

Novel methods to probe exoplanet atmospheres using ground based spectrophotometry

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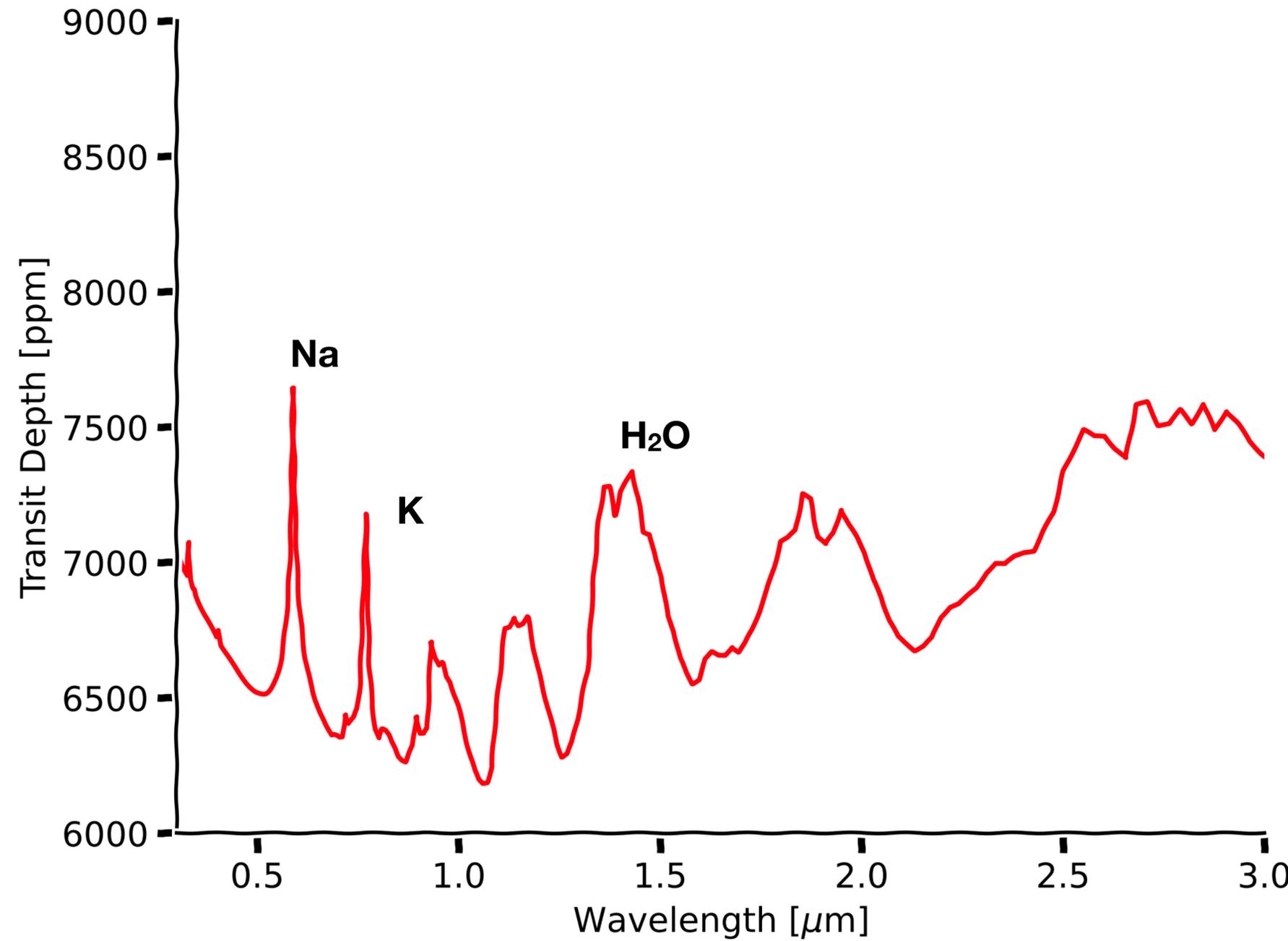
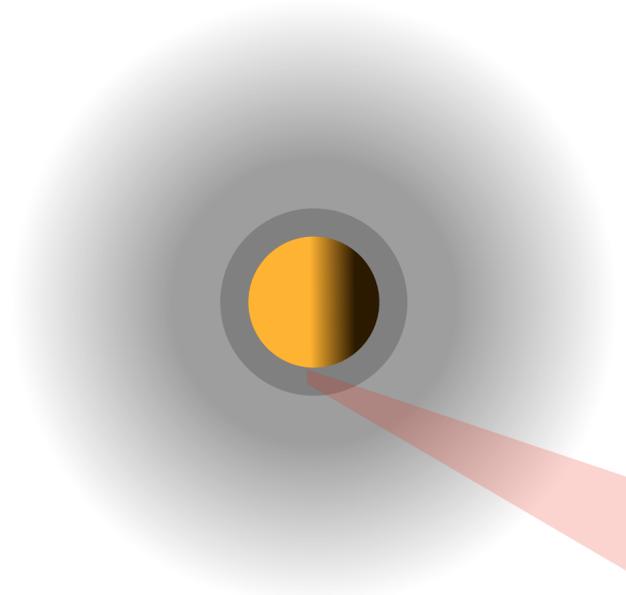


ESO Atmo 2021



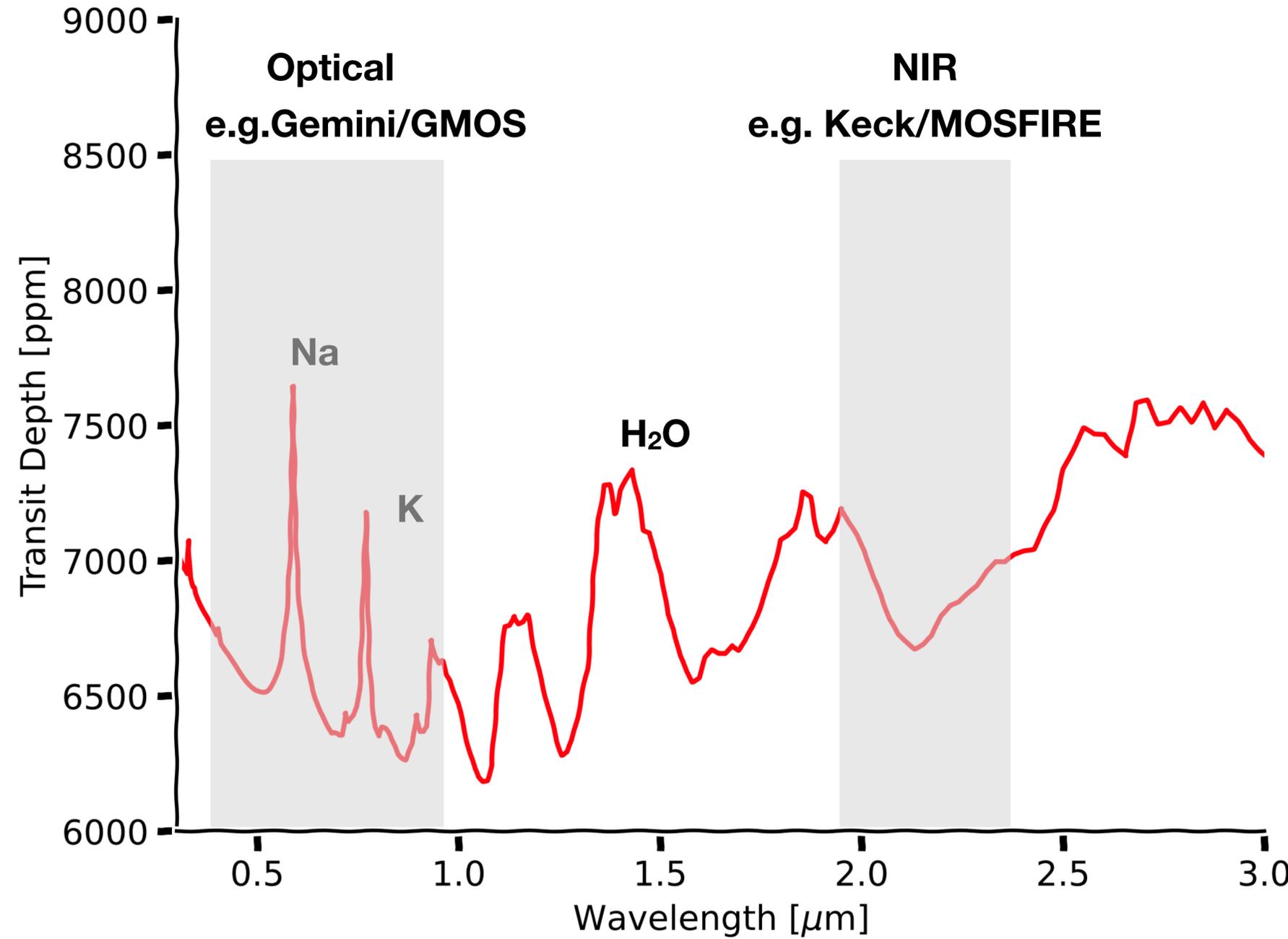
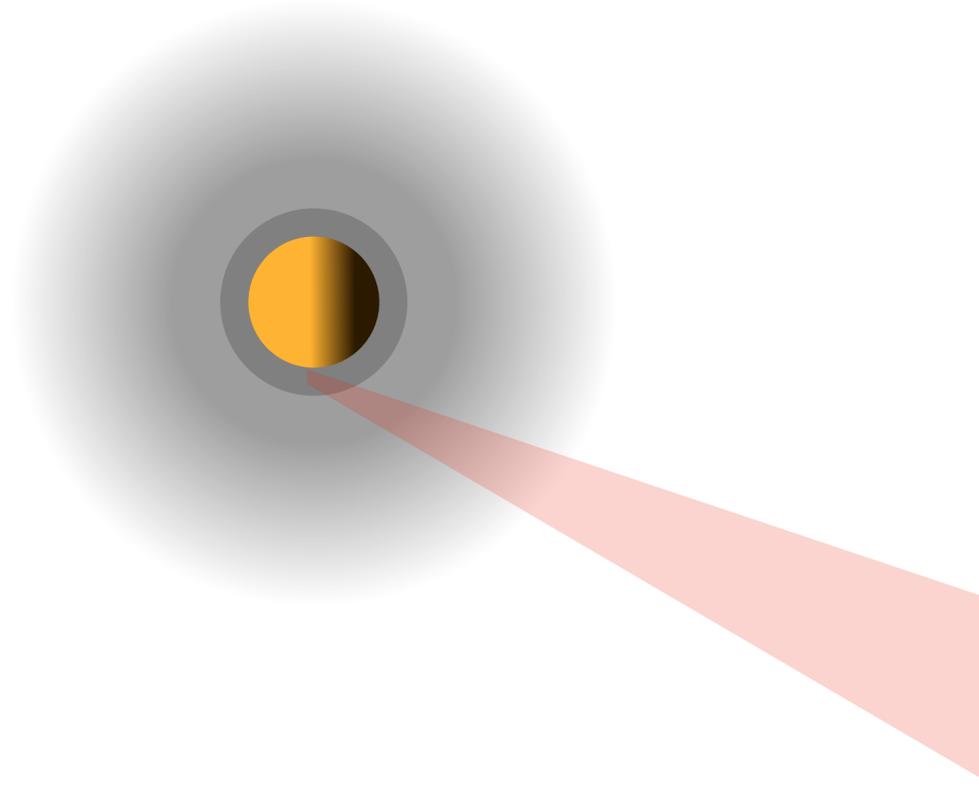
Probing Transiting Exoplanet Atmospheres

from ground, in low-resolution

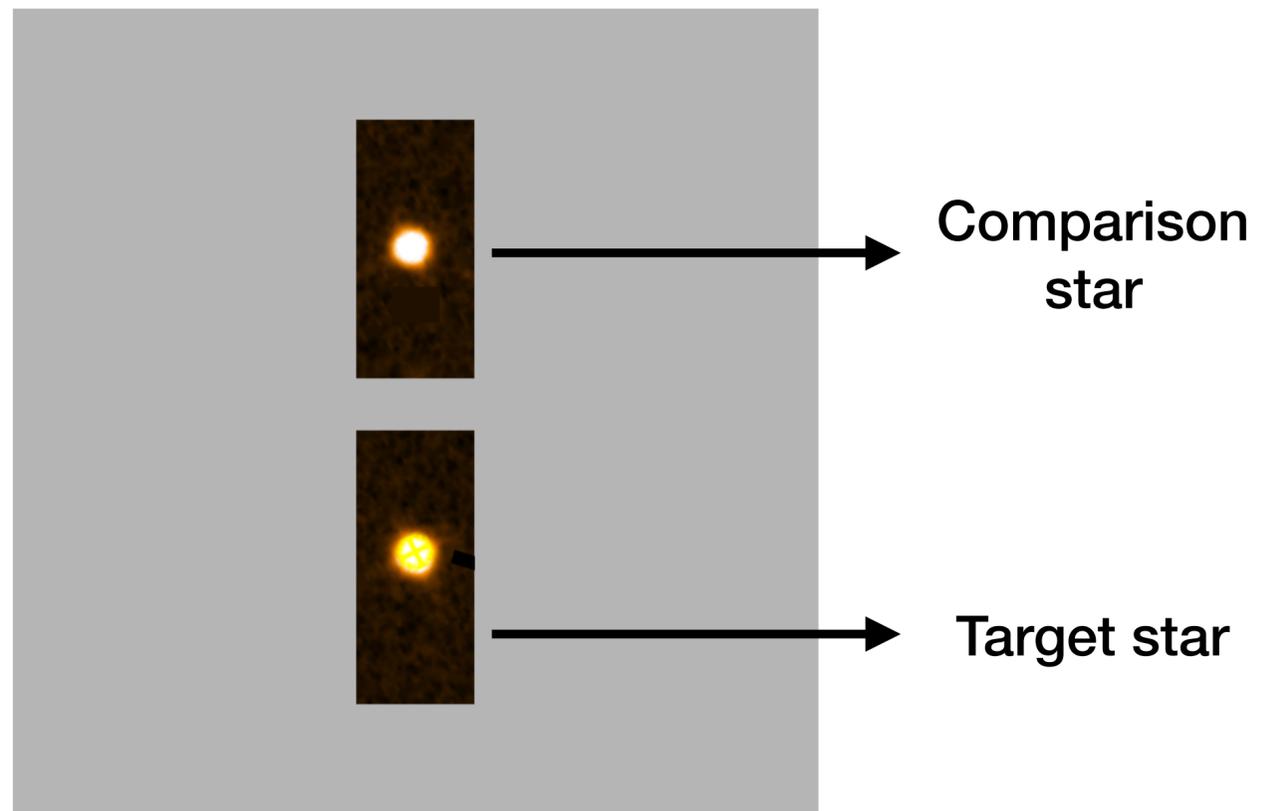


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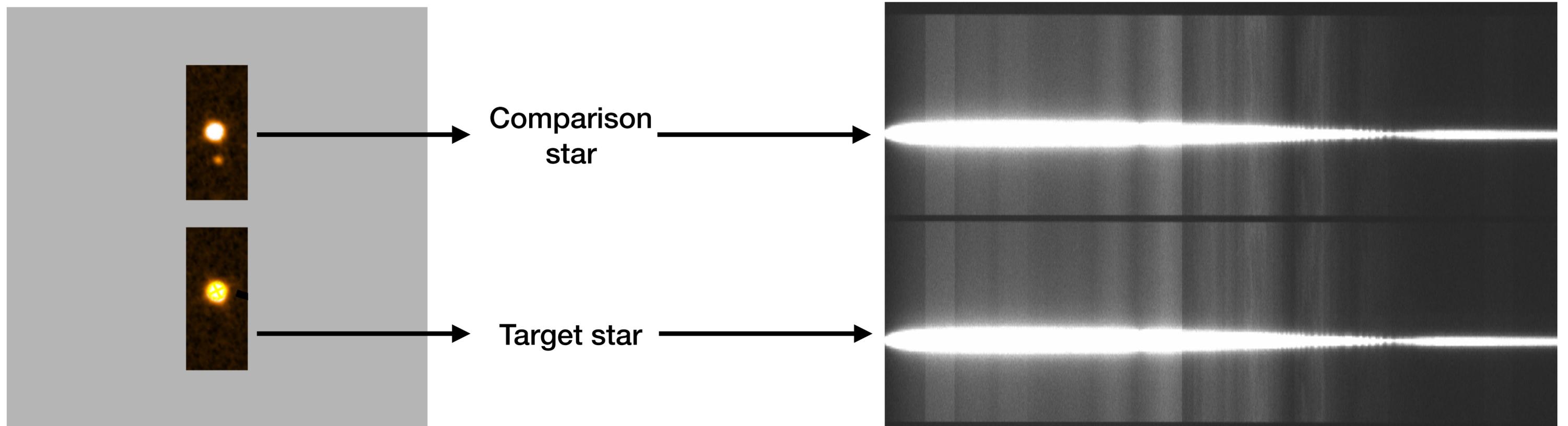
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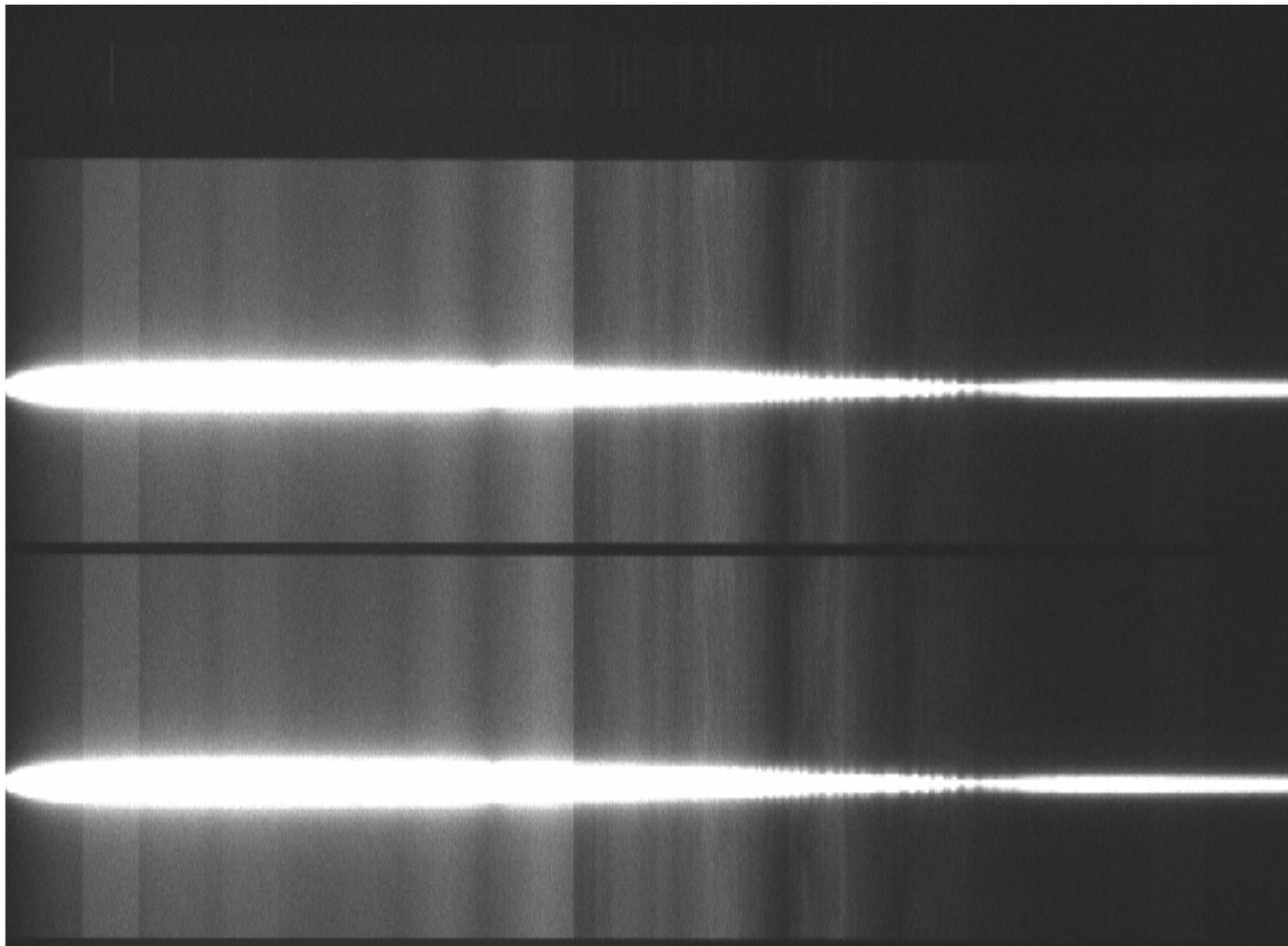
Traditional ground-based multi-object spectroscopy (MOS)



