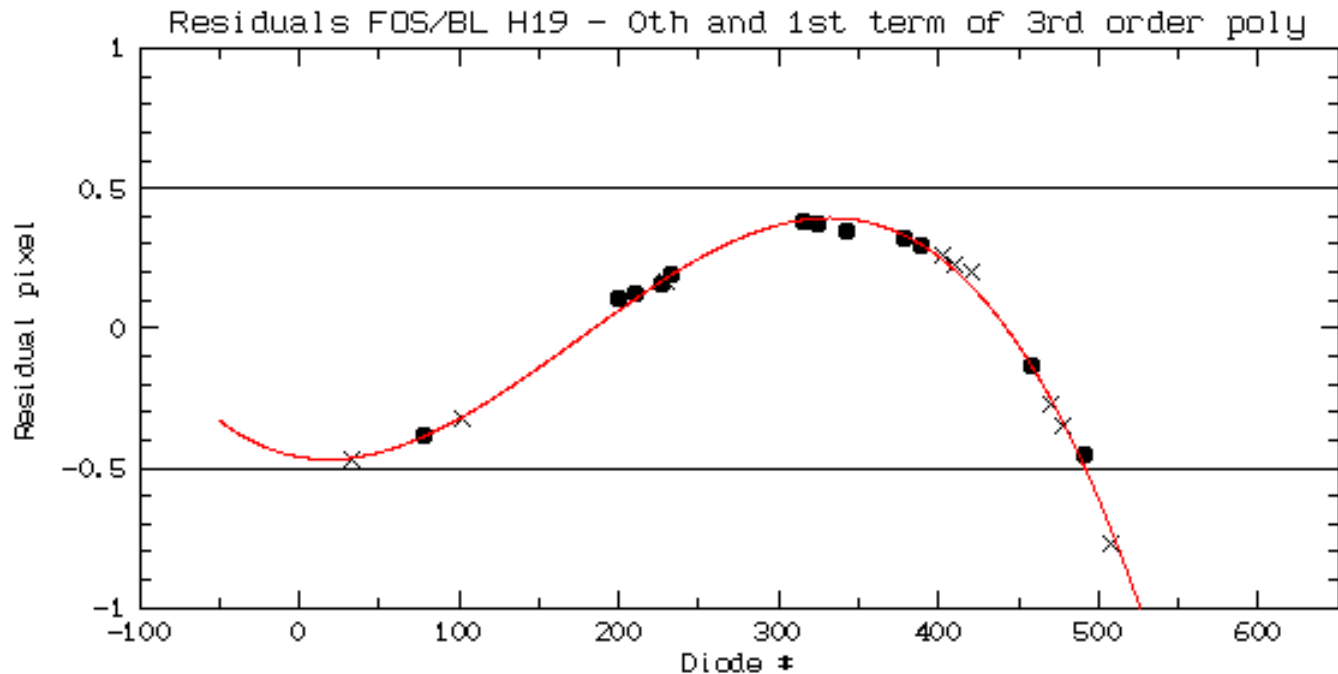


Demonstration Case

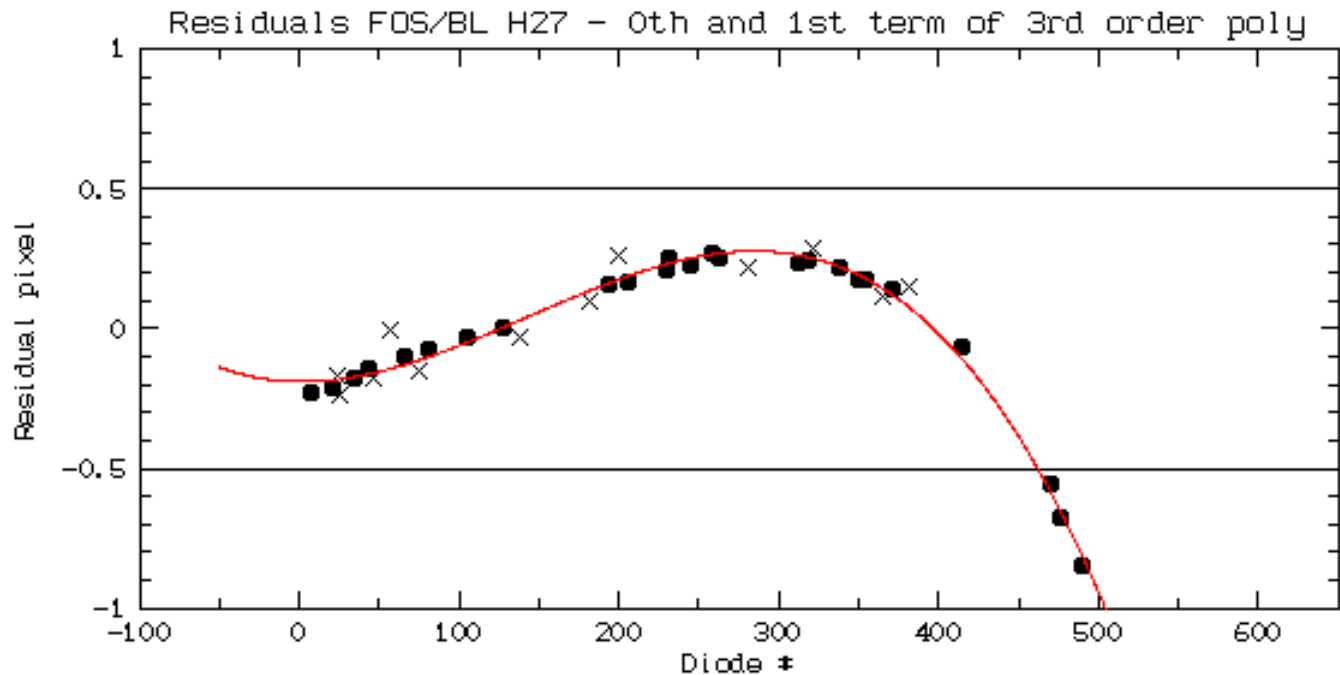
- **HST Faint Object Spectrograph (FOS)**
 - Relatively straight layout
 - Easy to grasp impact of “physical insight” on calibration
 - Obvious projection to FORSes ...

- **Case STIS (UVES, CRIRES, XShooter...)**
 - More complex (2D – echelles, multi-objects...)
 - But also “done” (in principle ...)

3rd order Poly - Dispersion Relation



3rd order Poly - Dispersion Relation



FOS Dispersion Model - Physical Principle

- **Relevant FOS Optical Layout in High Res Mode**

- wheel holds 5 different 1st order spherical gratings, 4 used per detector
- Imaged onto blue/red channel Digicon tubes

- **Physical Principles**

- ray optics equation from grating to photocathode
- $$z = (f - f_0) * [\alpha + \beta + \arcsin \{ s / m * \lambda - \sin (\alpha - \beta) \}] - z_0$$
- S-distortion in Digicon tube (off-axis aberration in E x B geometry)
- $$x = g * z + h * \tan (i + j * z) - x_0$$

