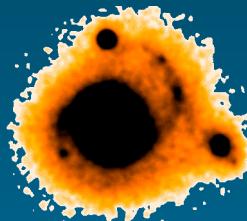


Case Studies of Quasar Host Galaxies



Julia Scharwächter

- Ph.D. at University of Cologne (Germany)
- Fellow in Chile since May 2005 (La Silla, NTT)

Guideline

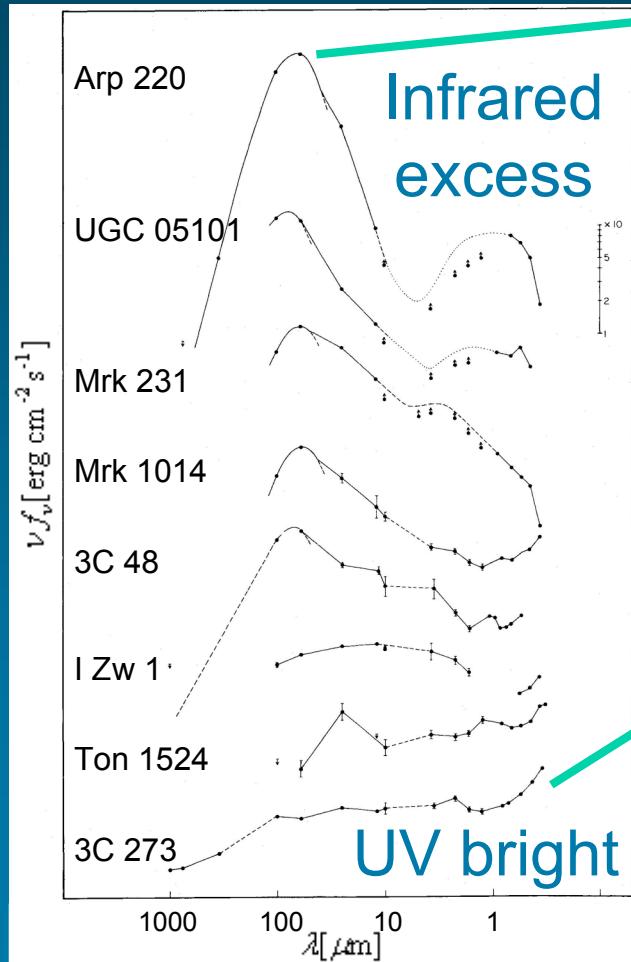
- Motivation
 - Hypothesis of a ULIRG-to-QSO evolution
- Part 1:
 - Multi-wavelength study of I Zw 1
[A. Eckart, E. Schinnerer, J. G. Staguhn, I. Saviane, J. Zuther, S. Pfalzner]
- Part 2:
 - First glance at SDSS J114203.40+005135.8
[V. D. Ivanov, L. E. Tacconi-Garman, J. Kotilainen, J. Reunanen, J. Zuther, A. Eckart, R. Schödel]

Evolutionary Hypothesis

Motivation

- Spectral energy distributions

[Sanders et al. 1988]



Ultra-luminous infrared galaxies
(ULIRGs)
Starburst and obscured AGN

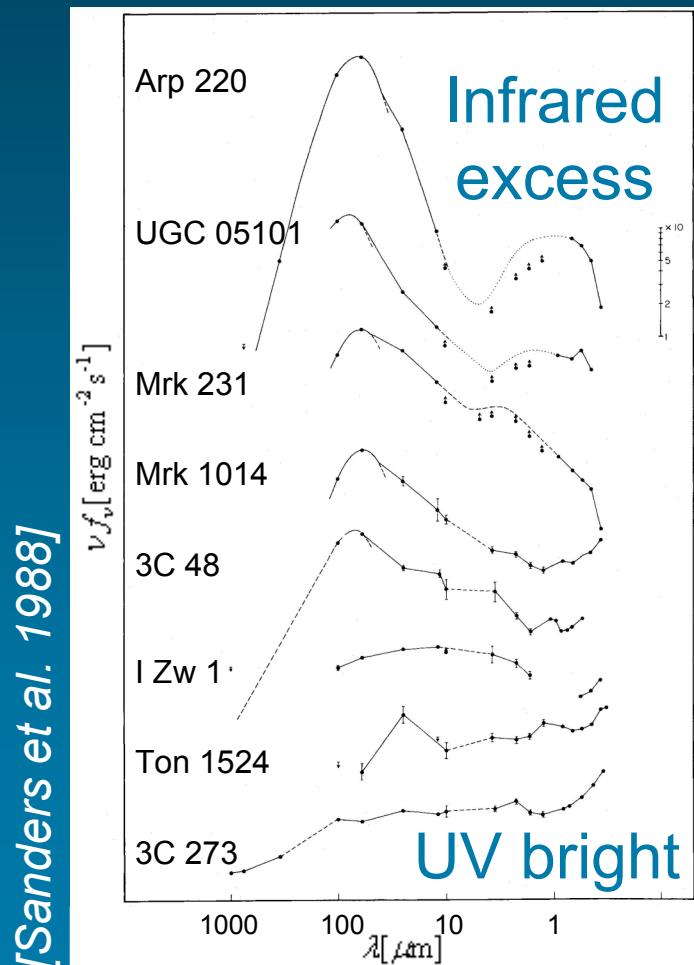
Smooth transition reflects
evolutionary sequence

Quasi-stellar objects
(QSOs)
Clear line-of-sight towards AGN

Physical Scenario

Motivation

- Sanders et al. 1988:



Galaxy merger causes
gas inflow



Starburst activity and
dust-enshrouded AGN

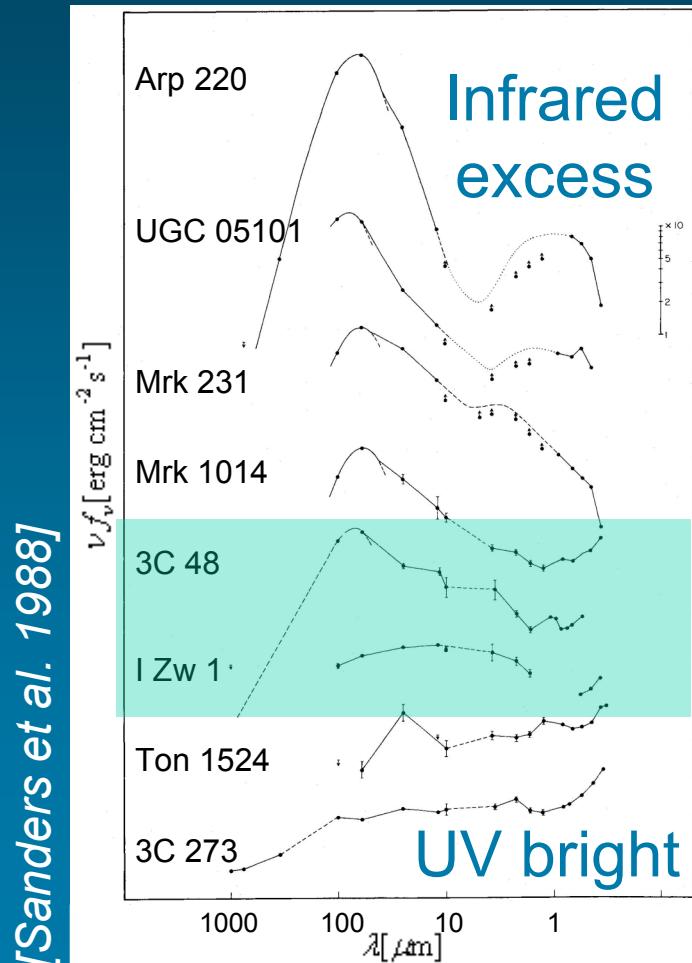


AGN blown free of
dust envelope

ULIRG to QSO Evolution?