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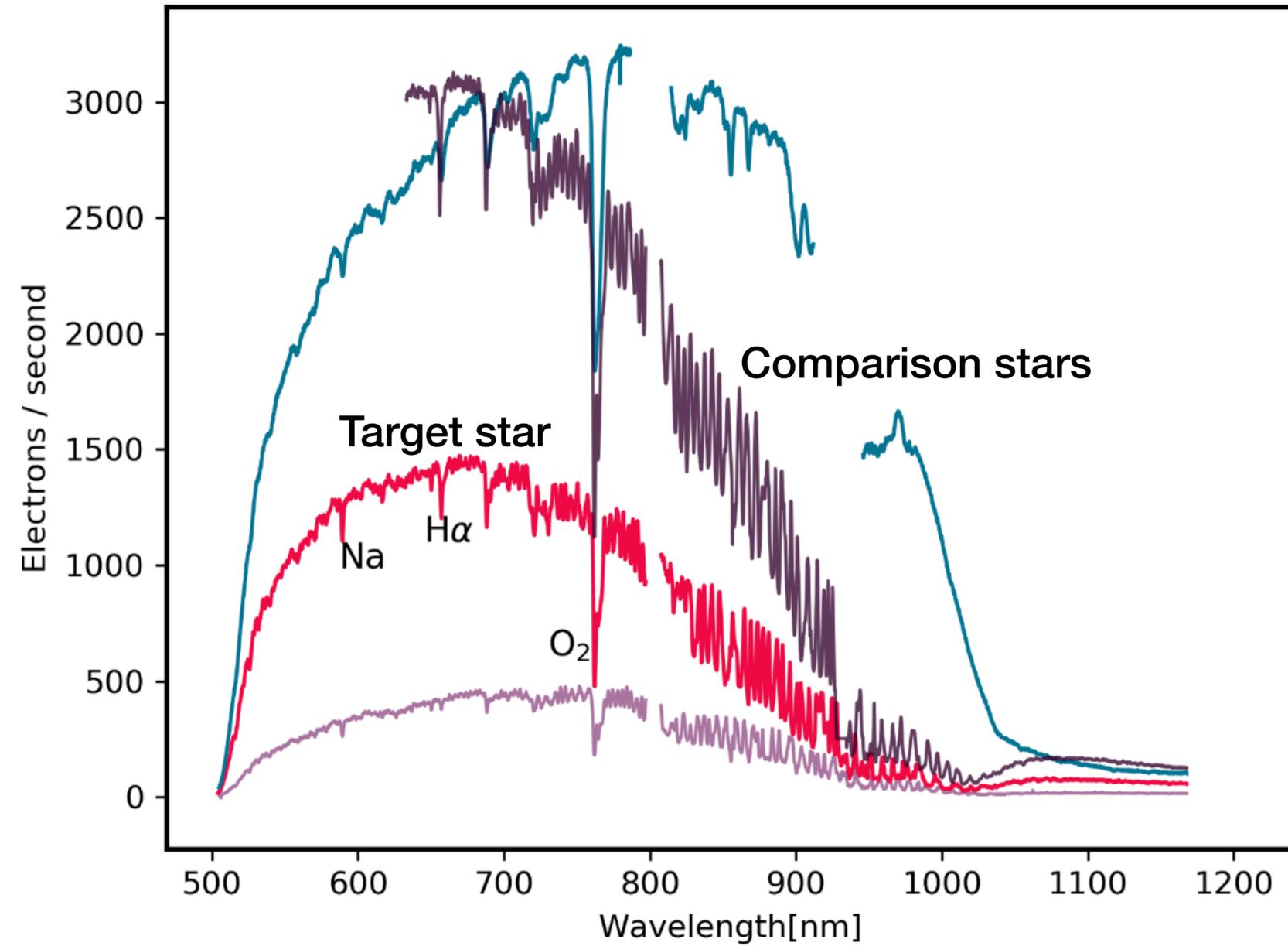


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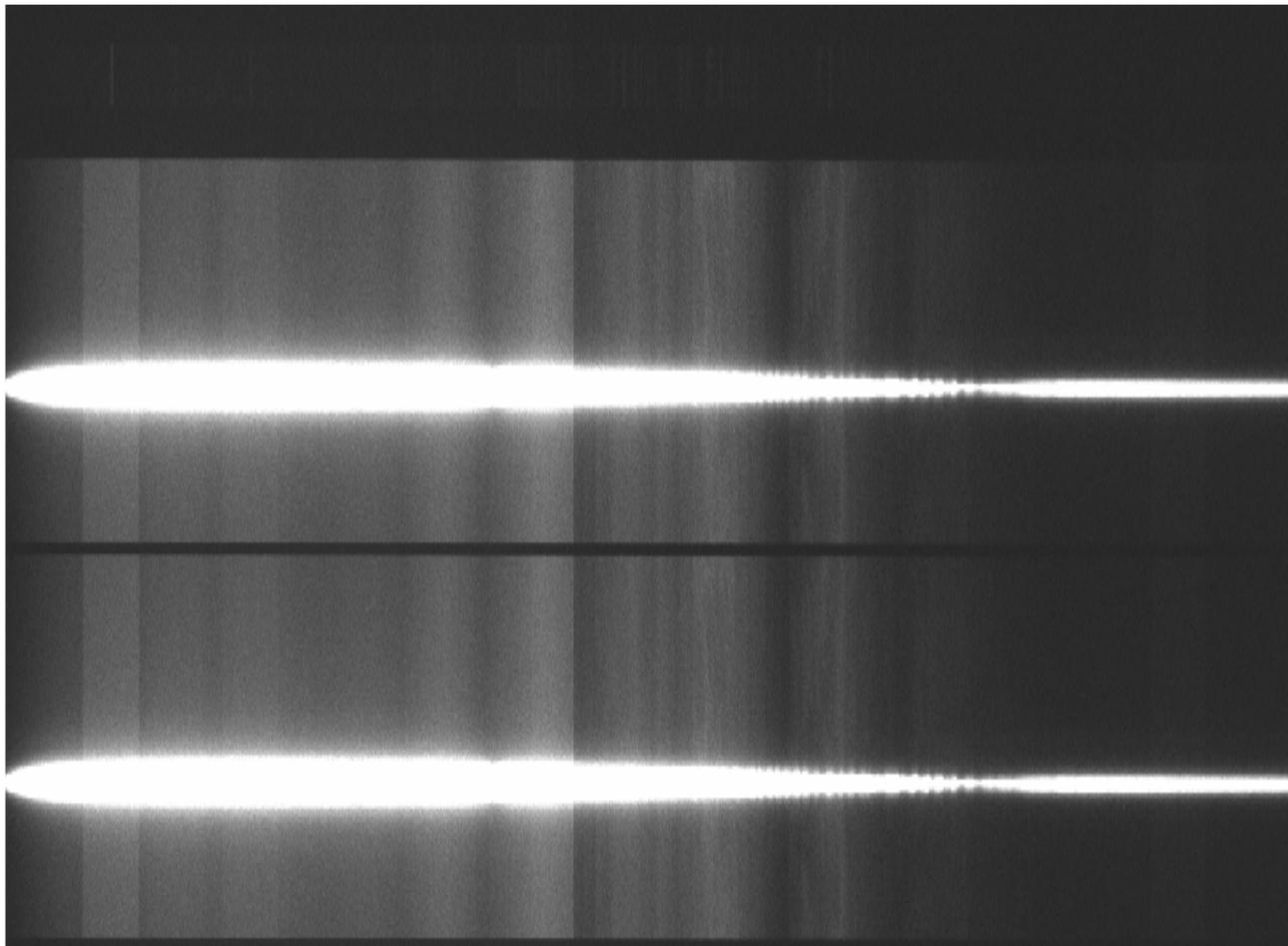


Comparison star

Target star

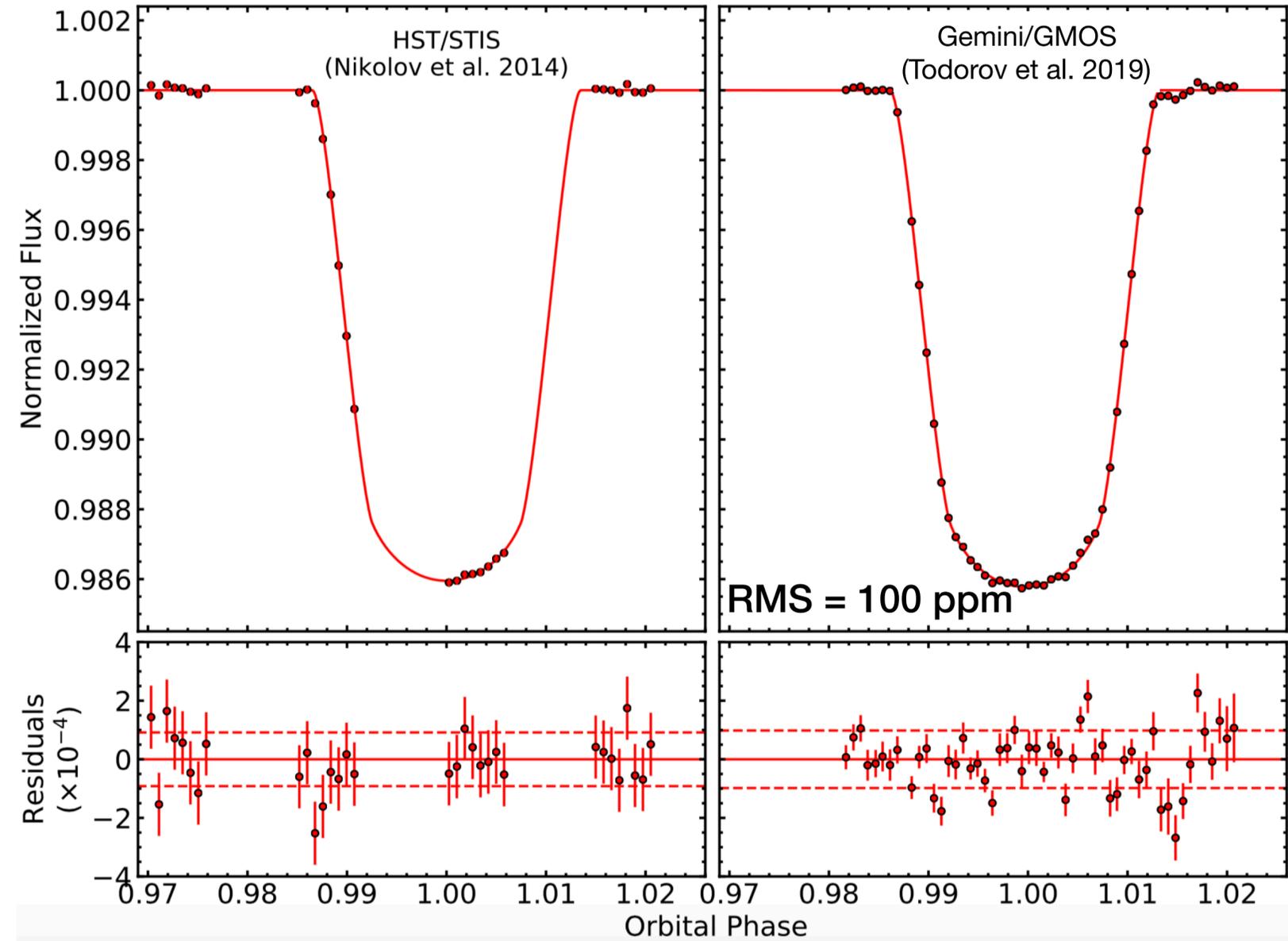


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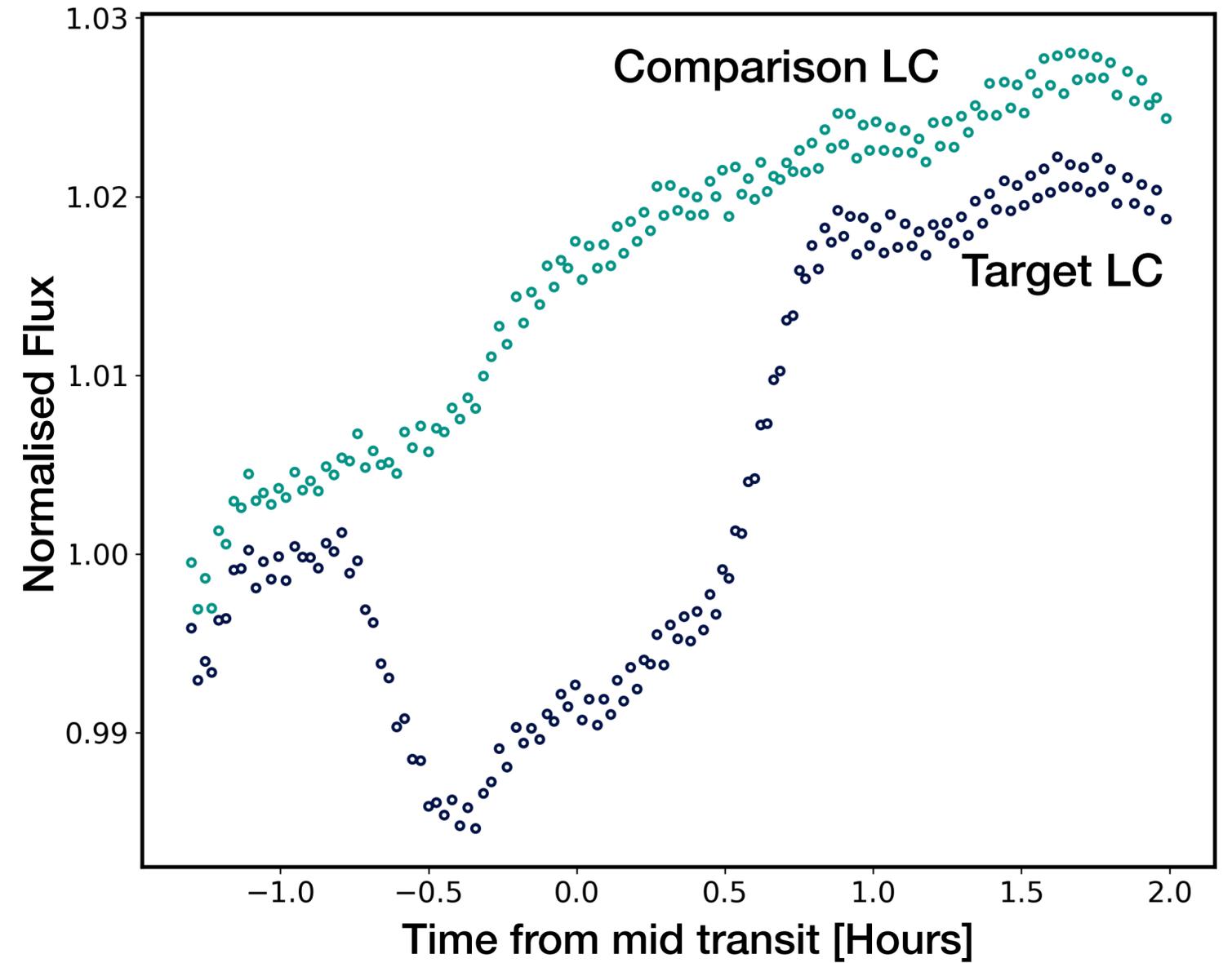
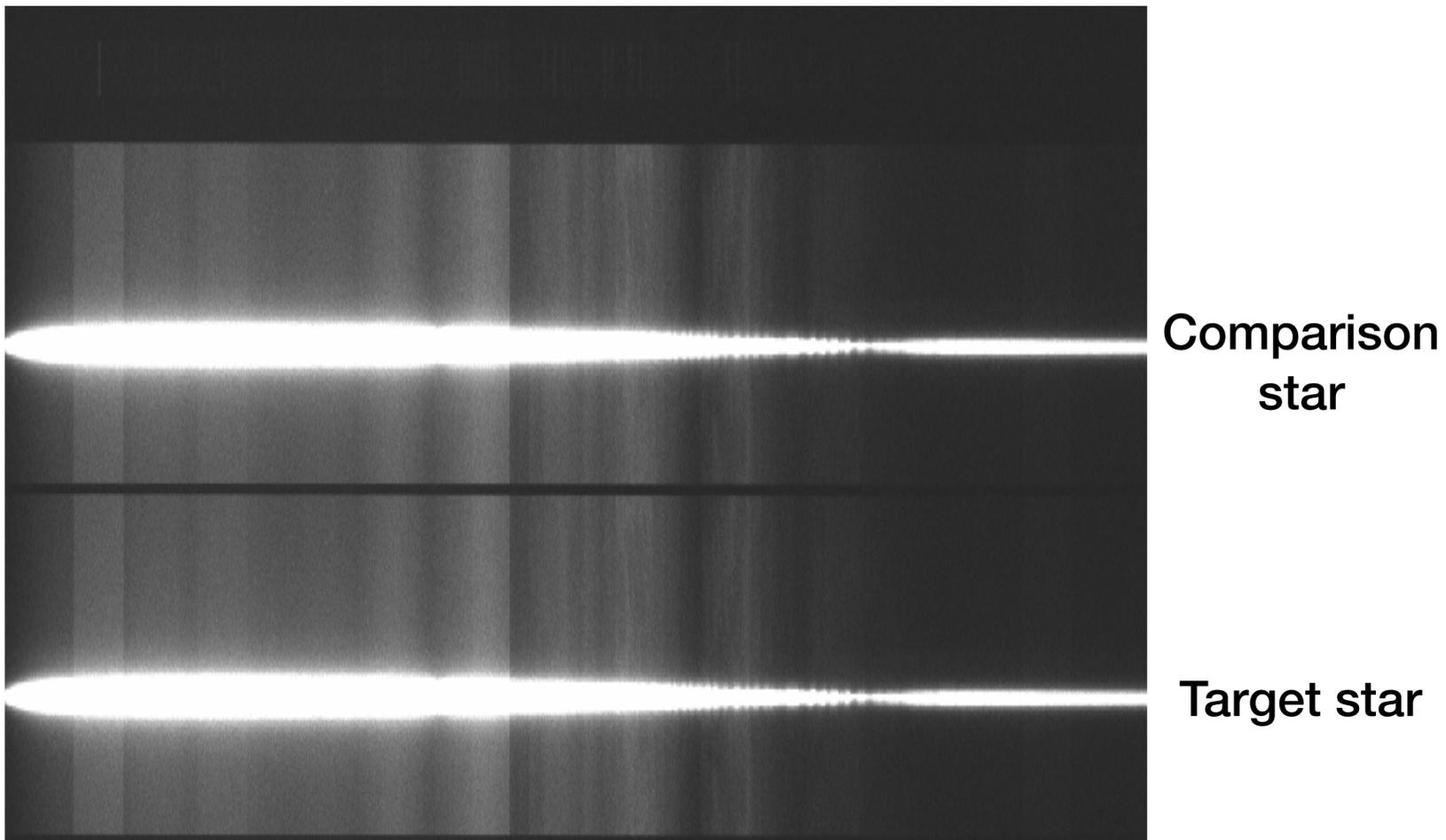