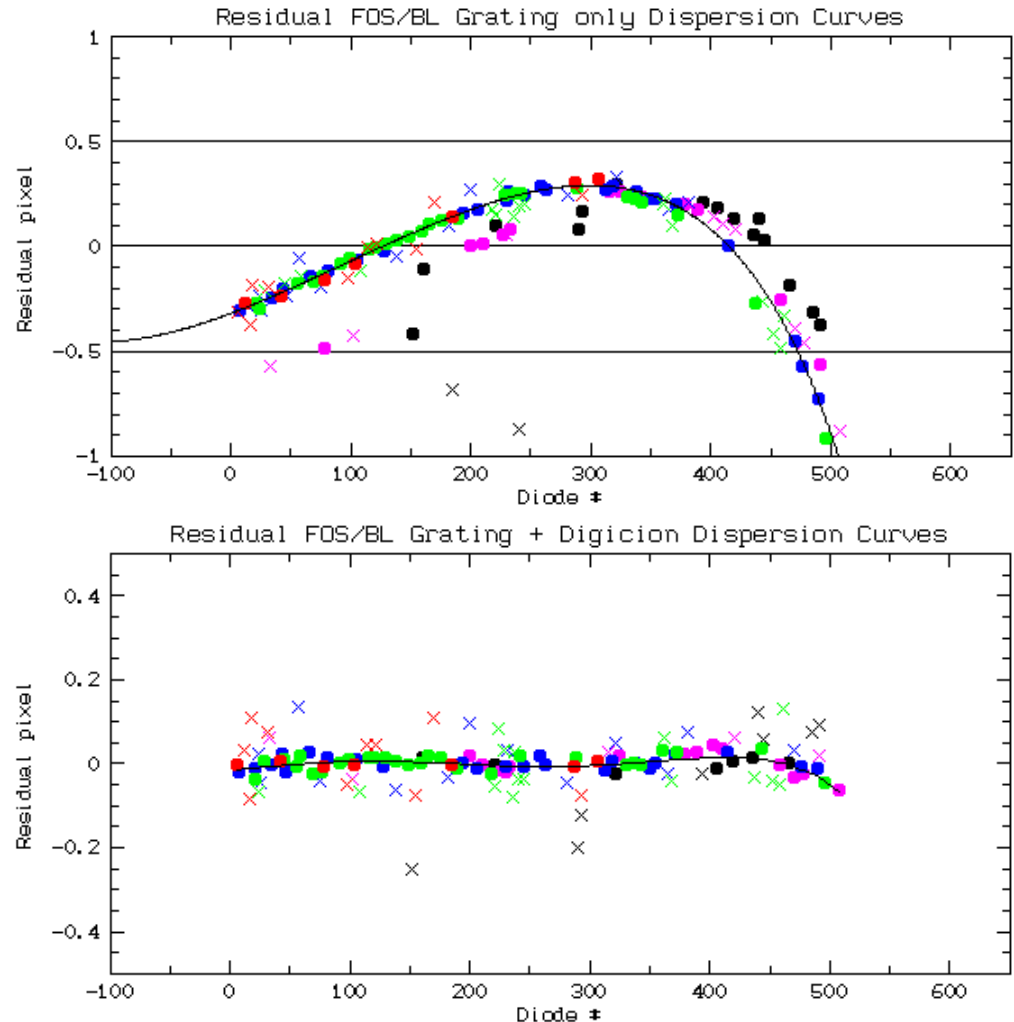
- **Restrictions on Parameters**

- common to all gratings on a given detector: $z_0, x_0, g, h, i, j, f_0$

- common to the red and blue channel per grating: $f, (\alpha - \beta), [s / m]^*$

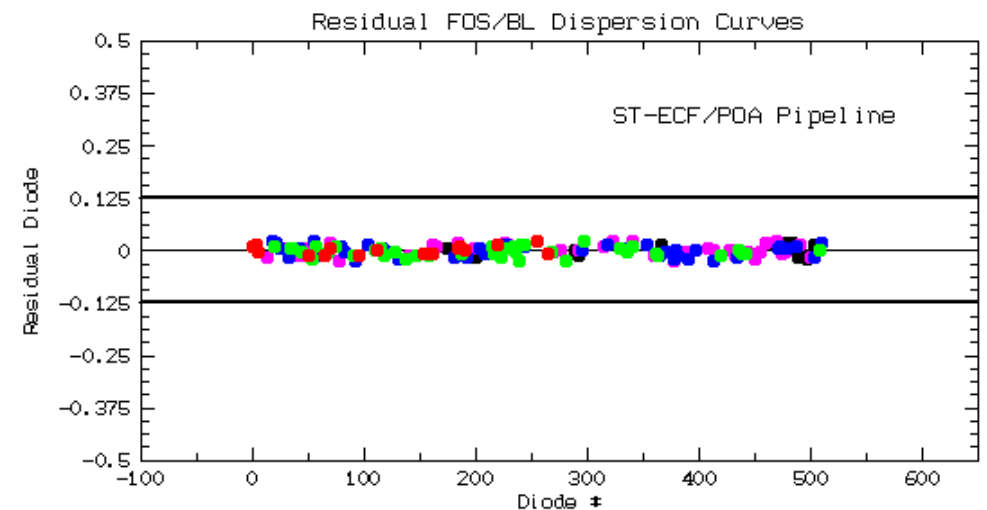
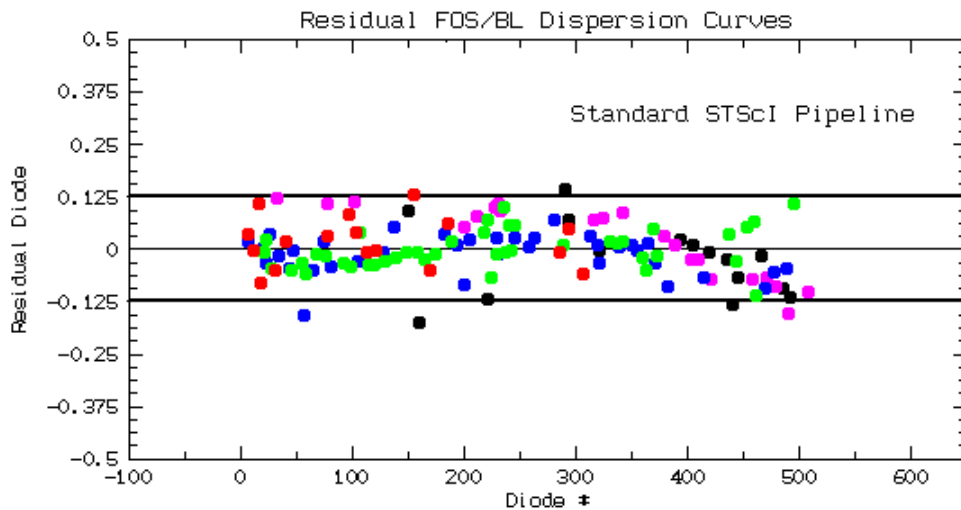
FOS Dispersion Model

- **Result (FOS BLUE)**
 - Assume that S-Distortion common to the 3 clean library list modes valid for NUV/FUV as well
 - Optimize common solution including the S-Distortion
 - Final residuals below 0.1 pix amplitude
 - Common pattern to all modes is pin-cushion distortion

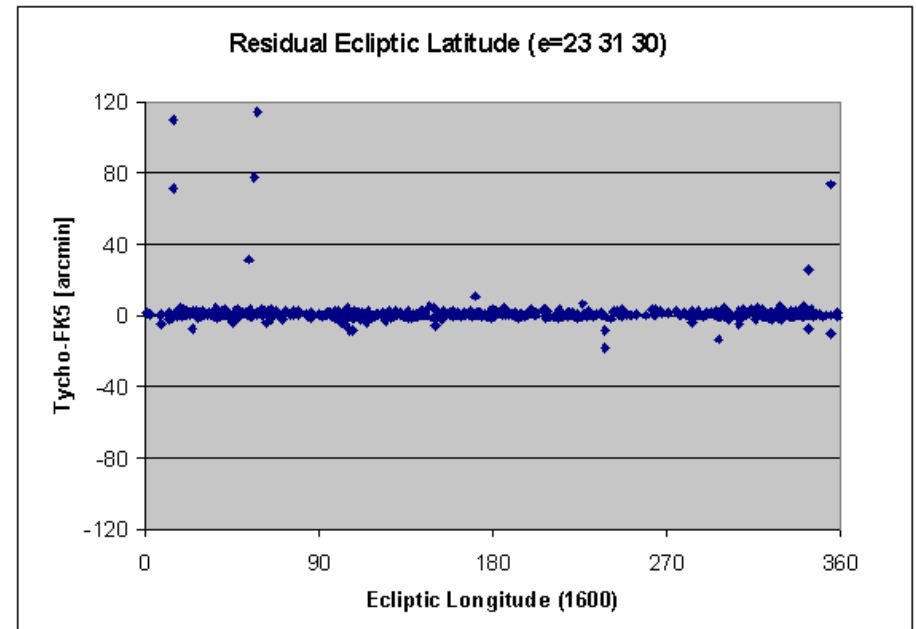
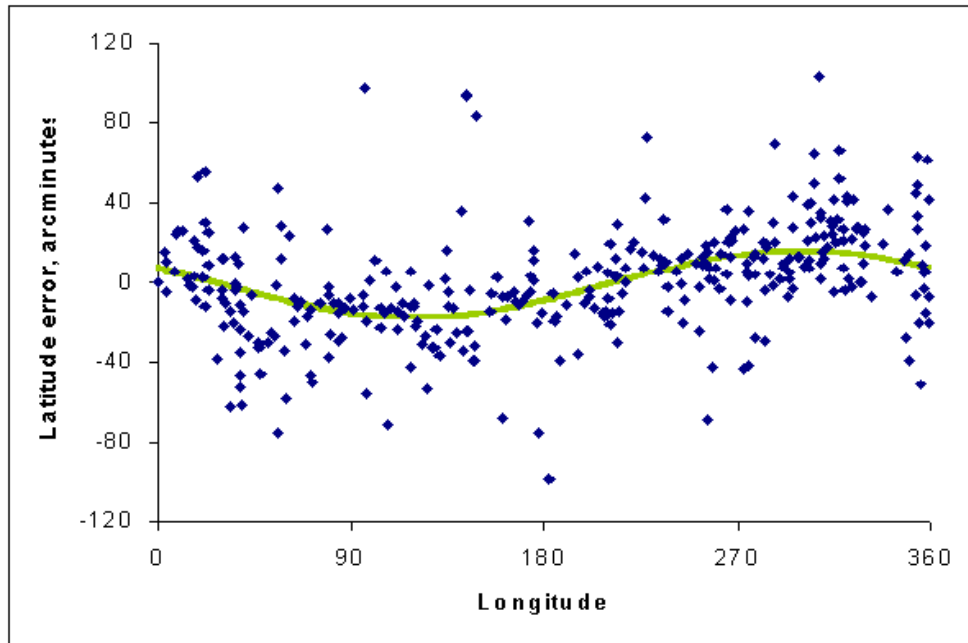