Motivation

- Transition objects

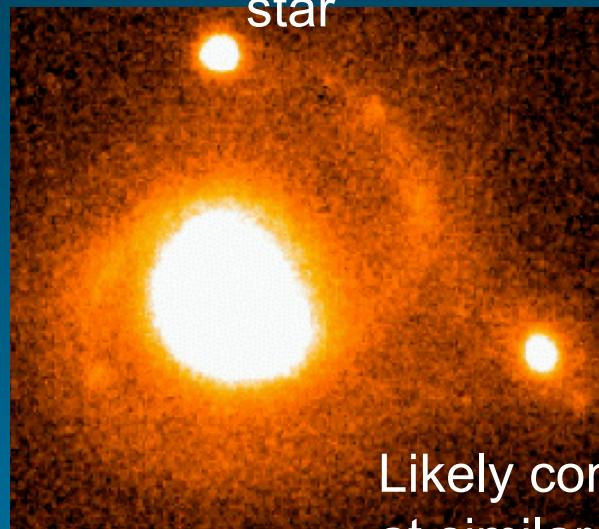


Pick “transition objects”
and study them in detail

- Signs for Mergers?
- Starburst properties?
- Gas as fuel for AGN?

In a nutshell

- One of the closest QSOs ($z=0.06$) and possible counterpart of high- z QSOs
- Prototype narrow-line Seyfert 1
- Direct spectroscopic evidence for nuclear starburst
[Schinnerer et al. 1998]



NIR images and spectra (ISAAC)
mm observations (BIMA, PdBI)

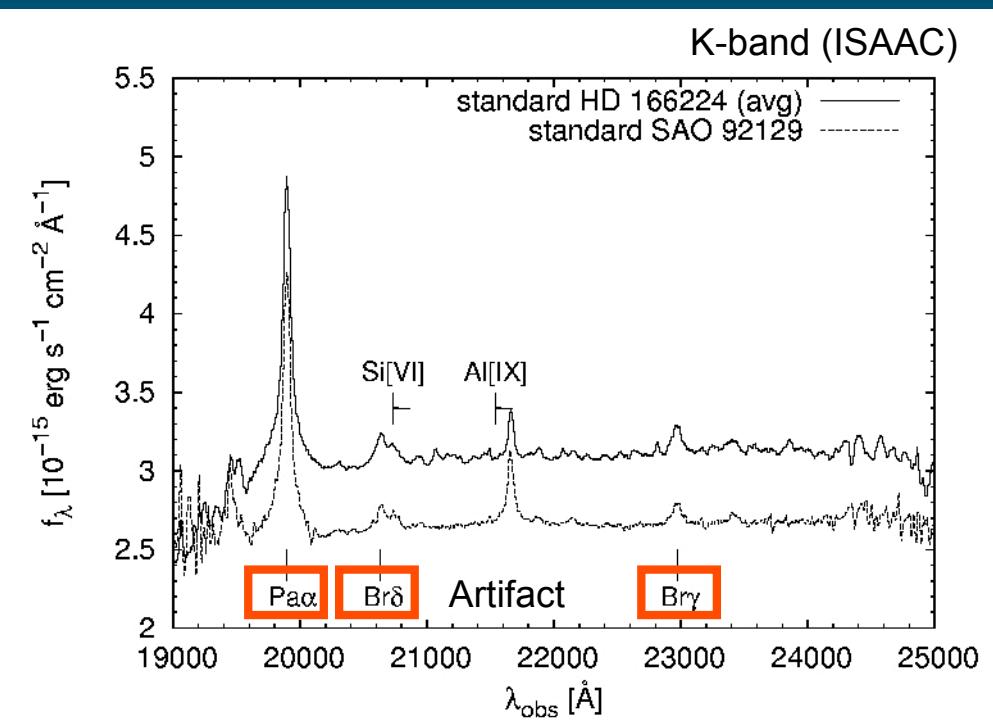
Nuclear Spectrum

[Schinnerer et al. 1998, Scharwächter et al. 2007]

Narrow-line Seyfert 1

- Small black hole mass and high accretion rate [e.g. Mathur 2000]

QSO in the formation?



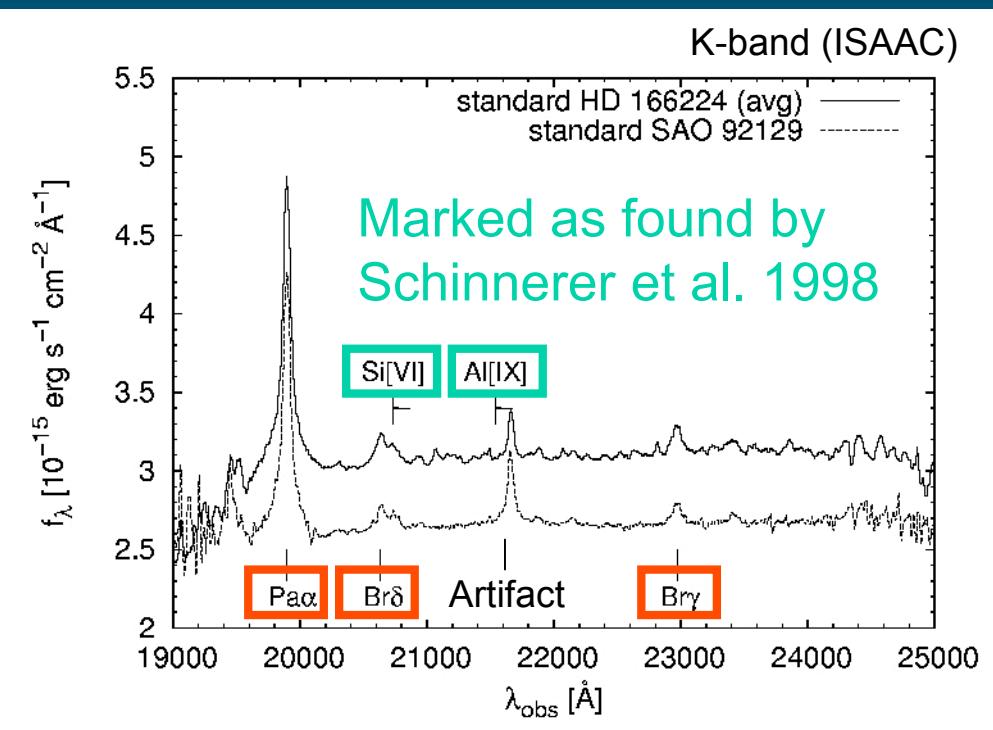
Nuclear Spectrum

[Schinnerer et al. 1998, Scharwächter et al. 2007]

Possible nuclear outflow:

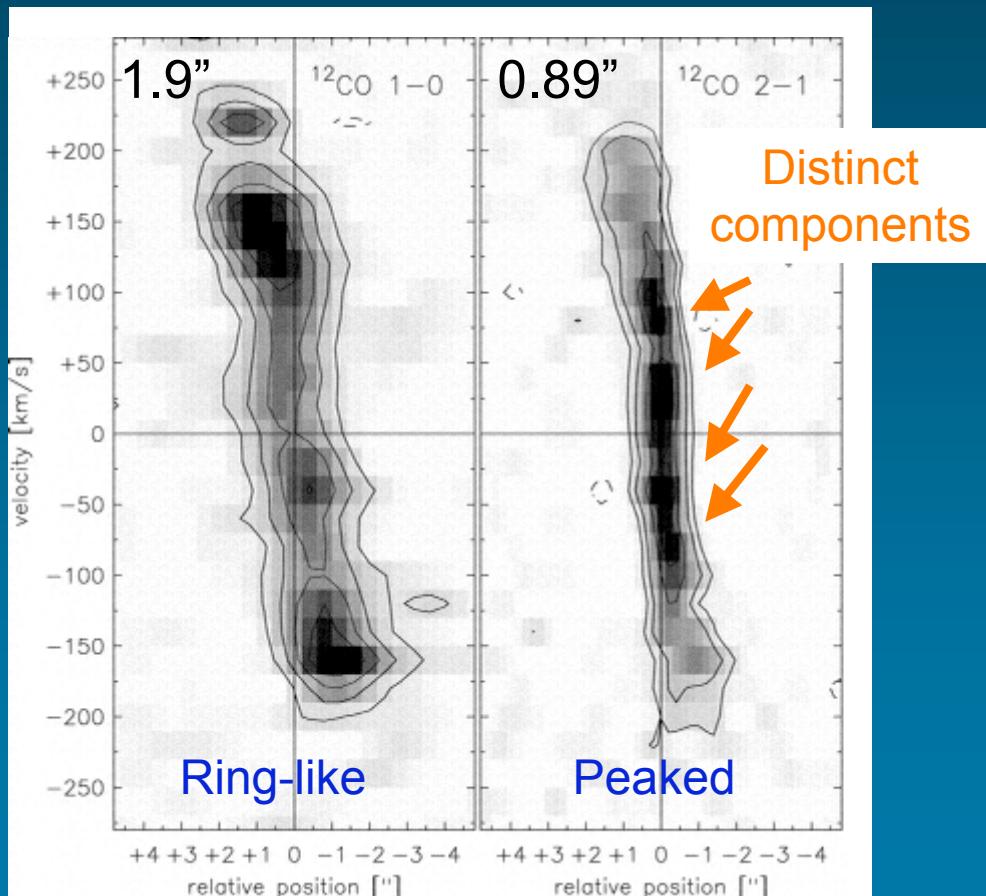
- Blueshifted high-excitation lines ($[SiVI]$ by ~ 1460 km/s wrt to I Zw 1)

Young QSO stage?



mm Observations

Plateau de Bure [Staguhn et al. 2004]



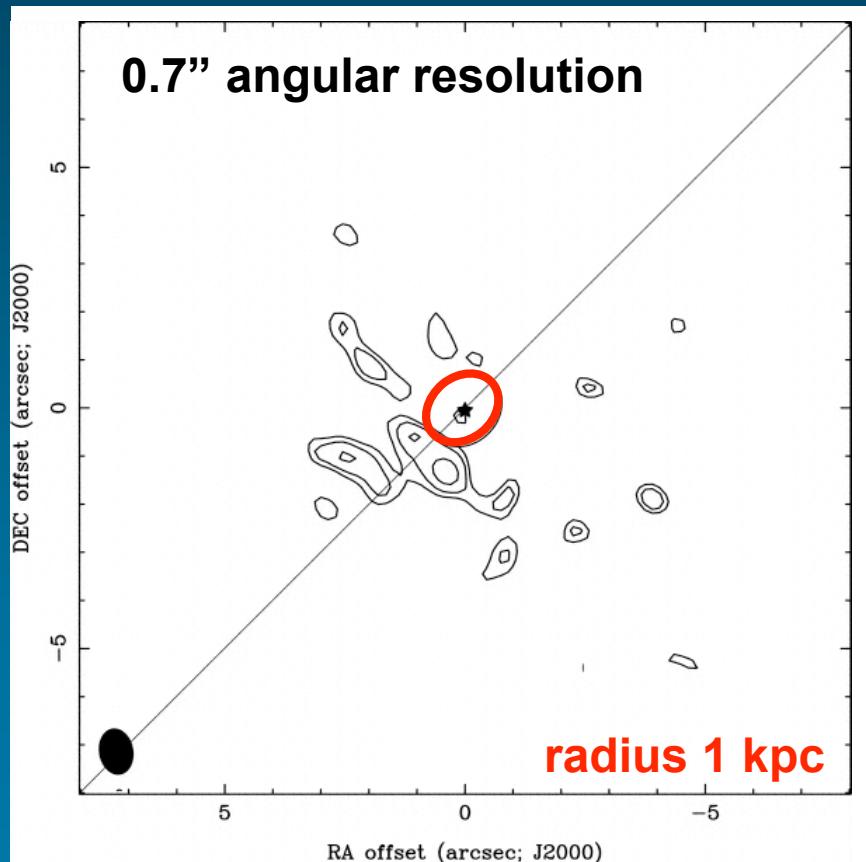
Position-velocity diagrams

- Disk: Cold and/or sub-thermally excited gas
- Nucleus: Warm optically thick gas with distinct components

Excitation by starburst
not by AGN

mm Observations

$^{12}\text{CO}(1-0)$ map (BIMA)



[Staguhn et al. 2004]

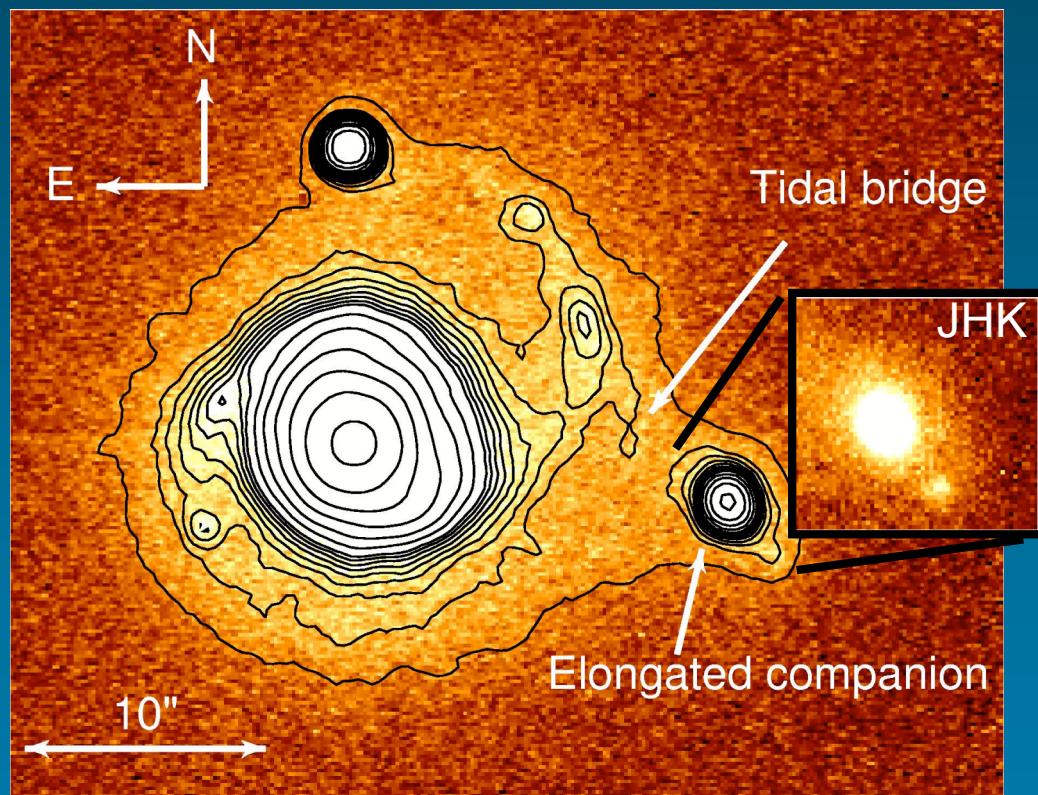
Circum-nuclear
molecular ring

- Likely location of nuclear starburst
- Typical in barred spiral galaxies
- In un-barred spirals possibly caused by minor merger

Minor merger?