Comparison
star

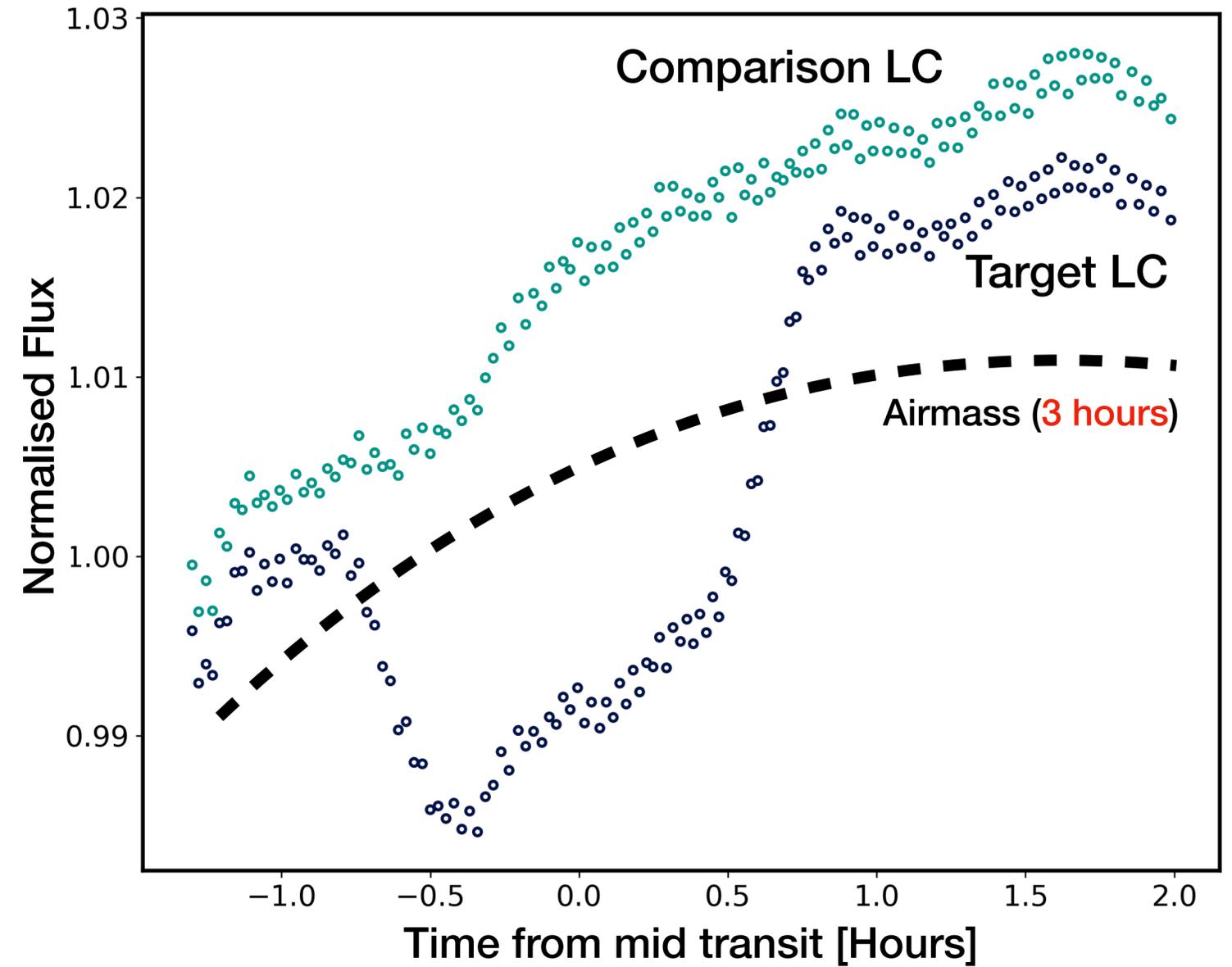
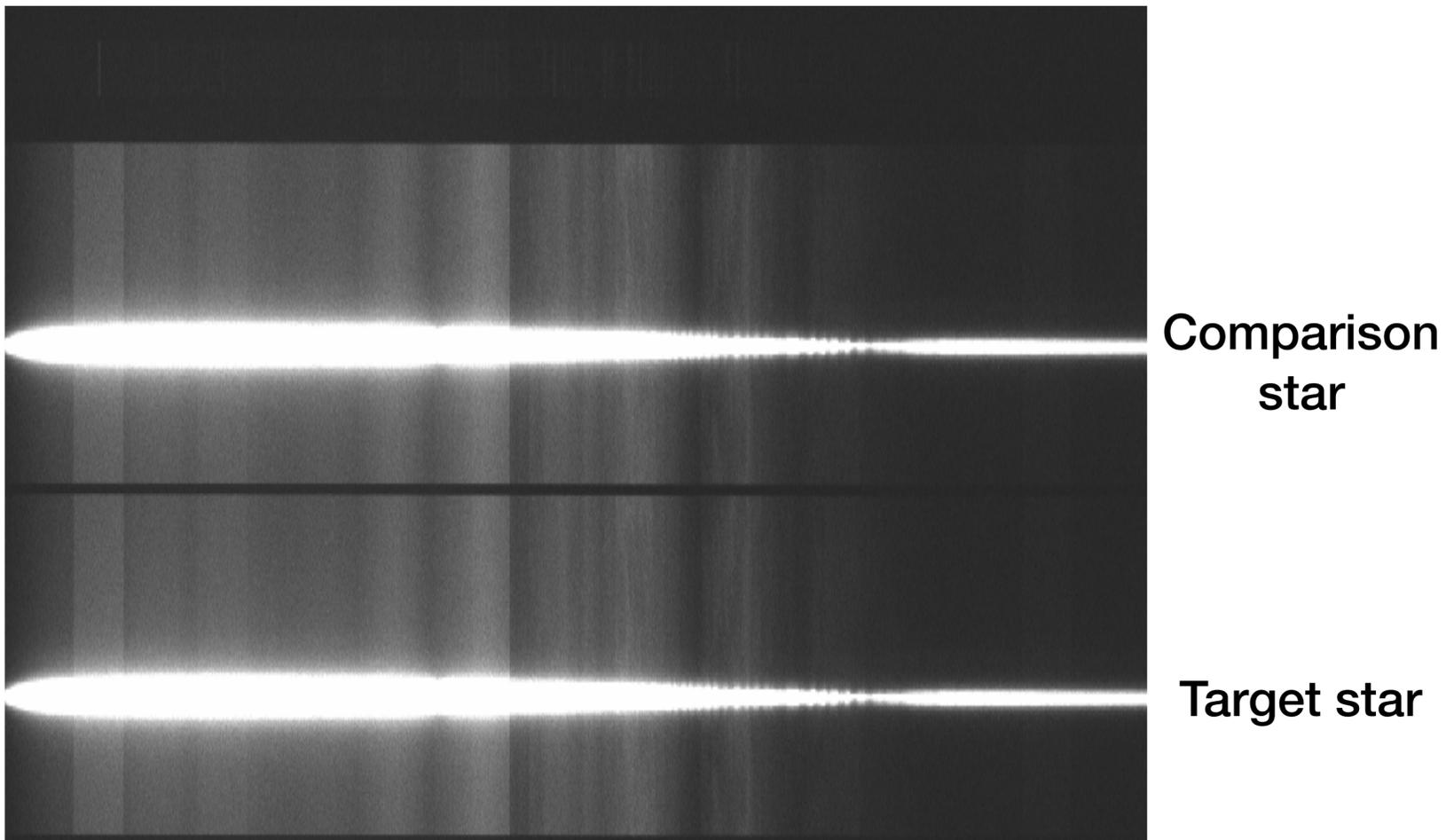
Target star



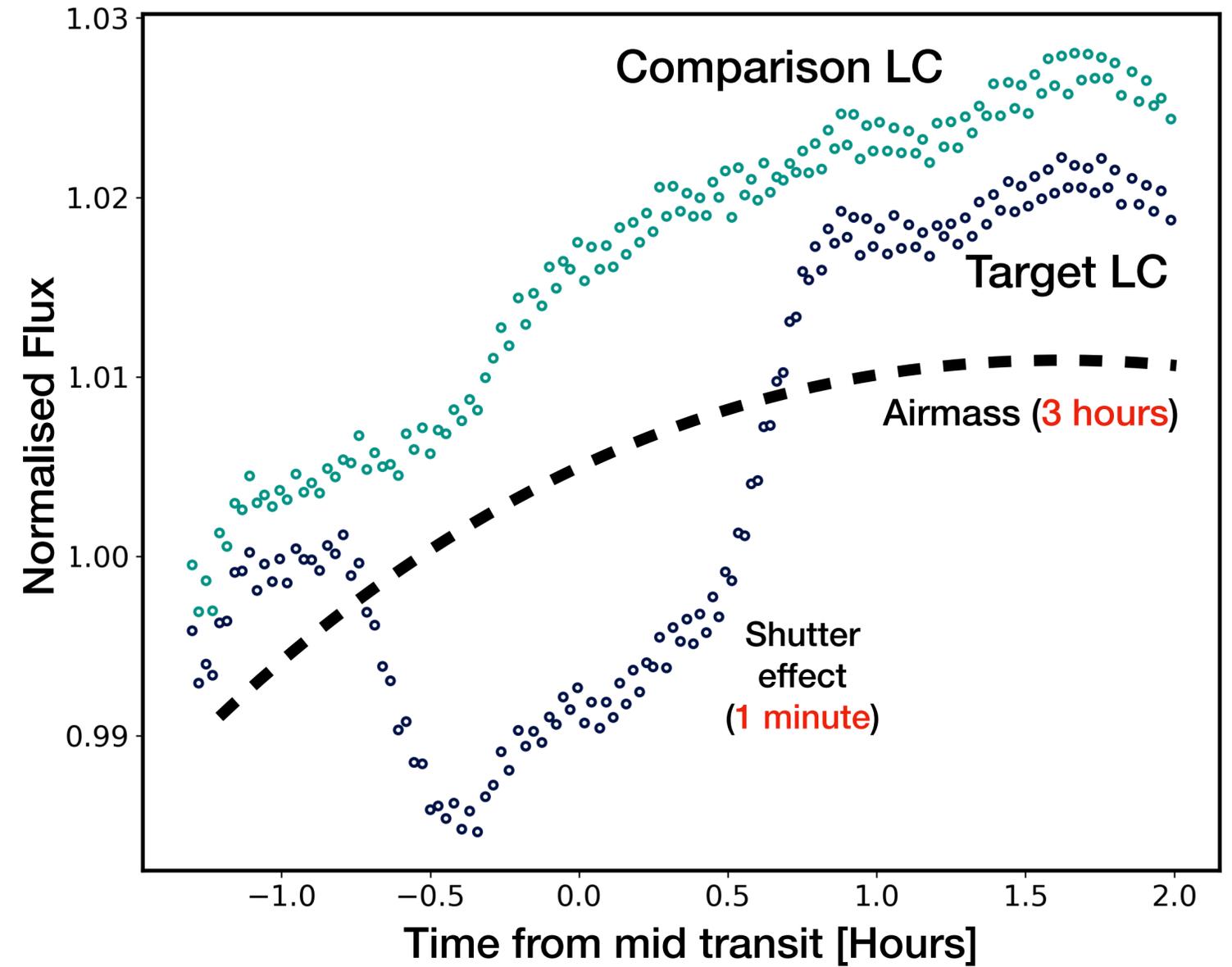
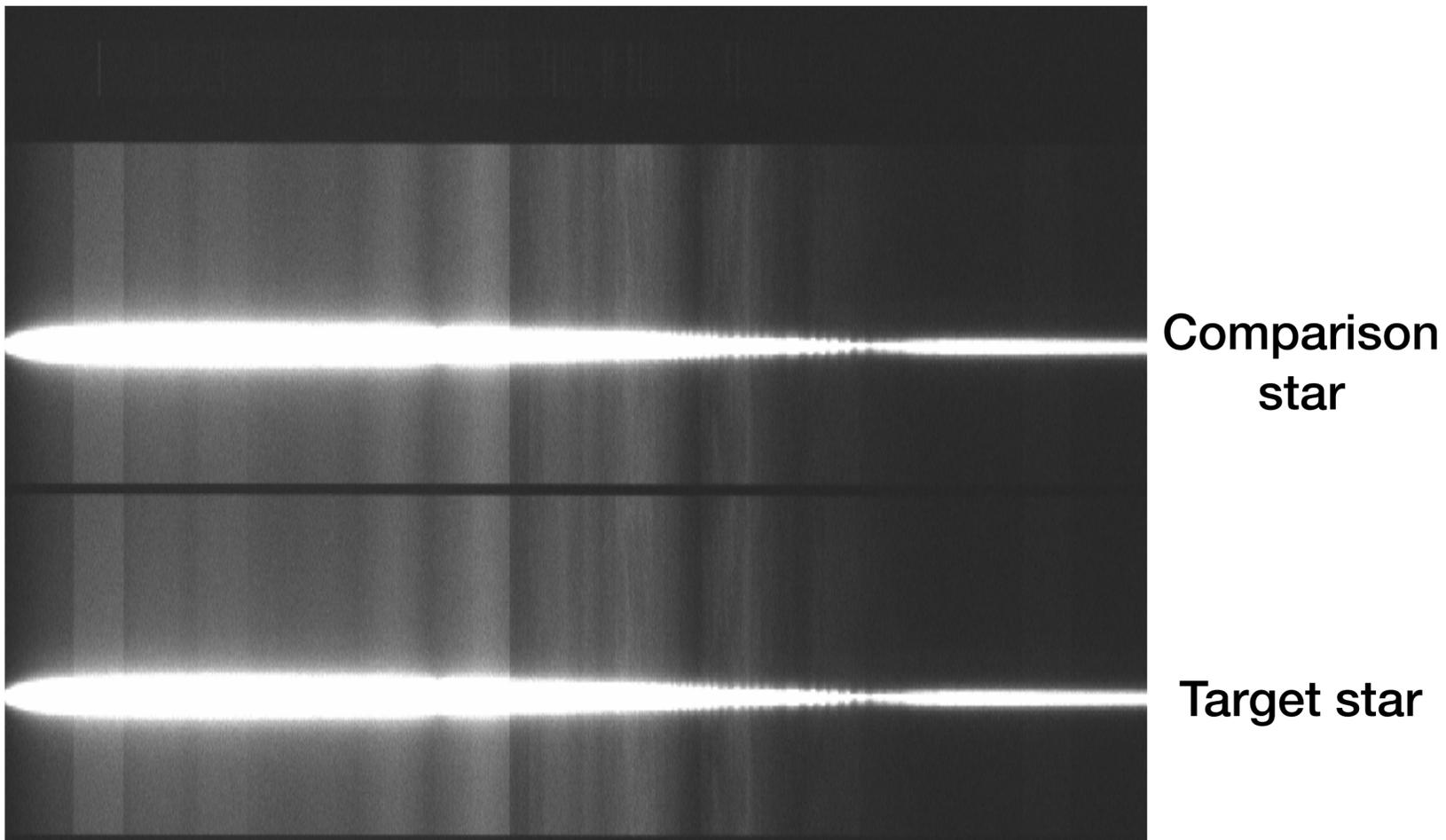
Revisiting ground-based multi-object spectroscopy (MOS)



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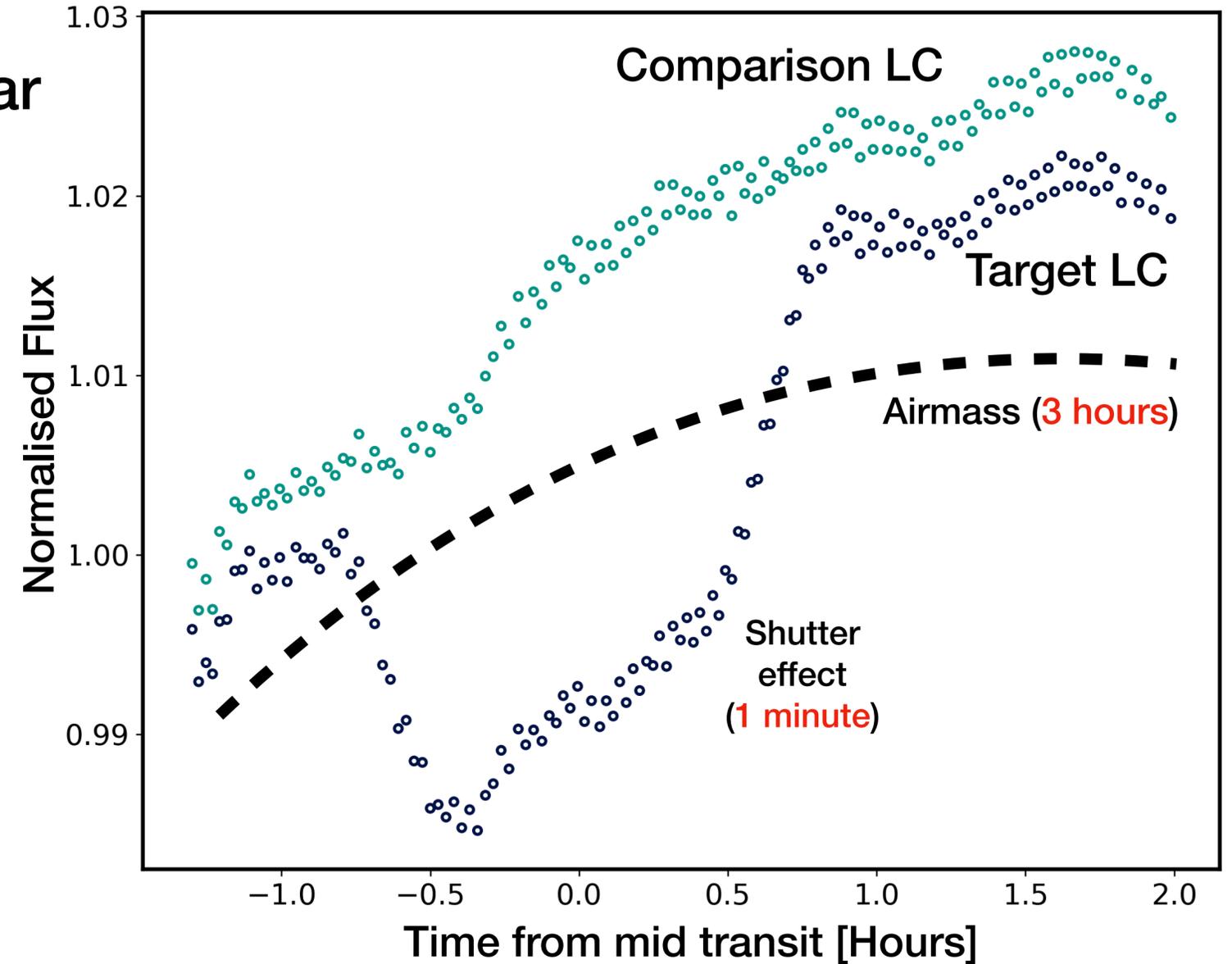