Dispersion Relations for FOS

- **Shown are residuals measured w.r.t. model solution**
- **FOS Dispersion Model valid for all gratings (differing colors)**
 - mode specific parameters: only grating constant, grating angle
- **Classical polynomial fit will fit all lines well**
 - whether or not they are blends, wrong identifications, too sparsely spread



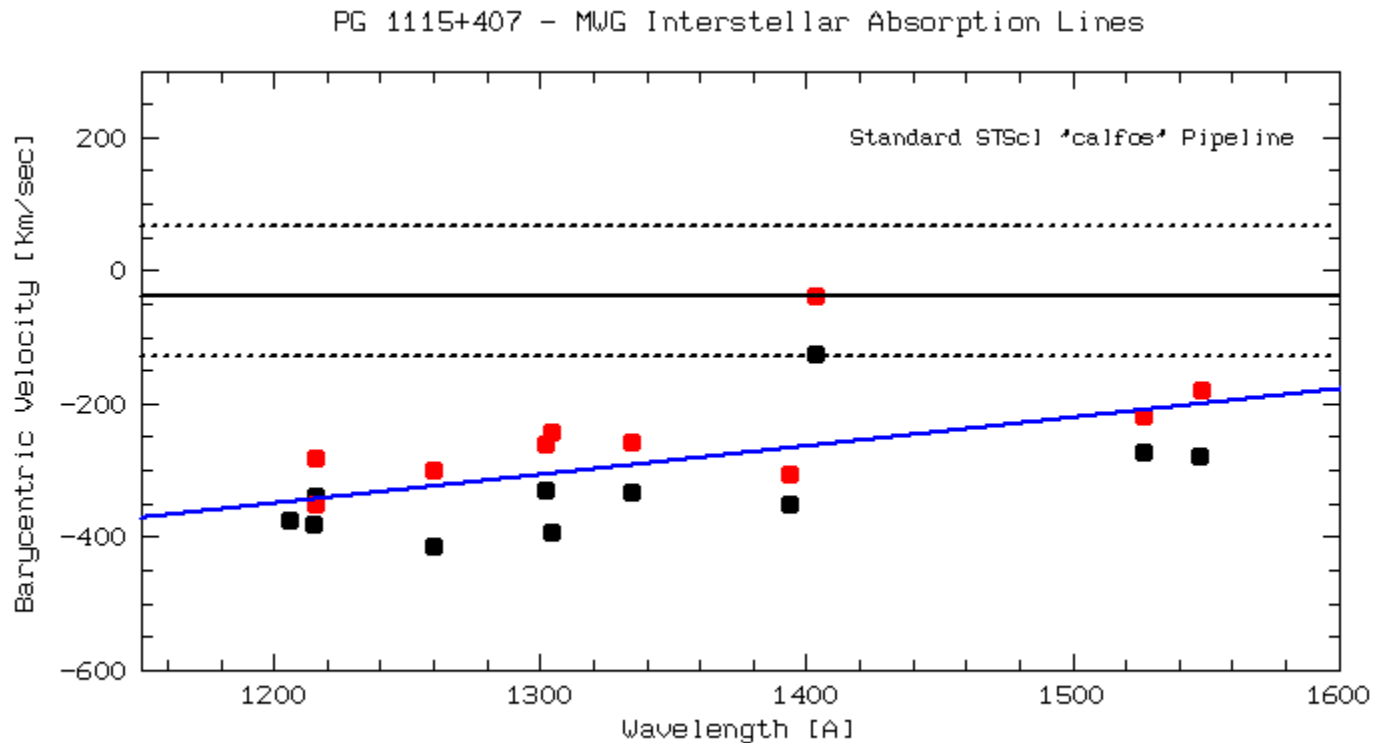
Hipparchos → Ptolemaios → Tycho



MW Halo absorptions in QSO spectrum

Standard “calfos”

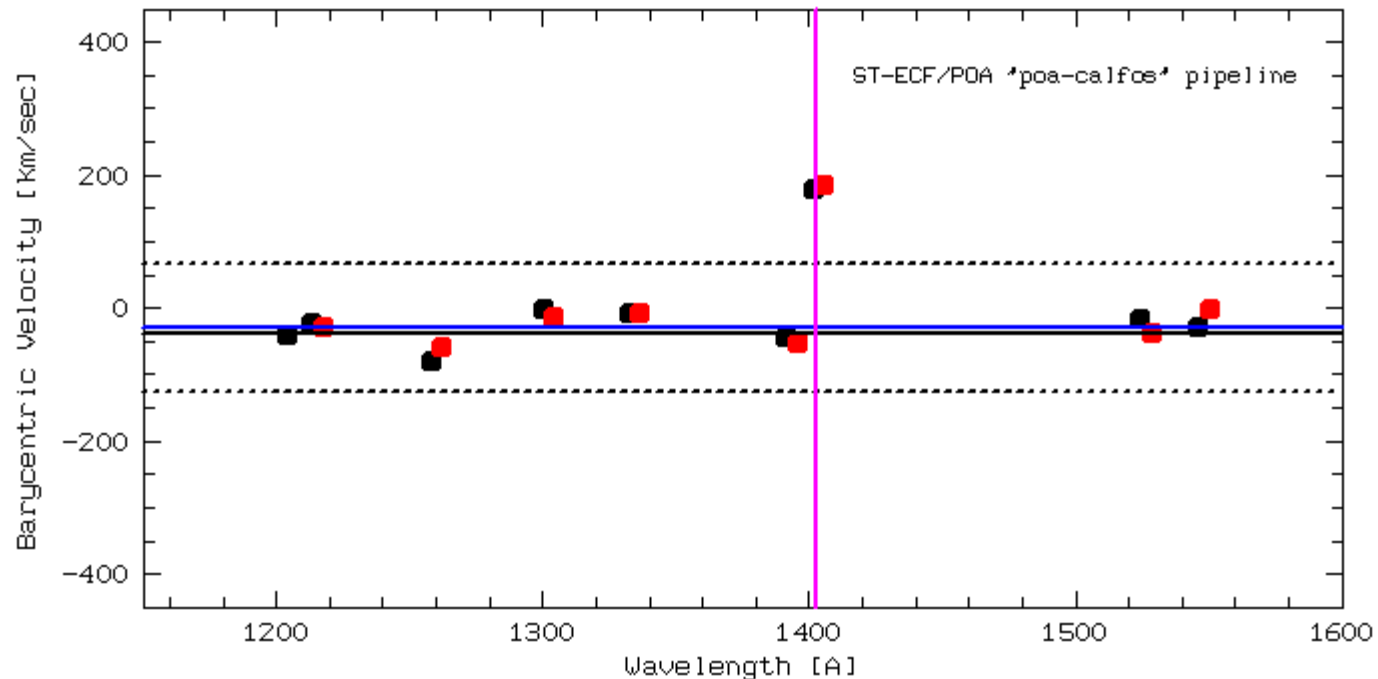
- 2 long exposures (blk,red) show repeat error
- Unphysical dependency of velocity on wavelength
- Only one absorption at 1403 Å seems to fit expectation



MW Halo absorptions in QSO spectrum improved with “poa_calfos”

- POA GIMP correction minimizes repeat error from 2 long exposures (blk, red)
- All but one line match central value of expectation (21 cm line)
- The one not fitting now is QSO Ly alpha redshifted (= NEW science)
- Result of POA GIMP correction in combination with physical model dispersion