Near-Infrared Imaging

- J-band image (ISAAC)

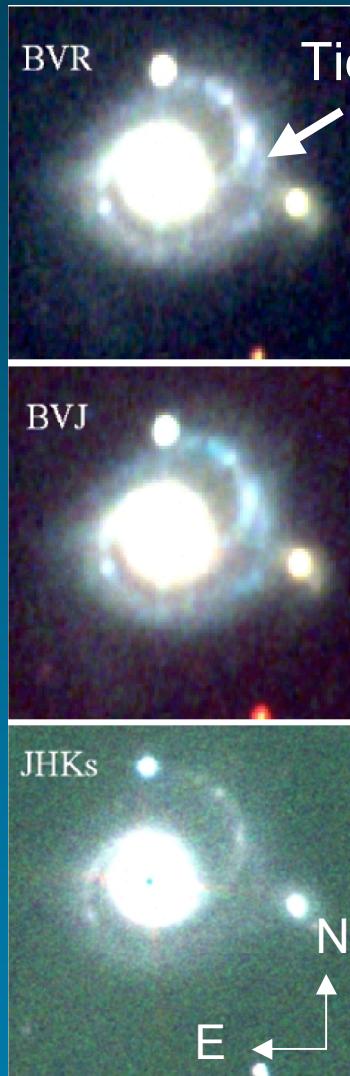


[Scharwächter et al. 2003, 2007]

Evidence for ongoing merger

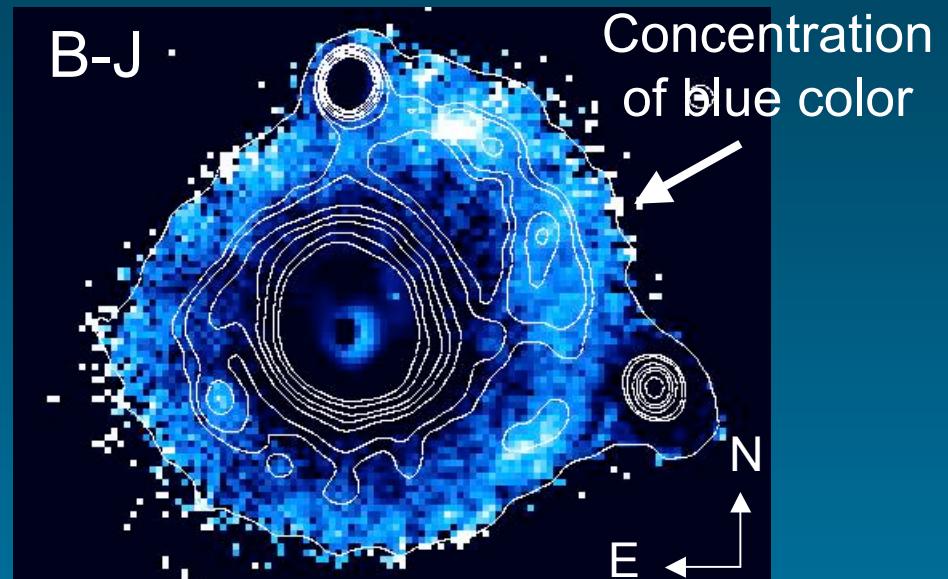
- Tidal bridge
- Elongated companion
- Apparent “tidal tail” of companion resolves into separate object

Optical-NIR Color Composites



ISAAC, EFOSC2

[Scharwächter et al. 2007]



Dark: red; Bright: blue

Possible indication of
minor-merger enhanced
star formation activity

ULIRG-to-QSO transition stage?

- Possible young stage of nuclear activity in I Zw 1 indicated by nuclear spectral characteristics → Young QSO
- Molecular ring as possible location of ongoing starburst → Aged ULIRG
- Spiral host galaxy: No recent major merger → Unlike typical ULIRGs!
- Further evidence for ongoing minor merger which may
 - Induce nuclear activity [e.g. Hernquist & Mihos 1995]
 - enhance Seyfert activity to QSO levels [e.g. Corbin 2000]
 - be unrelated to nuclear activity

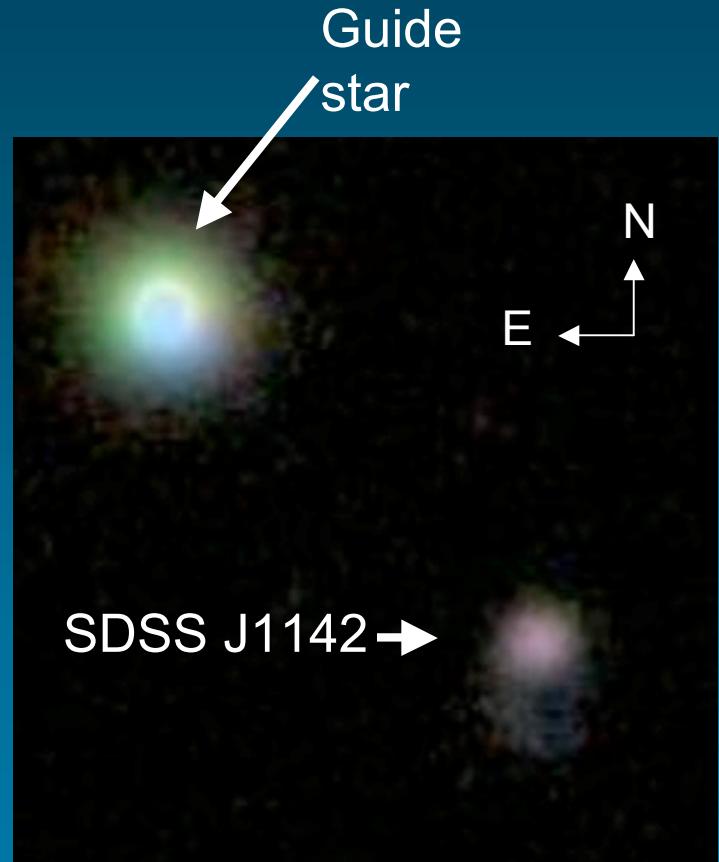
SDSS J114203.40+005135.8

SDSS J1142

In a nutshell

- $z=0.245$
- IRAS source
- Spectral characteristics of star formation activity
[Magliocchetti et al. 2002]
- Radio source
[Magliocchetti et al. 2002]

