

# Finding Local Dwarf Galaxies in HI

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@eteq

Yale University

Hubble Fellow

with

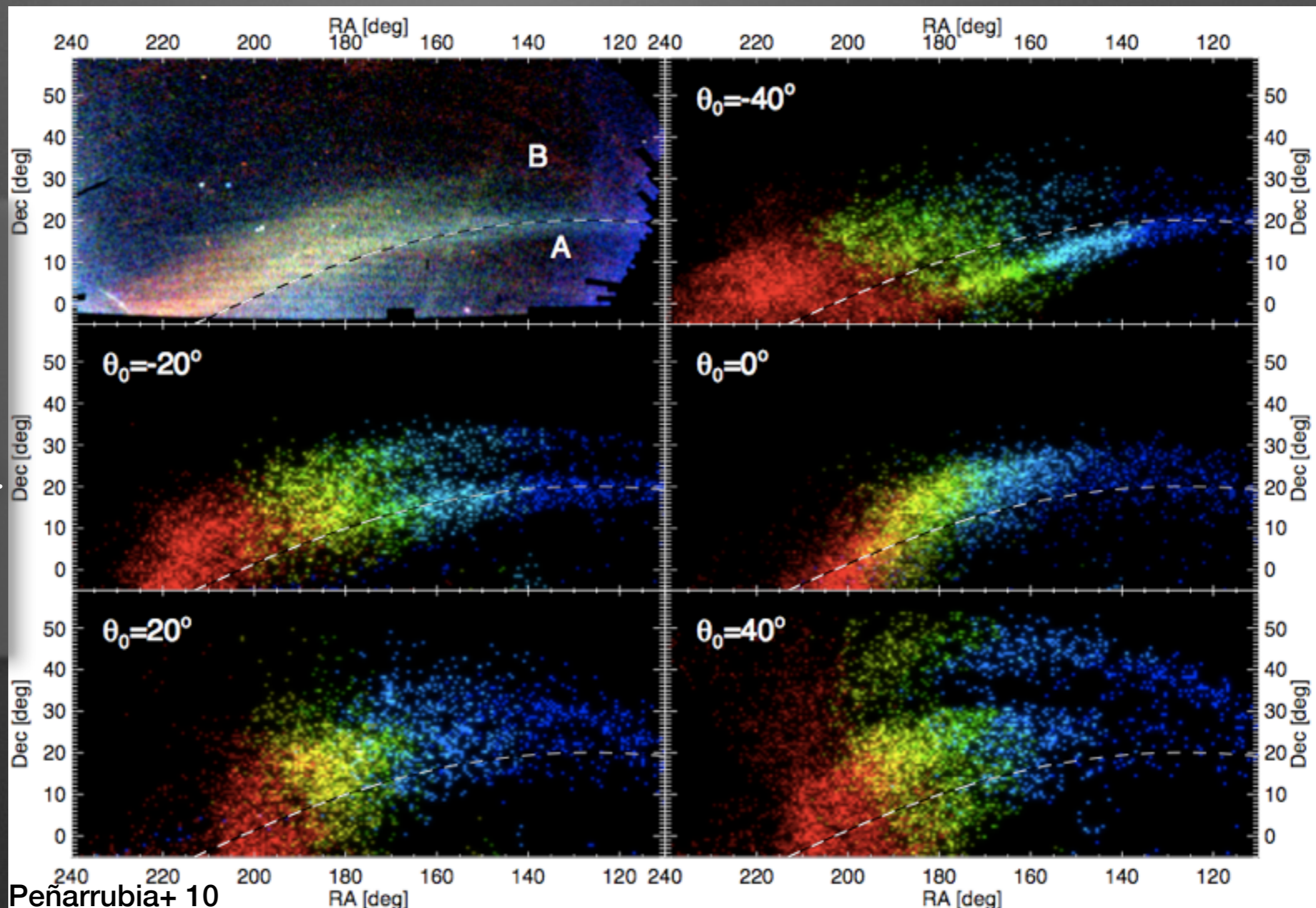
Marla Geha, Jana Grcevich, Mary Putman, Dan Stern, Josh Peek