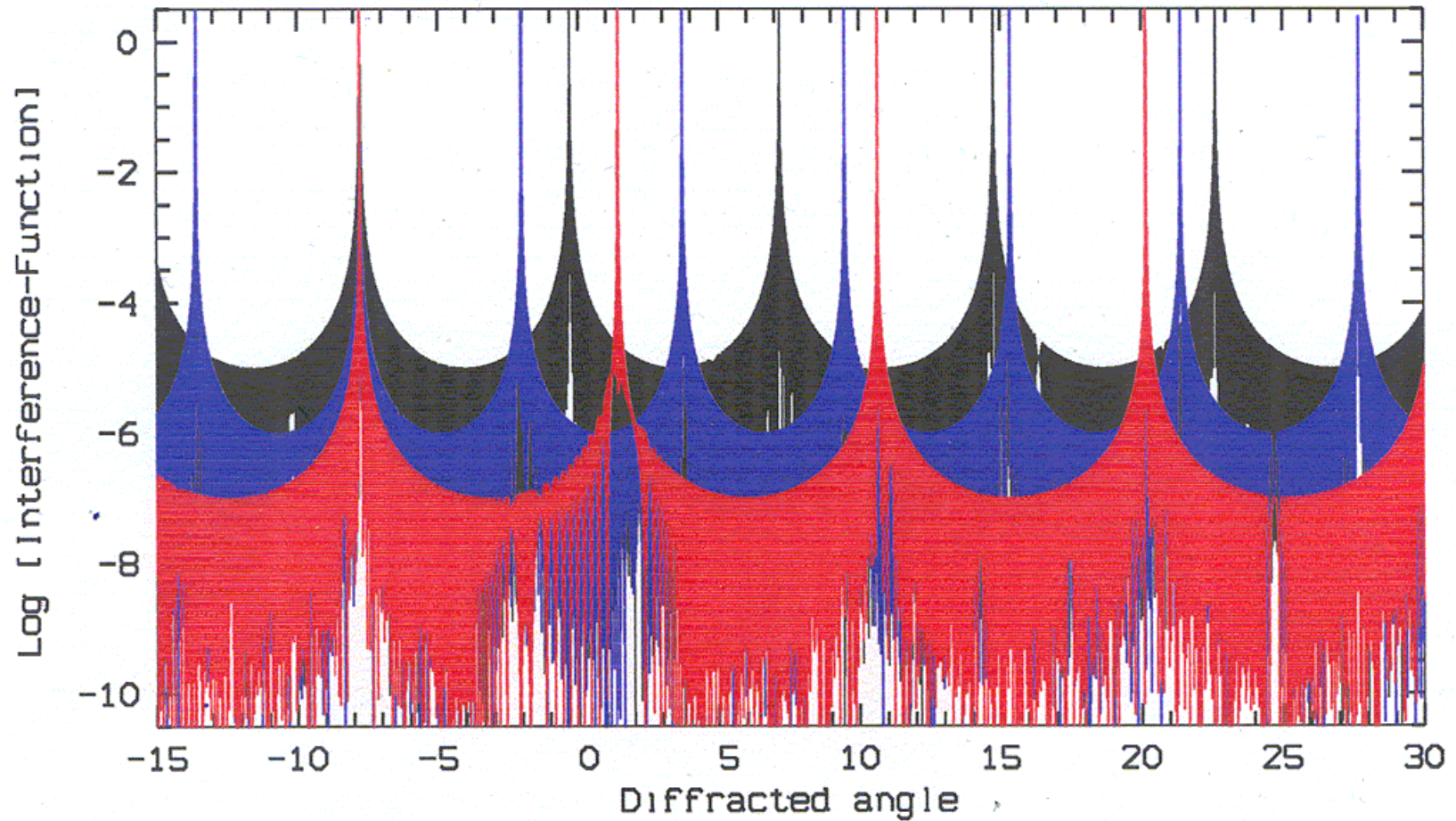
PG 1115+407 - MWG Interstellar Absorption Lines



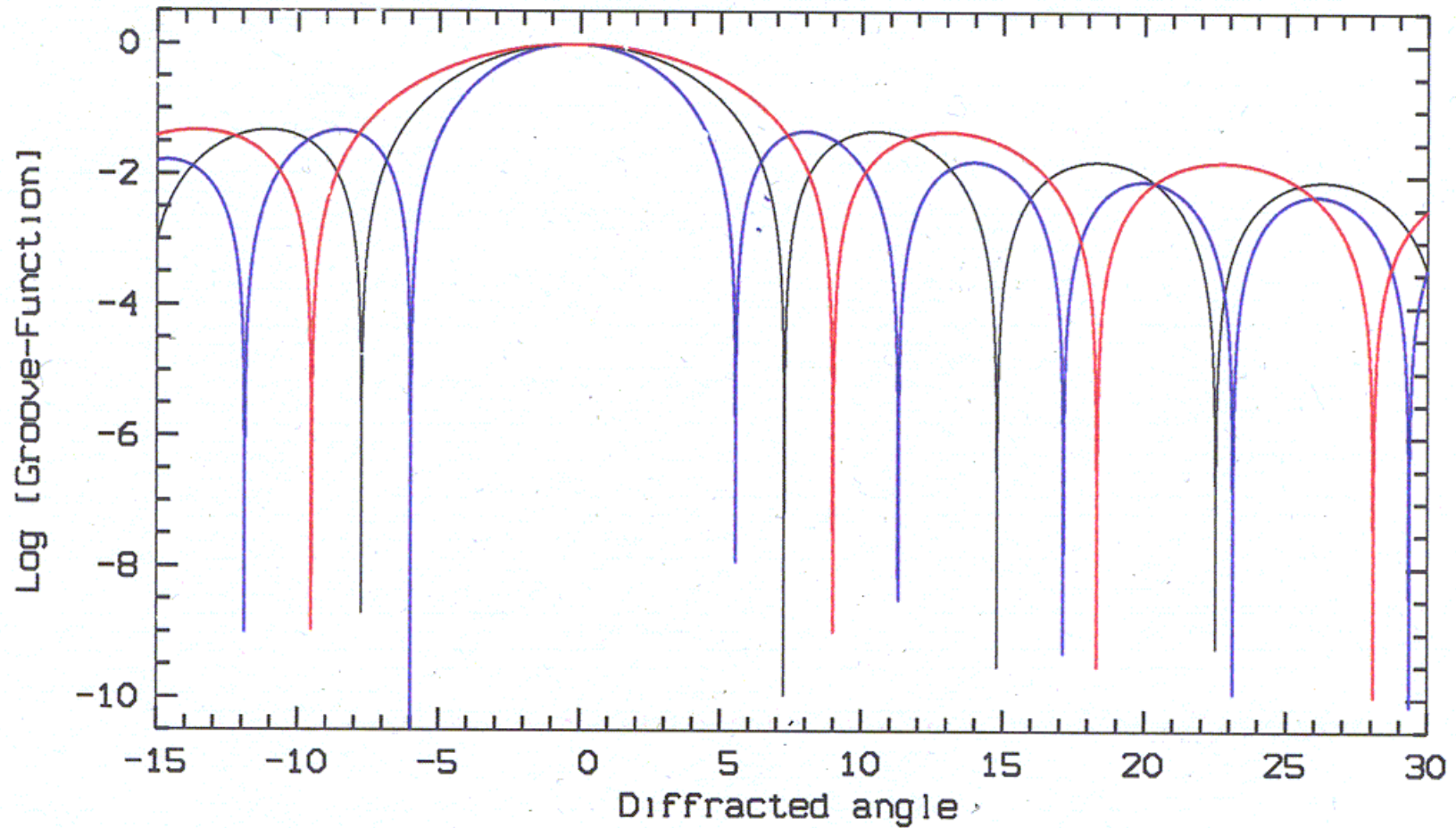
What about Flux Calibration

- **PM should only predict Blaze-Function**
 - Vignetting(s) will follow from optical path model
 - Mirror reflectivities etc. will enter as (accurate) laboratory measurables but are allowed to change as required by insight (measurement)
 - Combined model will be tuned so that StdStar comes out correctly
- **That is**
 - PM predicts the **SHAPE** of the flux calib curve
 - **SCALING** is the business of on-sky calibration (zero-points)