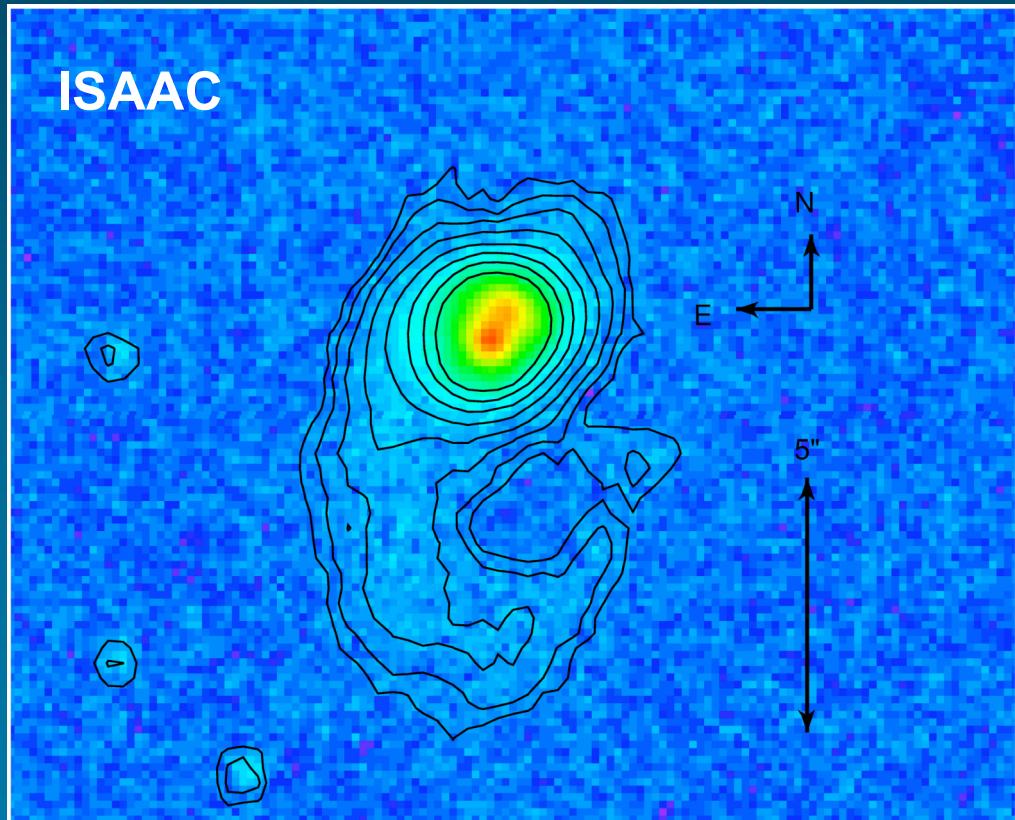
SDSS QSOs with
nearby guide stars



[SDSS DR5]

ISAAC Imaging

SDSS J1142



- First evidence for double structure in the central region

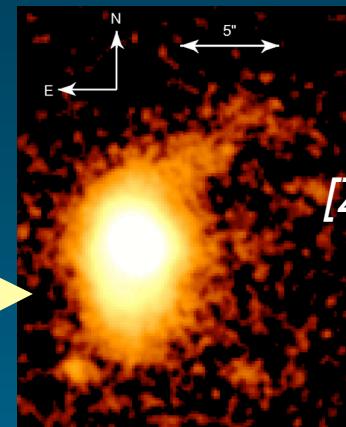
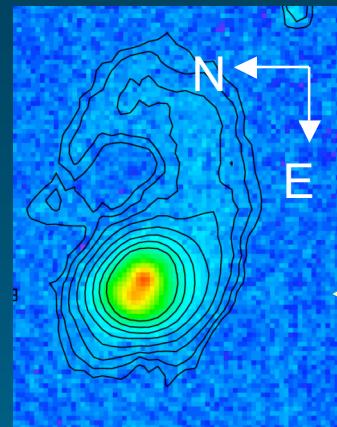
Two galaxy centers
in the late stage of
major merger,
or hot spot from
radio jet?

[Work in progress]

3C 48 Analog?

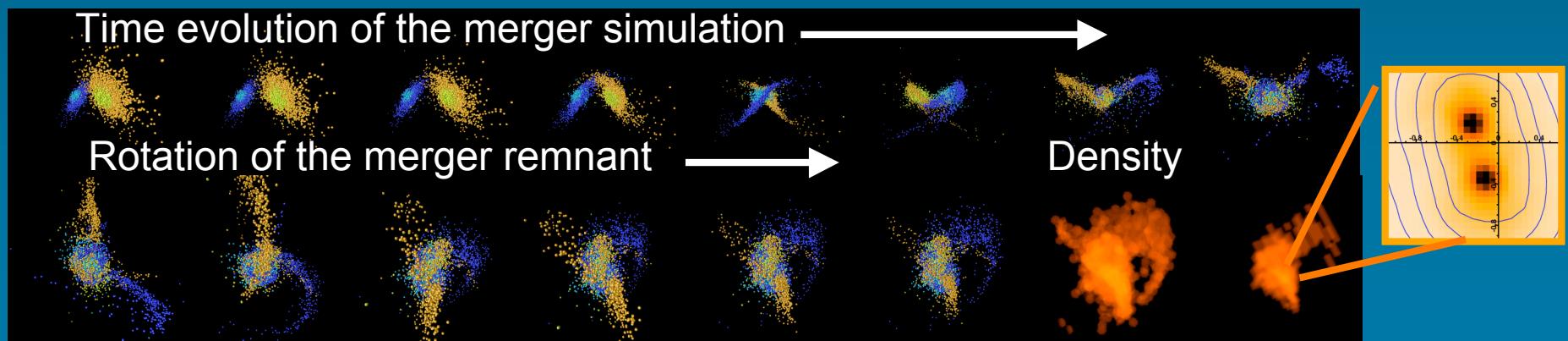
SDSS J1142

SDSS J1142
(ISAAC)



3C 48
(ISAAC)
[Zuther et al. 2003]