Trip-S ESO Workshop, Apr 13, 2015

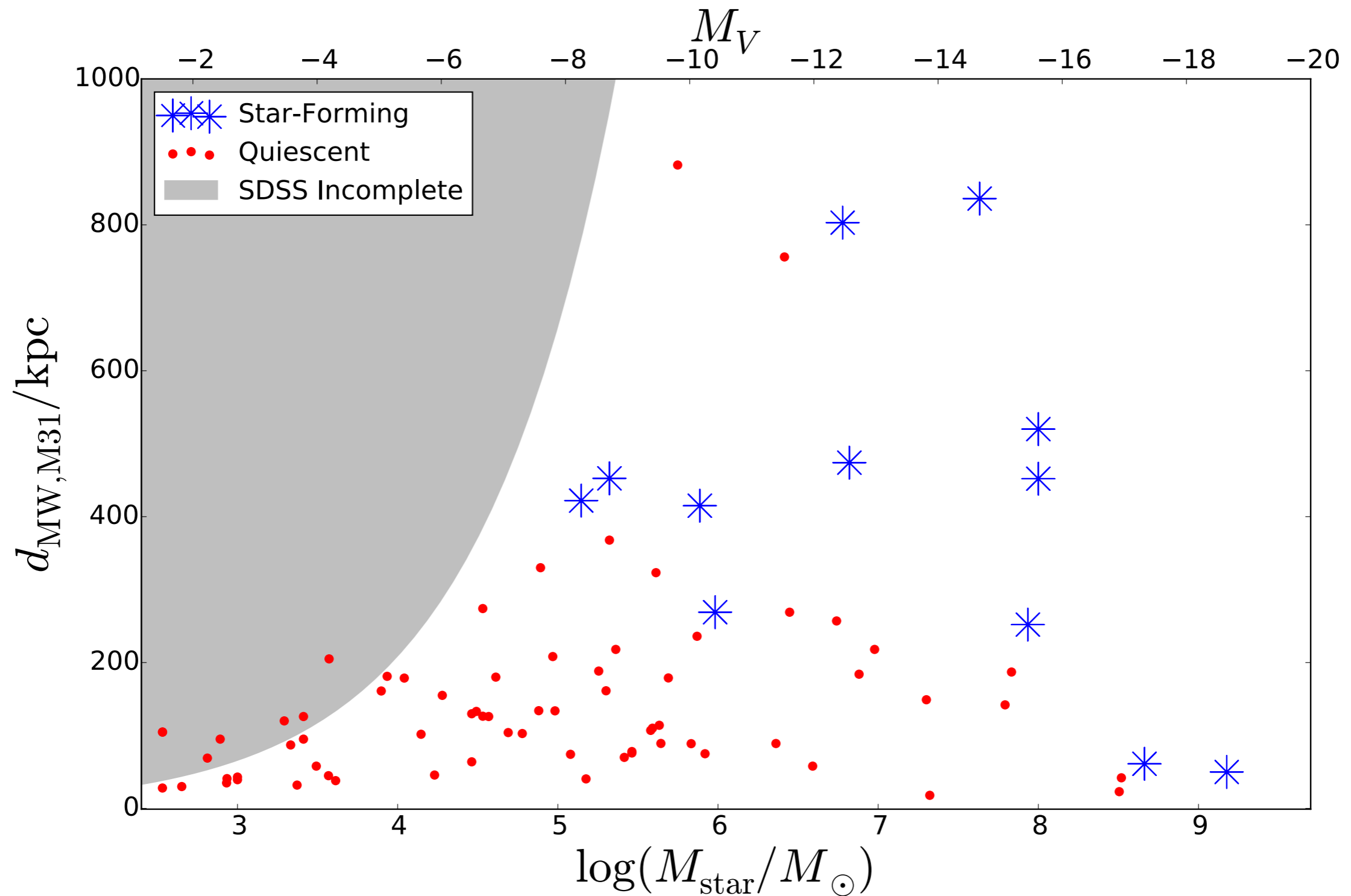


# “Initial Conditions” of Satellites Are Important

E.g.  
Existence  
and  
orientation  
of a disk  
matters

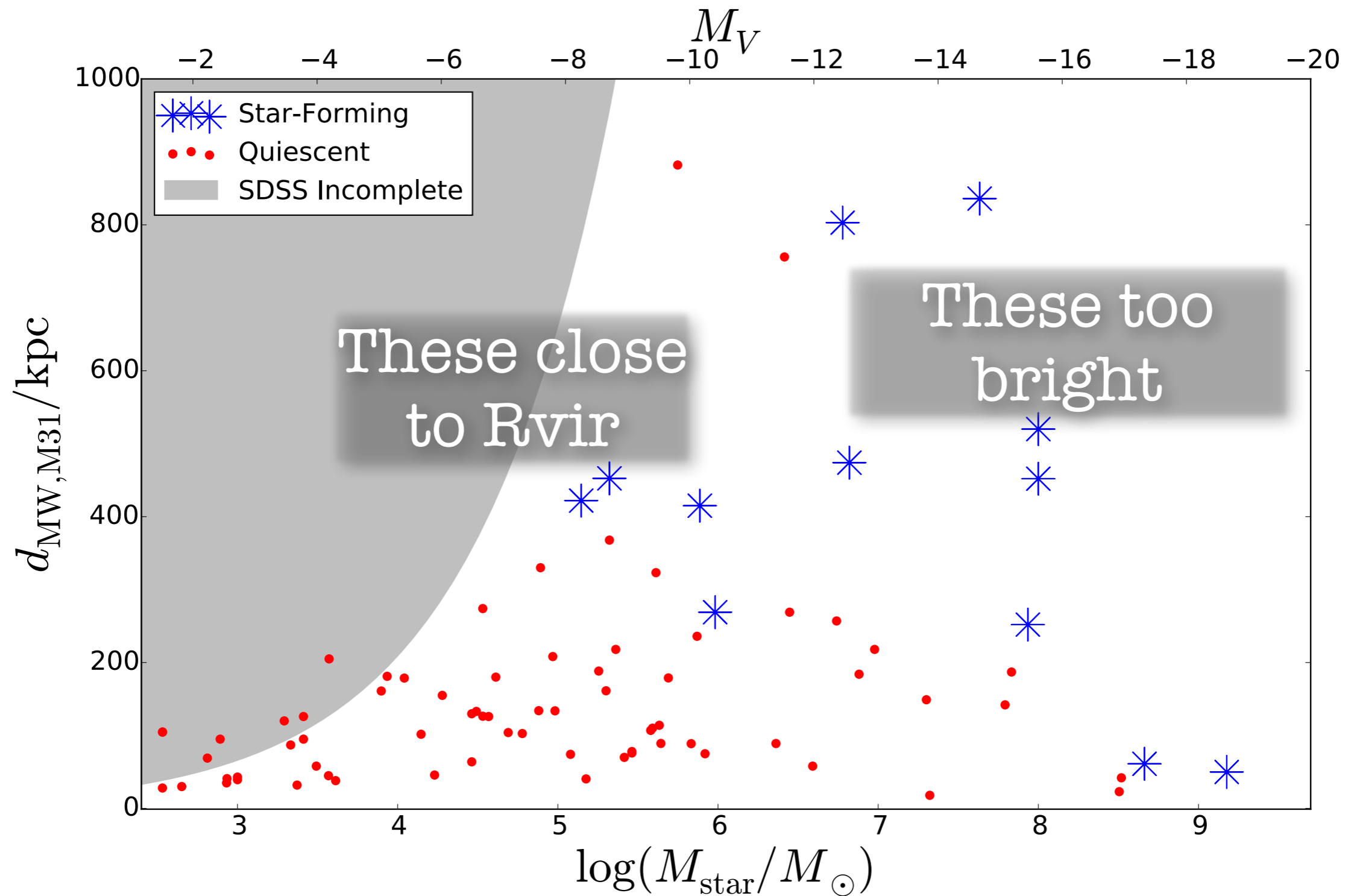


# So where can we look for progenitors?

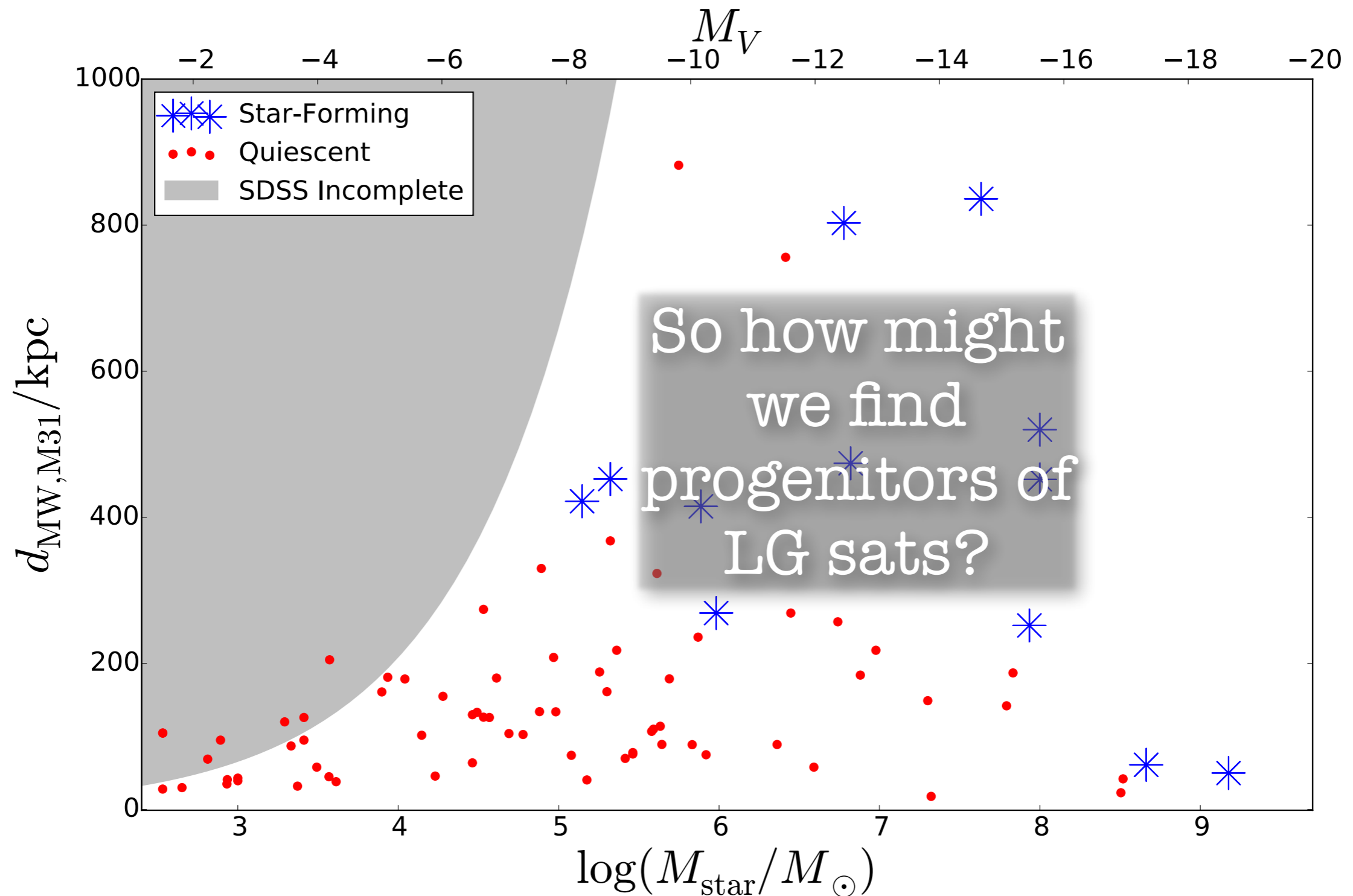




# So where can we look for progenitors?

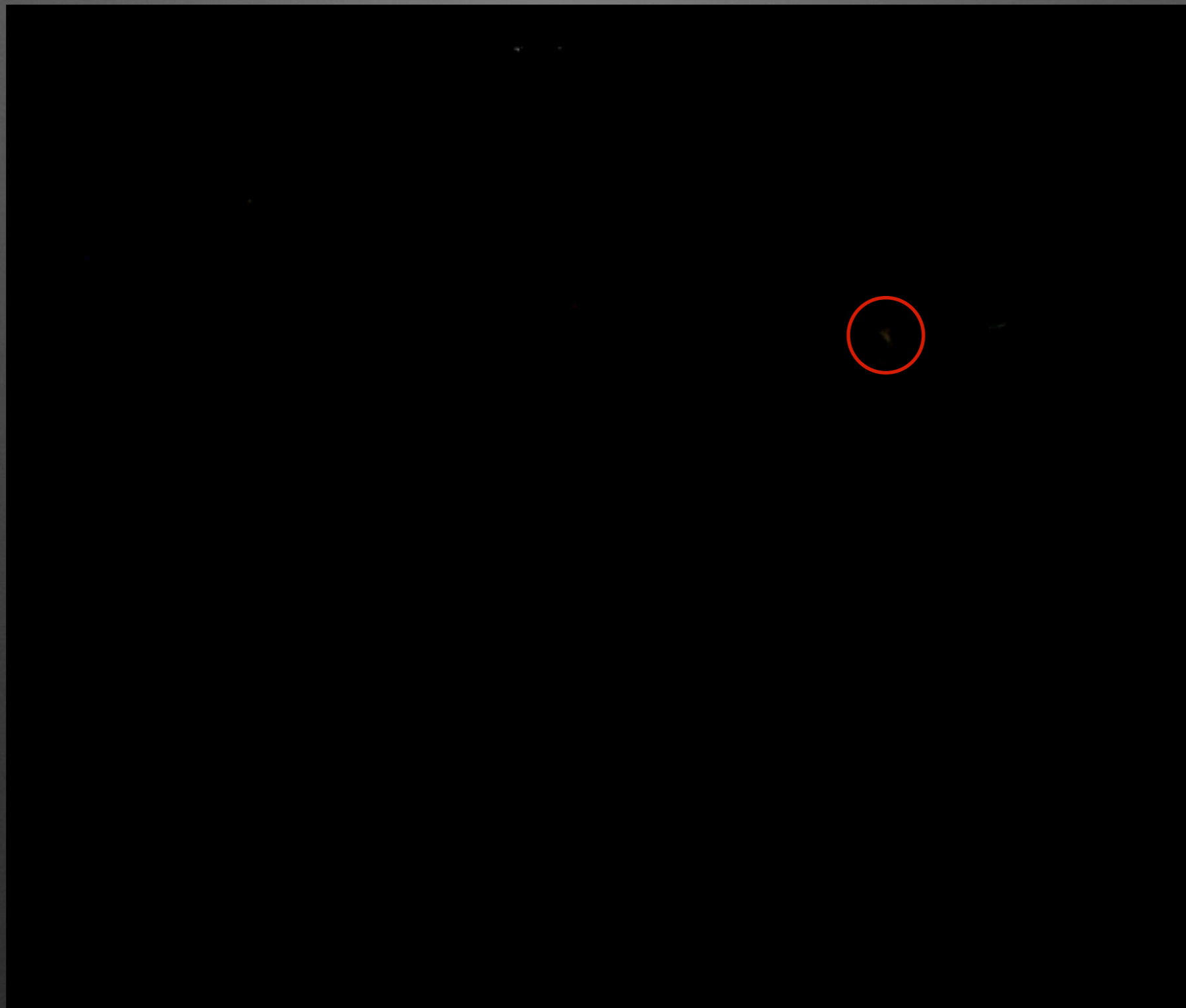


# So where can we look for progenitors?



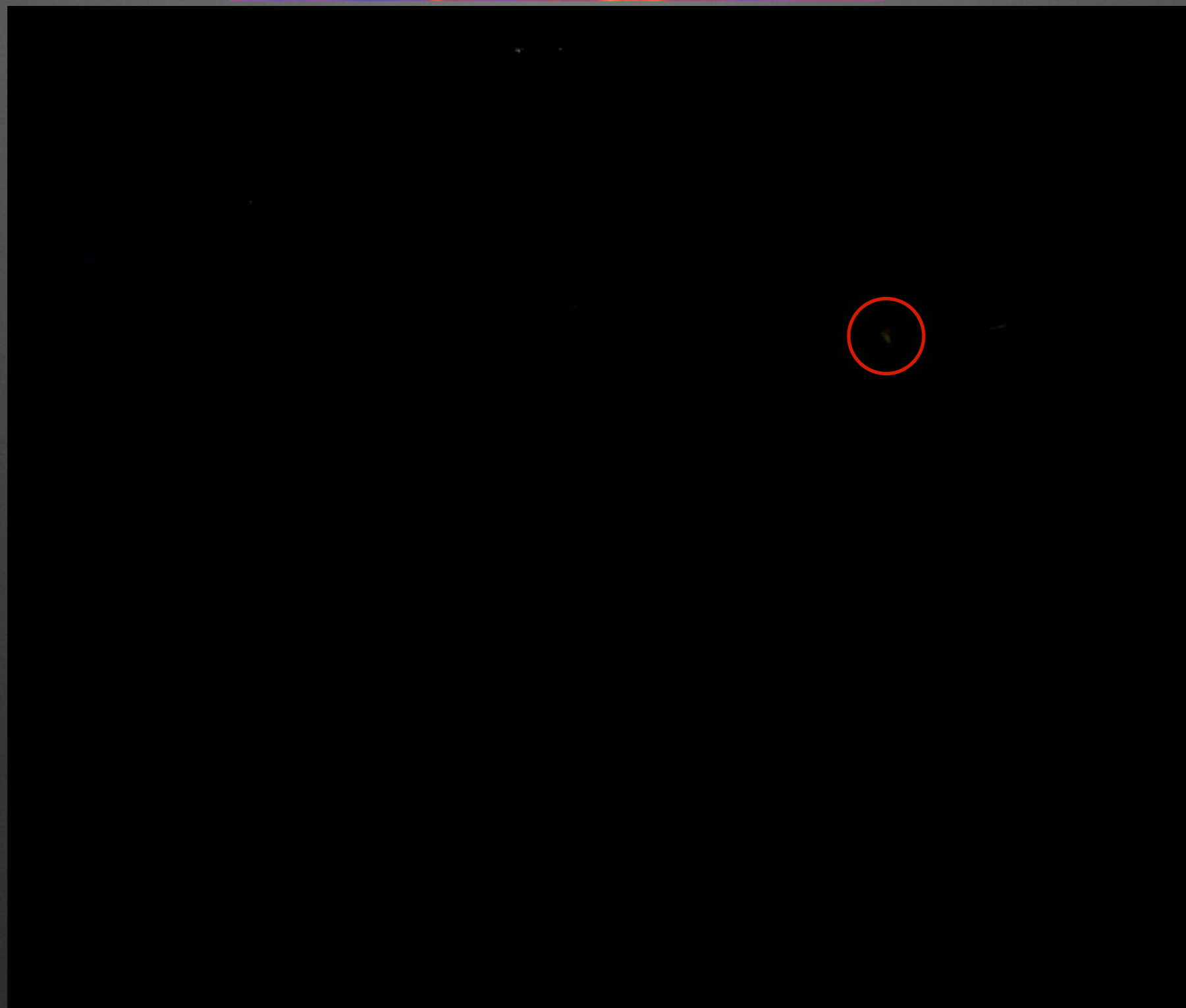


DR1, Peek+ 11



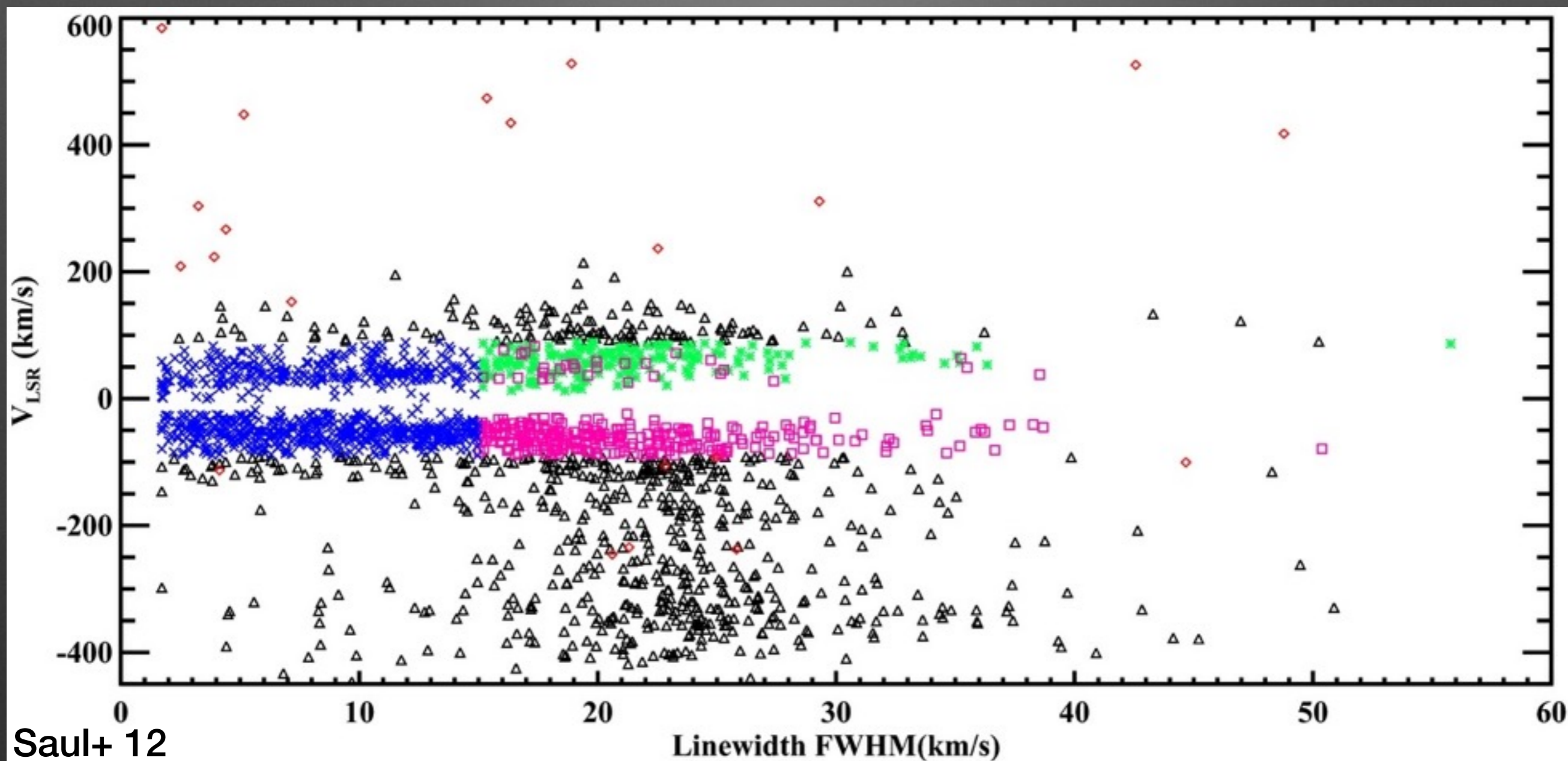


DR1, Peek+ 11





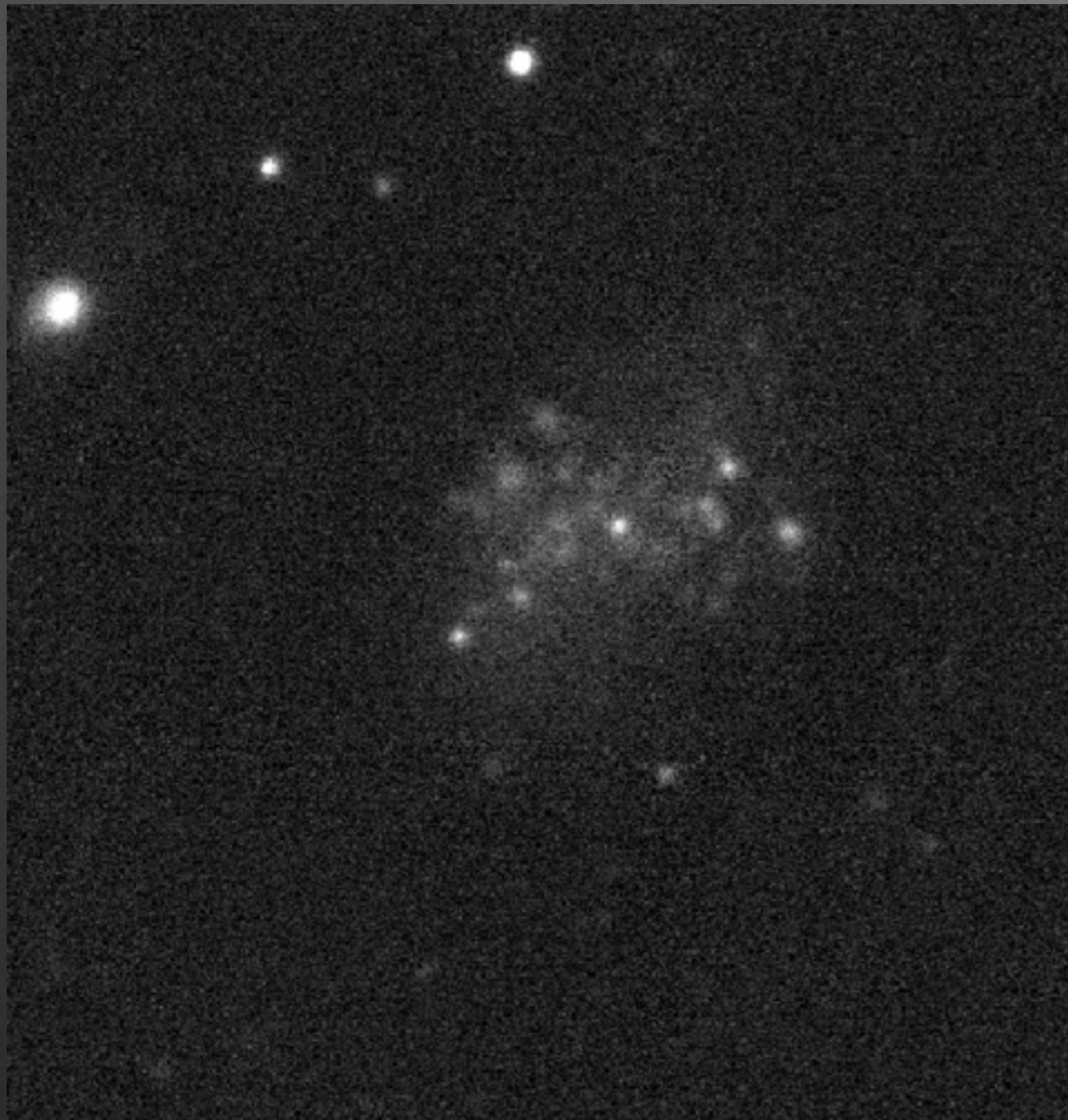
# Look for HI blobs that are not Galactic