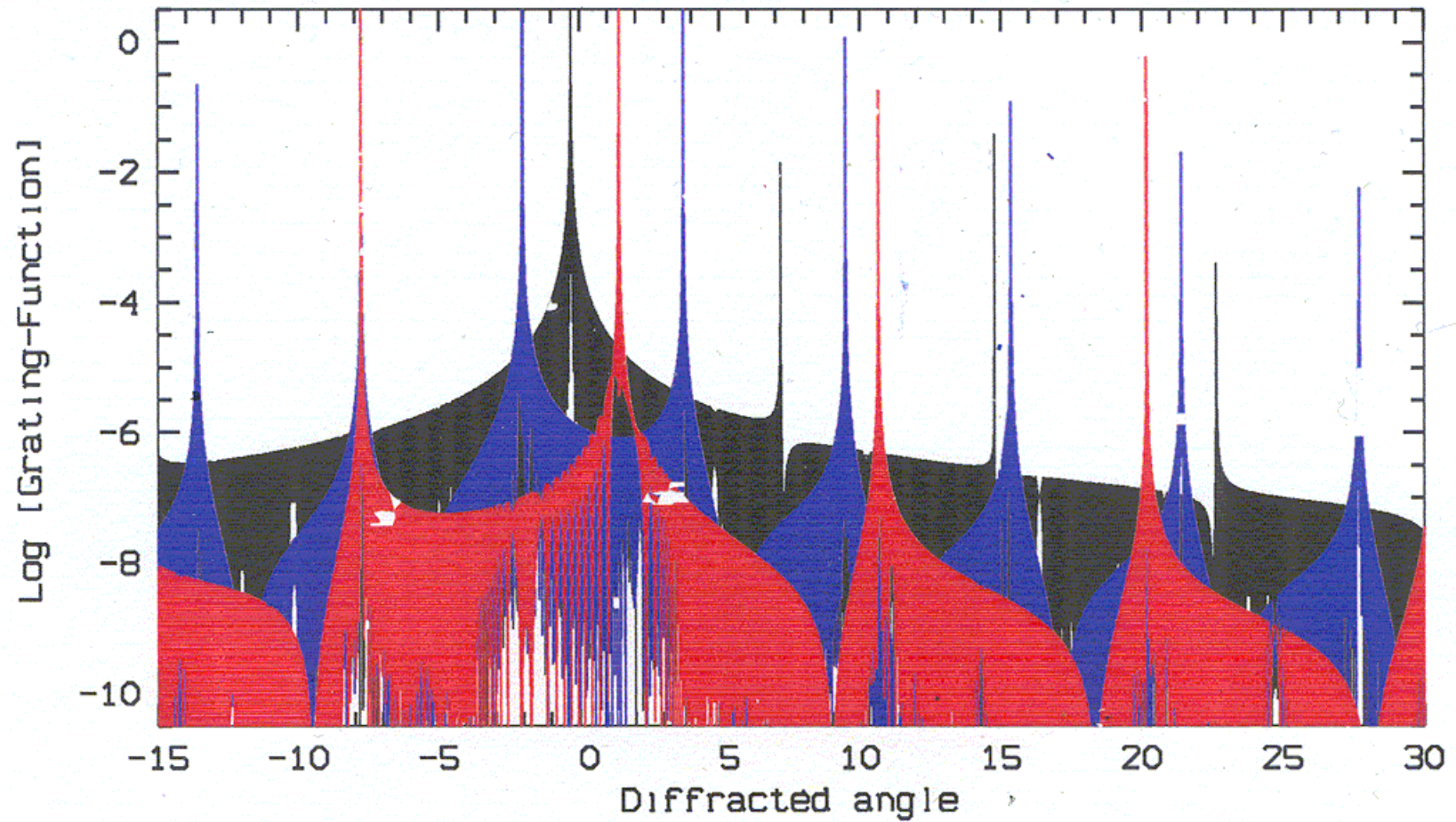
Grating - Interference



Grating - Diffraction

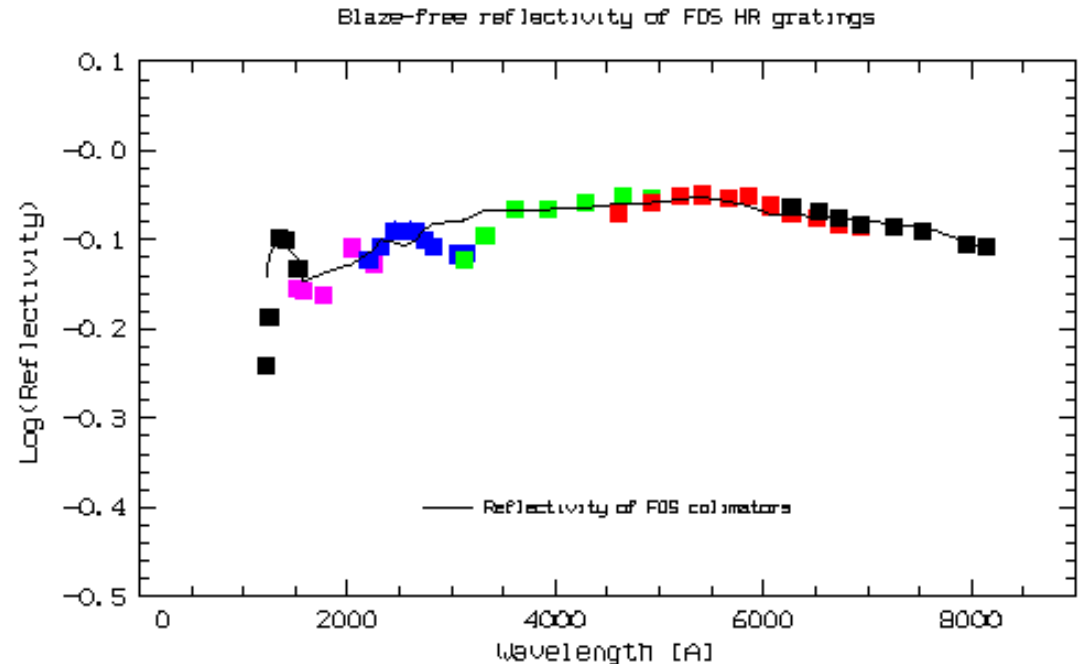


IF * DF

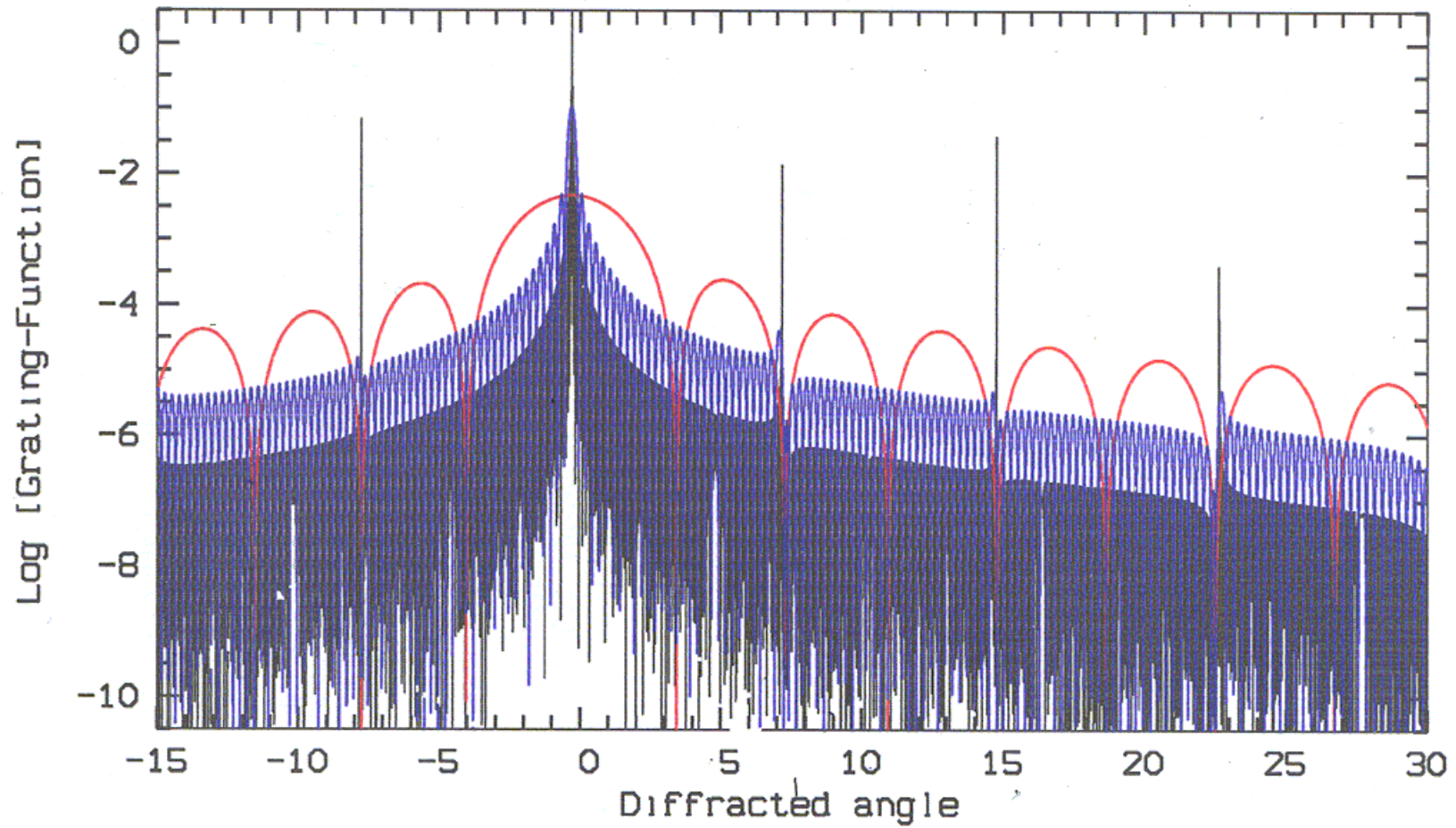


FOS Model - Prediction for Blaze

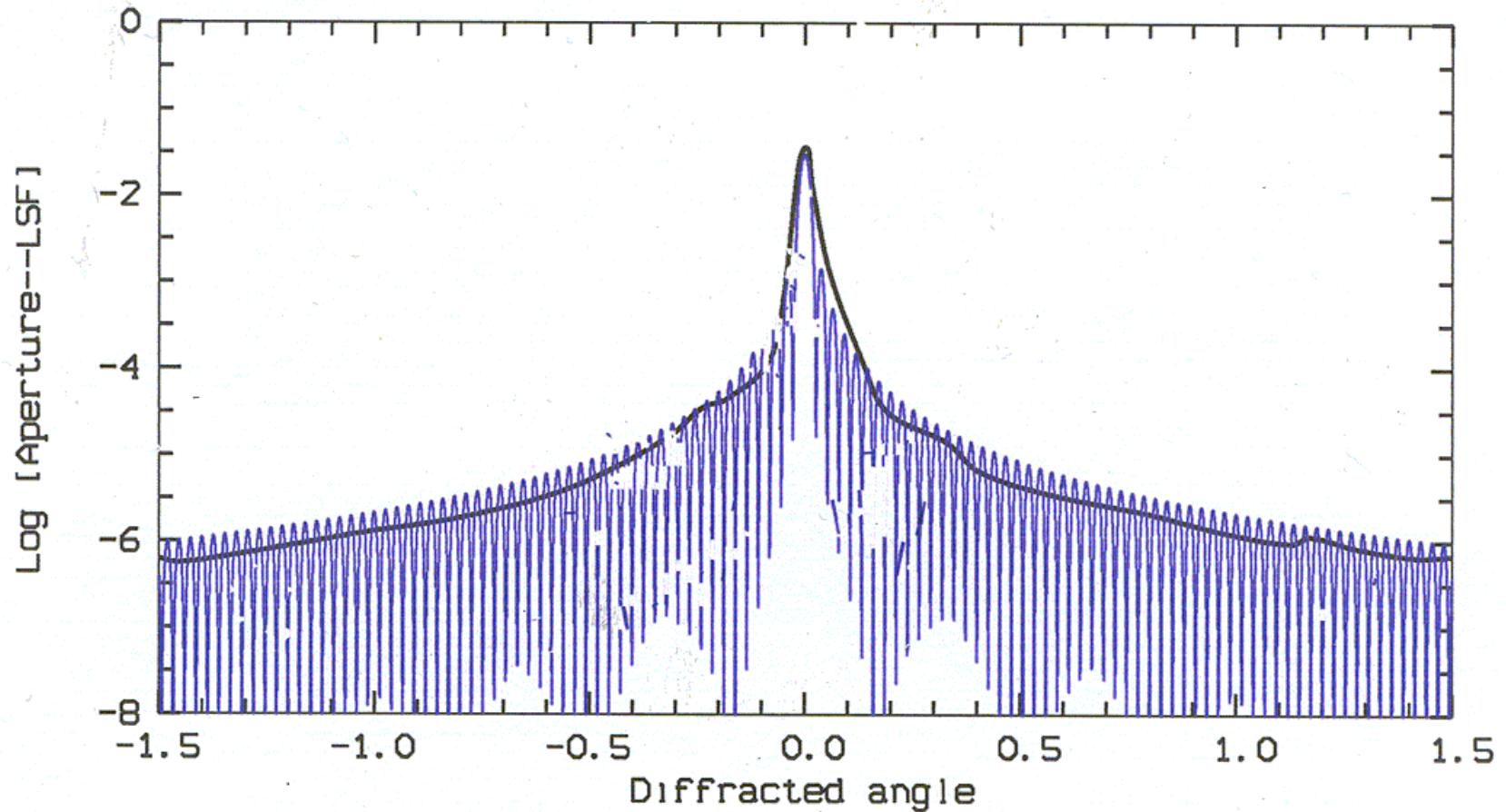
- **Predictive Power !!**
- **Blaze-Free Efficiency**