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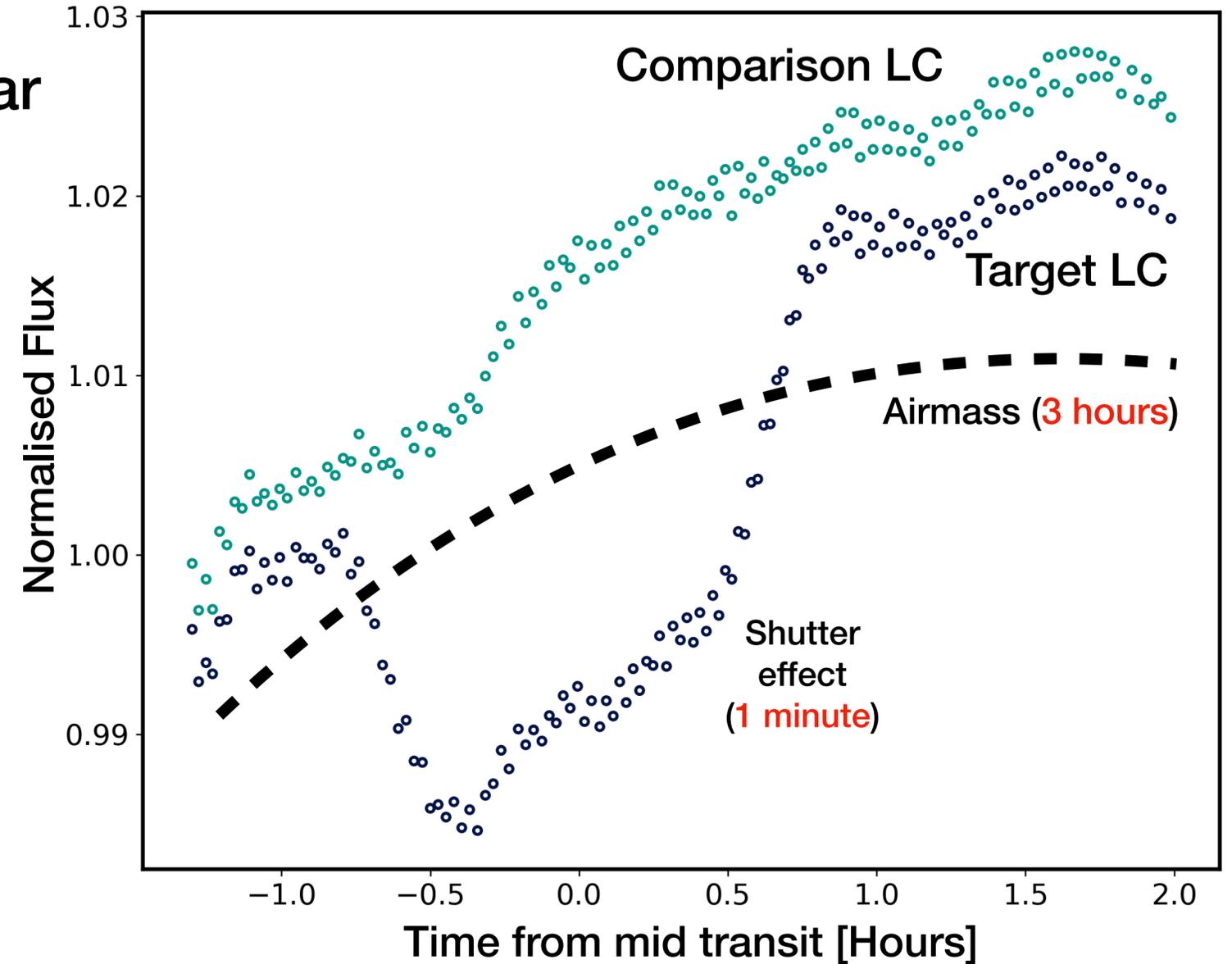
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Revisiting ground-based multi-object spectroscopy (MOS)

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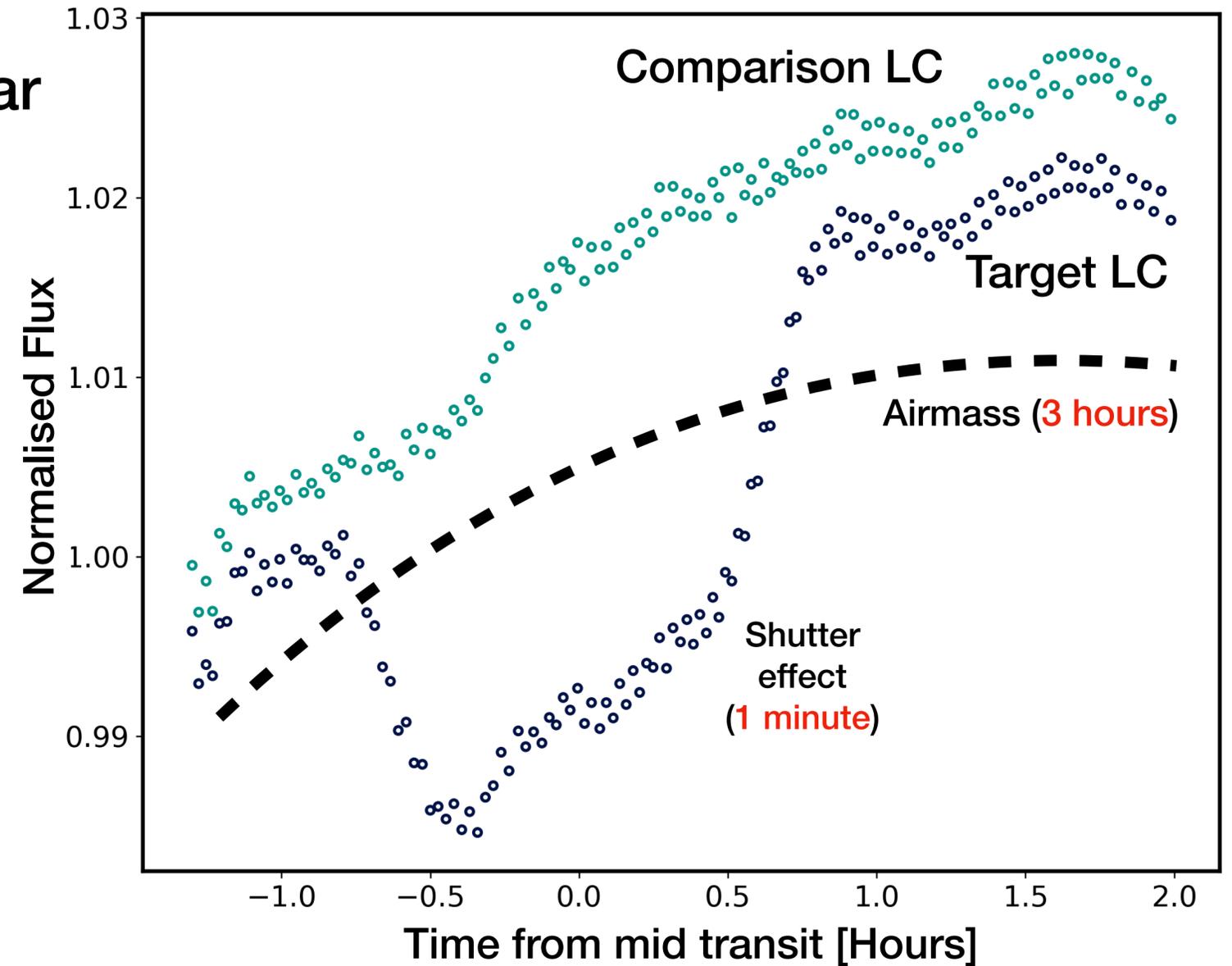


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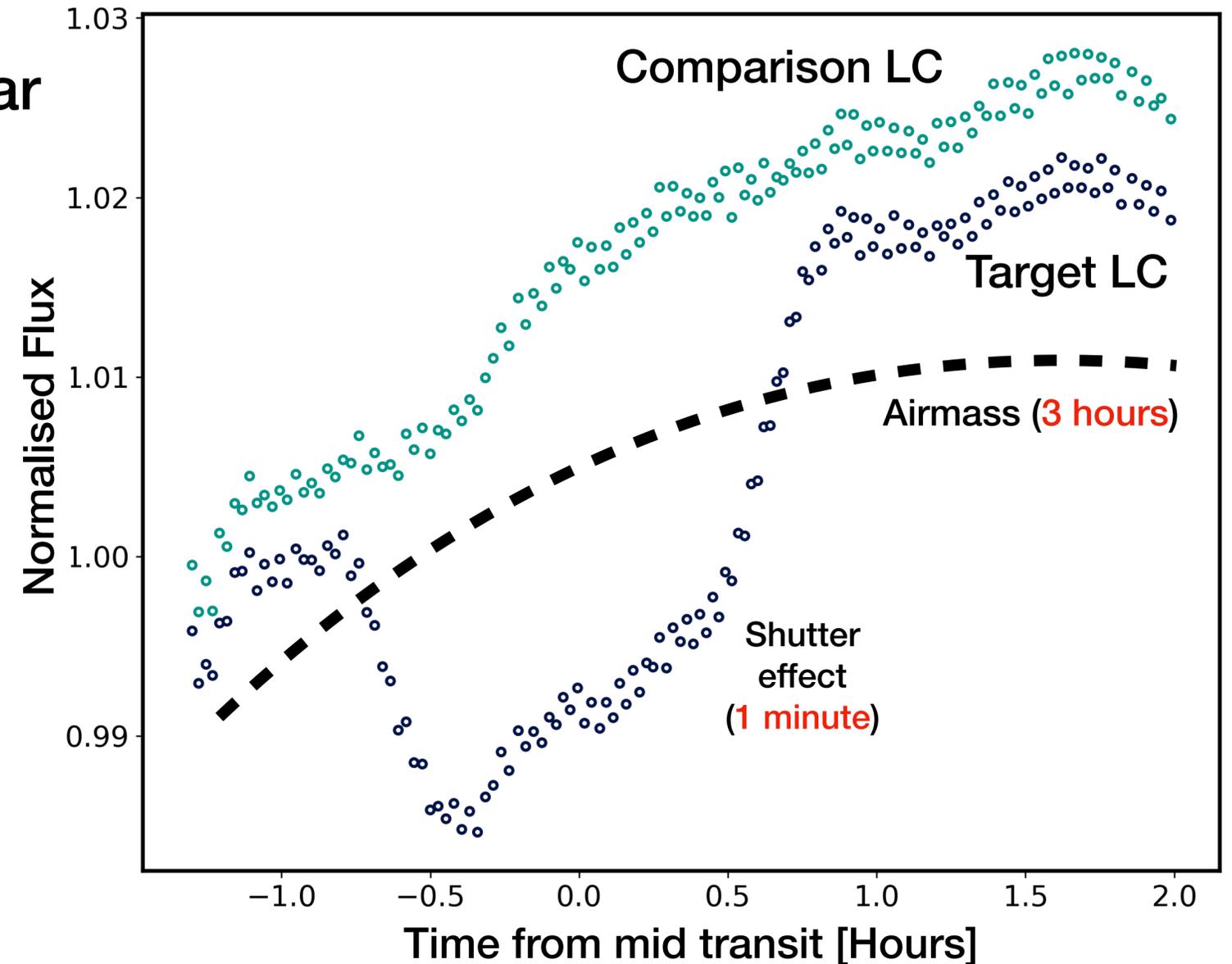
2. The operation of **normalising** can add **systematics!**



Revisiting ground-based multi-object spectroscopy (MOS)

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3. **Difficult to follow-up bright targets** with no nearby suitable comparison stars

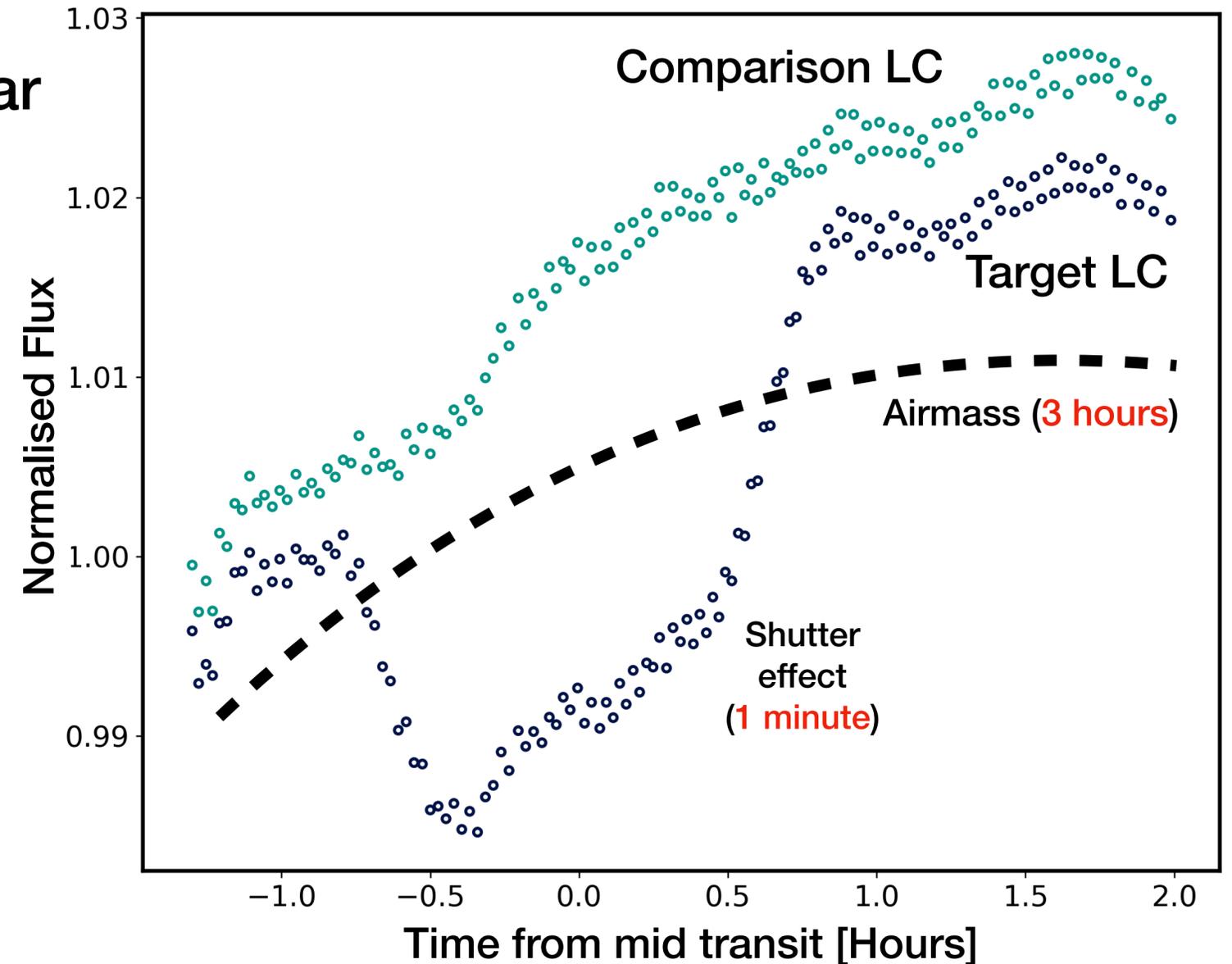


Revisiting ground-based multi-object spectroscopy (MOS)

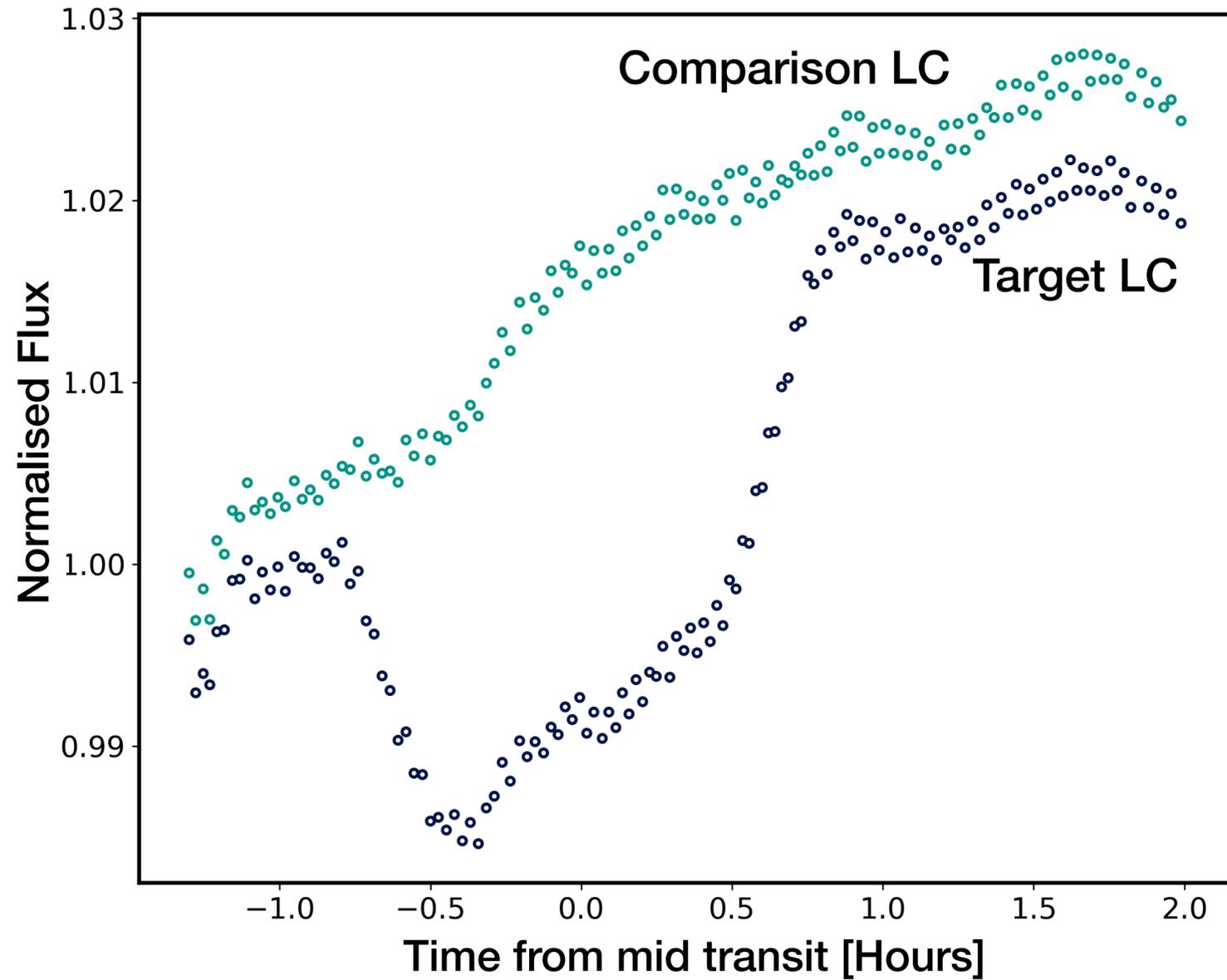
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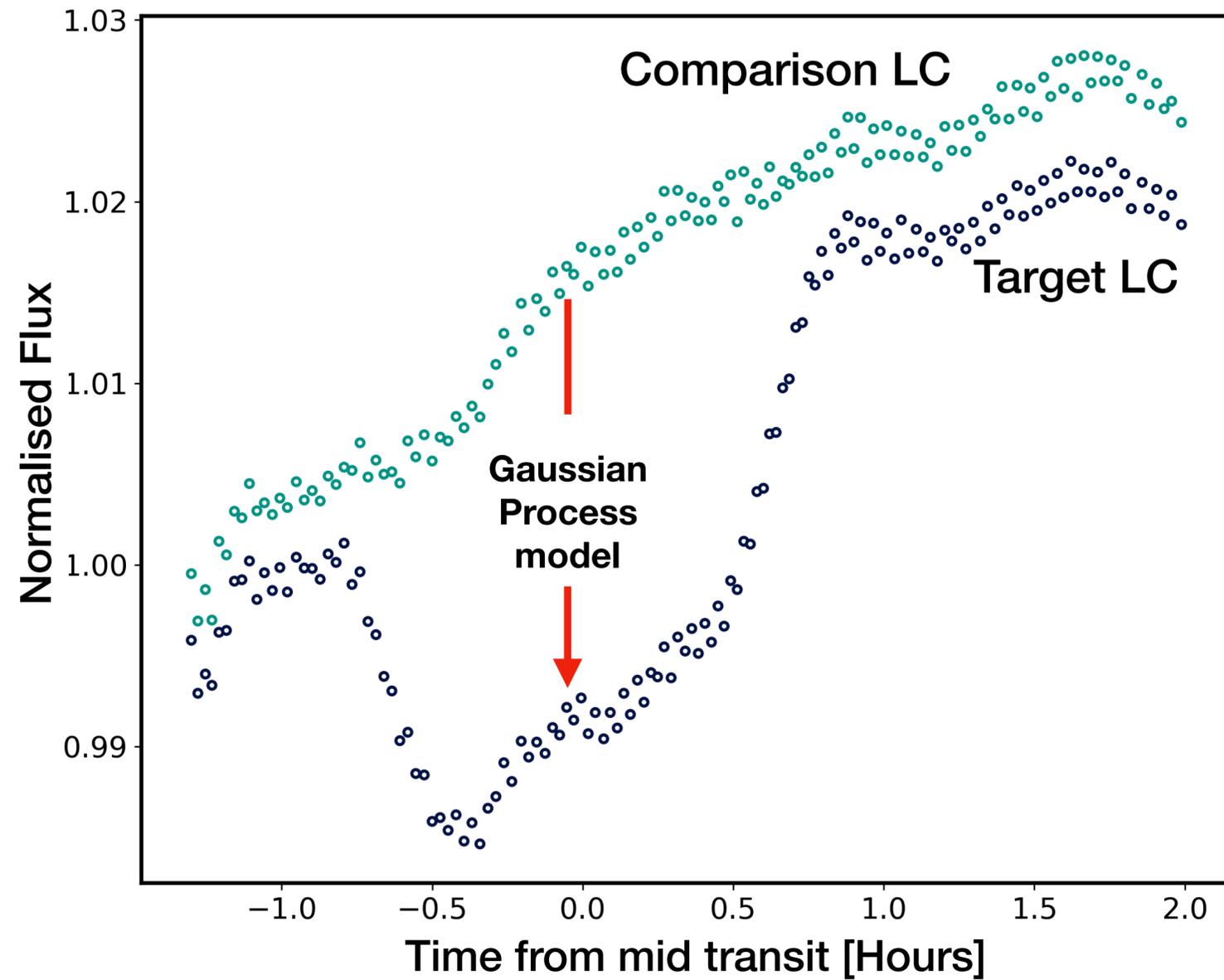
Need for a new method!



