

Figure 1: Globular cluster: 47 Tuc, V, very crowded field with charge transfer problems; 15 sec. CCD (ESO No. 1) exposure taken with the Danish 1.5 m at La Silla. (S. Ortolani, Padua Observatory.)

ceedings will appear later in 1989. Results of the analyses are also contained in the form of catalogues and other files in directory MISC\$DISK:[TESTIMAGES] on ESO's VAX system. Together with the test images themselves, this information will be made available to the general public on magnetic tape by the authors of this report. Copies of the tape can be obtained upon request.

Two of the test images are shown in the figures. They will be of particular interest to those investigating globular clusters and faint galaxy fields. Images with different exposure times are included. A few further images will be included in the future. These include simulated Hubble Space Telescope Wide Field/Planetary Camera (WF/PC) images — a typical globular cluster at distance 1 Mpc with an exposure of 300 seconds as seen using different filters.

The advantages of using this set of test images are two-fold. They provide a set of results which can be used for benchmarking future software packages and alternative algorithms devised by anyone working in this field. They also



Figure 2: Faint galaxy field: SA 68, contains very faint BV standards; 75 min. J. plate taken with Kitt Peak National Observatory 4 m. (D. Koo, Lick Observatory.)

provide a common point of interest for those who develop 2-dimensional software for stellar and galaxy photometry. One hopes that such systems can be used as building-blocks for systems of the future rather than the all-too-common restarting from first principles.

TEX and MATHOR3-TEX for Astronomy and Astrophysics Journal and Supplement Series

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H.-U. DANIEL

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J. BERGER and D. SAVARY

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For all subsequent issues authors themselves may send submissions to the series in either MATHOR3 or T_EX format. Les Editions de Physique will provide the appropriate "Supplements Series Macro Package" to any authors who wish to use plain T_EX for the preparation of their articles. Additionally, a special version of the MATHOR3-T_EX software package is now available to help authors and their staff who are not "T_EXperts".

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ESO, the European Southern Observaand operate an astronomical observatory in the southern hemisphere, equipaim of furthering and organizing collaboration in astronomy.... It is supported by eight countries: Belgium, Denmark, France, the Federal Republic of Germany, Italy, the Netherlands, Sweden and Switzerland. It operates the La Silla observatory in the Atacama desert, 600 km north of Santiago de Chile, at 2,400 m scopes with diameters up to 3.6 m and a New Technology Telescope (NTT) will become operational soon and a giant telescope (VLT=Very Large Telescope), consisting of four 8-m telescopes (equivalent aperture = 16 m) is tists make proposals each year for the use of the telescopes at La Silla. The ESO Headquarters are located in Garching, near Munich, FRG. It is the scientifictechnical and administrative centre of ESO, where technical development pro-La Silla observatory with the most advanced instruments. There are also rope ESO employs about 150 international Staff members, Fellows and Assotion, 150 local Staff members.

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