 - should be the coating
 - IF*DF @ peaks in 1st order = efficiency as $f(\lambda)$
 - predicted sensitivity (-HST)
 - compare to empirical flux cal
 - derive blaze free efficiency
- **Standard Stars wrong !!**
 - FUV/NUV/opt joints ~ few %



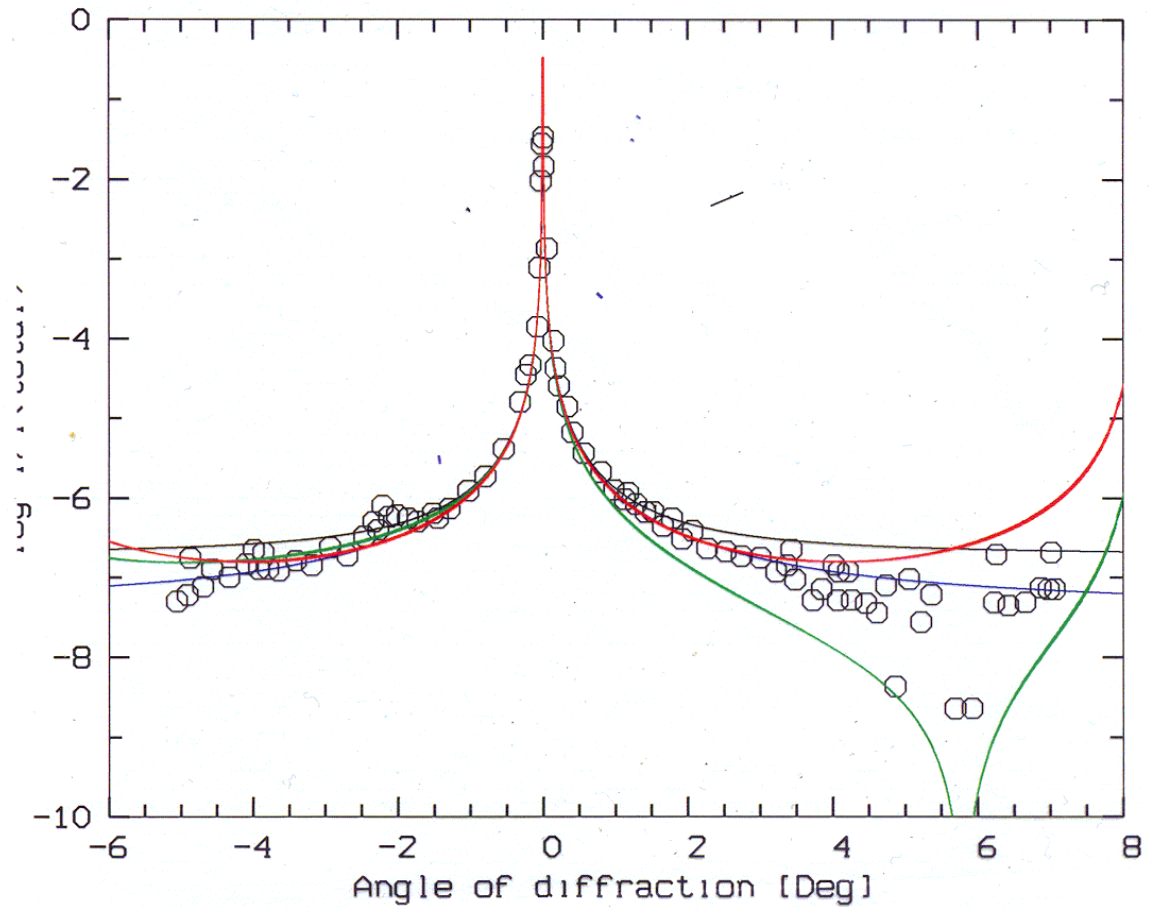
Line Profile - IF * DF (slit+collimator)