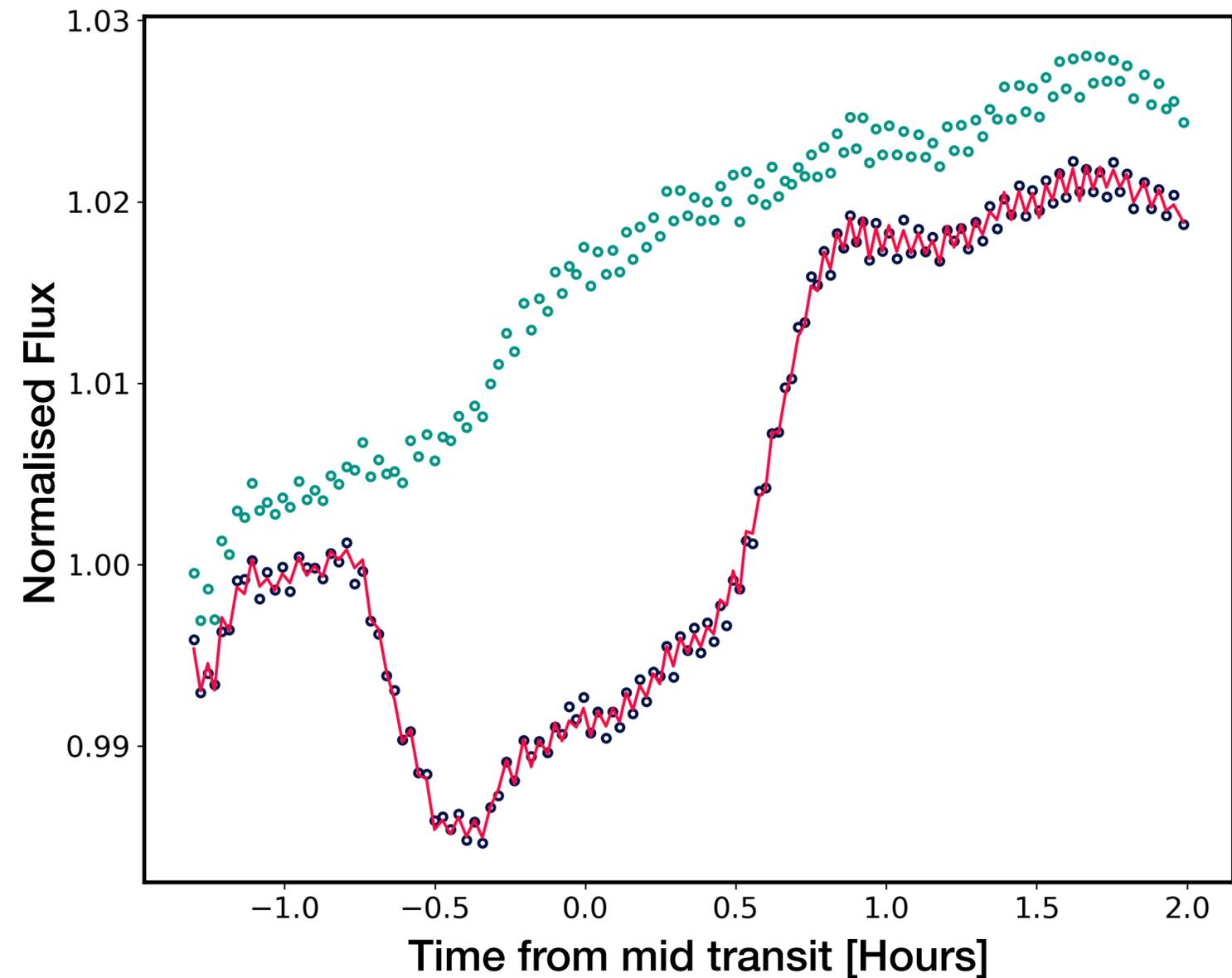
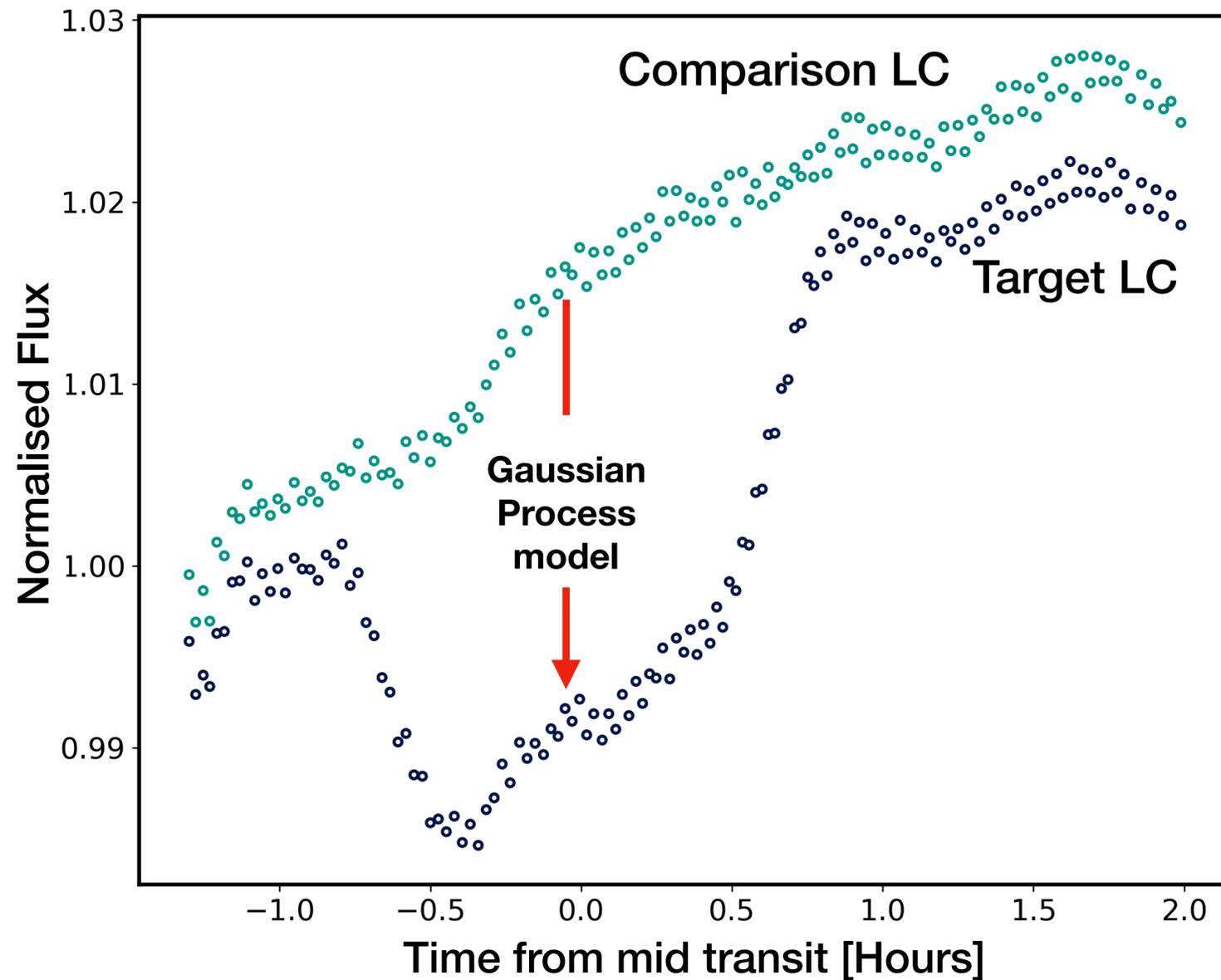
New method to correct for systematics in ground based light curves



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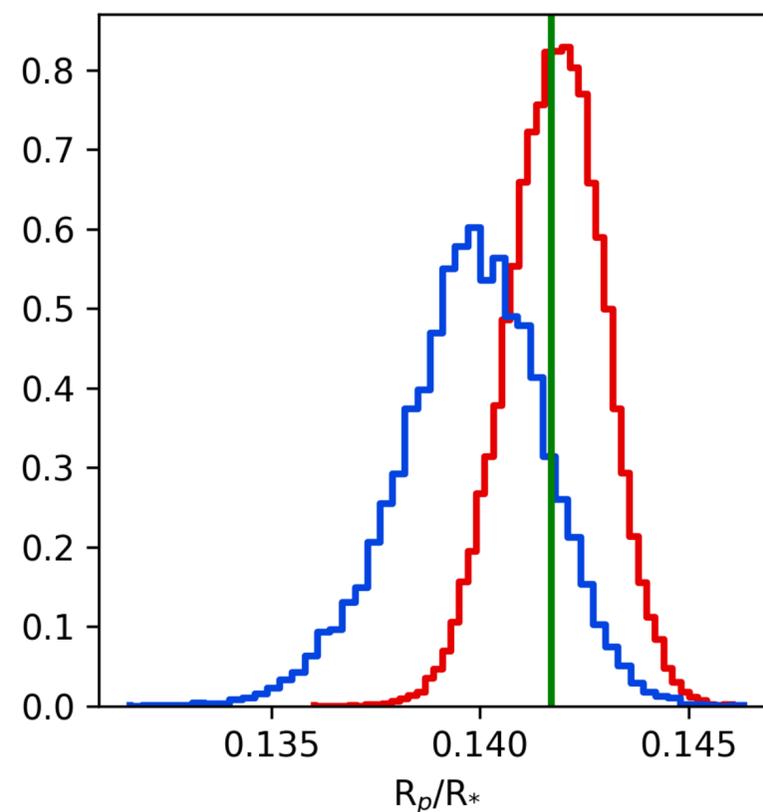
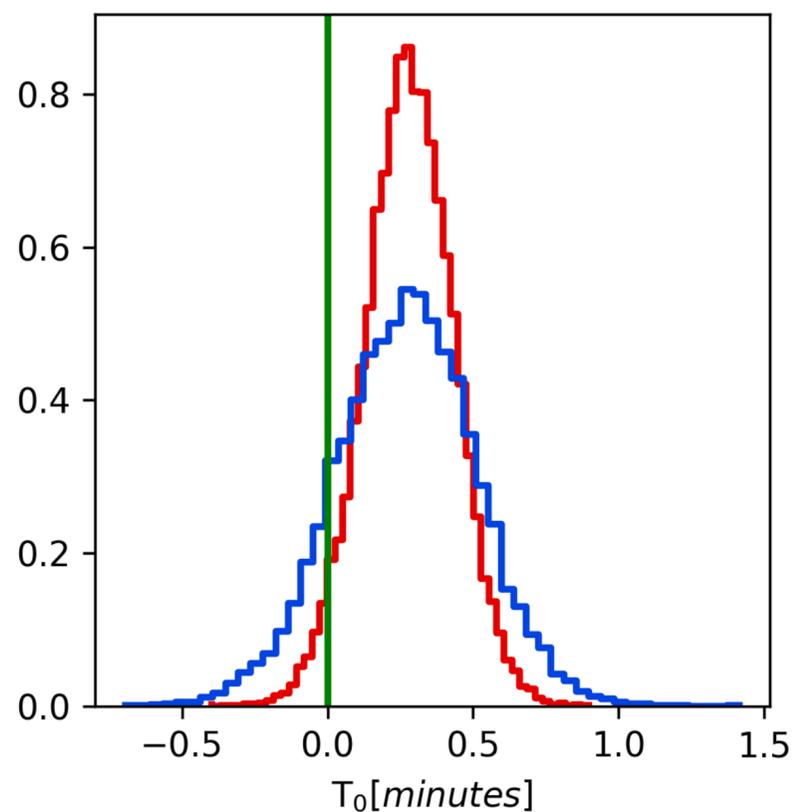
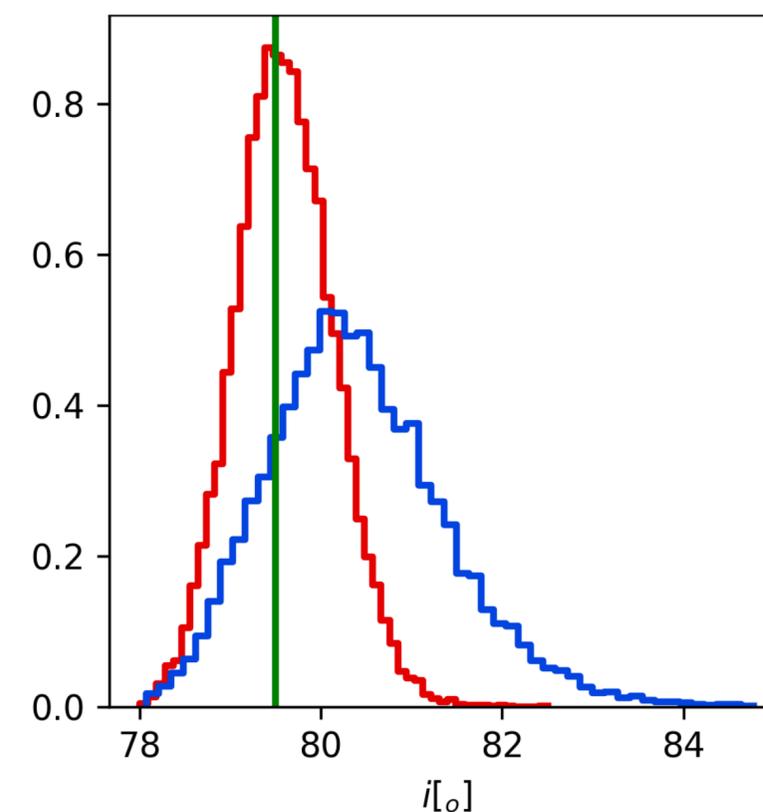
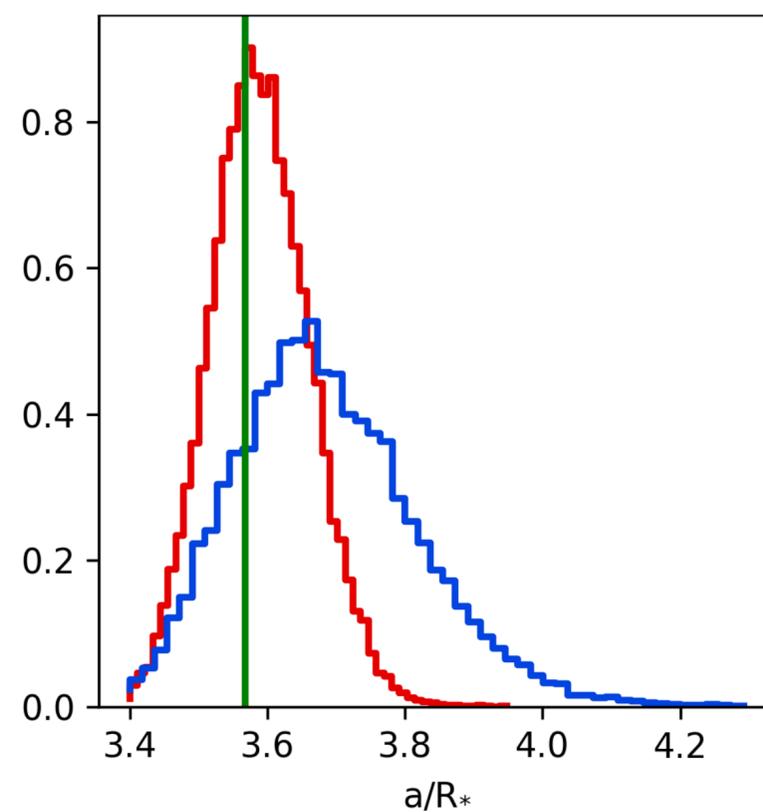
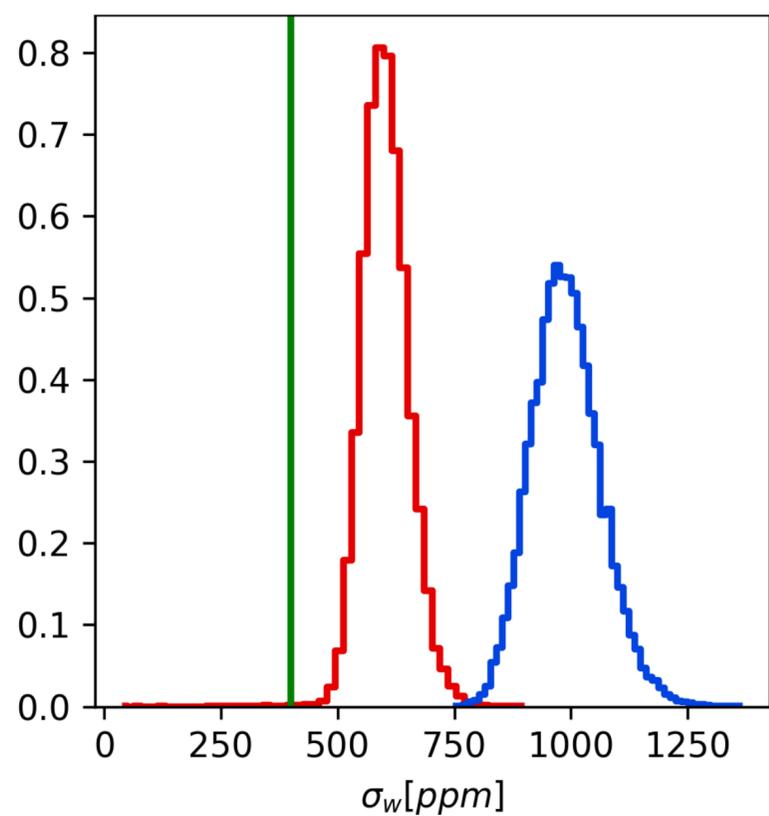