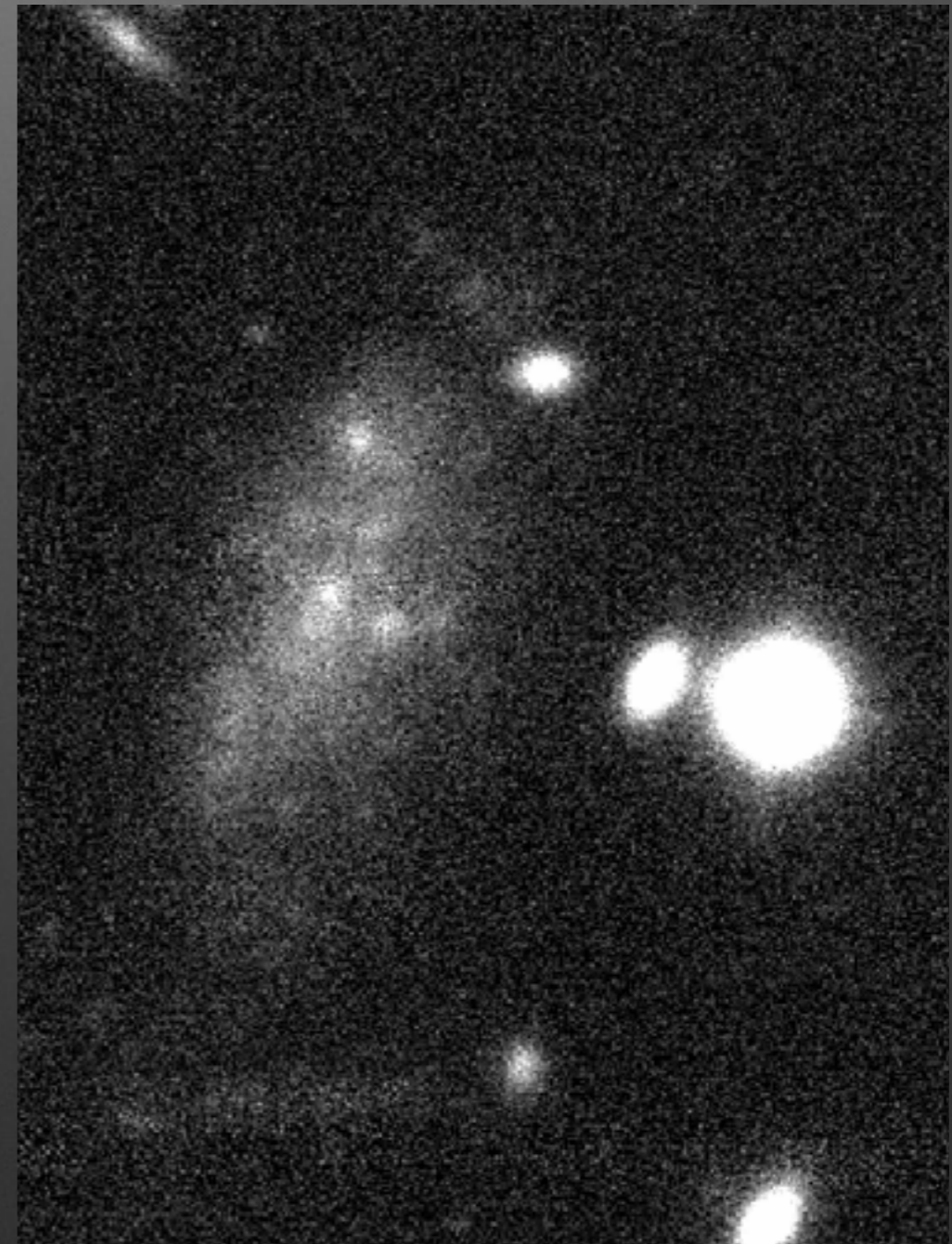


# Four (of ~ 30 candidates) have dwarfy optical counterparts

Pisces A



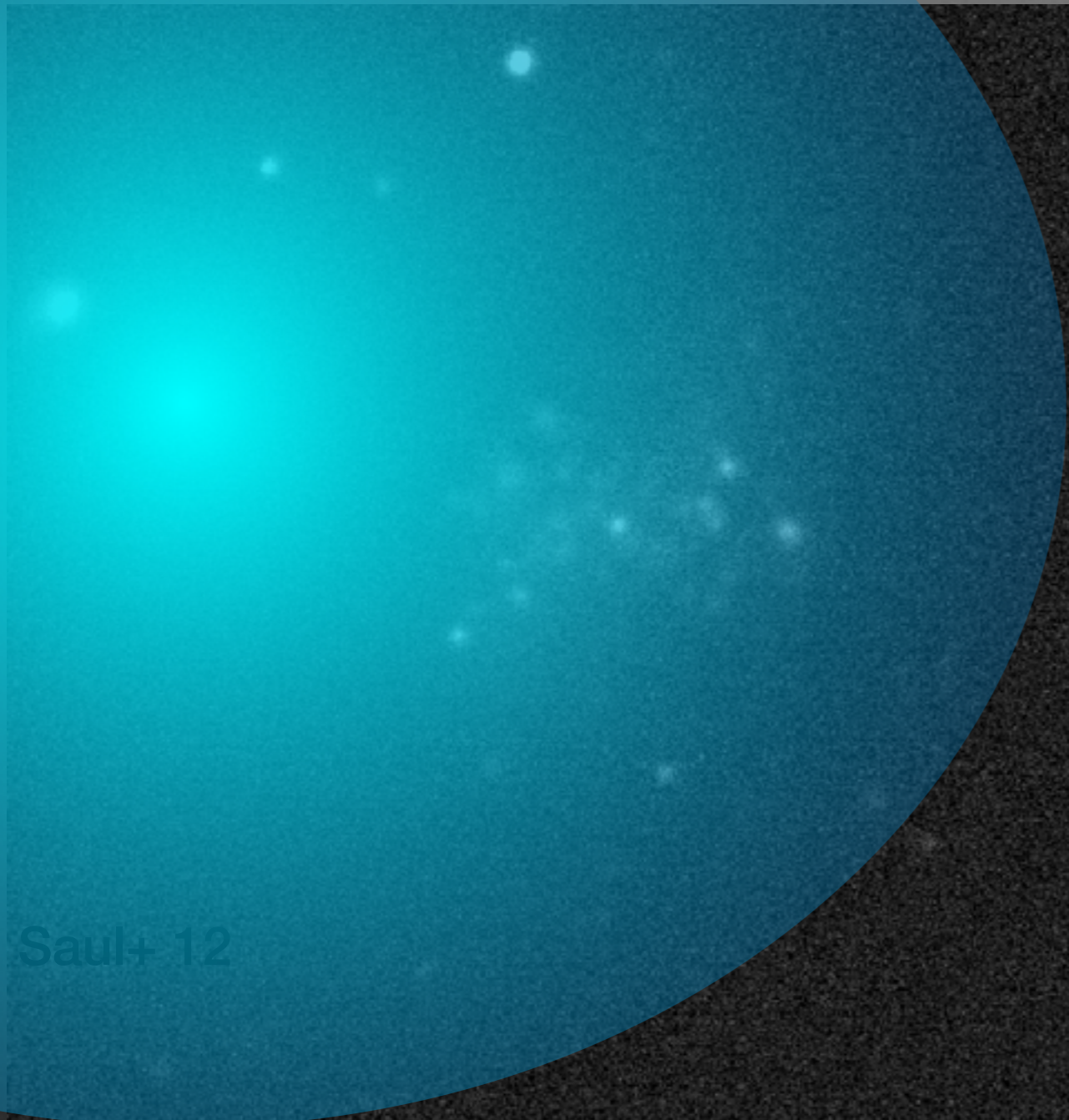
Pisces B



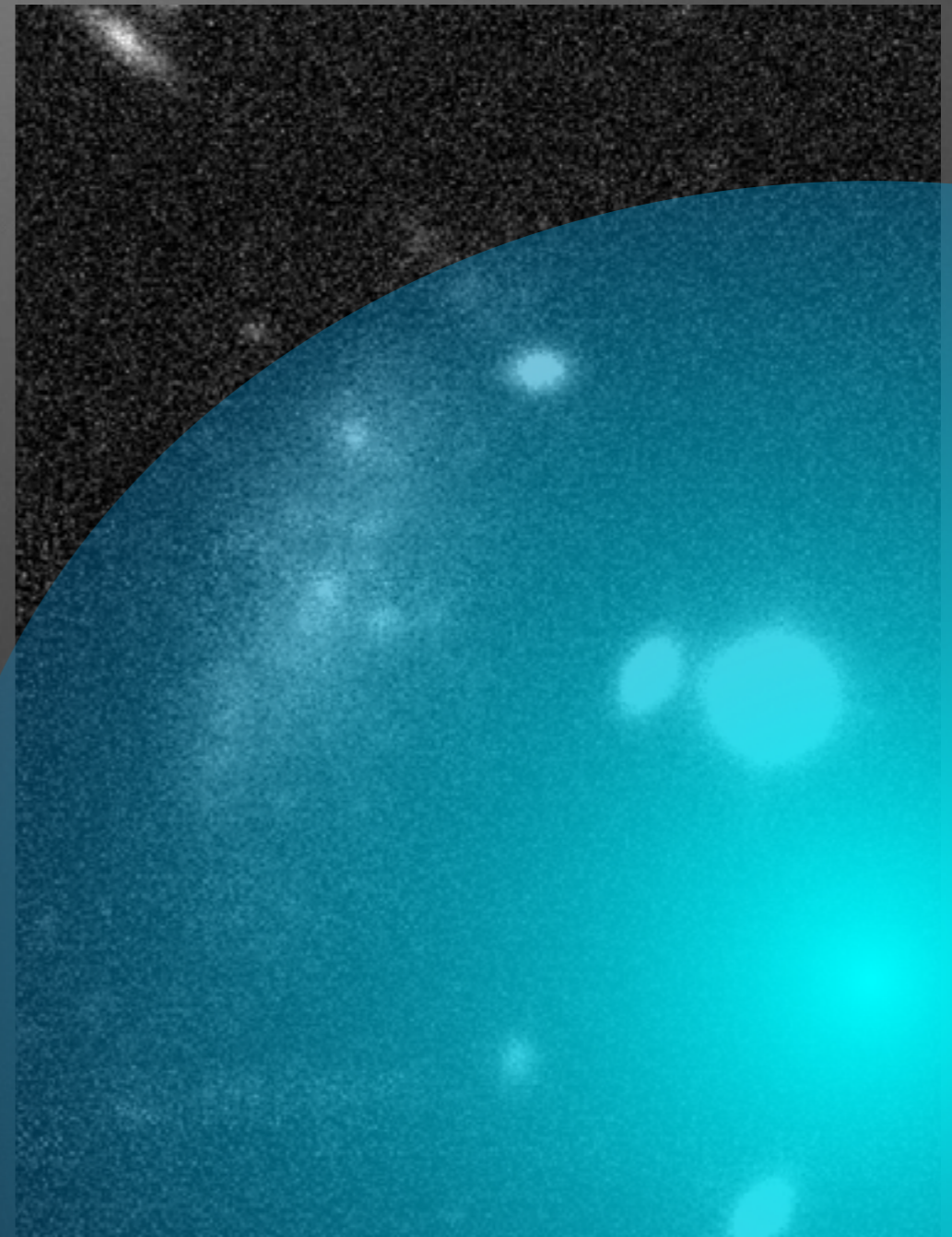
(See Sand+ 15 for other two)



# But are these the HI sources?



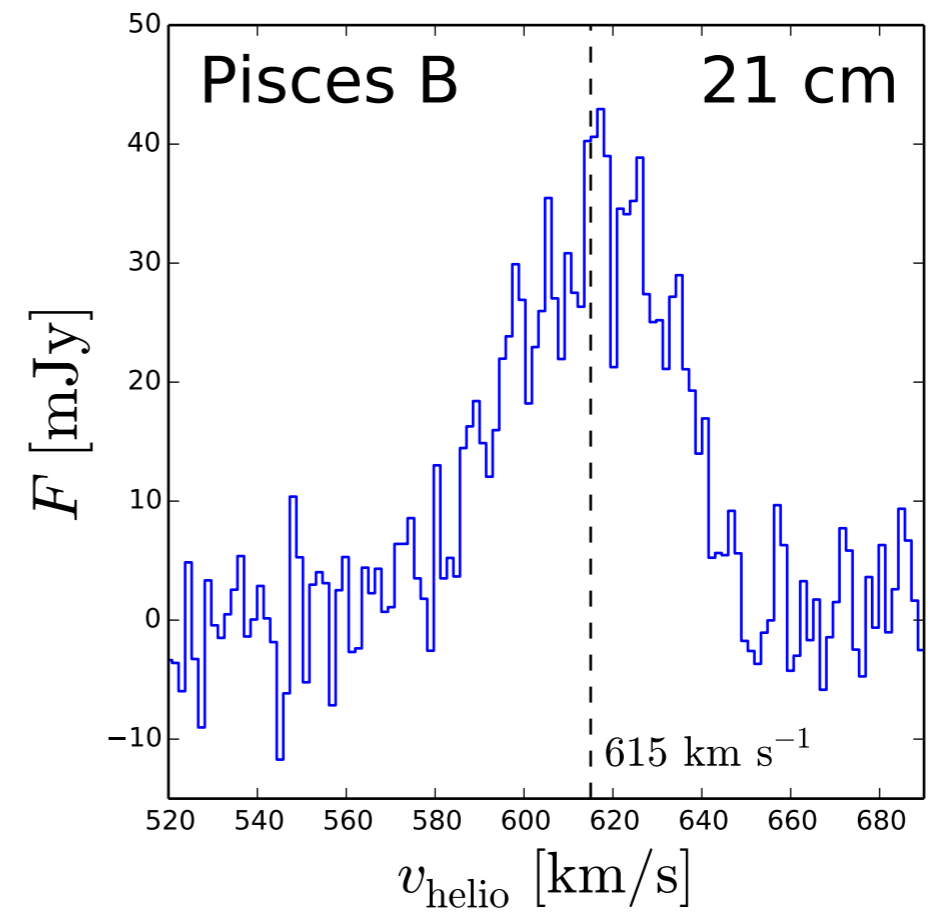
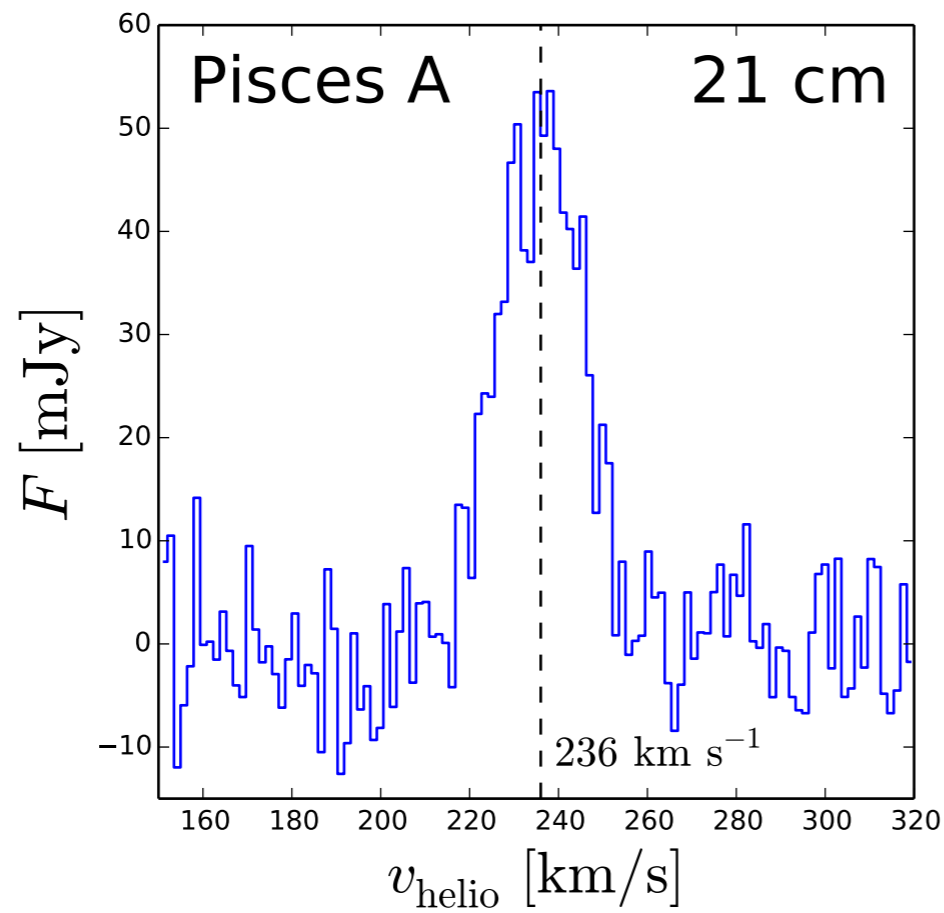
Saul+ 12



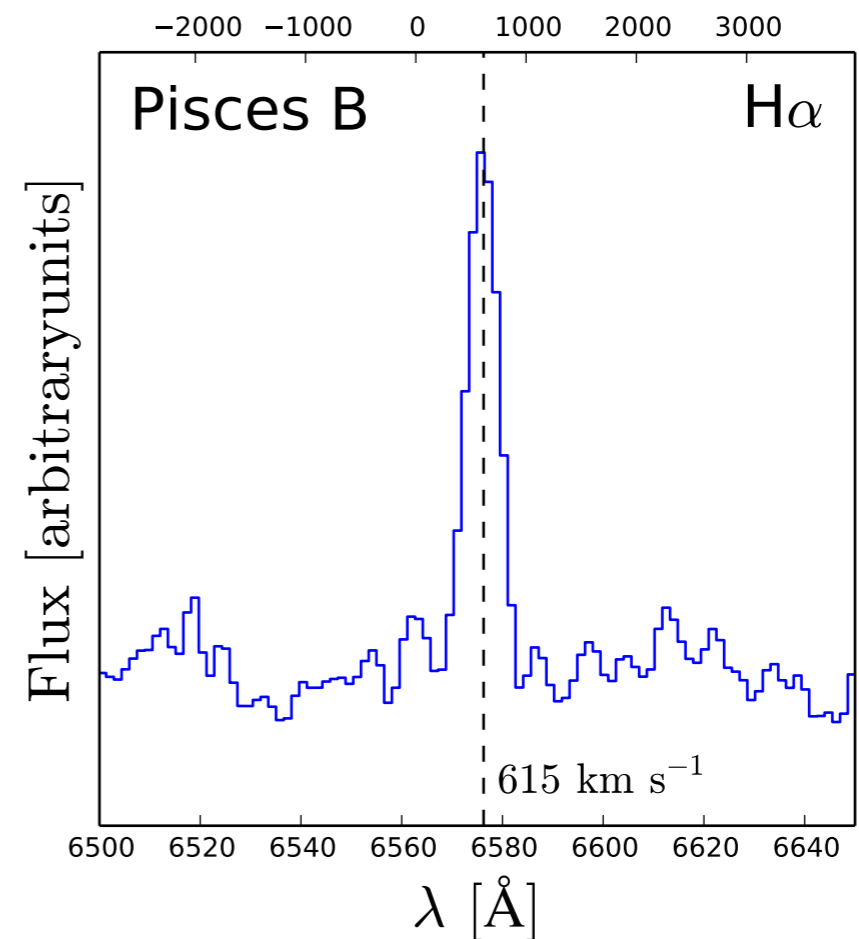
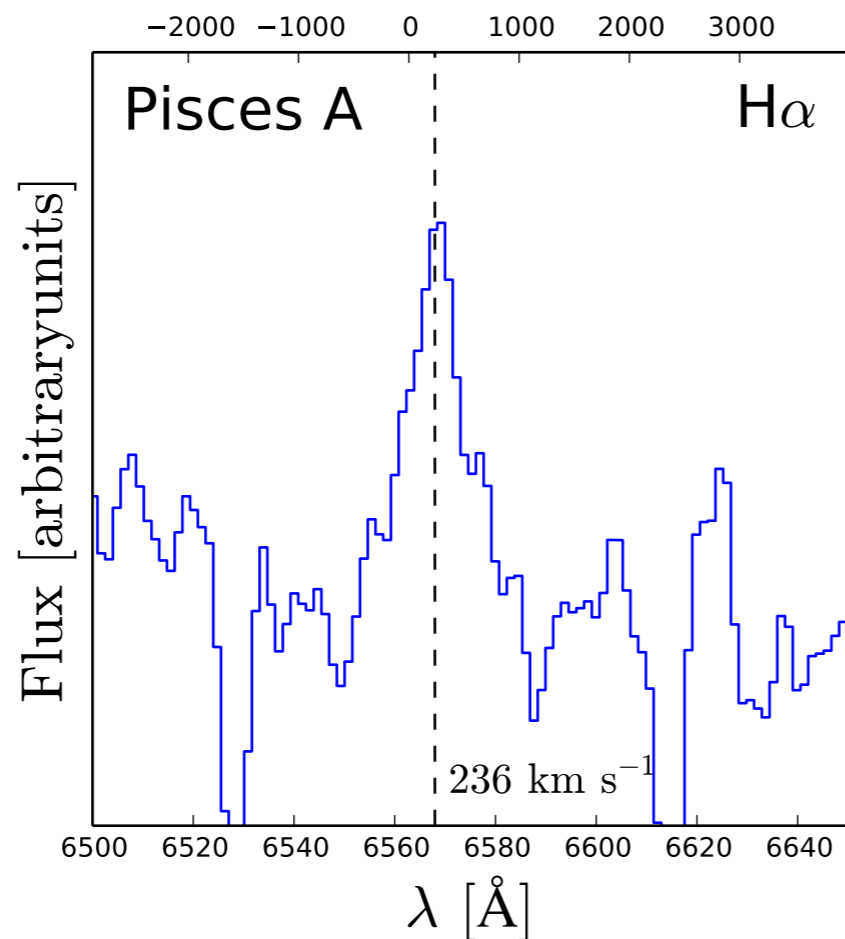


# The HI and H $\alpha$ velocities match

Radio

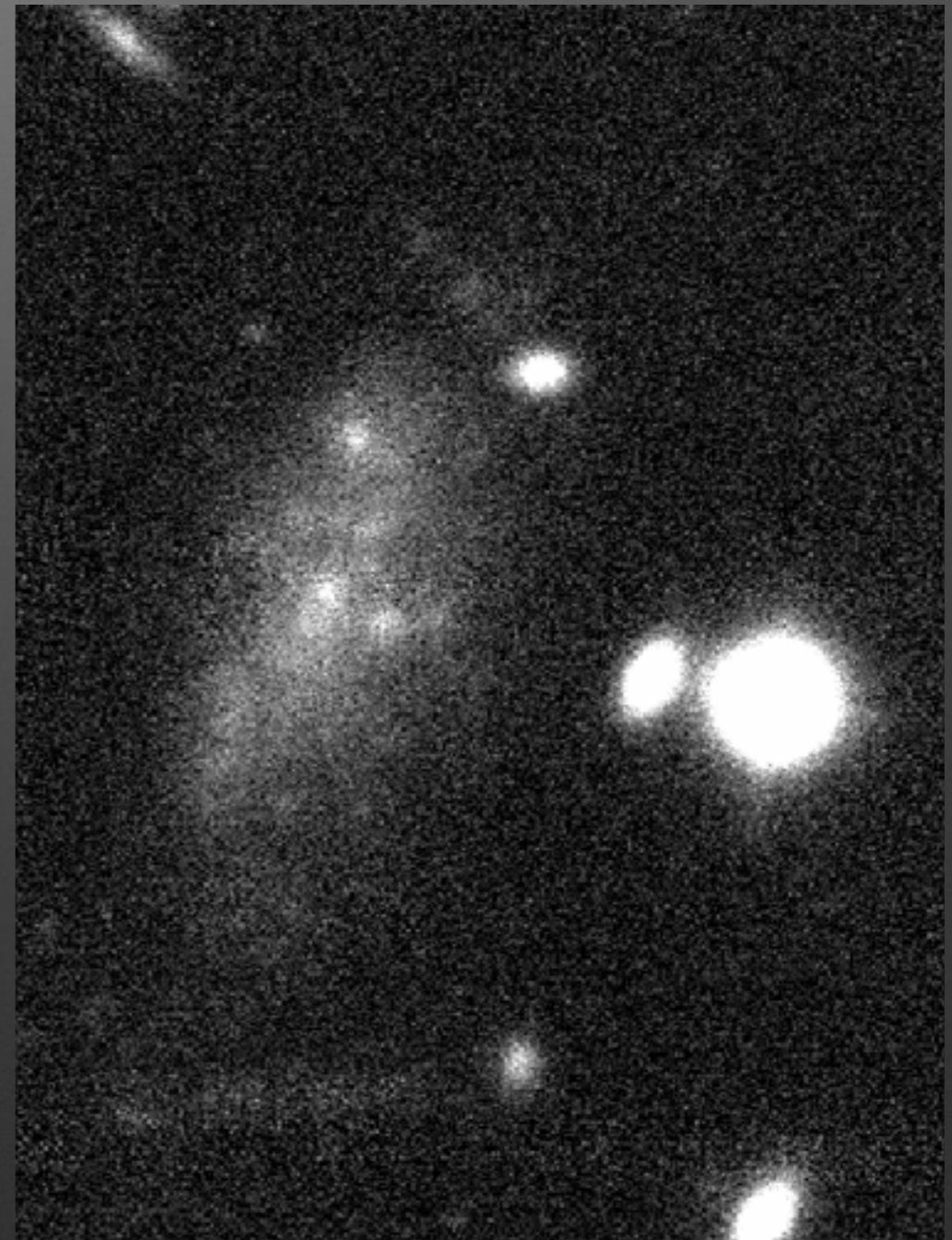
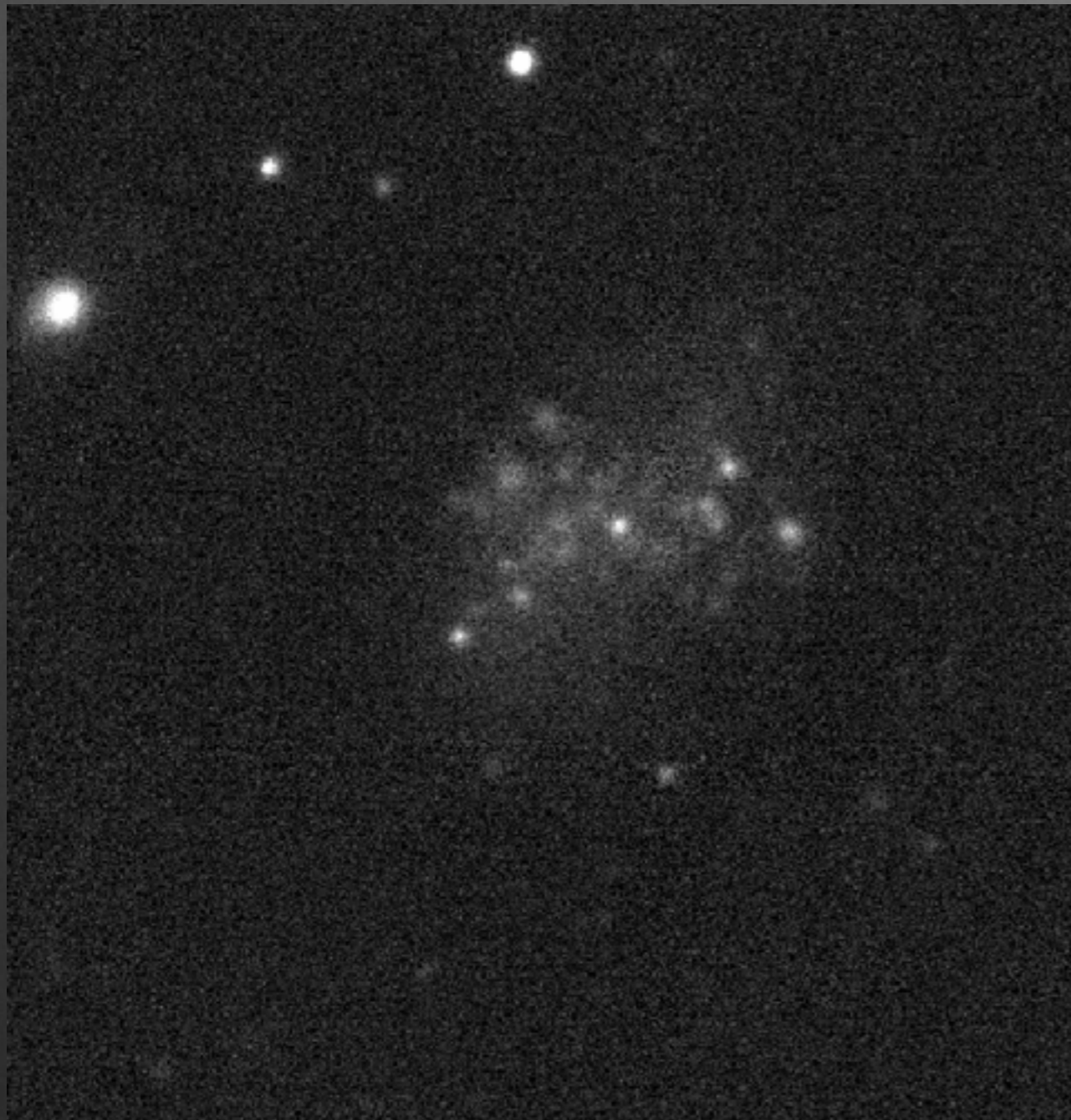