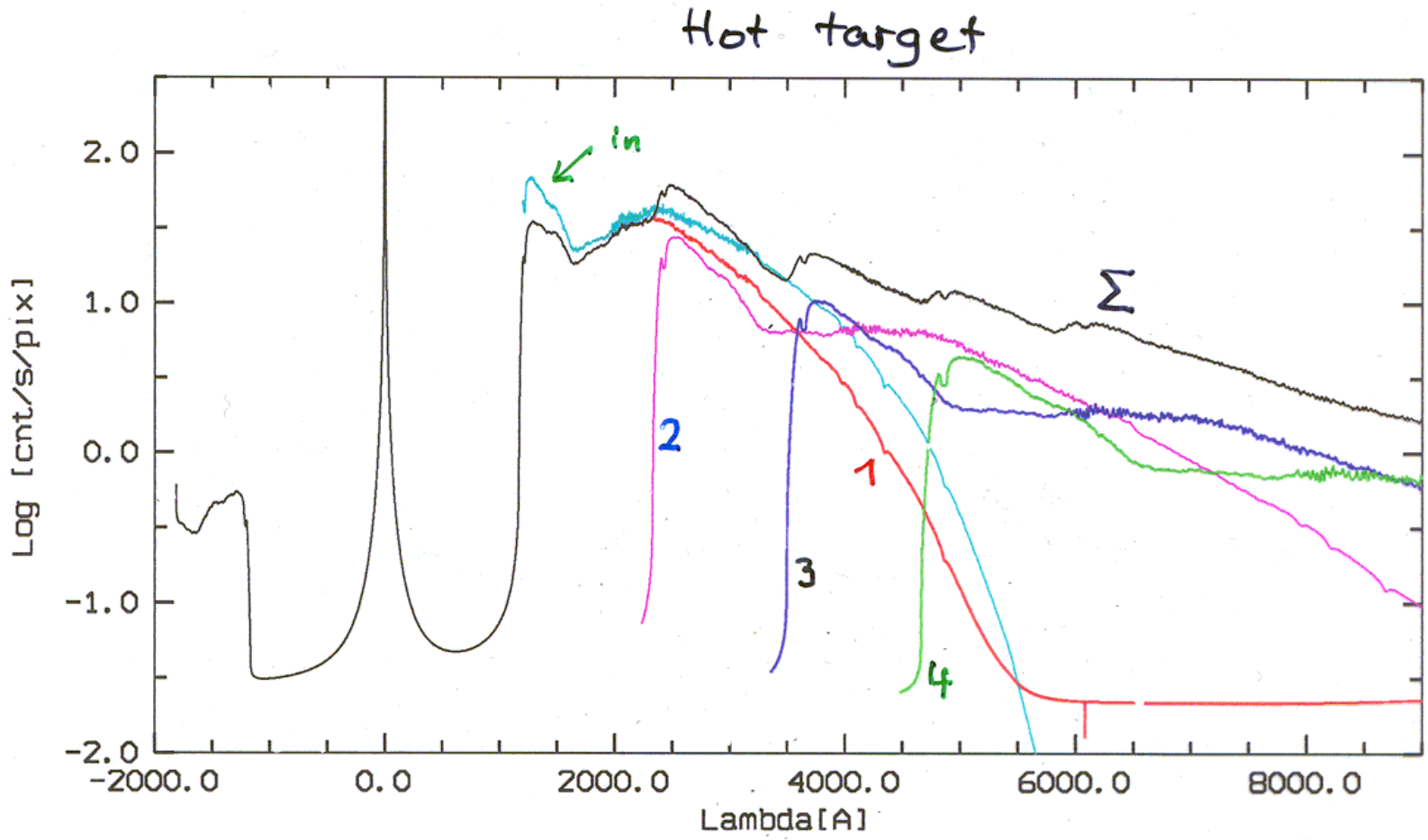
Line Profile - obs vs theory (-vign)



Line Profile - obs vs theory (2)



Full Throughput Model - Hot Target

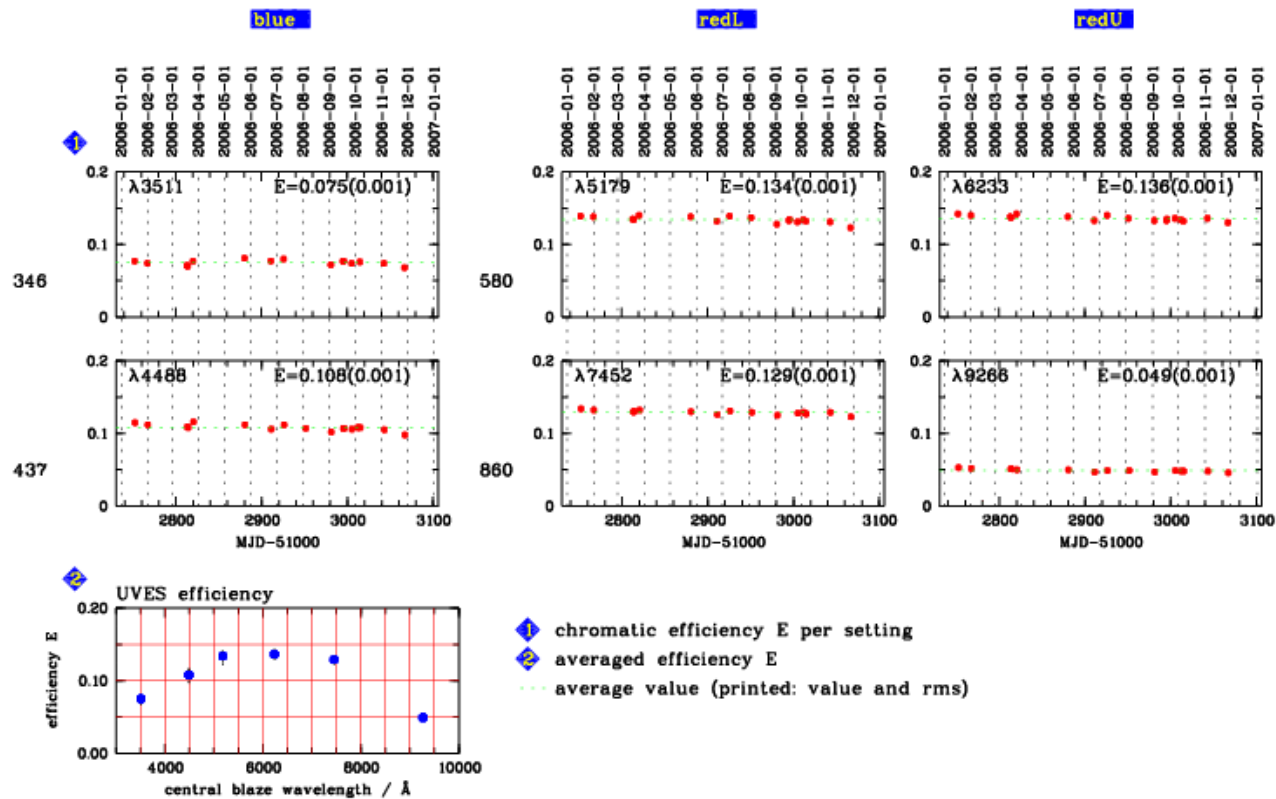


Roadmap

- **Very many “Observatory” pieces already in place**
 - Cal Plans , DQC and data base
 - Know how of “how to deal with data” (extraction, peculiarities)
 - Build up of extensive data base on detector performance
- **Know how, building blocks for “PhysModels” + Lab Standards**
 - many individuals already carry parts of the “company knowledge”
 - Parts of original IPMG integrated into ESO-INS (Bristow, Kerber)
 - PhysMod based calibration part of CRIRES and XShooter Projects
- **Required: Consolidation – Sustainability - Development**
 - Critical review of reference data and processes
 - Clearing station to achieve coherent view
 - Injection of advanced calibration concepts into instrument design

Stability of Instruments

KUEYEN/UVES trending: STD-NIGHT 2006
 instrument efficiency (from night-time STD stars); last date: 2006-11-27



Yes, agreed in principle, but ...

- **Non-stability of instruments on ground**
 - PM strengths → meaningful parameters → insight
 - and see DQC trending → eg. UVES, FORSes usually very stable
- **On ground we have an atmosphere**
 - PM strengths → decoupling of instr. / atmosph. effects
 - Atmosphere becomes a separate “controllable” item
- **Our instruments are too complicated**
 - we designed them, so we have the insight
 - more complex → more substantial insight helps

Atmosphere (terrestrial)

- **Key: Separate Instrumental Stuff from Atmosphere**
 - Required: PhysMod based Calibration
- **Atmosphere - becomes another “model item”**
 - At least from 320 to about 850 nm (the DoD knows that also beyond)
 - Extinction → 3-4 components → Form well known
 - Actual scaling should be controlled by LOSSAM / Std.Star expos
- **Sky Brightness - calls for a scaleable model as well**
 - F. Patat’s talk on Tuesday morning

Atmosphere (terrestrial) cont.