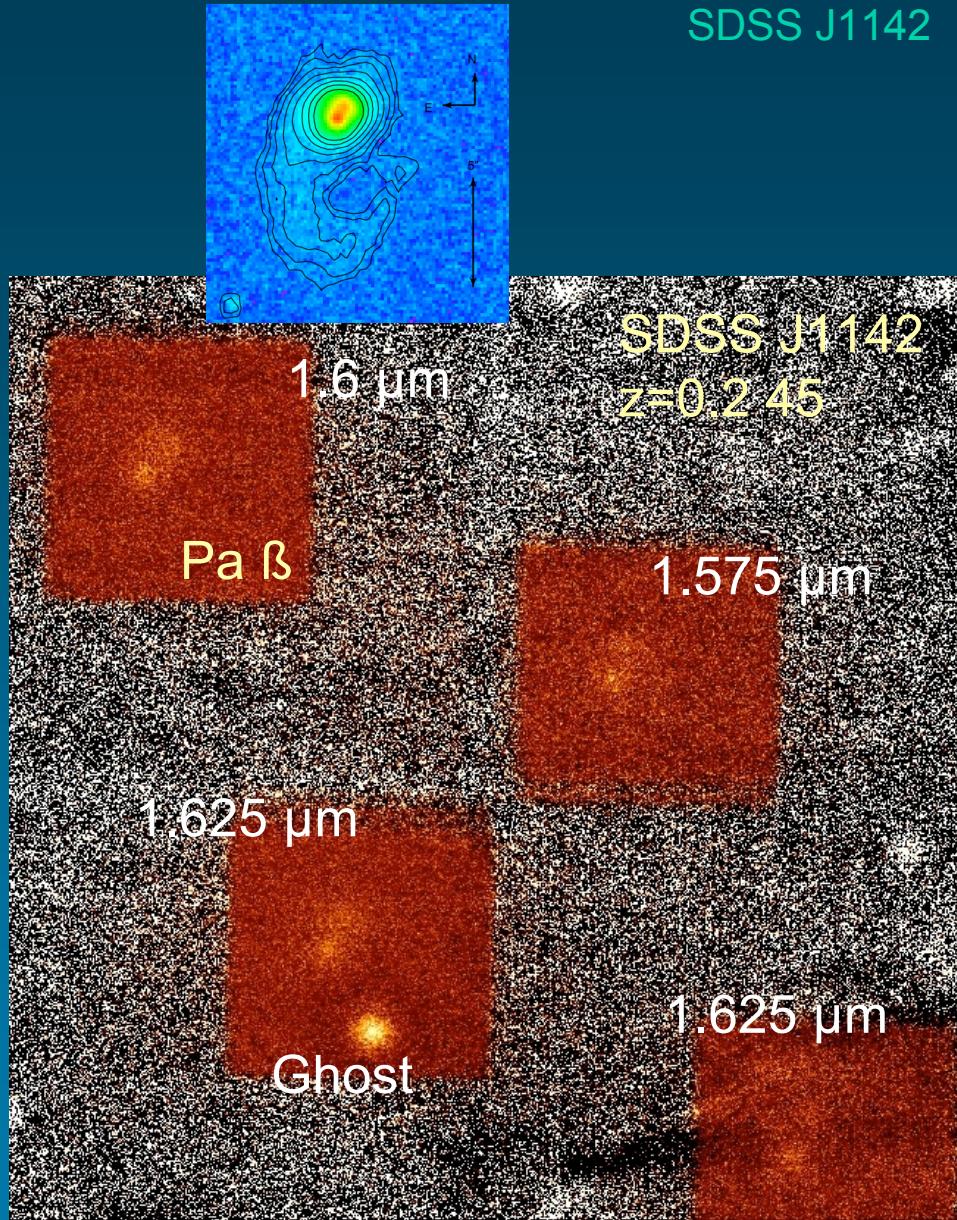
- Multi-particle merger simulation for 3C 48 [Scharwächter et al. 2004]



NACO Imaging

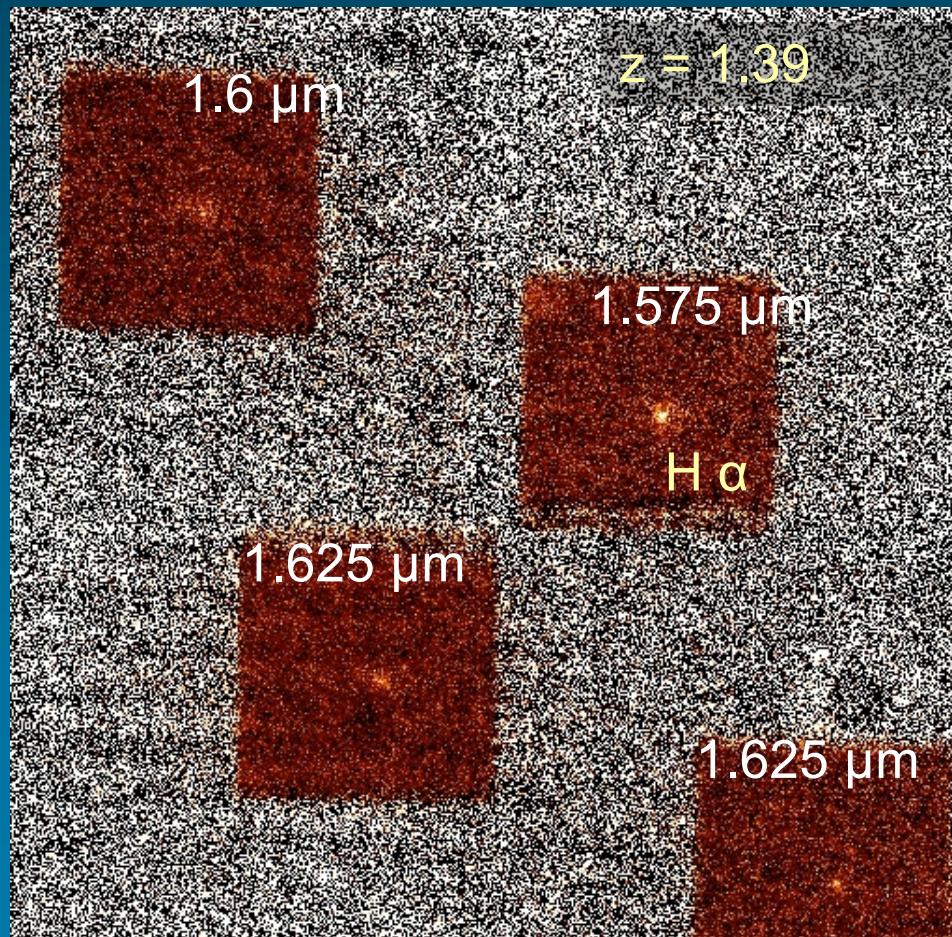
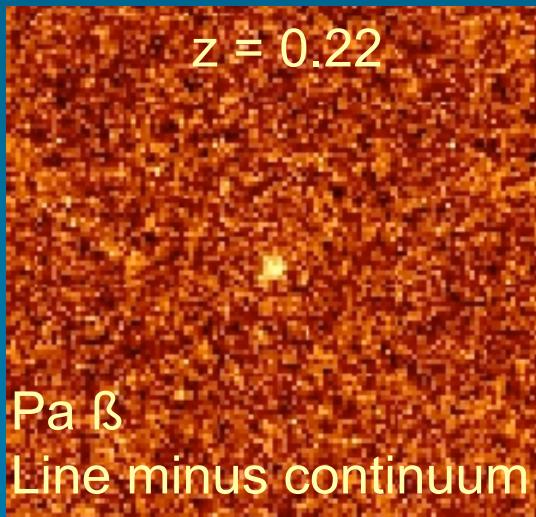
Simultaneous Differential Imager (SDI)

- Emission line gas in QSOs at high angular resolution (“tunable redshift” approach)
- SDSS J1142: Double structure clearly resolved
- Low S/N → so far...