New method to correct for systematics in ground based light curves



Also see Gibson et al. 2011

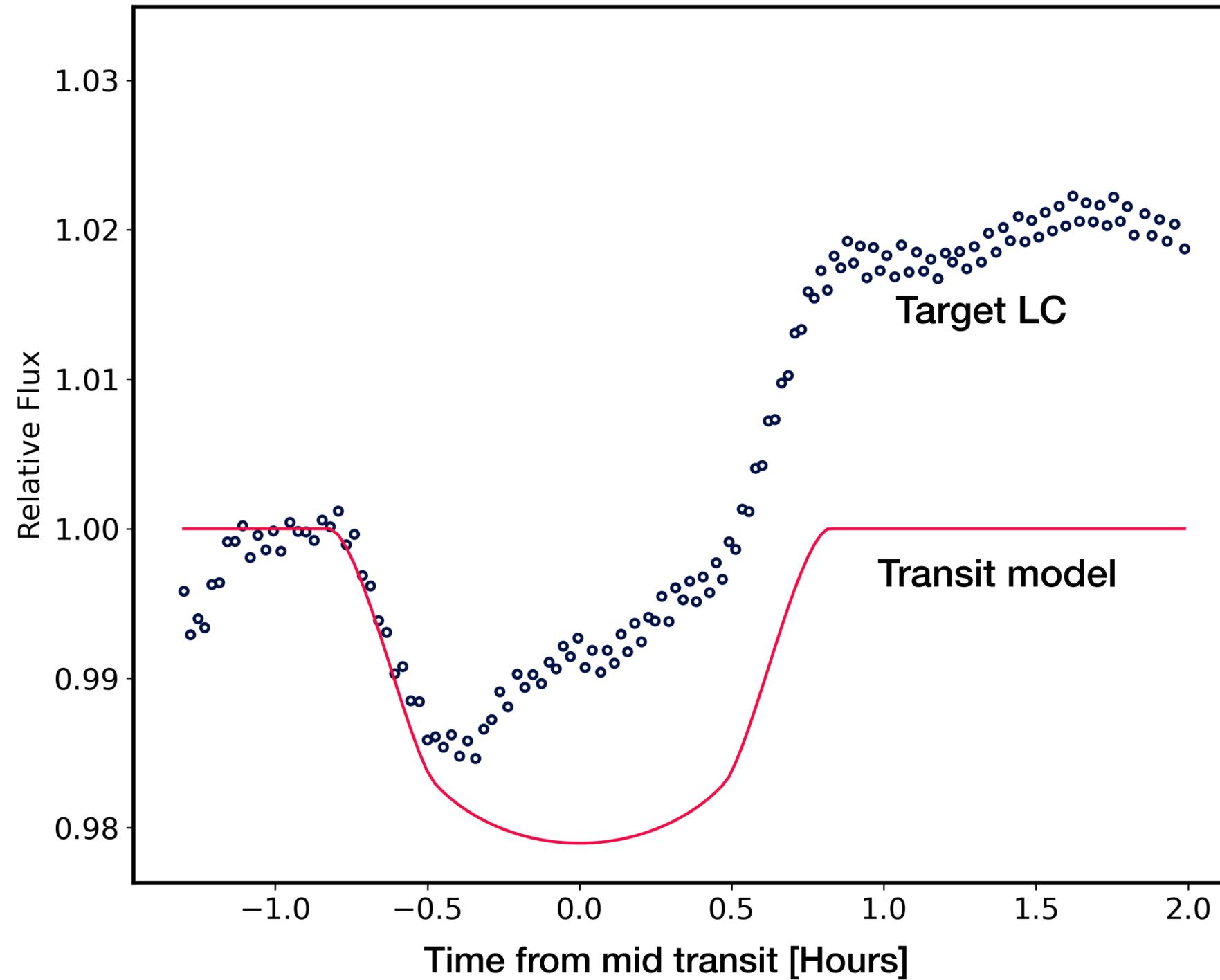
New method improves accuracy and precision of transit parameters



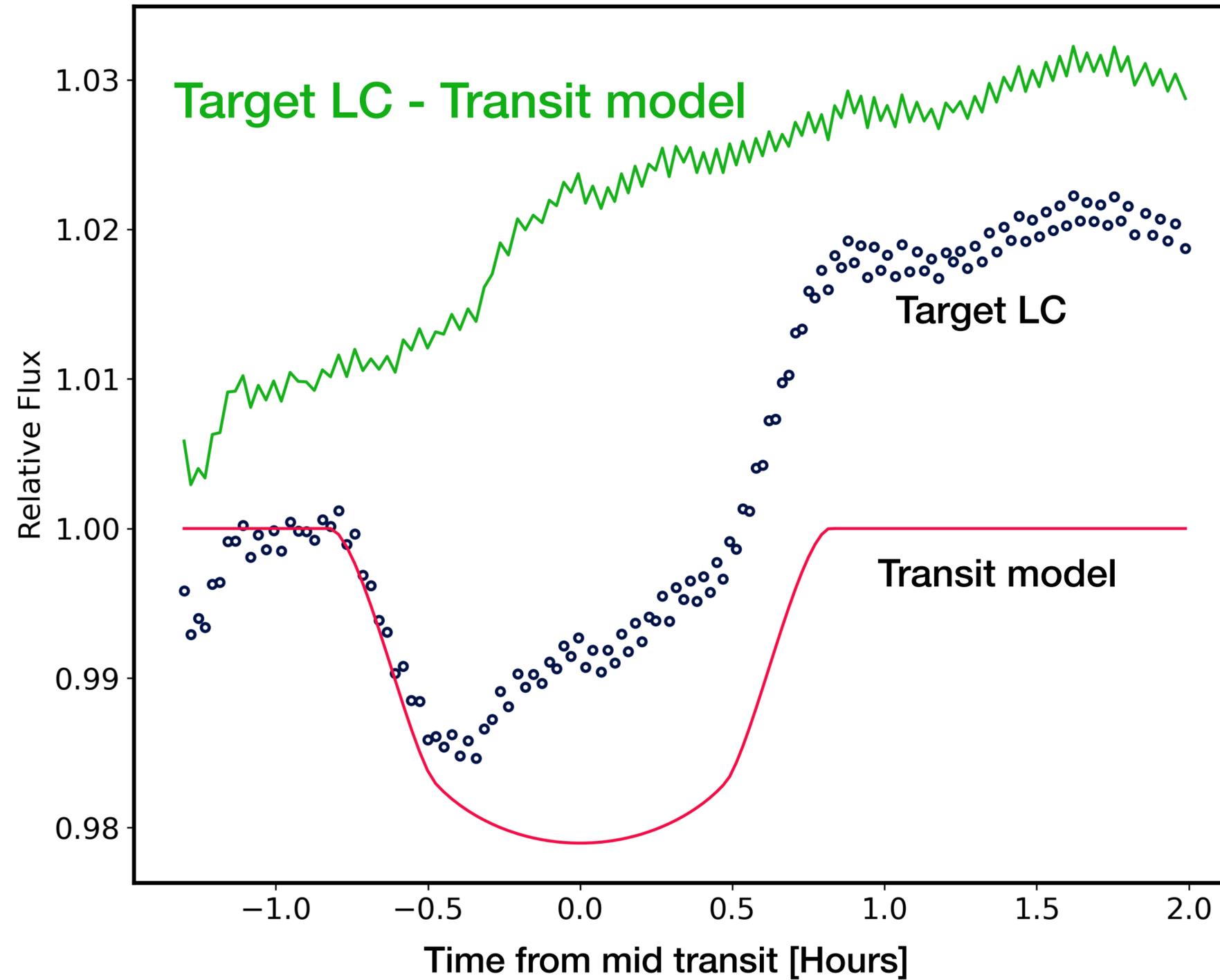
**New method:
Target LC**

**Conventional method :
Target/Comparison LC**

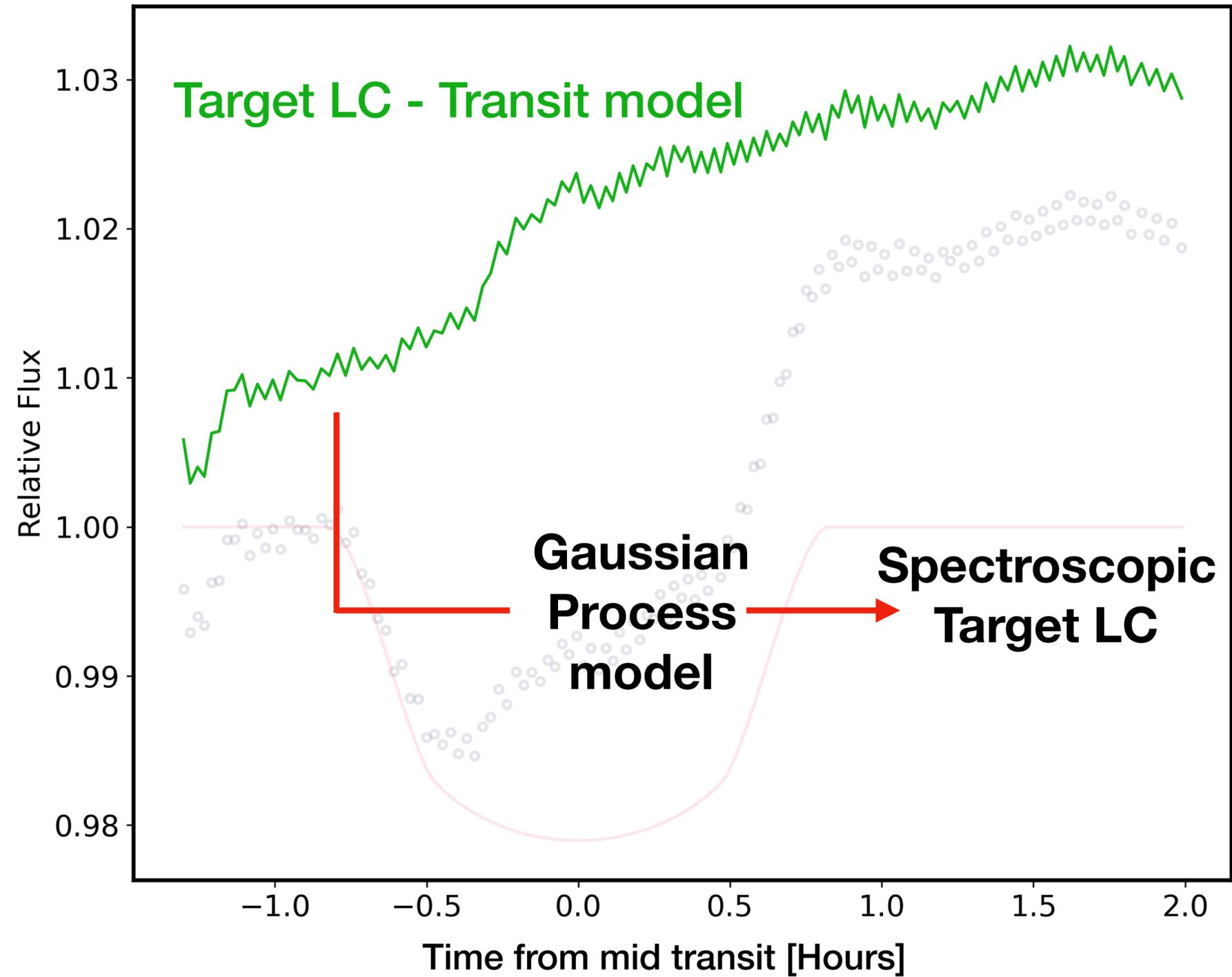
Extracting transmission spectrum using only the Target star