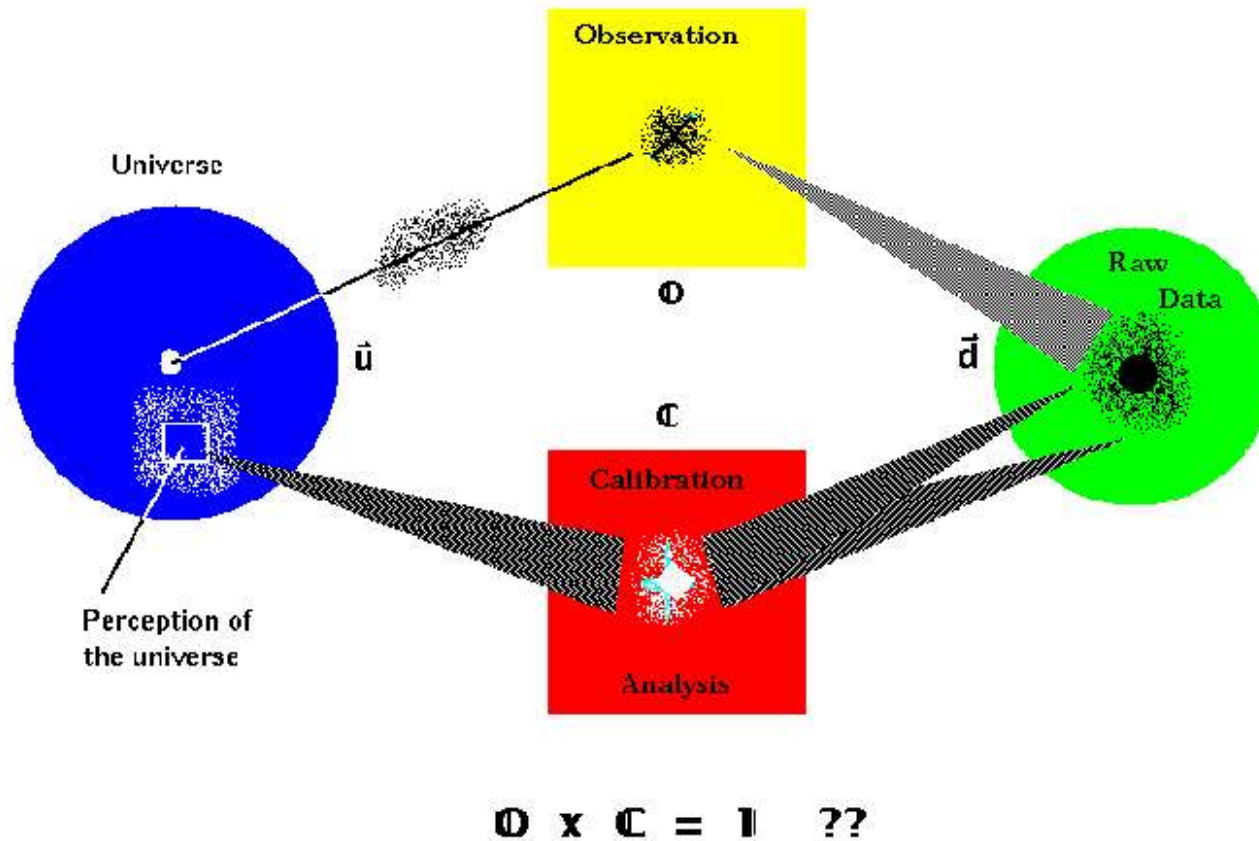
- **b.t.w. - I'm proud that**
 - everyone observing LS or PO is using “ATMOEXAN.TBL”
 - I constructed it in 1983
 - by matching a 3 component **Physical Model**
(Rayleigh, Ozone, Aerosols)
... to the sparse data points of Tueg (1977 = Messenger 11)
- **But**
 - **Anyone ever checked it for PO ?** – it was for LS altitude !!
 - It does not include dust (Vulcanoes, Copper-Mining ...)
 - Also, meanwhile we got “Globals” (Warming, Dimming etc ...)

-
- So much for those that still pretend that

**“Physical Models can not be of much use
at a ground based observatory ...
... because we have an Atmosphere.**

**Michael - you know, we prefer to use
good old ATMOEXAN instead ”**

Reminder: Observation Information Loop



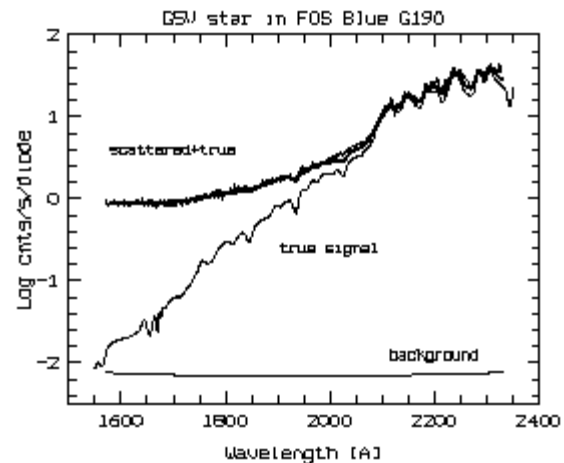
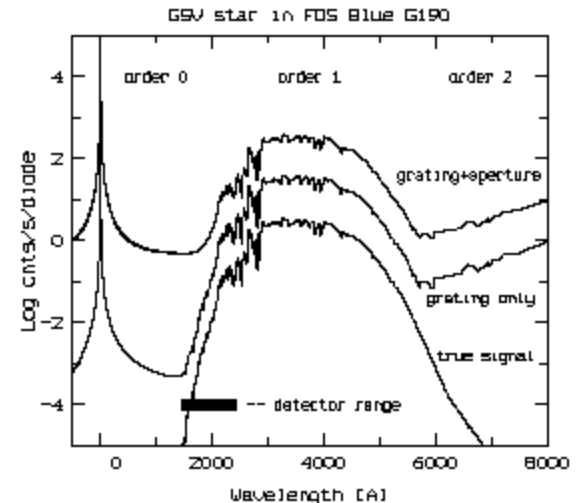
Definitions & Examples

- **First principle model** *Ray trace spectrograph model*
 - prescription based entirely on physical laws
 - very high predictive power
 - required to isolate effects while building physical models
- **Physical model** *UVES/STIS model (ray + dist.)*
 - prescription primarily based on physical / engineering insight
 - empirical “fudge” only as unavoidable (tolerable) substitute
 - sufficient predictive power for predictive calibration, forward analysis
- **Empirical model** *ETC, polyn. dispersion fits*
 - no physical insight required / inserted , can not be inverted
 - no predictive power outside data range / when params change

FOS Scatter Model

Test on Data

- **Recall**
 - wanted to predict the observed raw data for a cool target at UV wavelengths
 - Test for a Solar Analog
 - pass Kurucz model of Sun through the FOS model
 - compare with observations of solar analog 16 Cyg B
 - @ 160 nm signal is 1% of scattered (red) light
 - still the prediction agrees to better 5 % with actual data



Concepts to Get Around the Info Loop

- **Canonical Concept** *Empirical backward analysis*
 - empirical calibration relations → re-scale raw data → interpretation
- **Advanced Concept** *Predictive Calibration*
 - instrument models → noise free calibration relations
 - first principles → predictive capabilities outside “standards” range
 - Analysis: like empirical concept → “backward” (scaled raw data)
- **Superior Concept** *Forward Analysis*
 - can simulate raw data with sufficient accuracy and detail
 - evaluate theoretical target models in raw data domain
 - obtain likelihood estimates for range of potential target properties

Calibration in Context

- **Determine relation between output and the value of the input quantity, a reference standard (ISO 9000)**
- **Traceability** – establish accuracy by an unbroken line to higher standards. For each step evaluate uncertainty.
- **Quality Control** – monitoring, stability
- **Data Reduction – or better “Preparation”**
 - removal of instrumental signatures, extraction, “resampling”
- **Why do Calib and Reduc appear to be so intermingled ?**

FOS Calibration Issues - Physical View

- Geometric and Physical (wave) Optics
 - dispersion relations bad Pt-Ne/Cr line catalogue and polynomial fits
 - reflectivity, sensitivity, LSF, grating scatter, blaze functions
- Electron optics (**S-Distortion**, **GIMP**, **YBase trim**)
 - x (dispersion direction) --> **lambda zero points**, flat field shifting
 - y (spatial direction) --> **vignetting** (=absolute flux scale), **color terms**
- Thermal, mechanical, electronic, environmental items
 - **wavelength scale zero points** (bending), **flux scale** (pointing, vignetting),
 - **dark level** (solar cycles, particles in geo B-field, un baffled stray light)