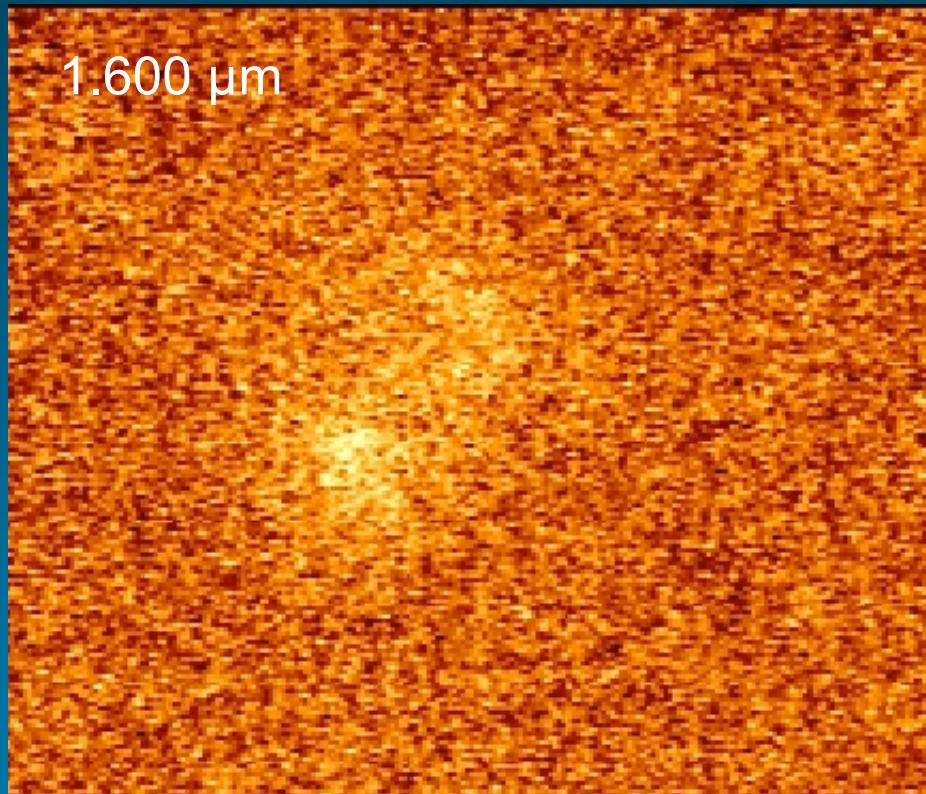
Ongoing Program

- ... P80 program
(UT 4 astronomers!!!)
- Laser guide star facility:
improved Strehl, larger
sample flexibility



Tentative First Glance

SDSS J1142



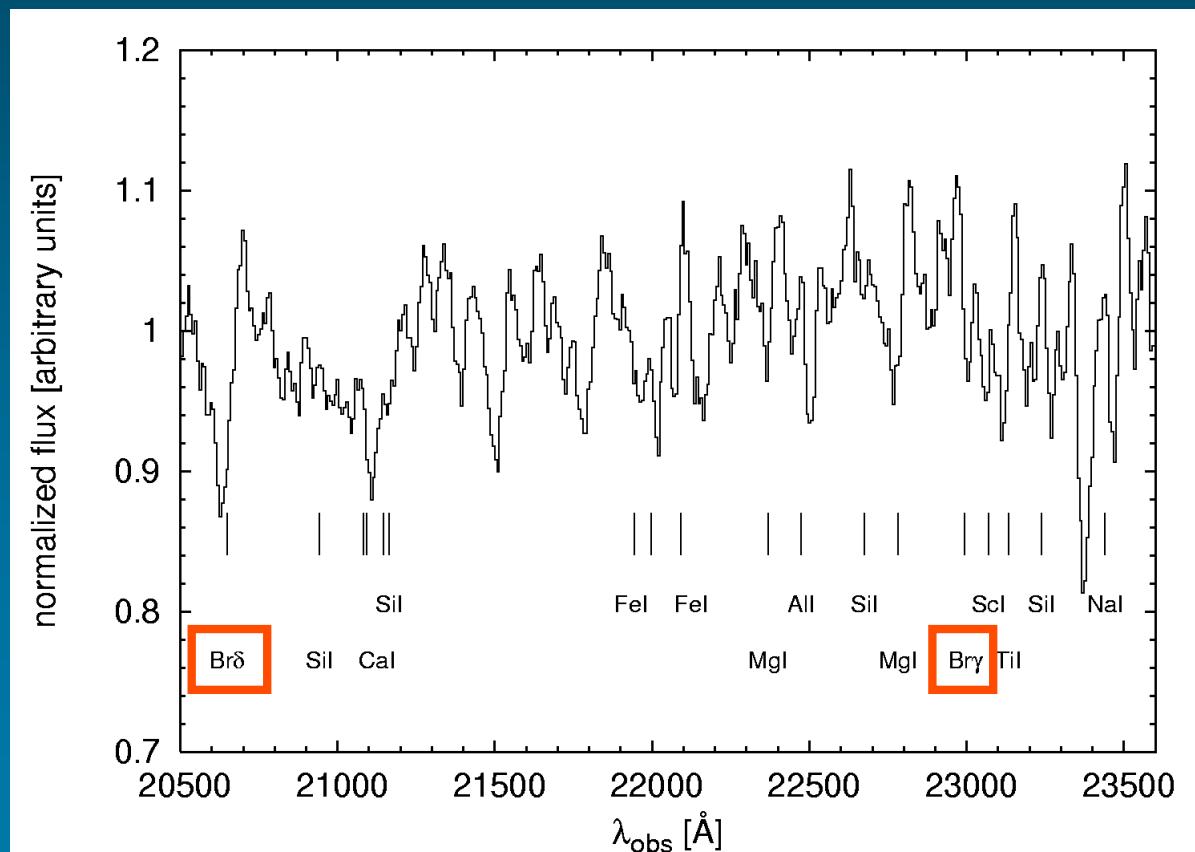
Role in the Evolutionary Sequence^{SDSS J1142}

- Major merger remnant
- Indications for star formation activity

Interesting object for
future detailed case study

The Likely Companion

K-band spectrum (ISAAC)



No emission lines,
perhaps
absorption lines

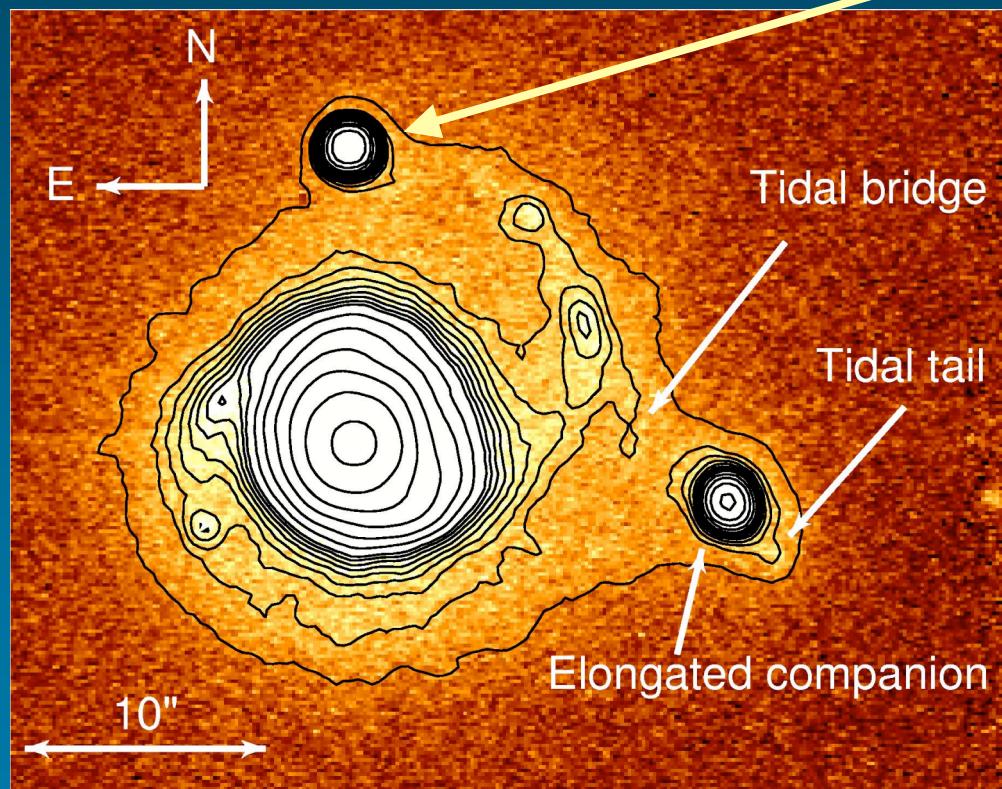
► Old evolved
stellar
population

*[in agreement with
Canalizo & Stockton
2001]*

Gas-poor
dwarf elliptical

The I Zw 1 System

J-band image (ISAAC)