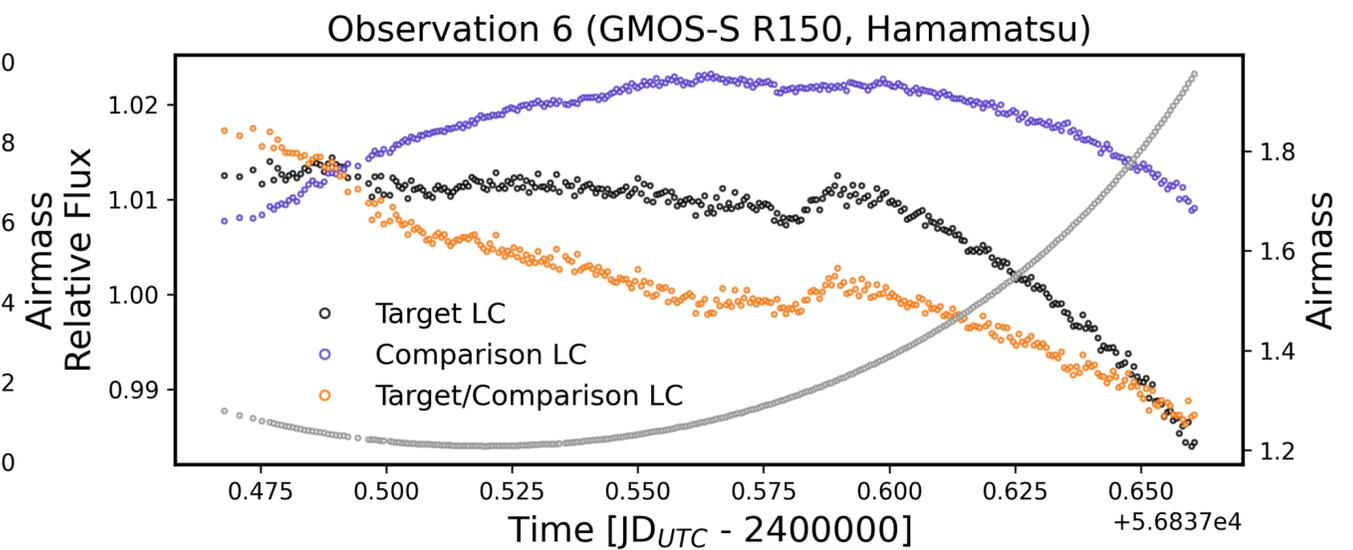
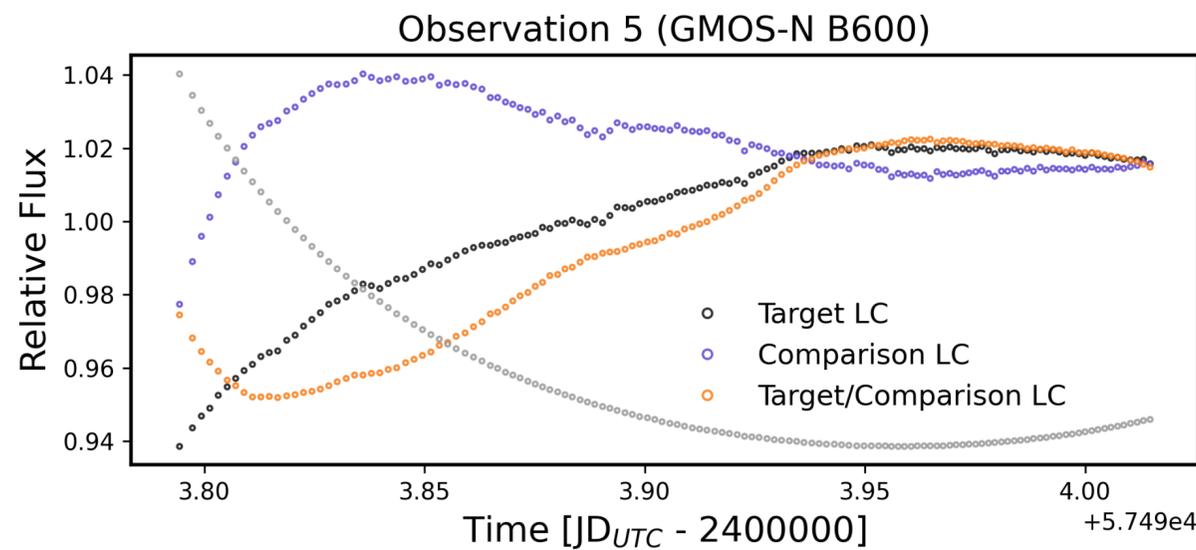
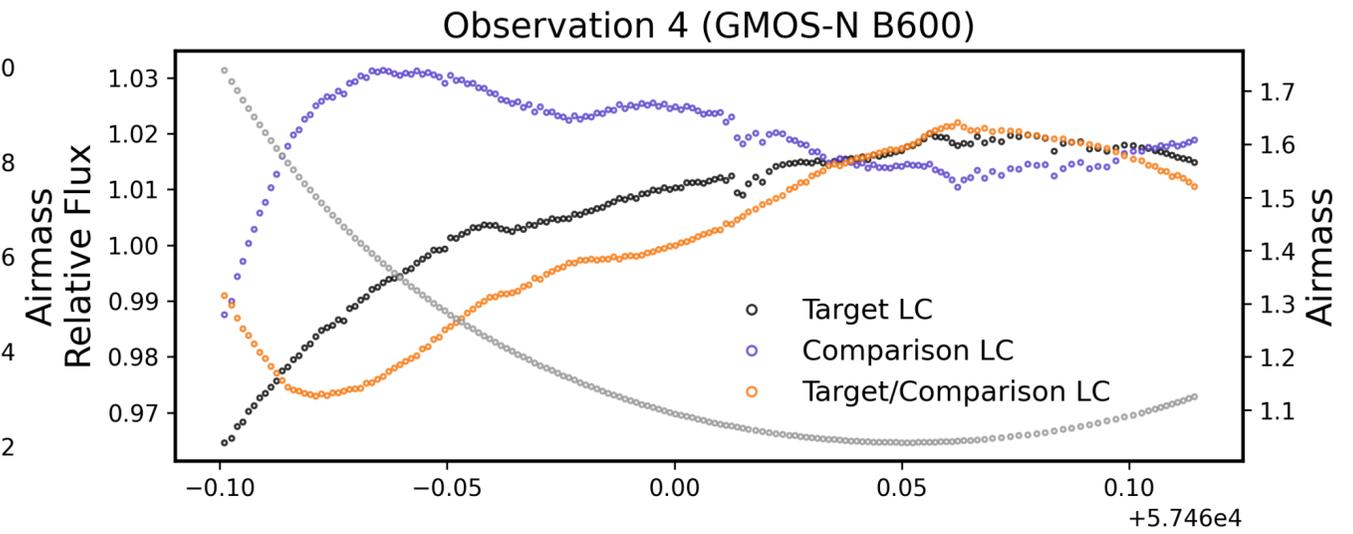
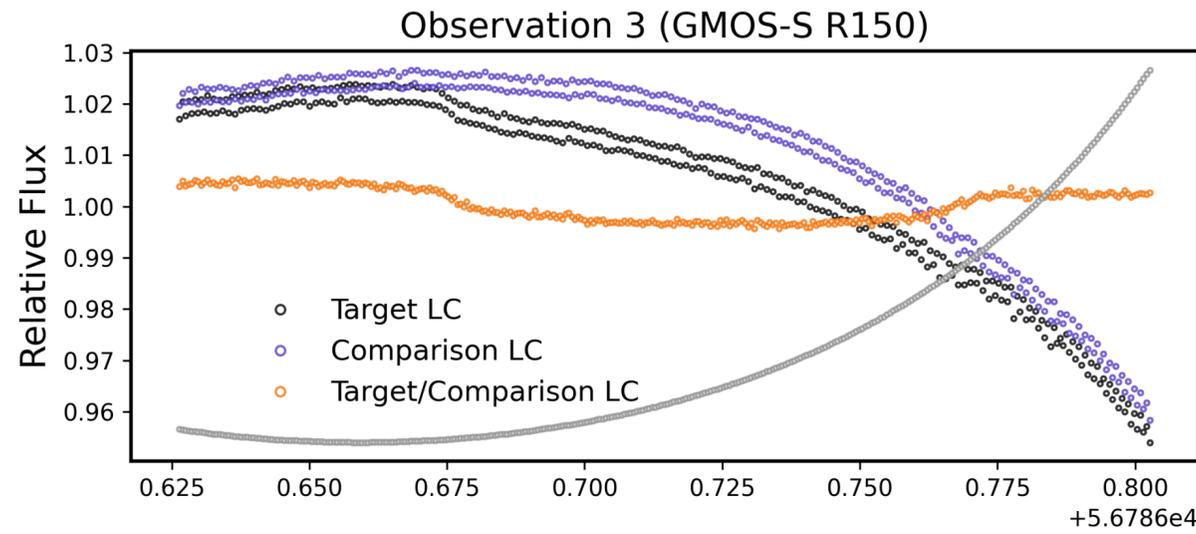
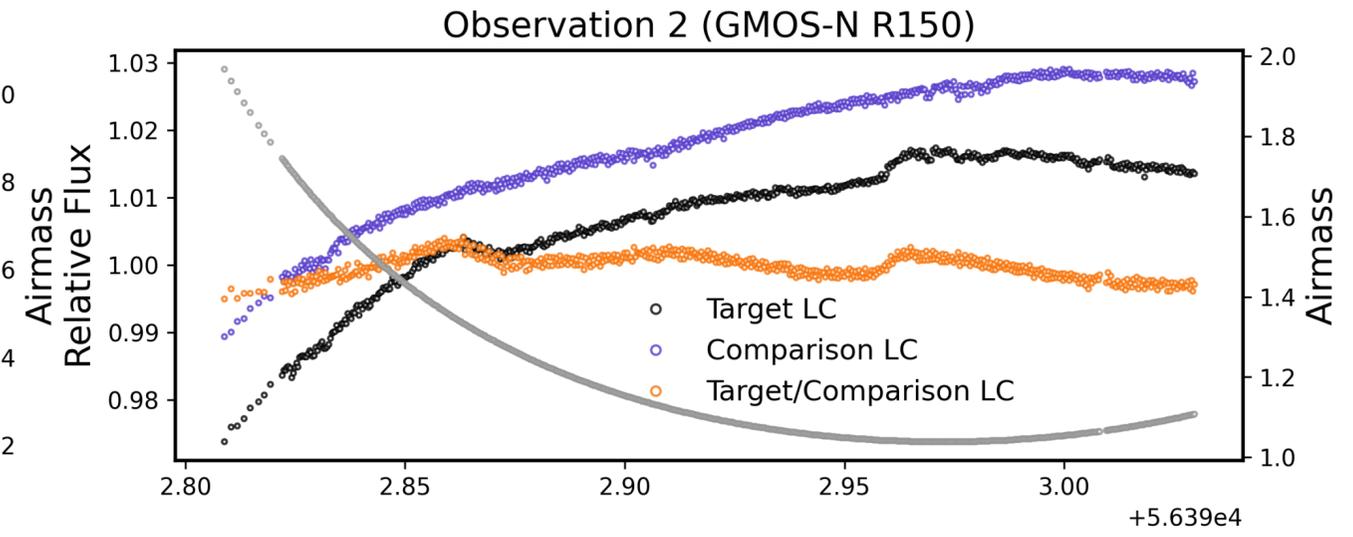
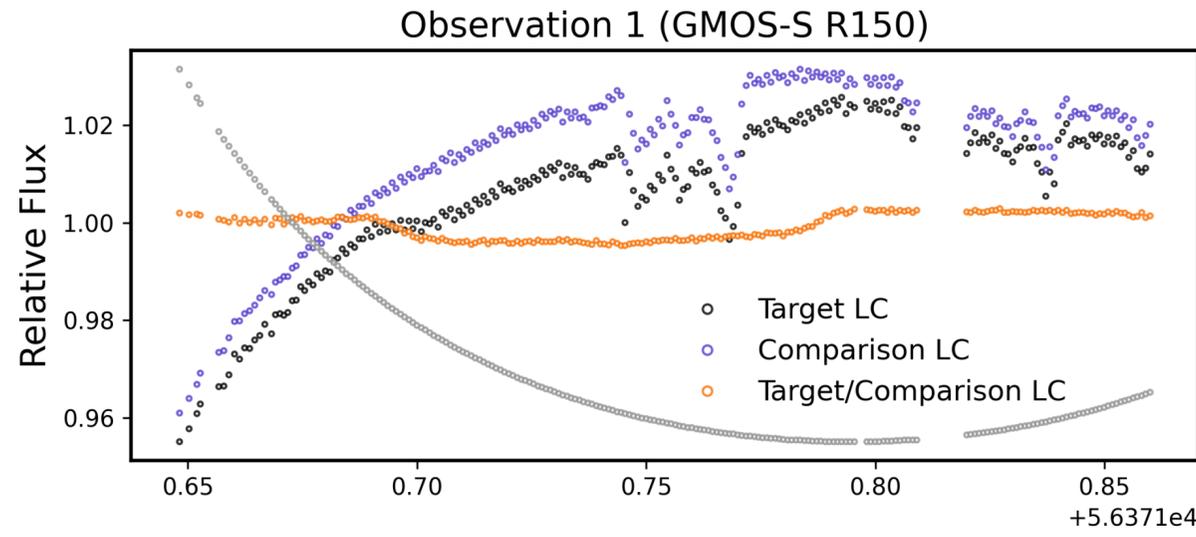
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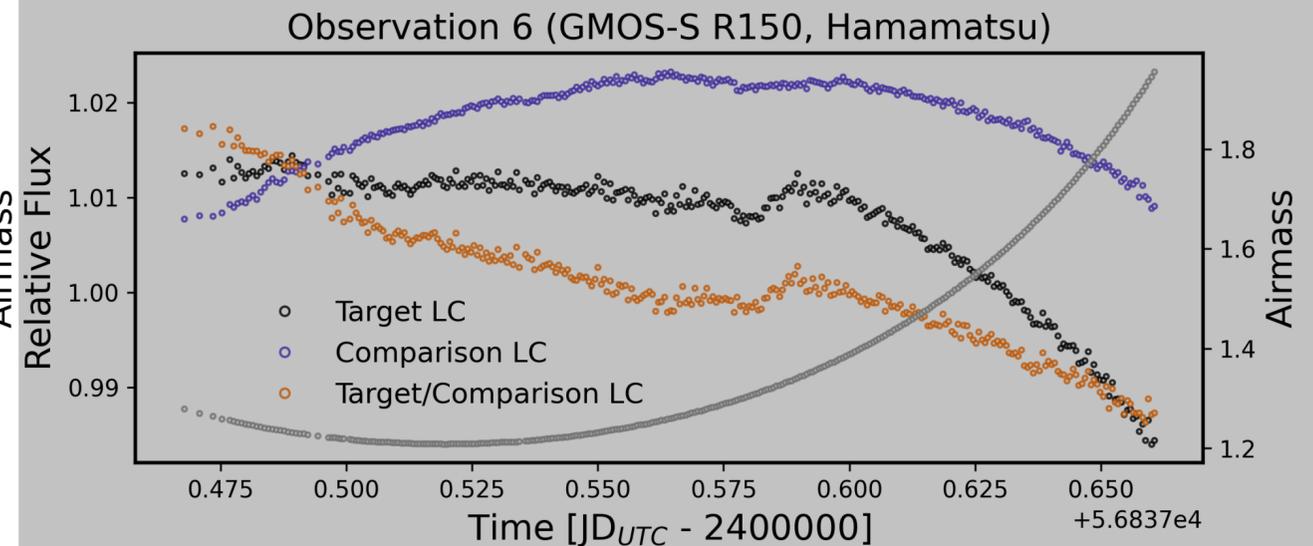
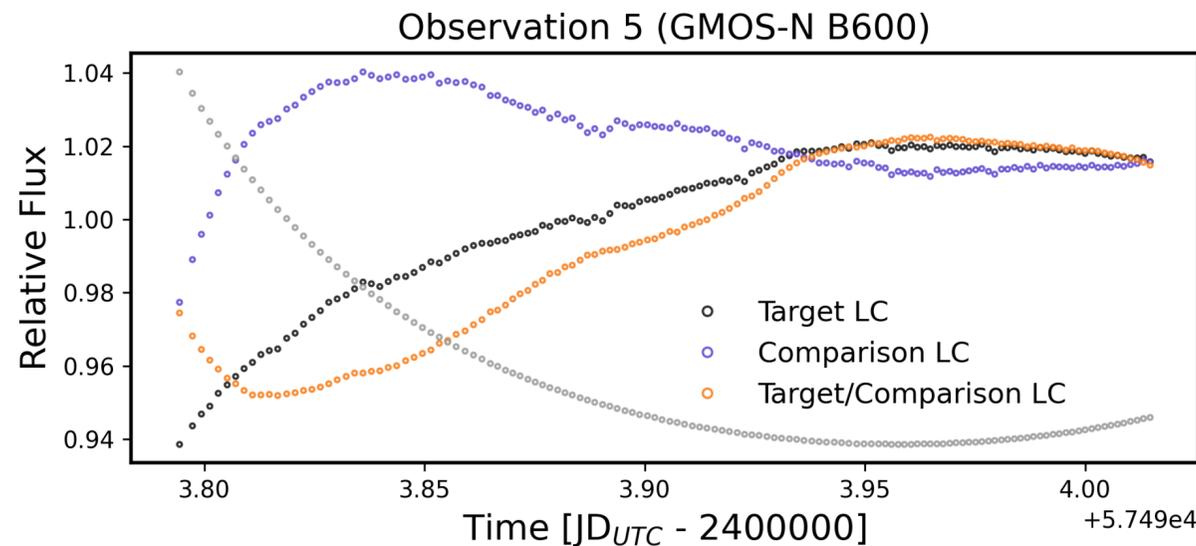
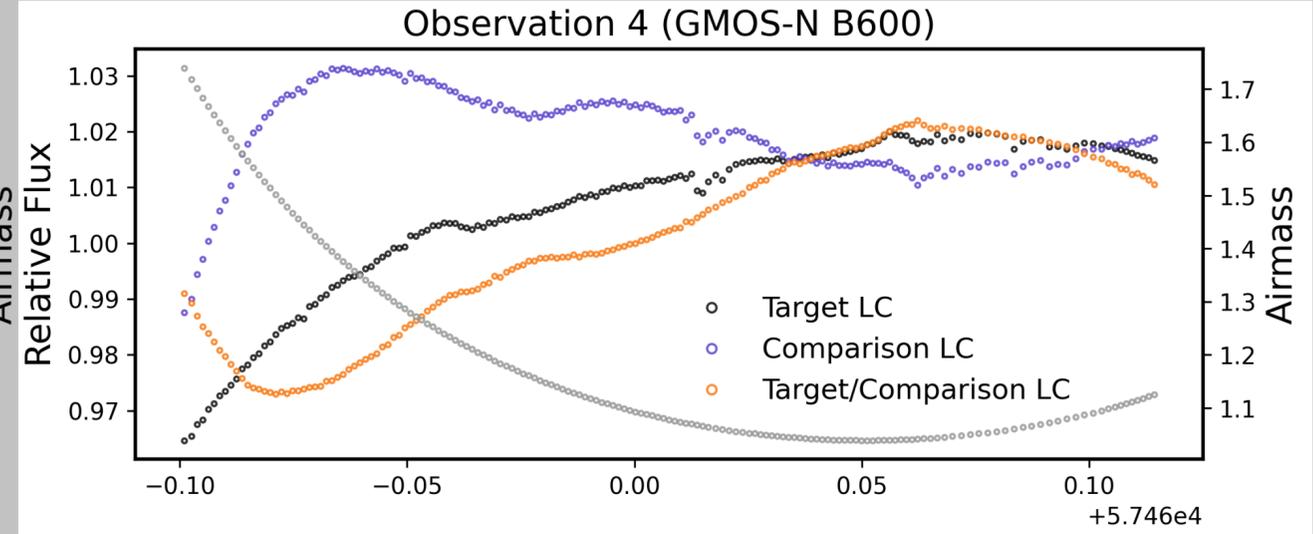
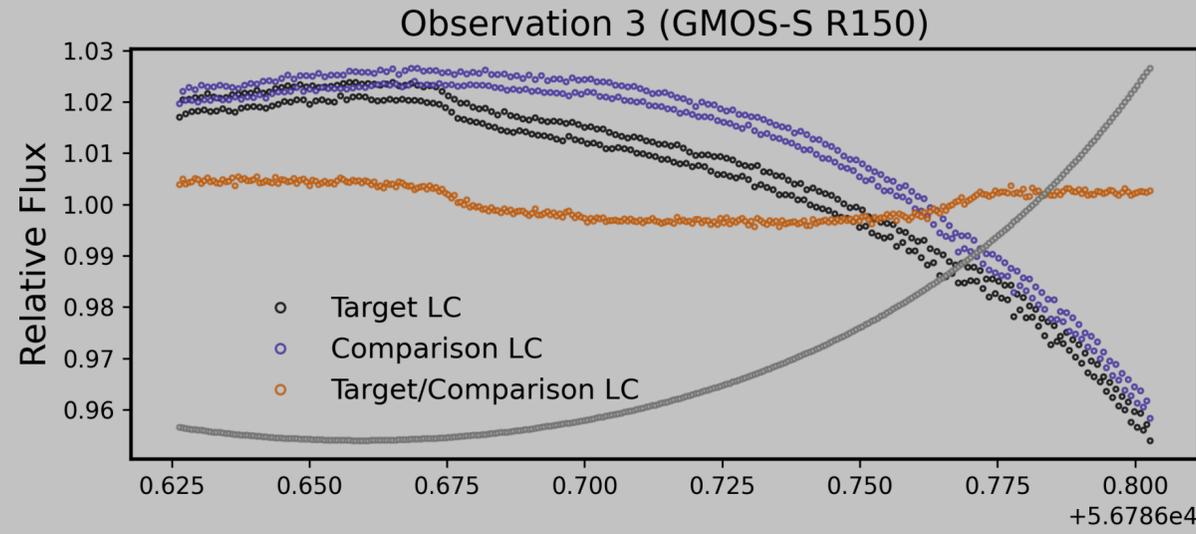
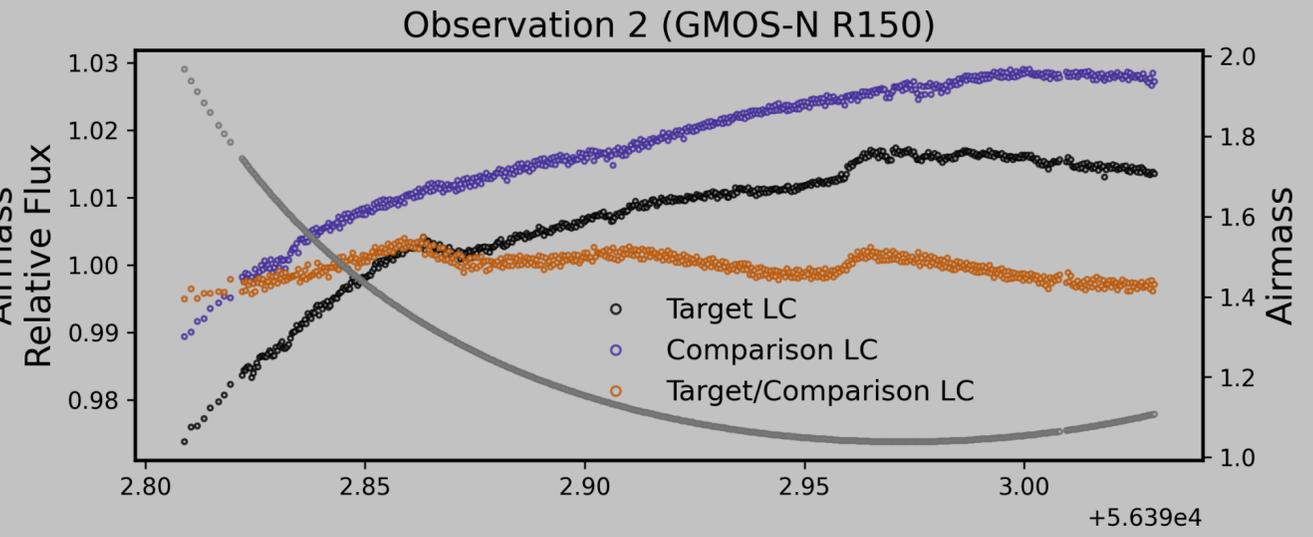
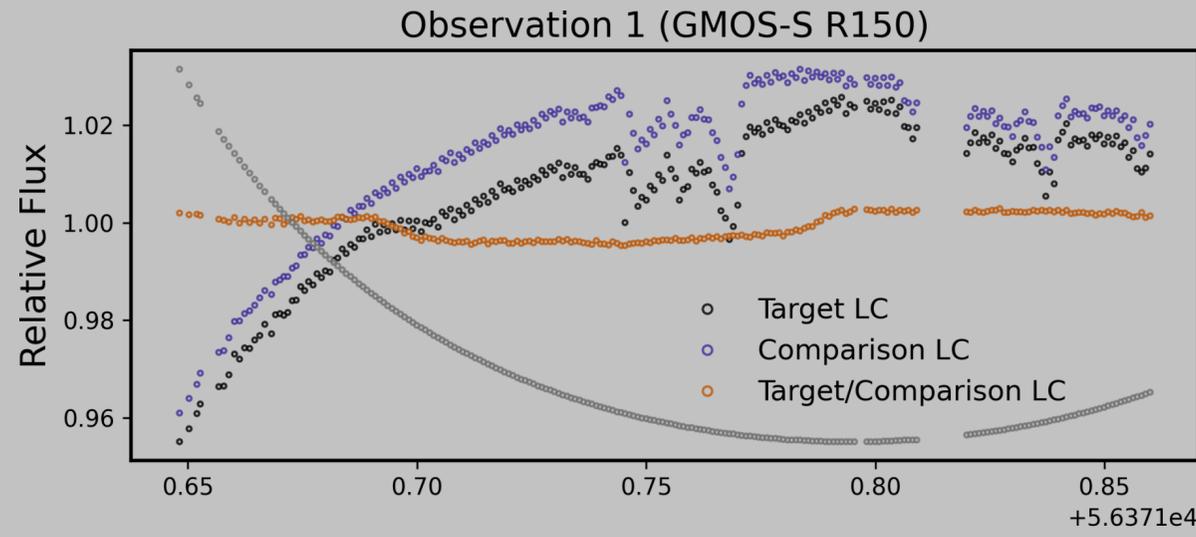
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Gemini/GMOS view of the warm Neptune HAT-P-26b