Optical



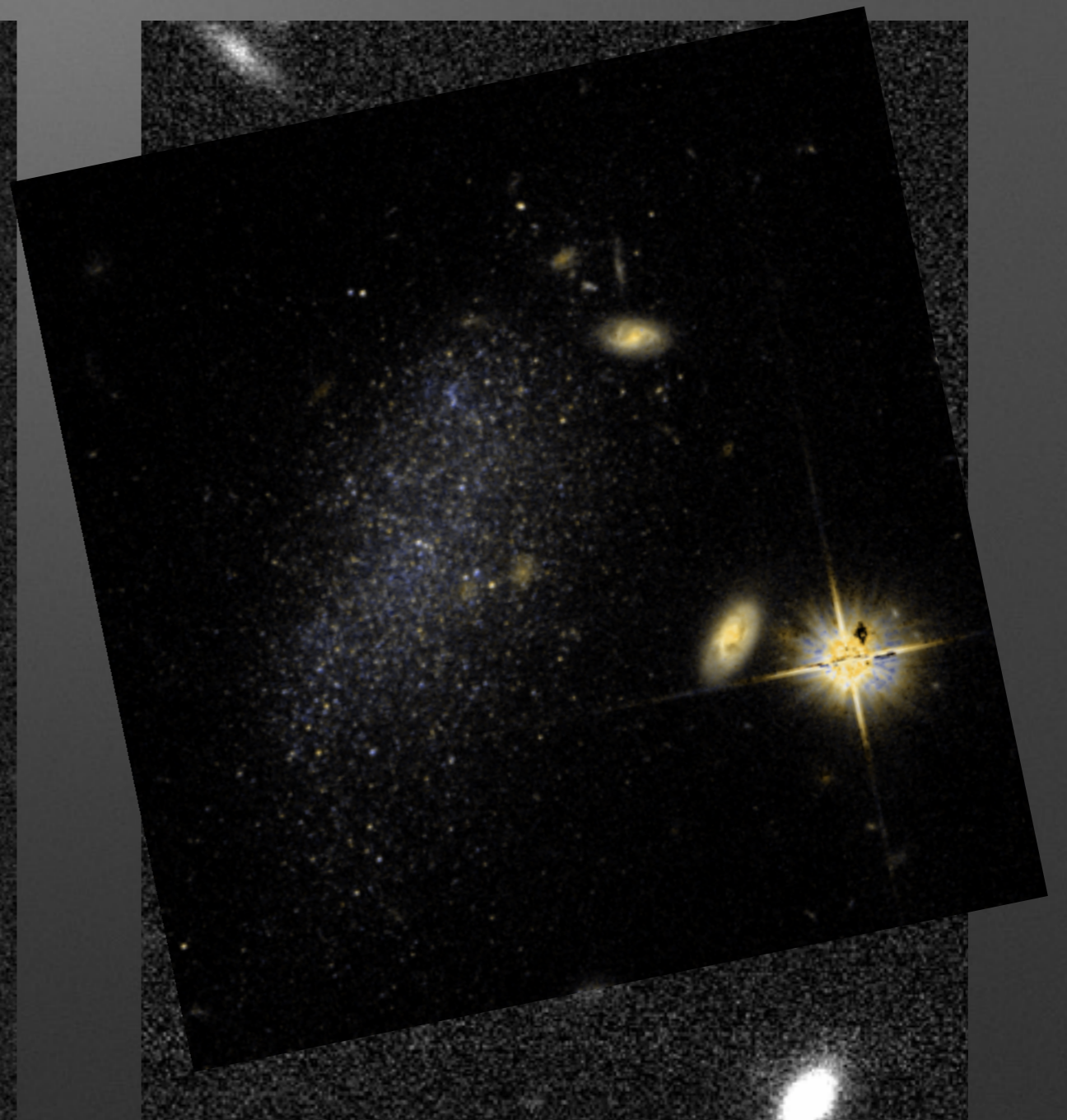
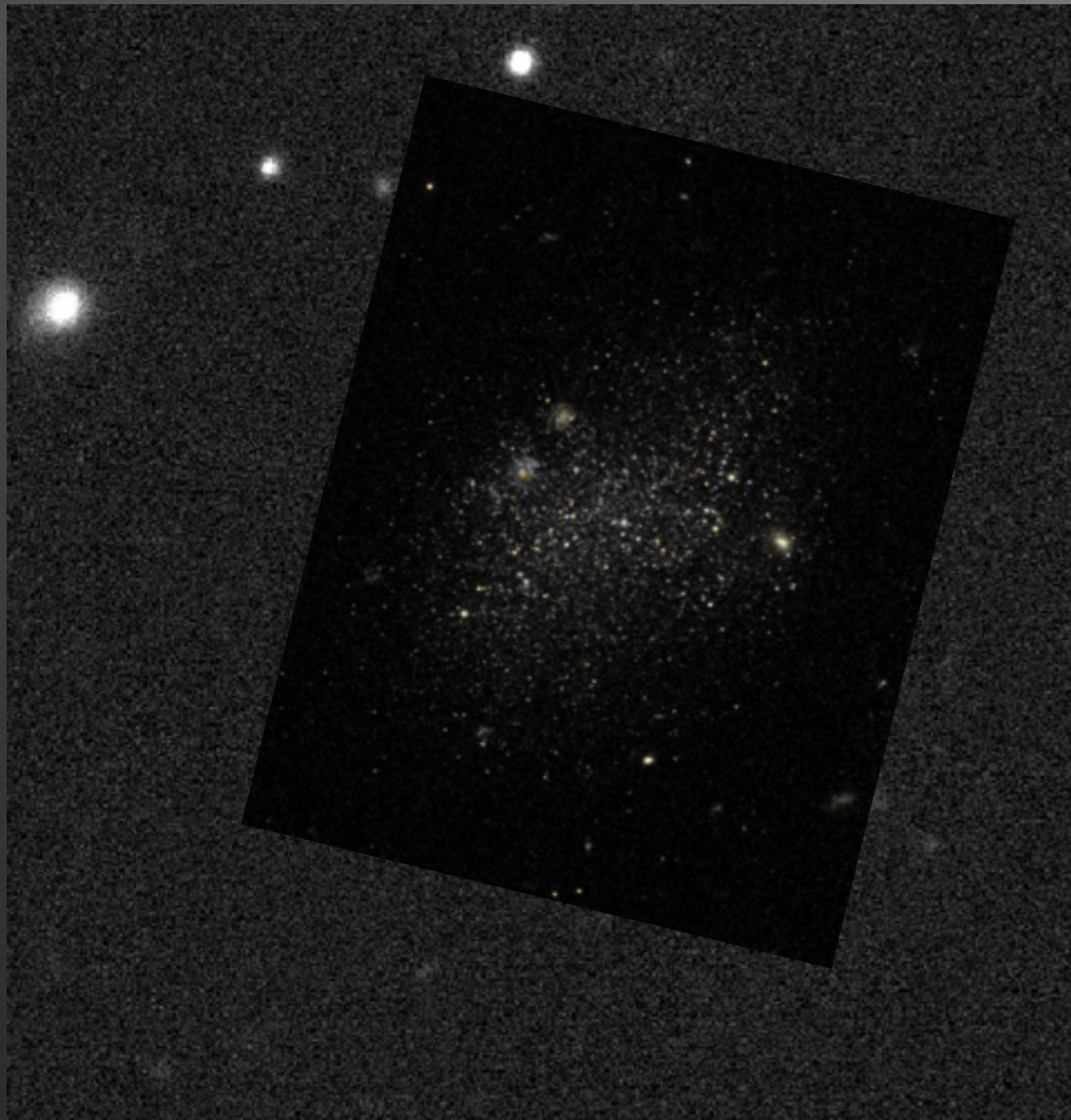


# HST imaging closes the case



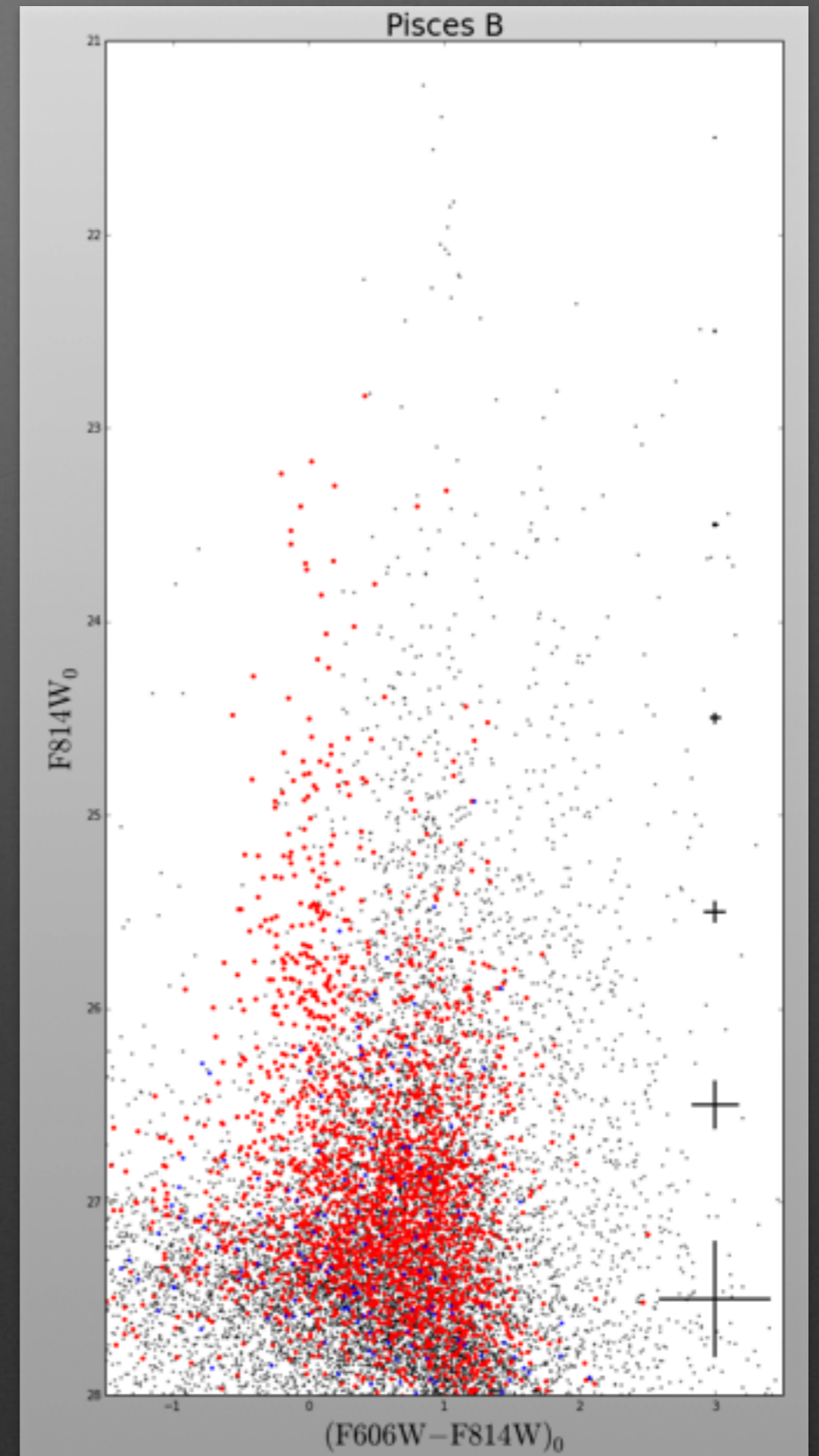
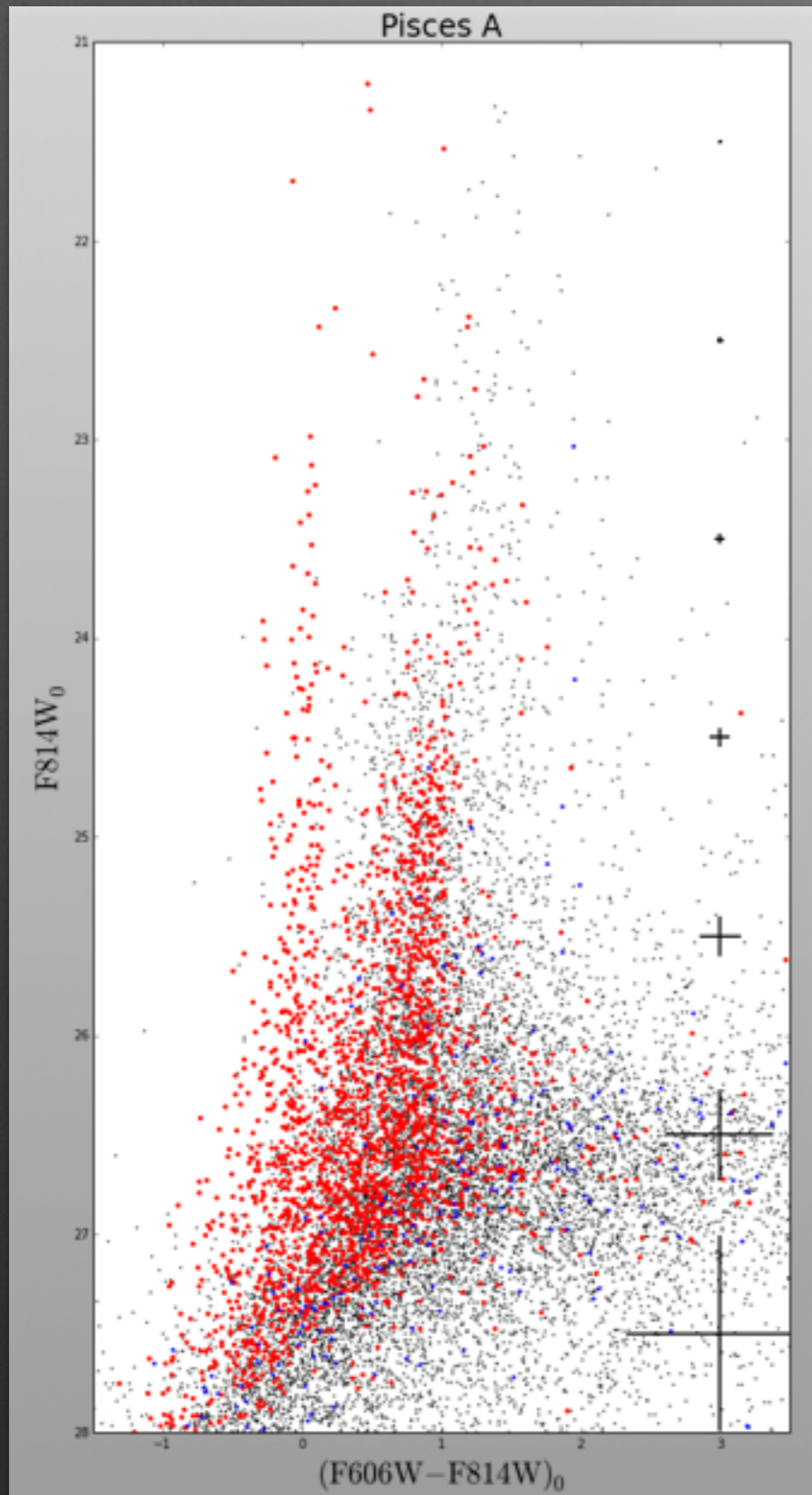


