

Stellar systems in high density environments

> NGC 6744

> > NGC

ESO Garching, Germany 27 June-01 July 2011



Virgo III Groups



Galaxy Formation and Evolution

- Hierarchical accretion/merging
 - Matter clumps through gravitation
 - Primordial gas starts forming first stars
 - Stars produce heavier elements ('metals')
 - Subsequent generations of stars: more metals
 - Massive galaxies assembled from smaller units
 - Galaxy encounters still occur
 Minor accretion to major merging
 - Galaxies continue to evolve



Super-massive central black hole also influences evolution

Observational Approaches

Study very distant galaxies > Observe evolution (far away = long ago)

Objects faint and small: little information

- Study nearby galaxies
 - Light not resolved in individual stars
 - Objects large & bright: structure accessible
 - Infer evolution through archaeology
 - Fossil record is cleanest in early-type galaxies

Study resolved stellar populations
Ages, metallicities and motions of stars
Archaeology of Galaxy and its neighbors









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This Conference

What can we learn from nearby clusters

- Star formation history in massive galaxies
- Dependence on local density
- Role of the central black hole
- Observations
 - Multi-wavelength imaging
 - Integral-field spectroscopy
- Theoretical models
 - Match predictions to observations

Forward look

Opportunities for VLT and E-ELT



E-ELT

Site selected and access guaranteed

Design essentially ready

- Resizing to 39.3m reduces risk and contains cost
- Retain adaptive design and unique exoplanet science
- Can be built in 10-11 years: opportunity to be first
- Viable funding model
 - Entire programme affordable for 15 MS if baseline funding scenario is implemented
 - More and more MS ready to commit
- ESO internally ready for construction by 1 Oct
- Everything on track for green light in Dec 2011