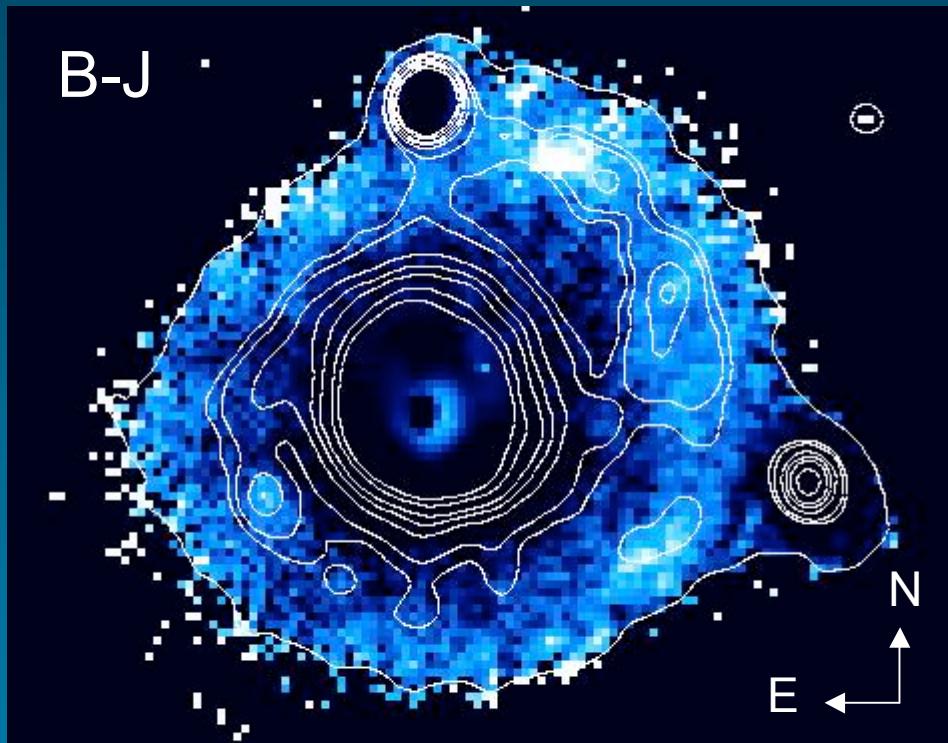
[Scharwächter et al. 2003]

Foreground star
[Stockton 1982]

As confirmed by
ISAAC spectrum

Companion galaxy
at approx. the
redshift of I Zw 1
[Canalizo & Stockton 2001]

Minor merger evidence



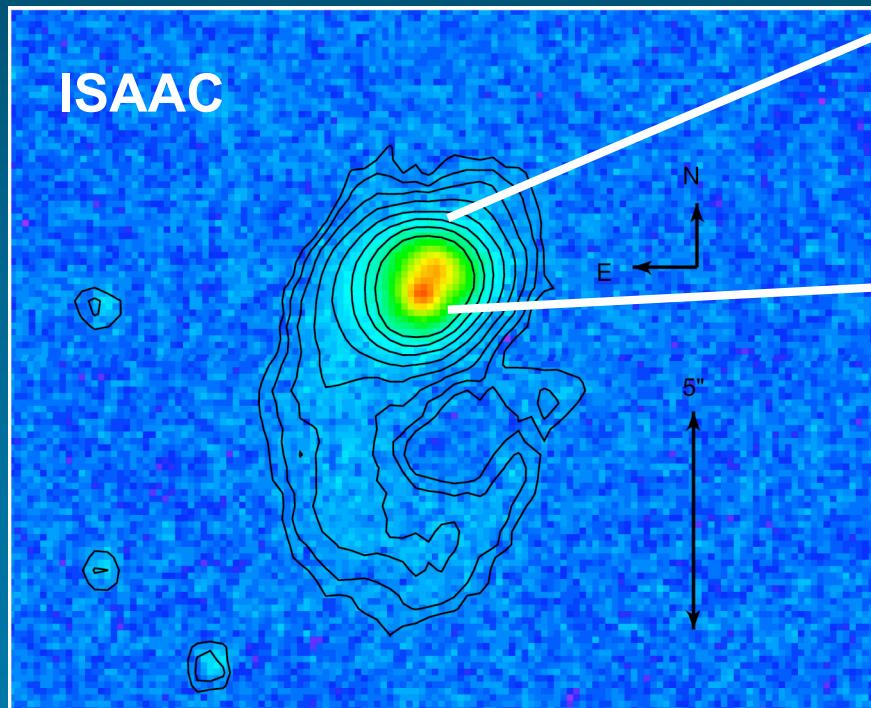
Bright: blue; Dark: red

- Concentration of blue color in western part of host
- ▶ Possible indication of star formation enhanced by tidal interaction

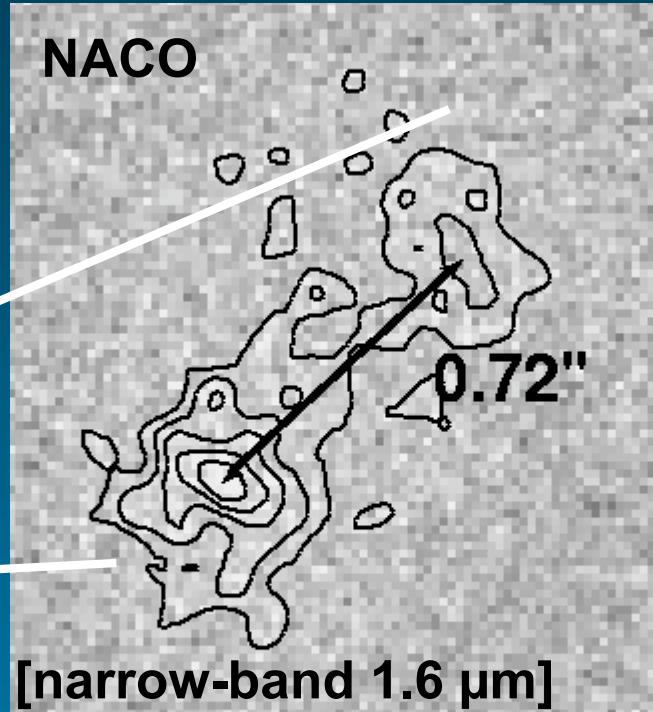
ISAAC Imaging

SDSS QSOs

- First evidence



[Work in progress]



- Hot spot
- Second galaxy nucleus

Collaborators

Andreas Eckart¹, Valentin D. Ivanov², Jari Kotilainen³,
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Juha Reunanen⁵, Eva Schinnerer⁶, Rainer Schödel¹,
Johannes G. Staguhn⁷, Lowell E. Tacconi-Garman⁸,
Jens Zuther¹

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² ESO Chile

³ University of Turku, Finland

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⁵ Leiden Observatory, The Netherlands

⁶ MPIA Heidelberg, Germany

⁷ NASA/Goddard Space Flight Center, USA

⁸ ESO Garching

My life before Chile...

Solingen (Germany)  40 km

Cologne: Studies (University of Cologne), Ph.D. thesis (supervised by Andreas Eckart) 12 000 km

ESO Chile in May 2005 (duty station: La Silla)



„Kölsch“ – Beer in 0.2 l glasses



Carnival – just
started on
11/11 at 11:11