# HST imaging closes the case





# HST yields CMDs

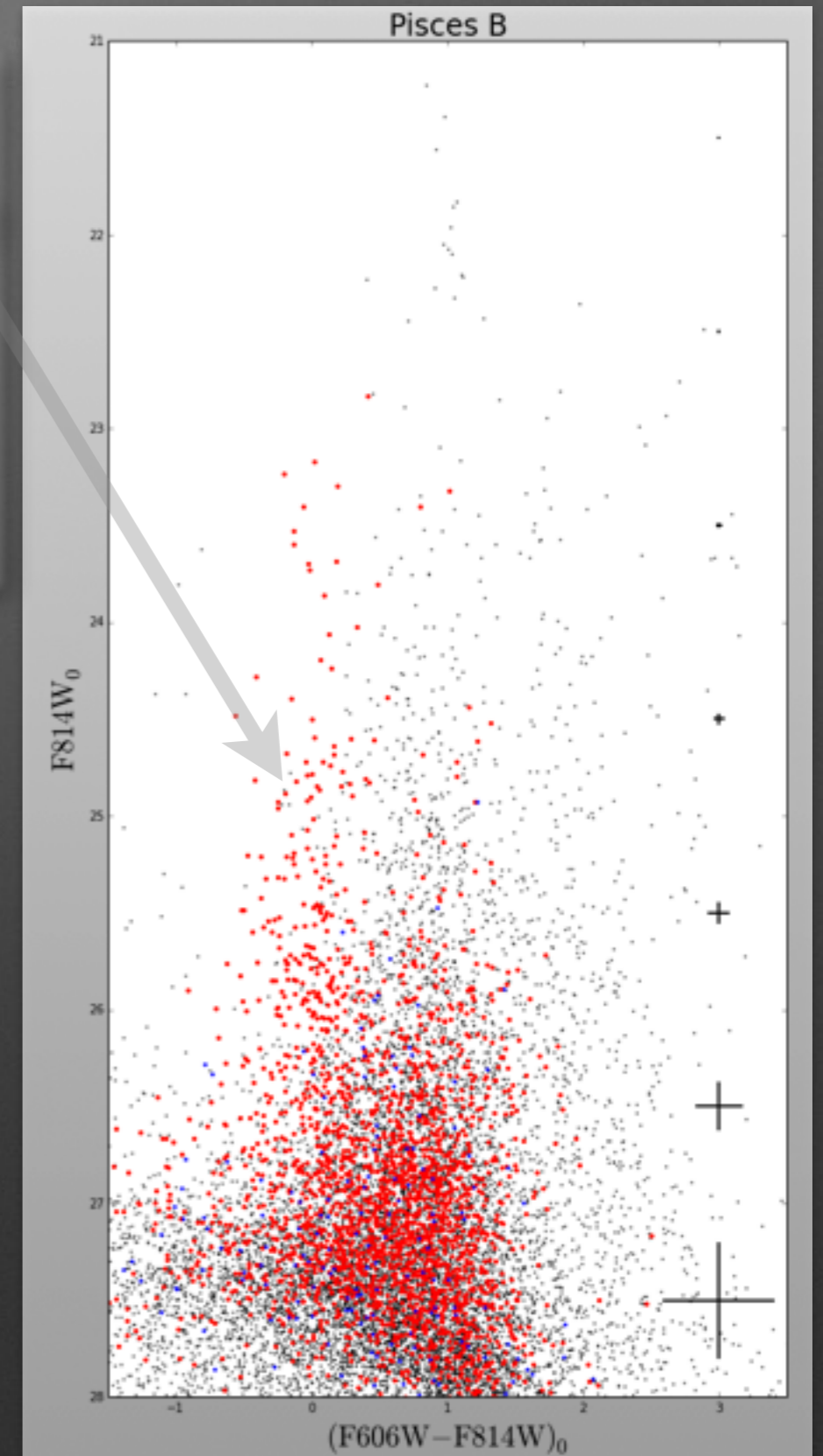
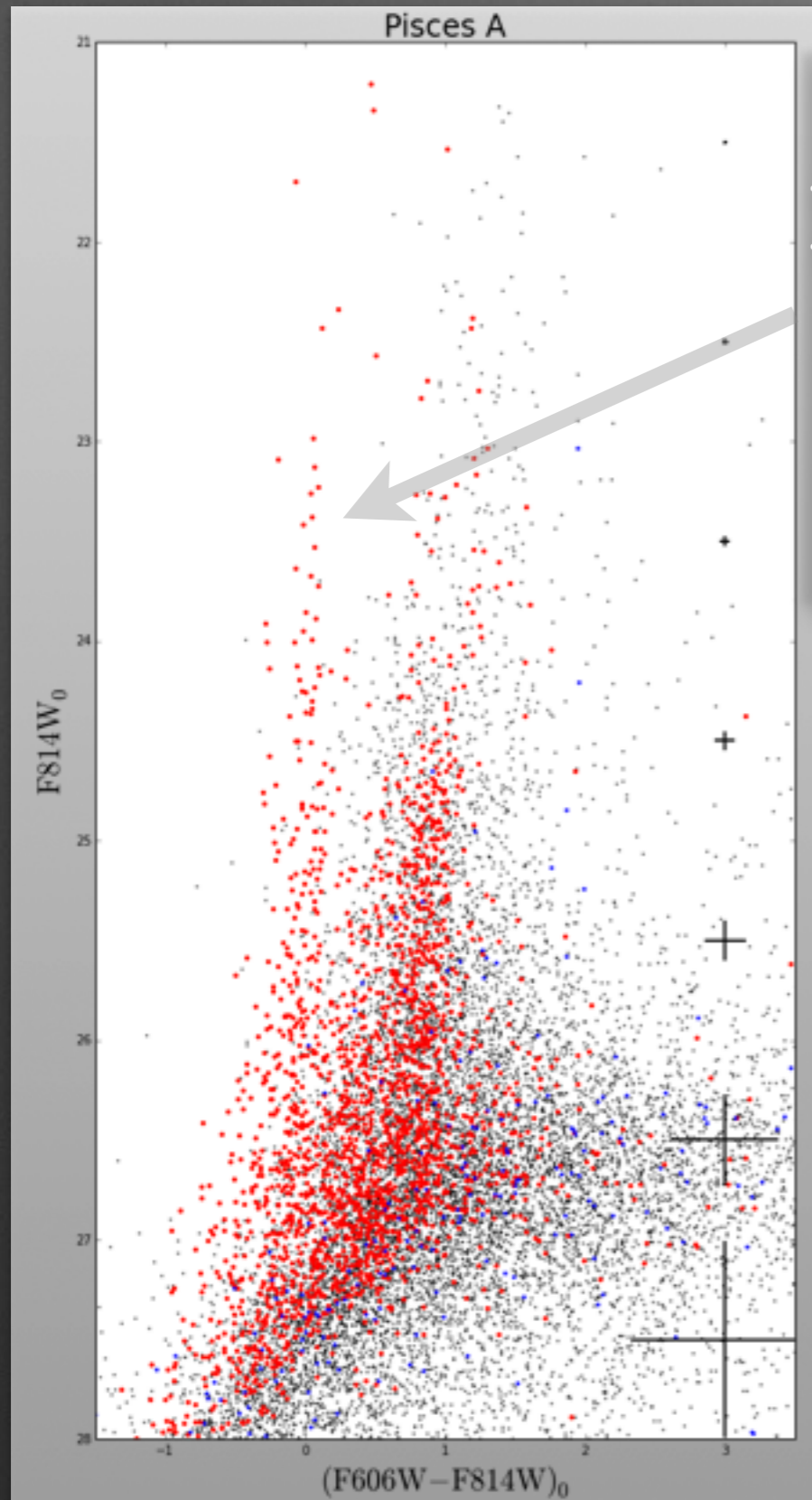




# CMDs show star formation,

## RGB

Strong  
Blue Plume  
= lots of  
light from  
recent SF

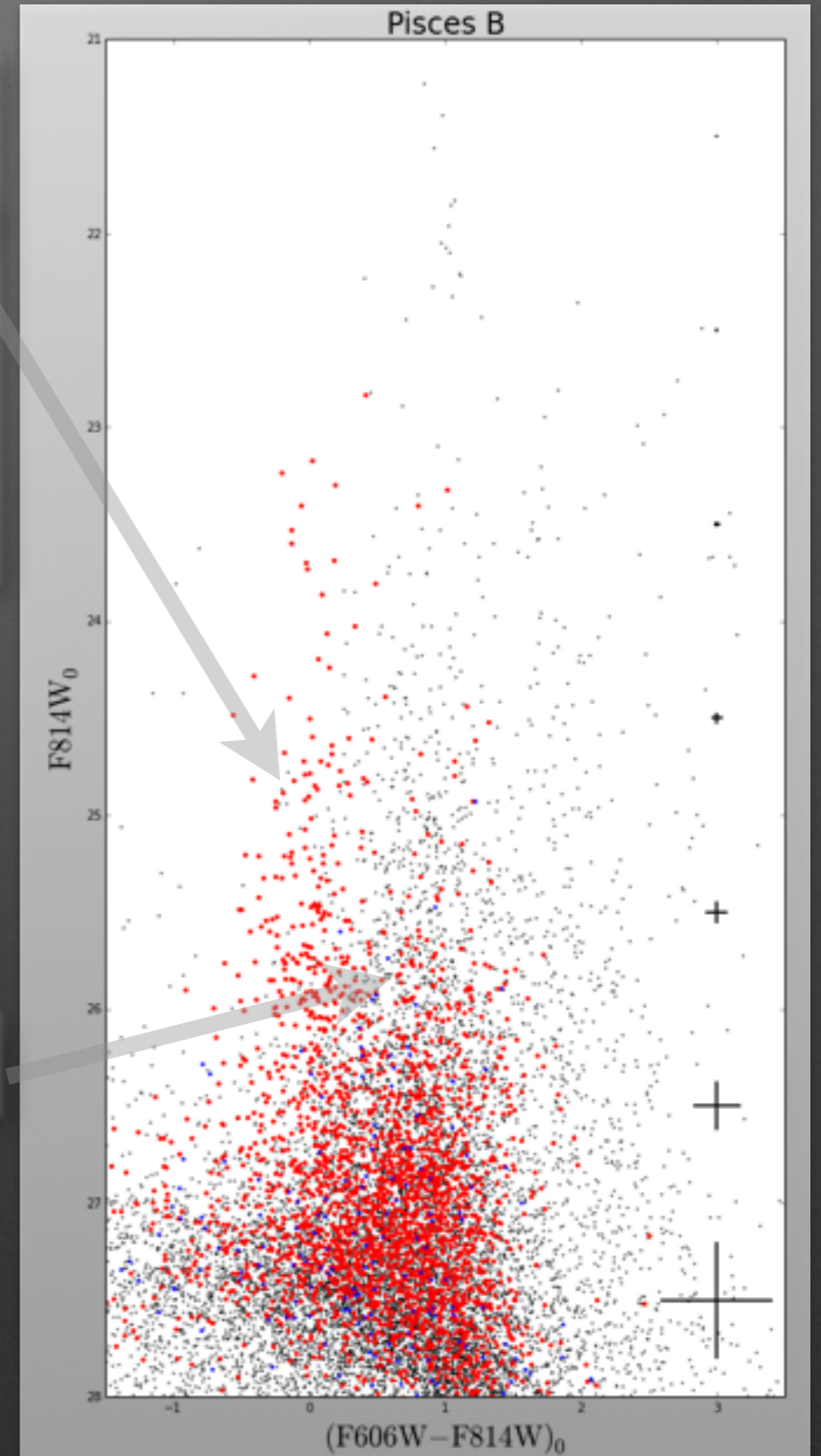
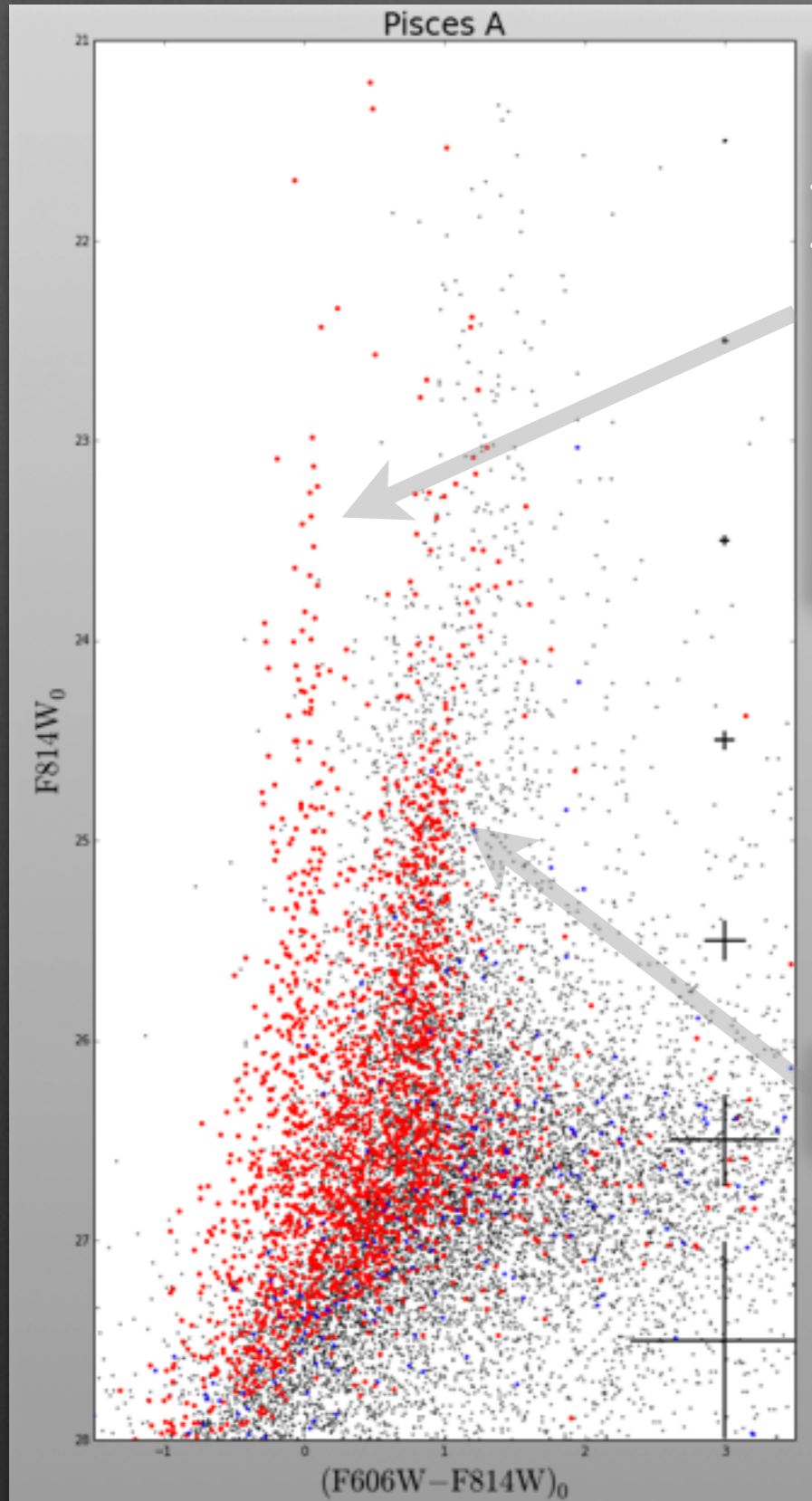




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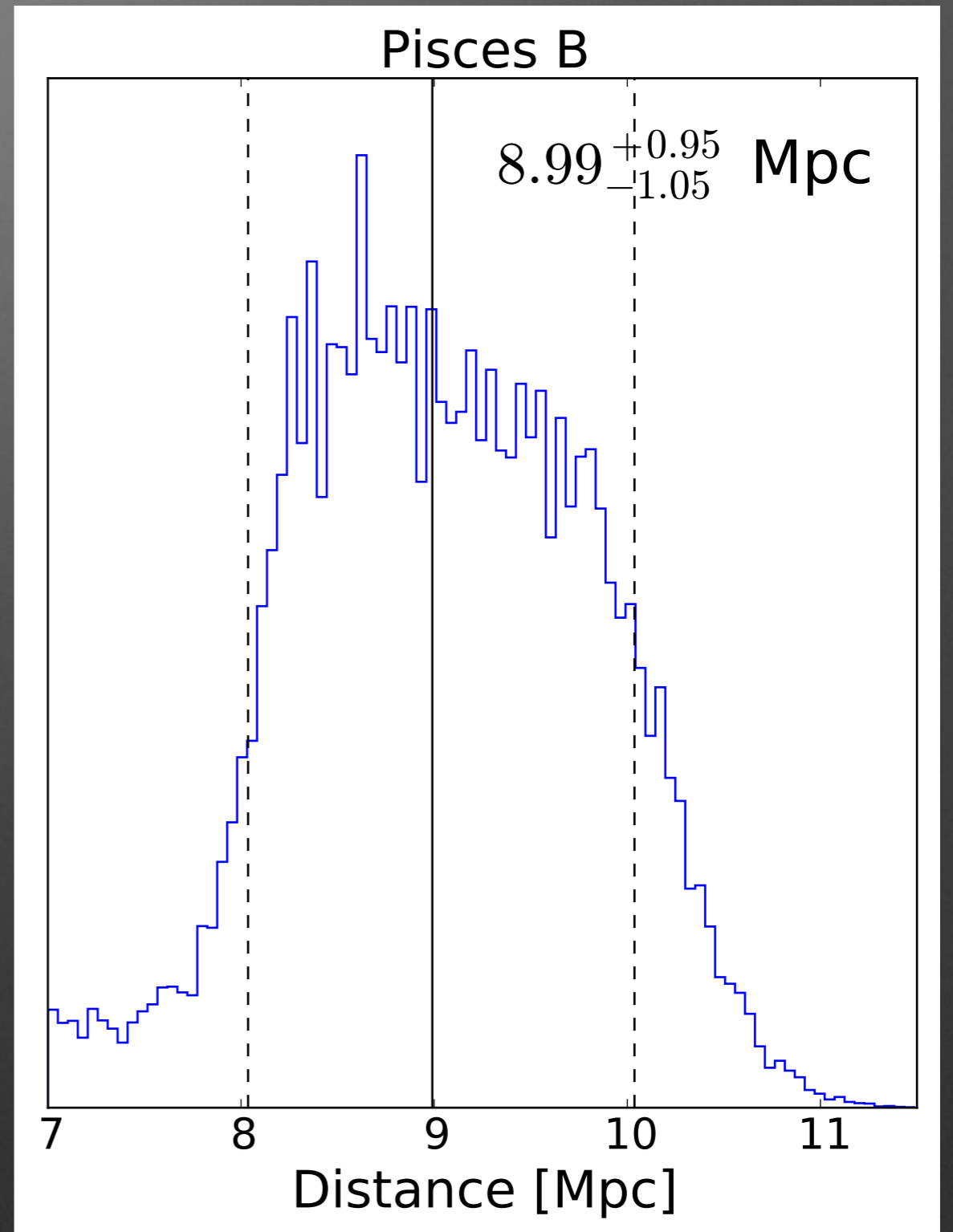
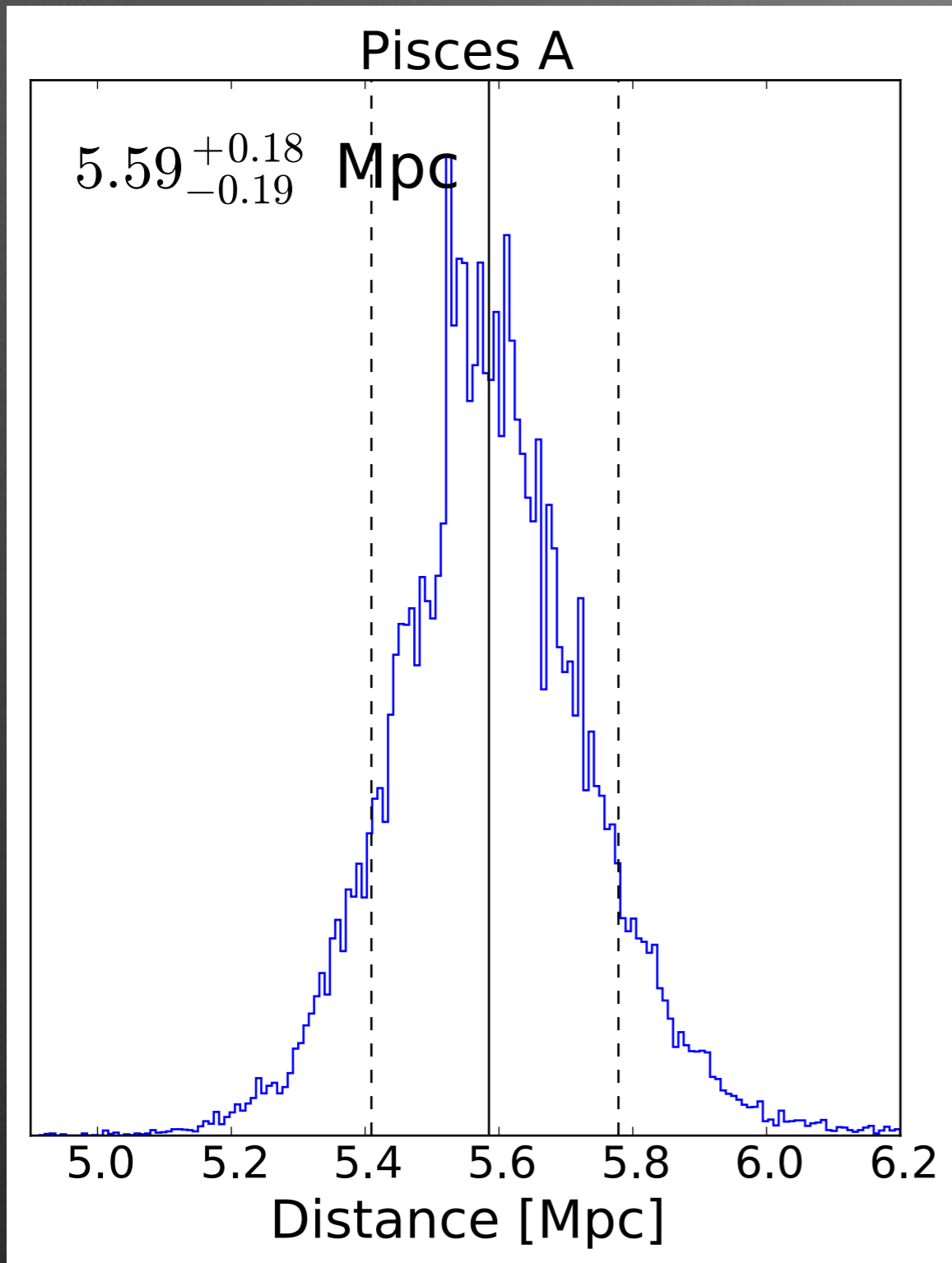
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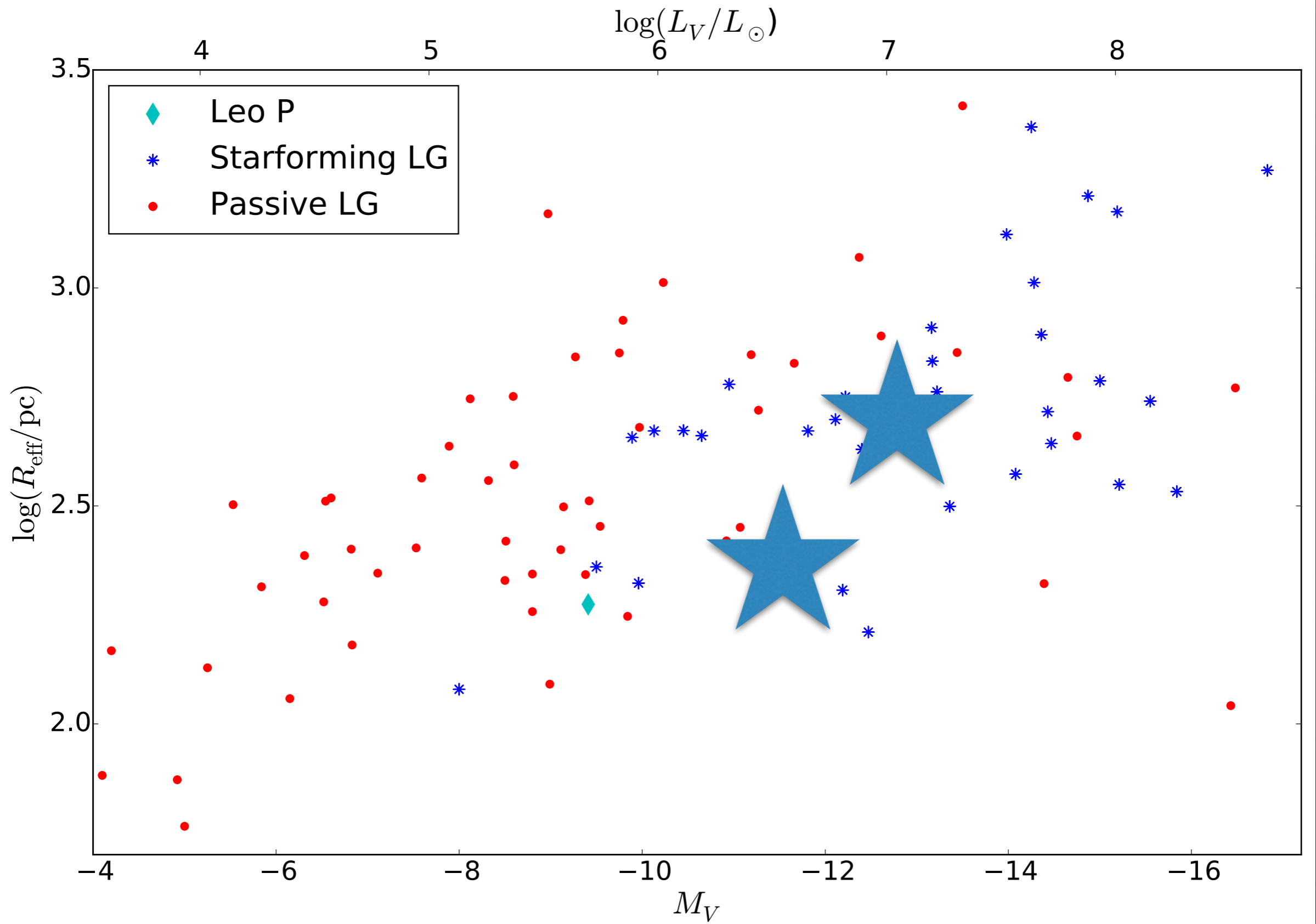
Clear RGB

# RGB yields distance

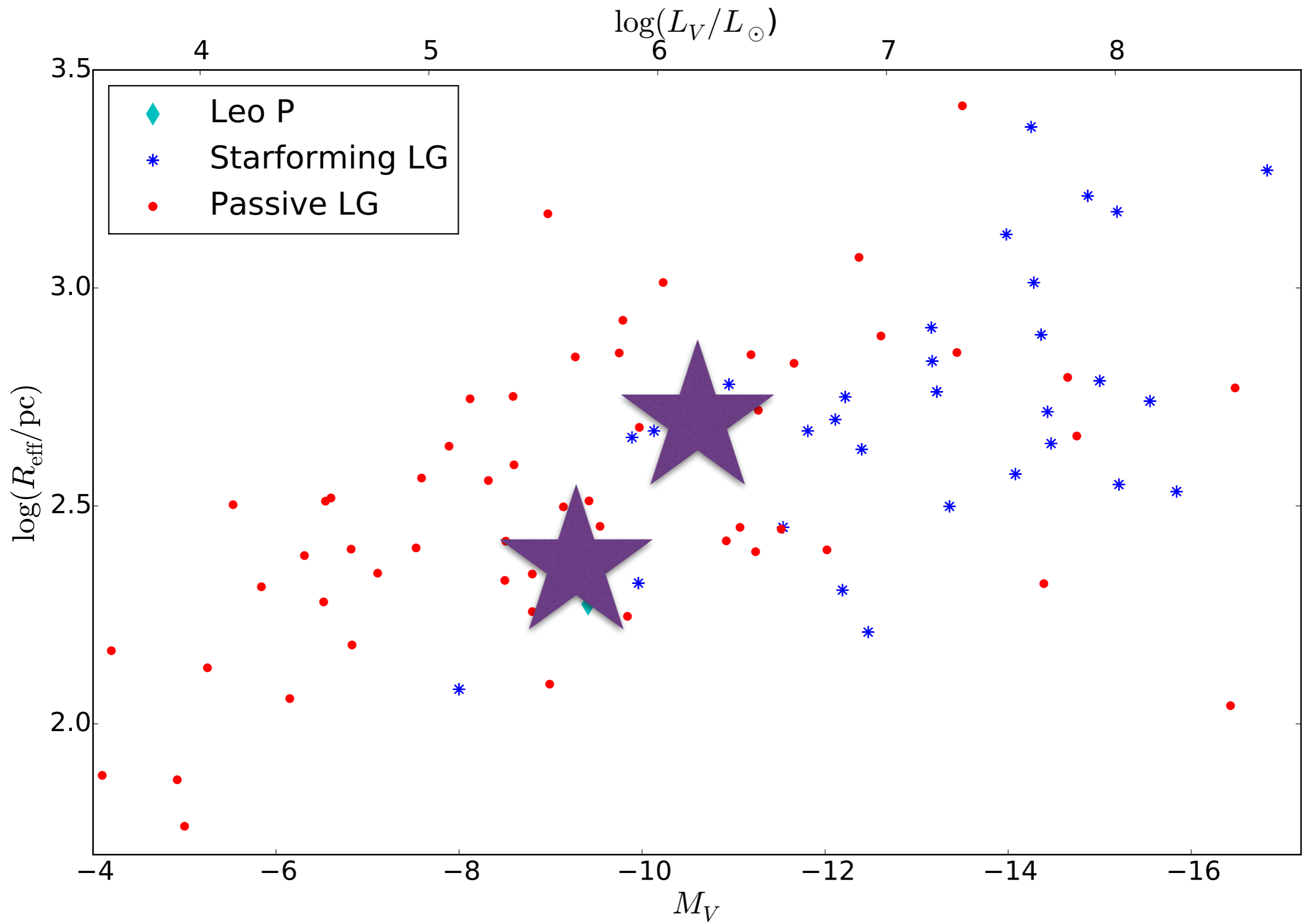




# Resemble SF LG dwarfs



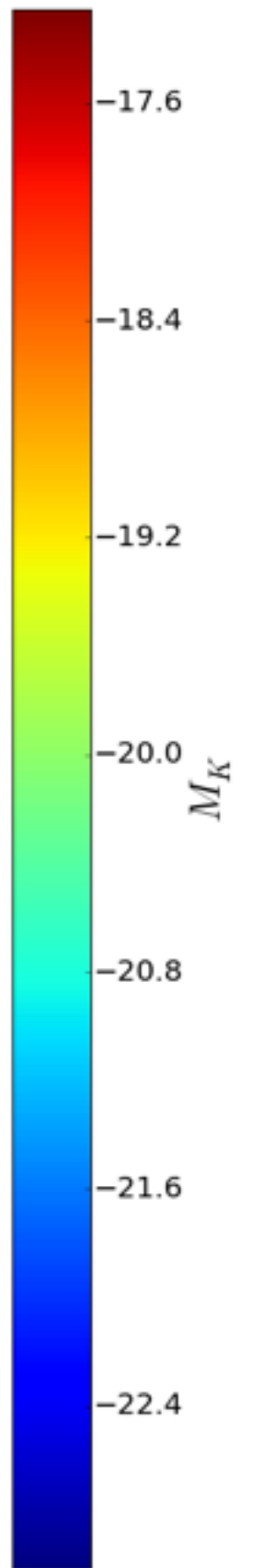
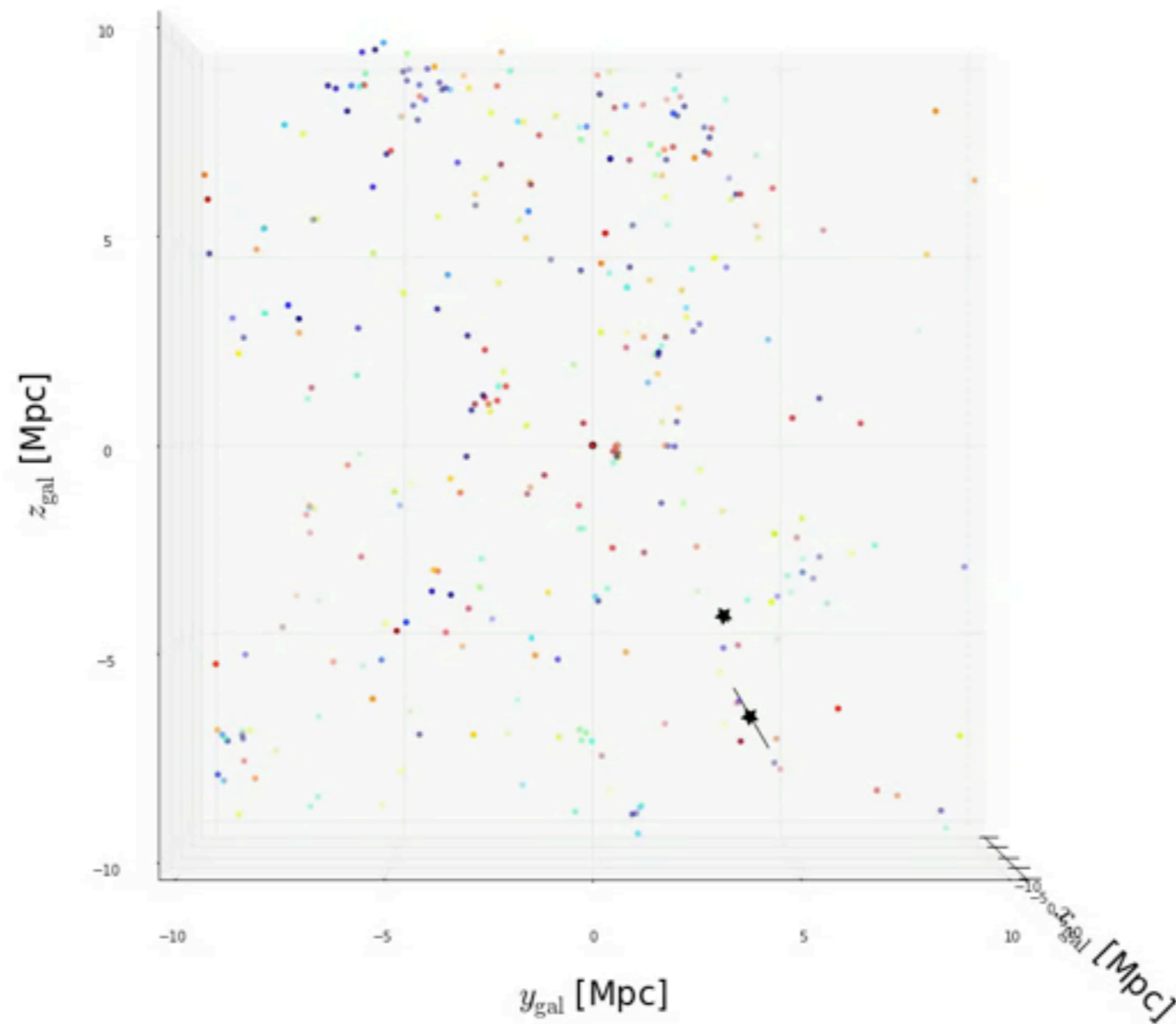
# Faded to passive, ~LG dSphs





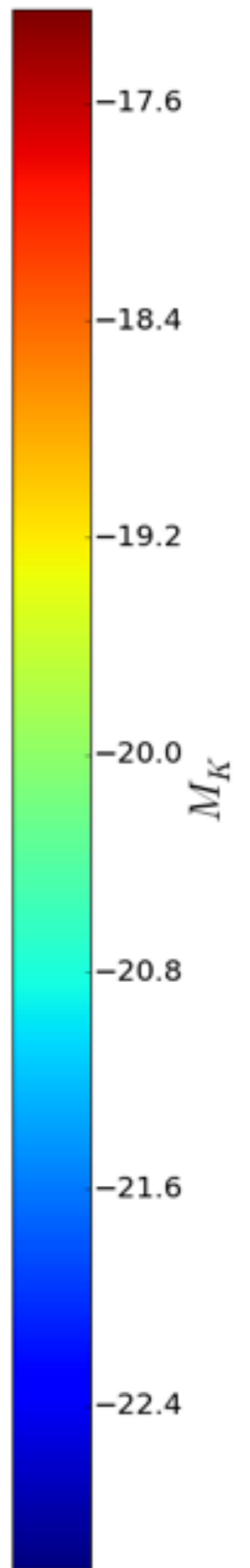
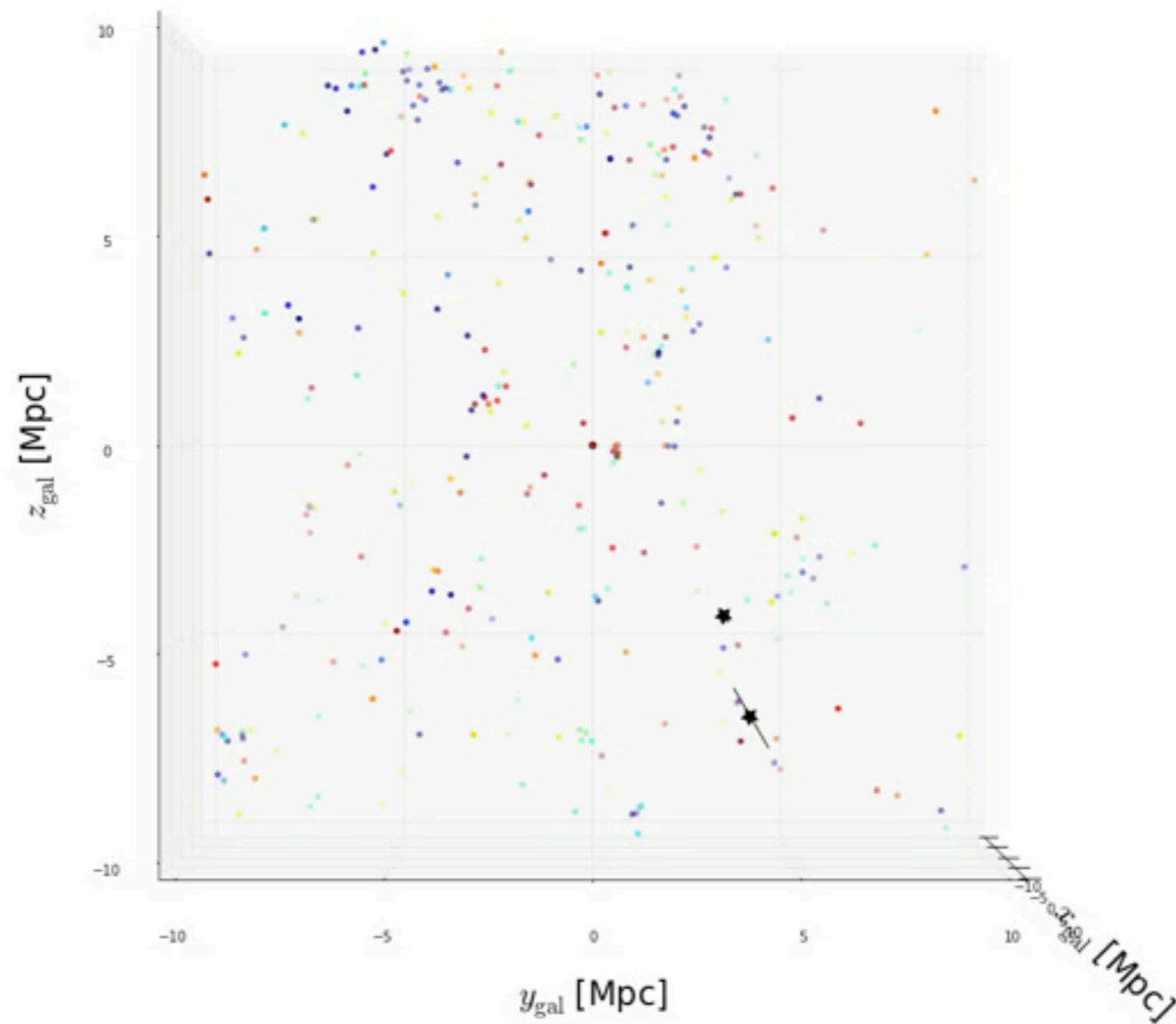
# At the edge of filament/void

Voids →  
delayed  
evolution



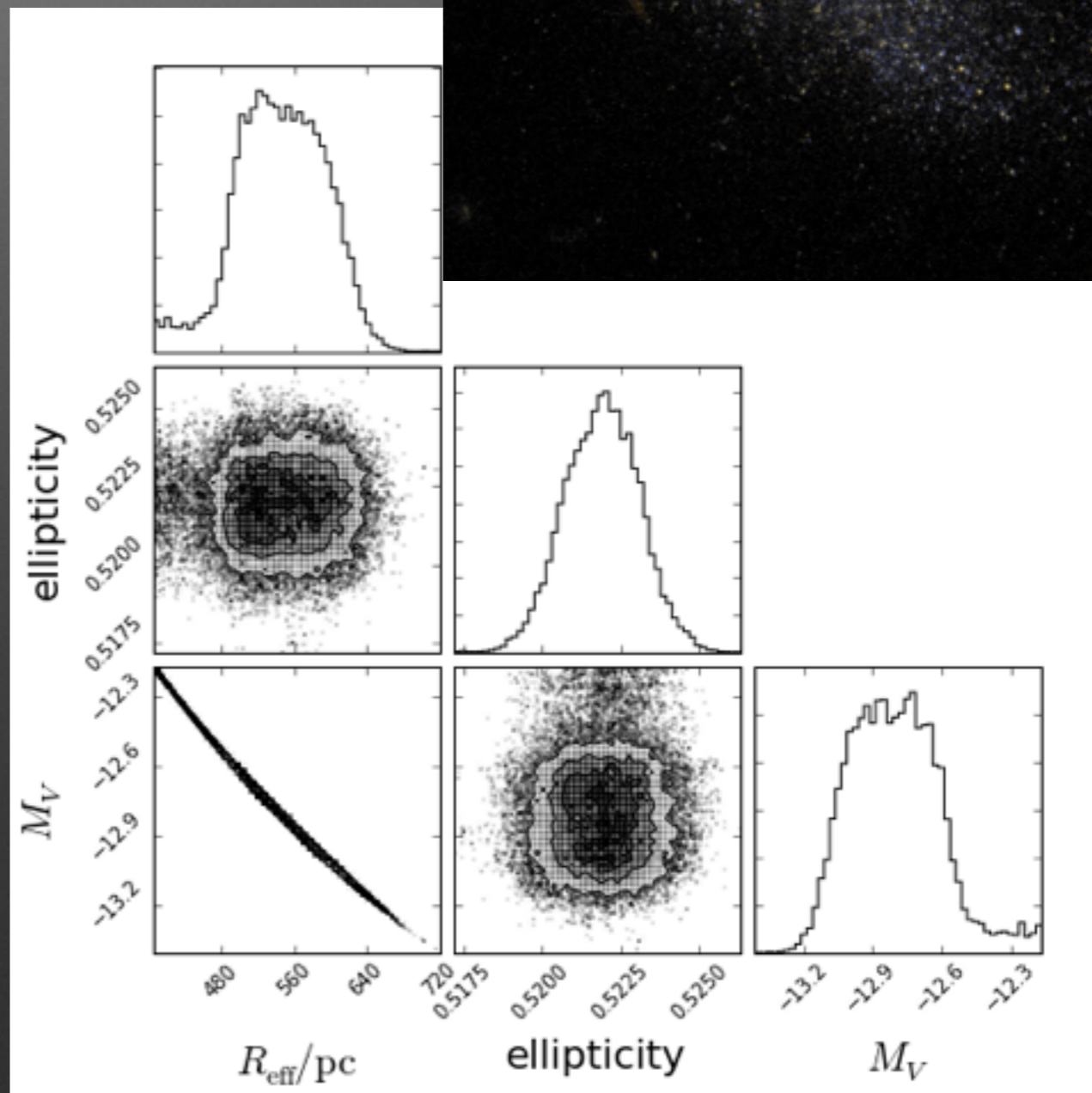
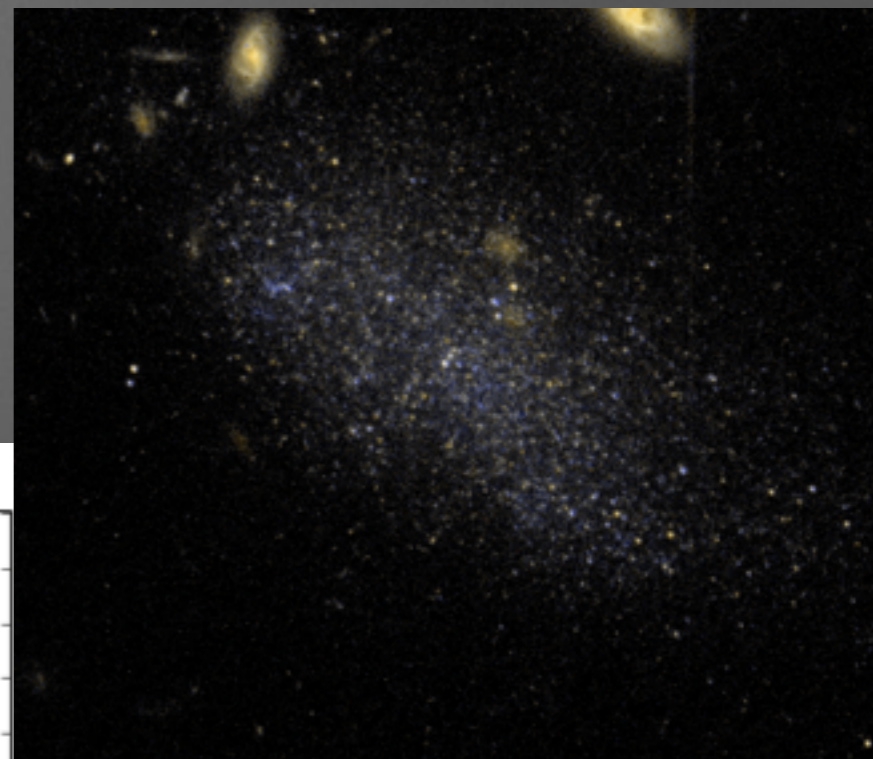
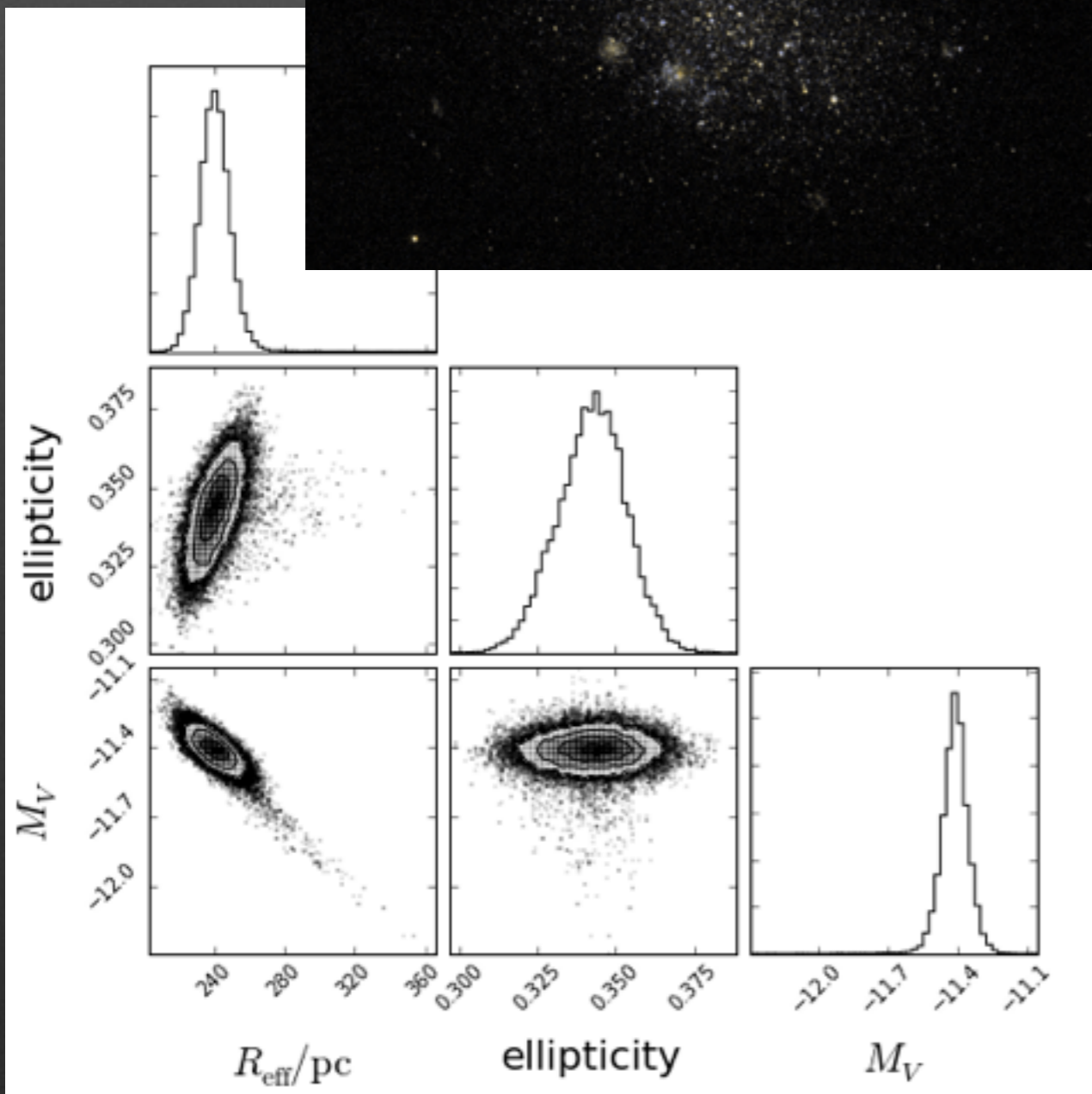
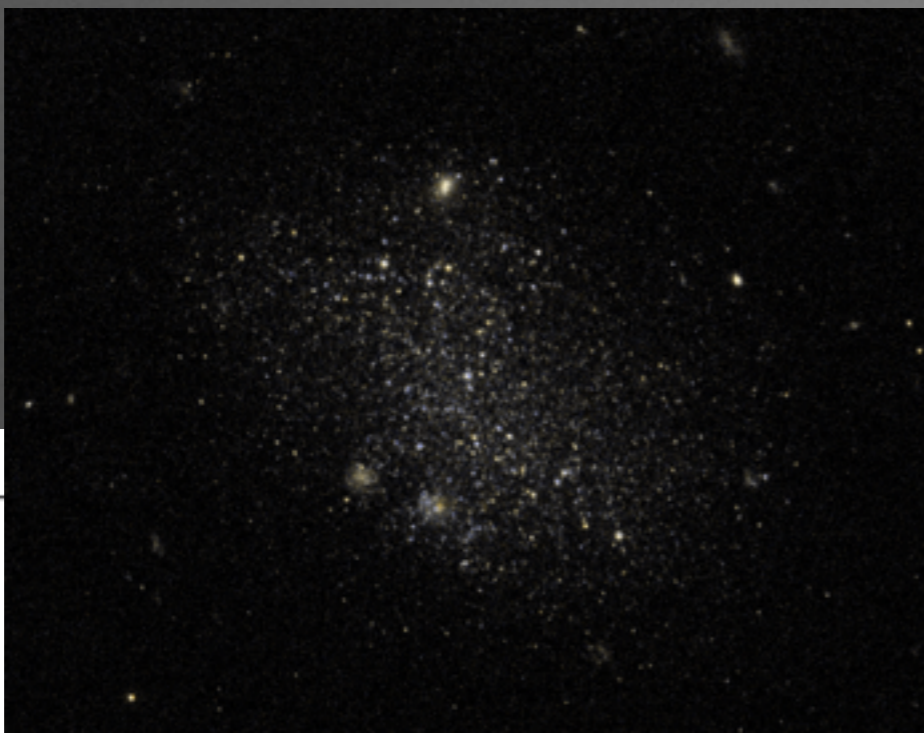
# At the edge of filament/void

Voids →  
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evolution





# These may be the best ICs?



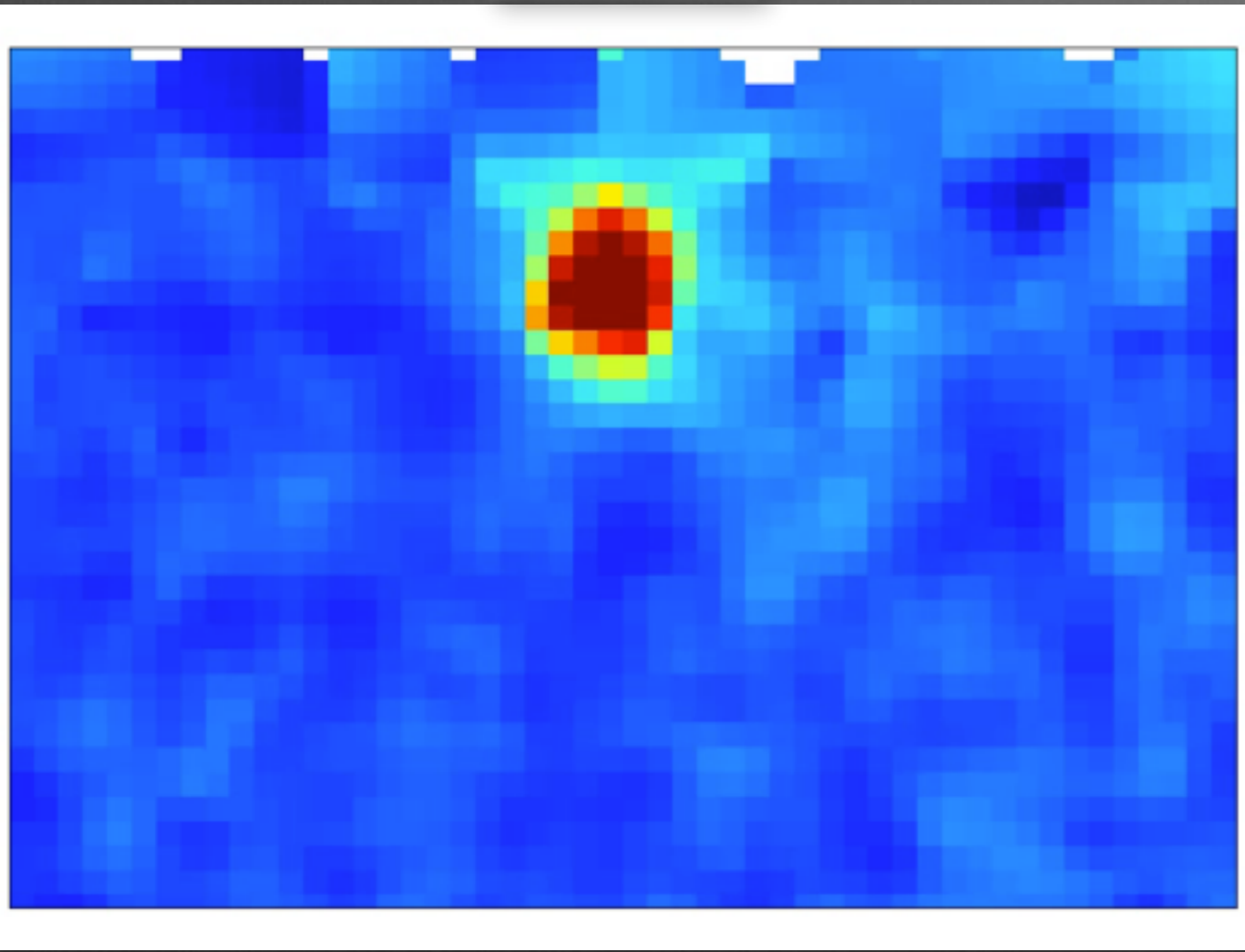
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# Is there hope for finding closer/fainter dwarfs in HI?

Leo T

Other Random Spot

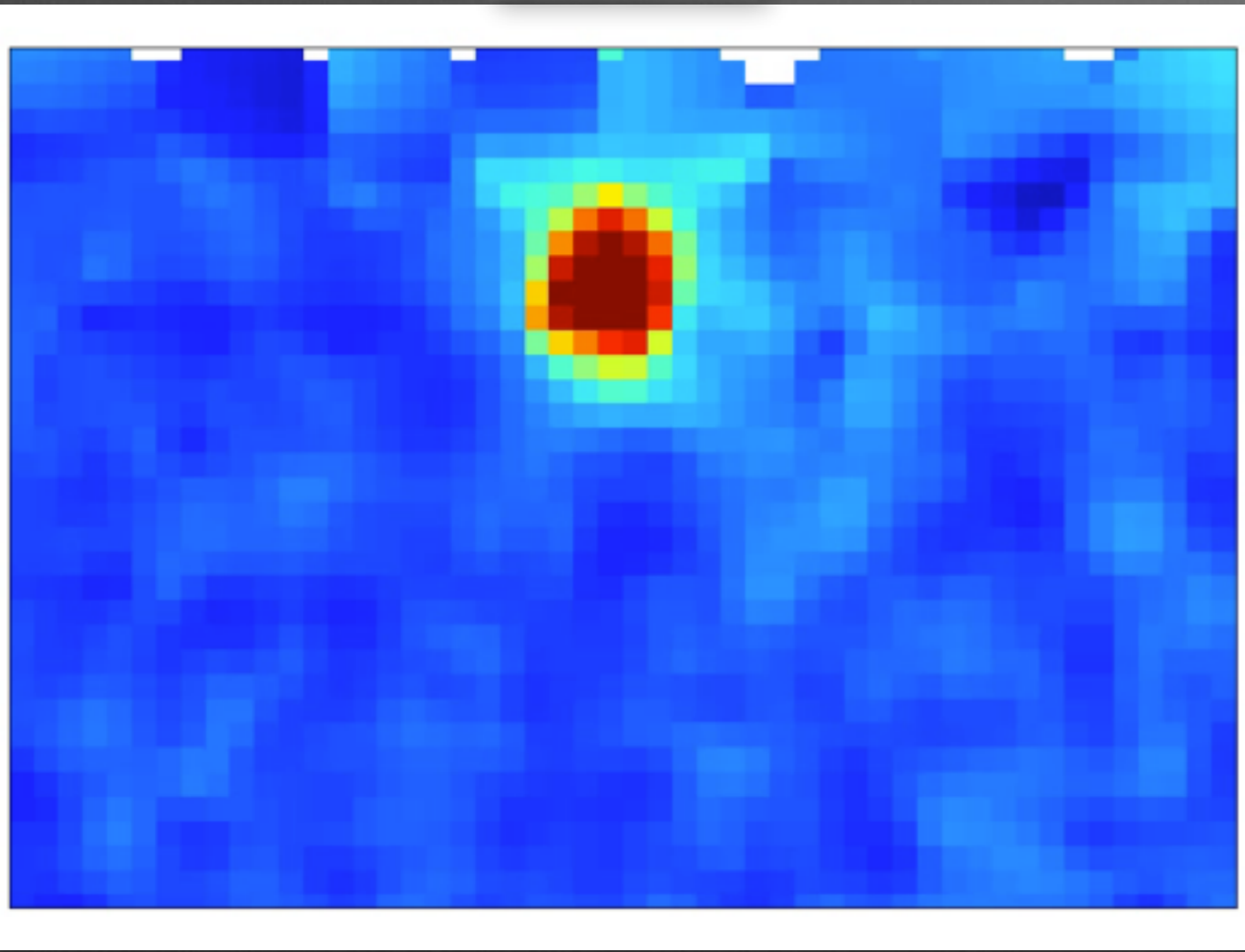


So yes, but need to go beyond current HI search techniques

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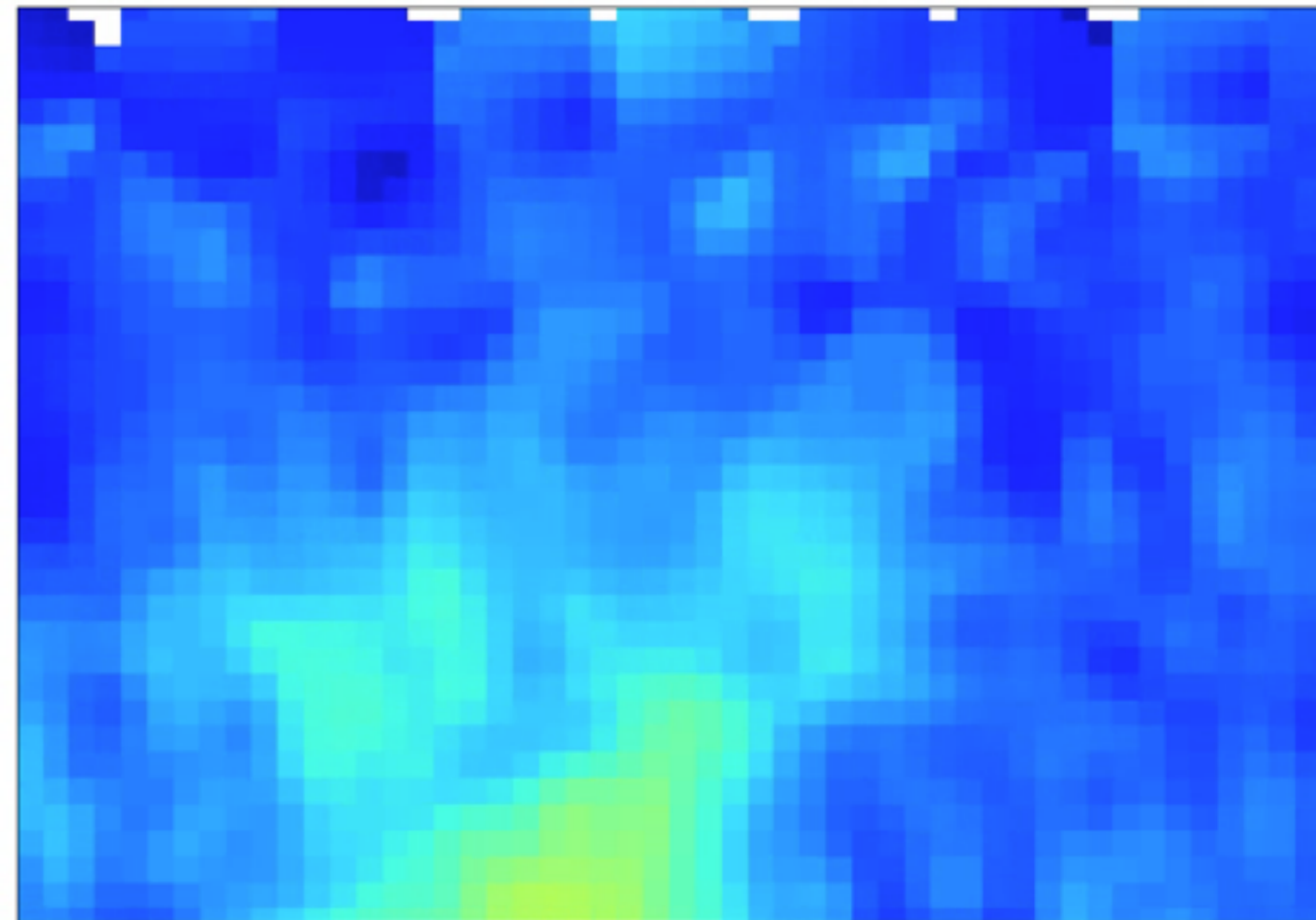
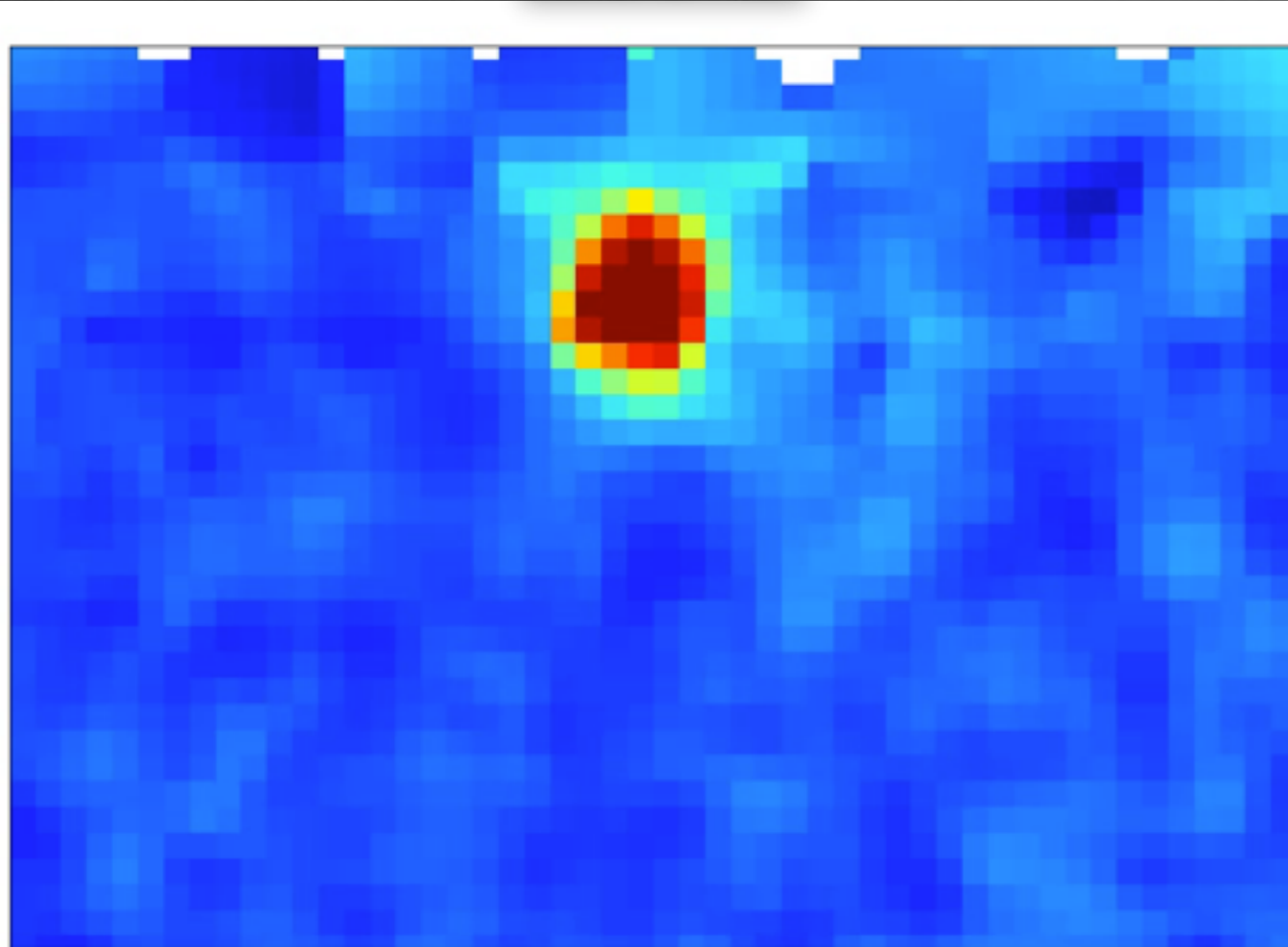
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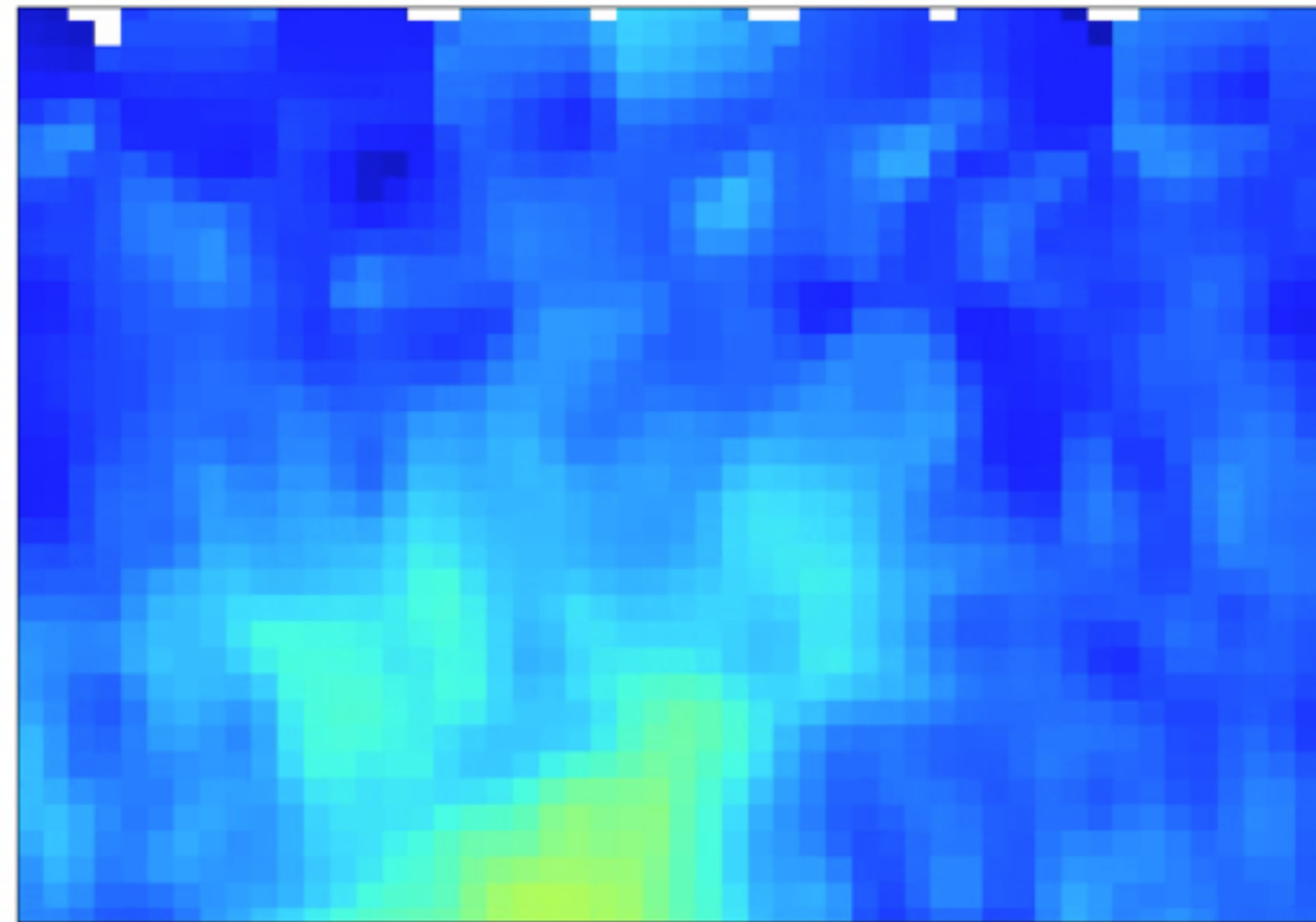
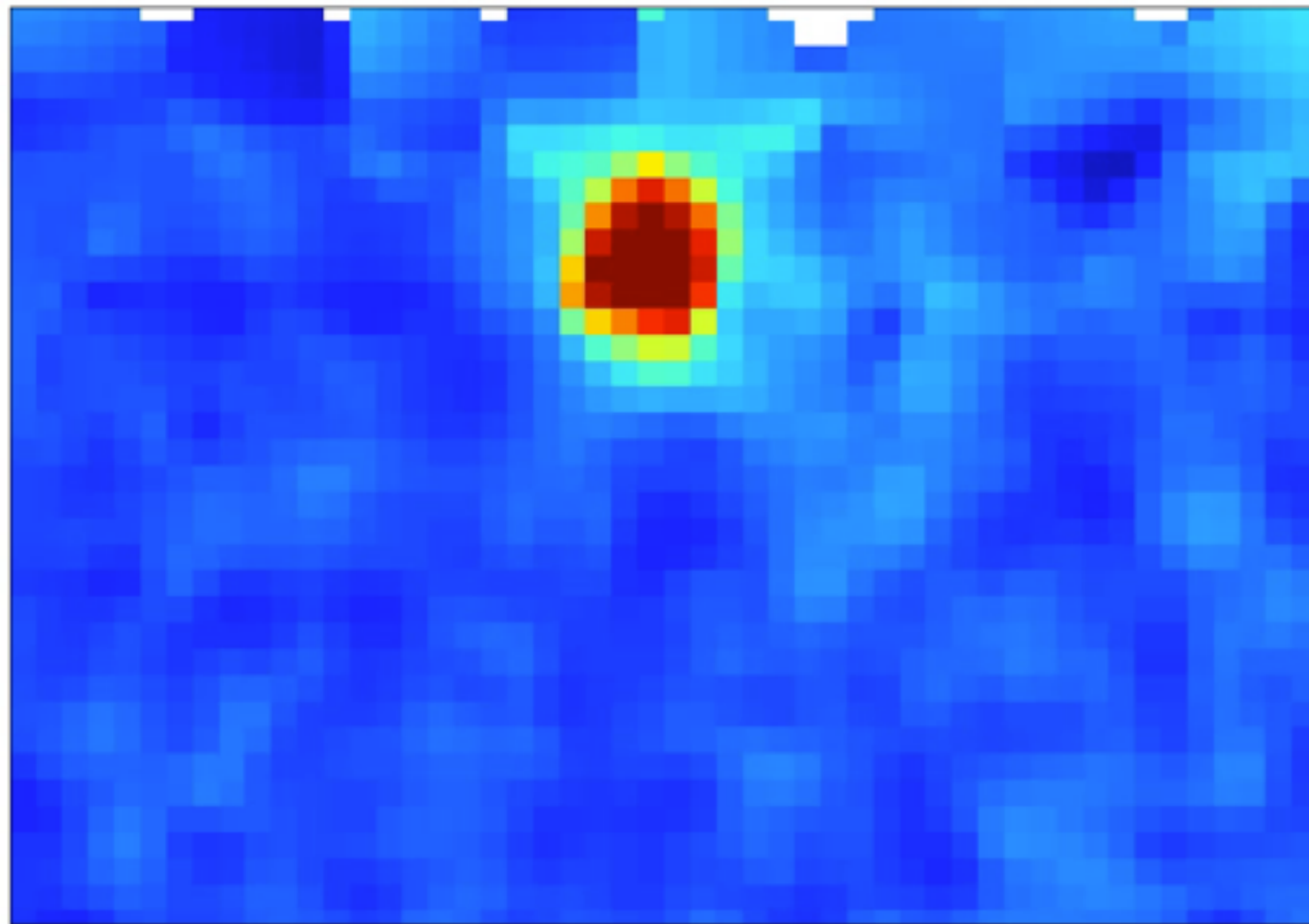


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So yes, but need to go beyond current HI search techniques



# Conclusions

- HI surveys can find dwarf galaxies comparable to Local Group satellites, but so far only in Local Volume
  - Such dwarfs provide context as possible progenitors of passive LG dwarfs
  - (Also begs the question: what are the HI clouds *without* optical counterparts?)
- Could be plenty more to be found, but need to separate from the Galactic foregrounds