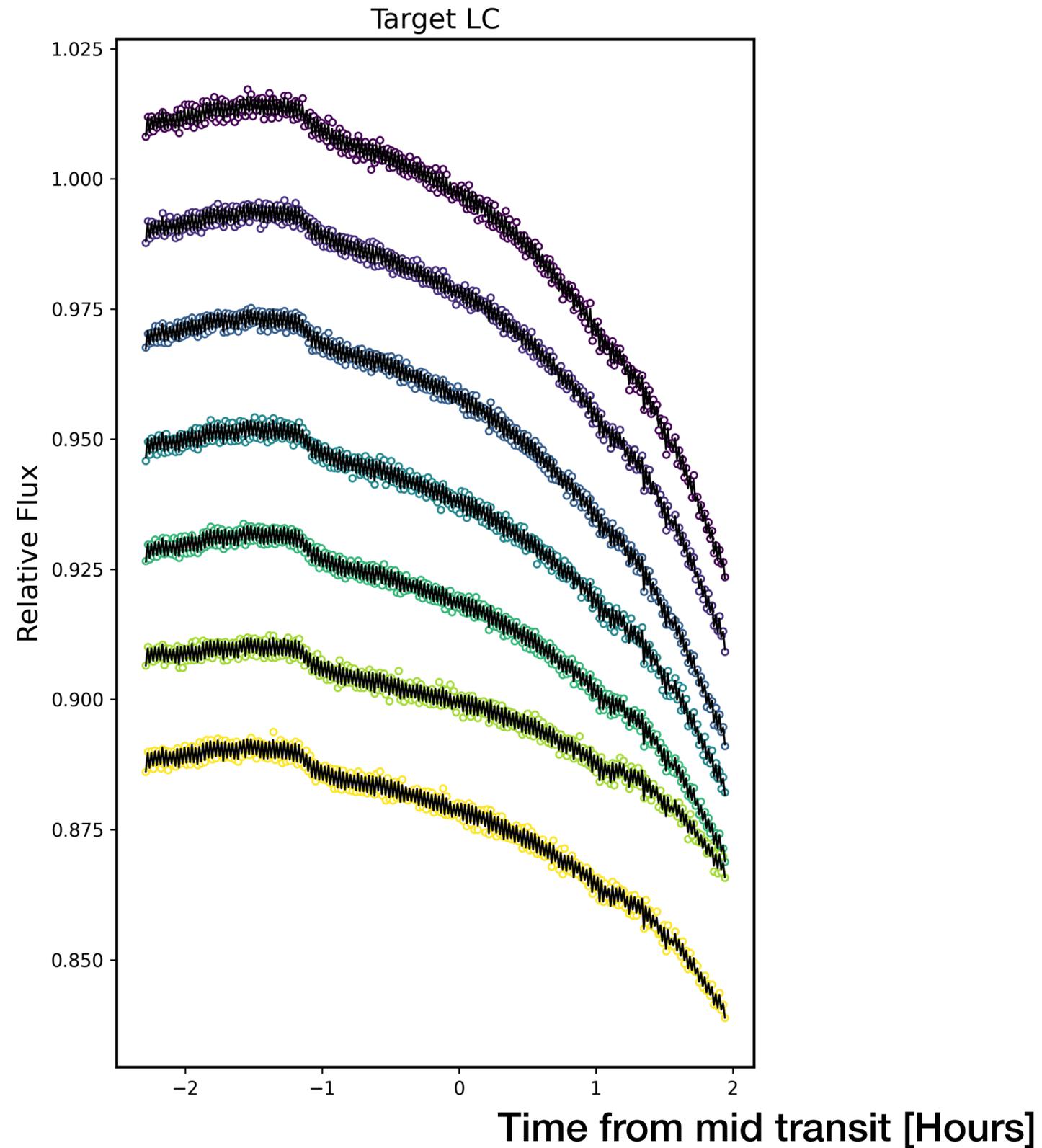


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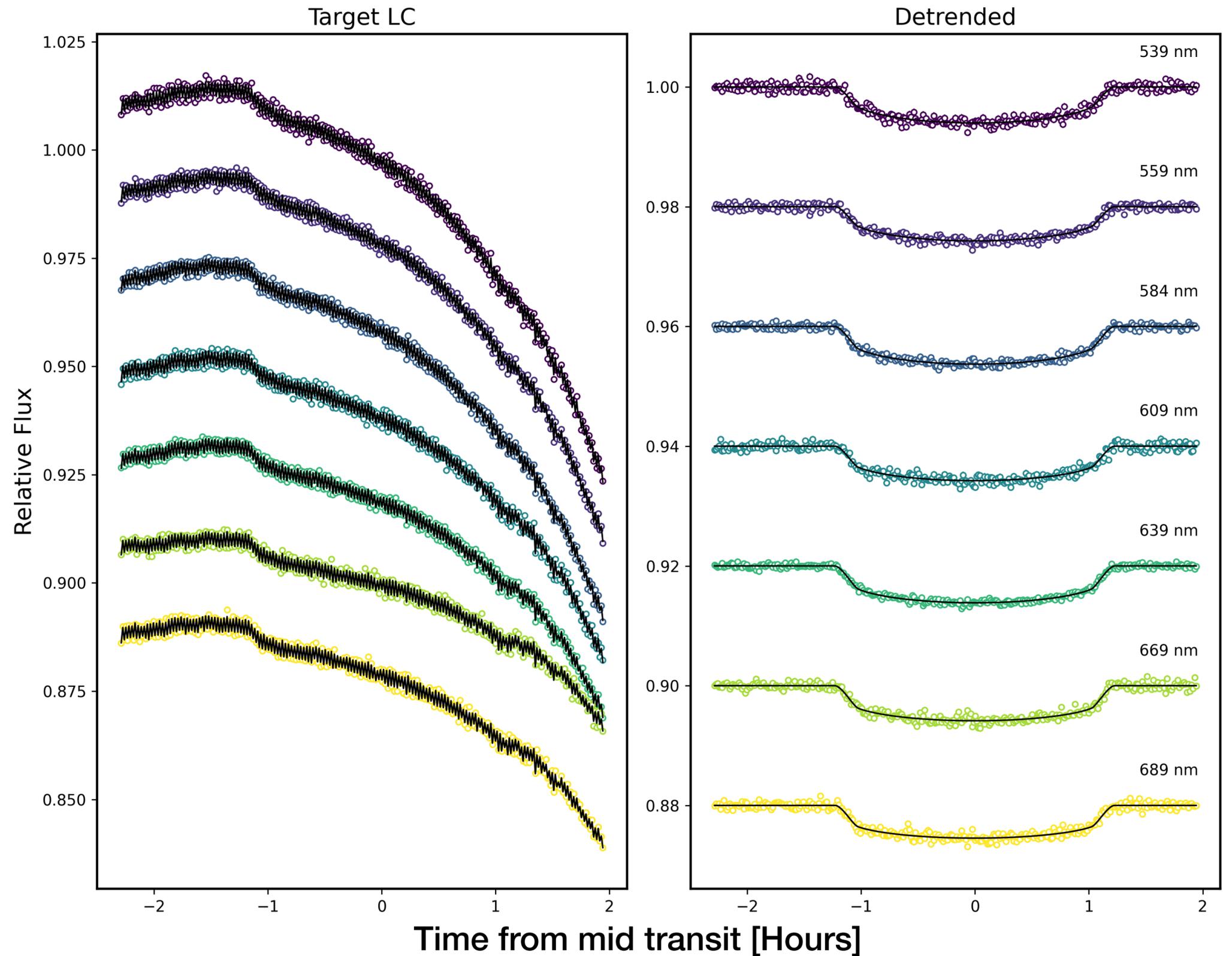
Extracting transmission spectrum using only the Target star

Correcting for
high and **low frequency**,
time and **wavelength**
dependent systematics

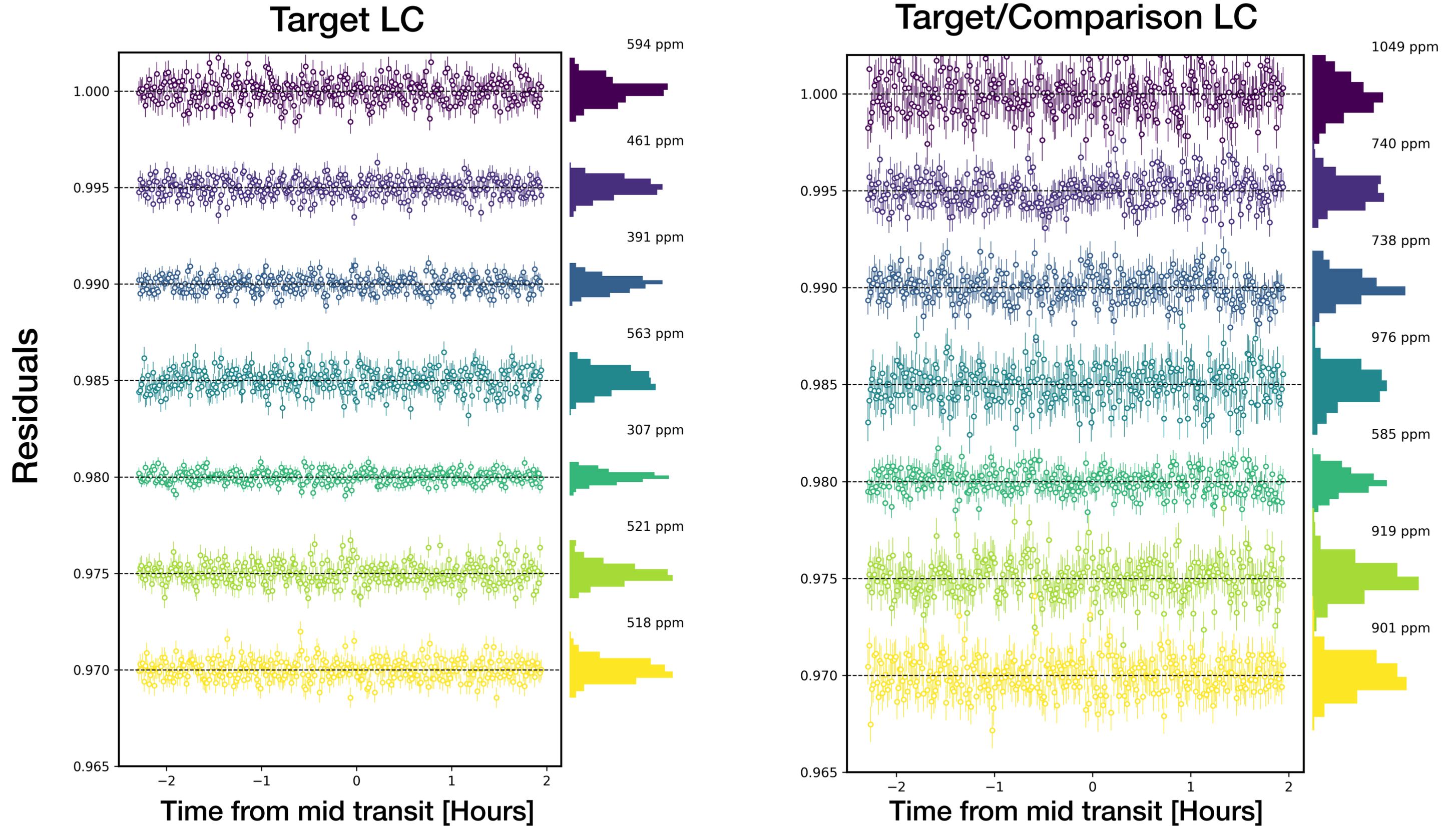


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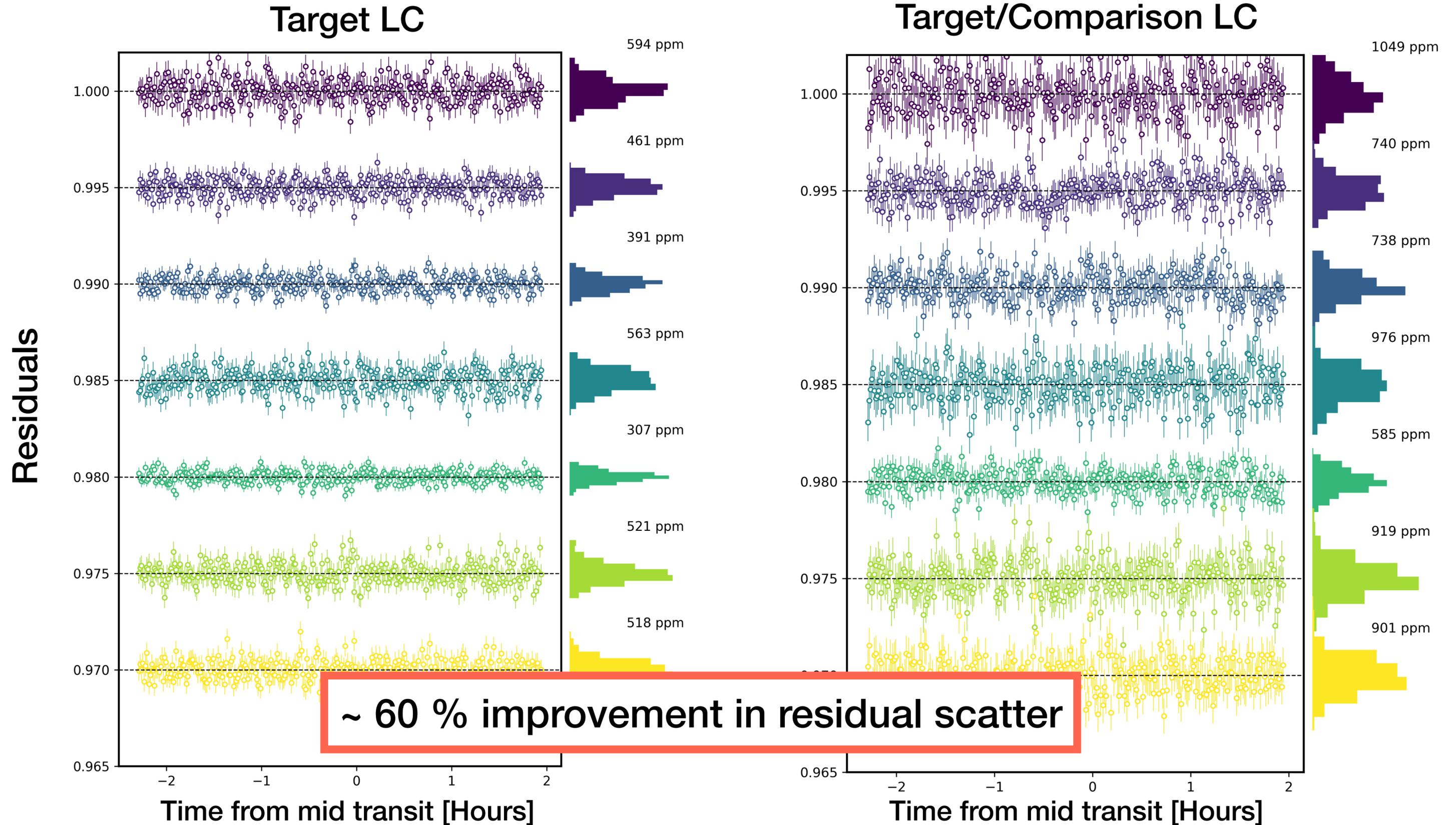
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Extracting transmission spectrum: new vs conventional method

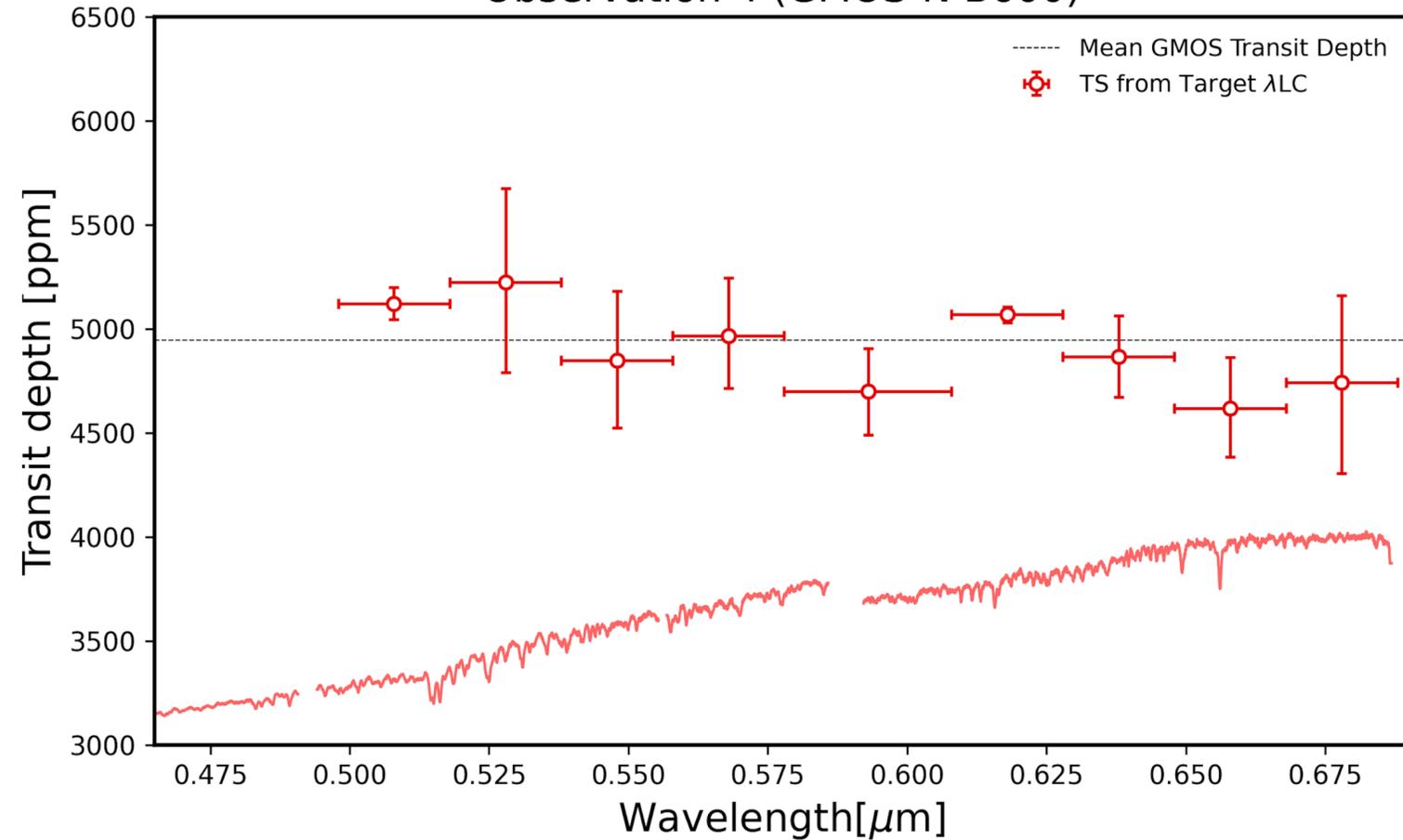


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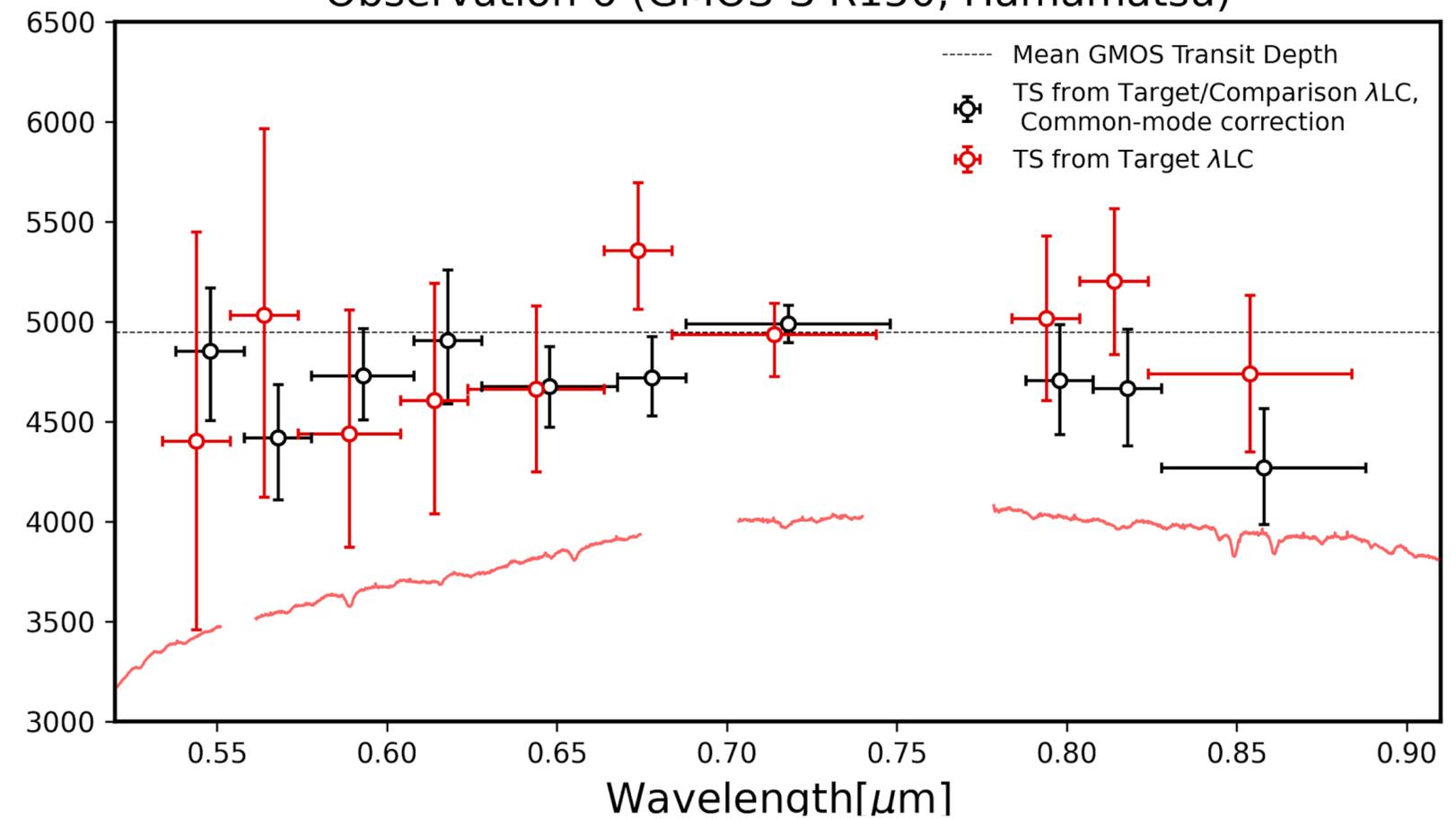


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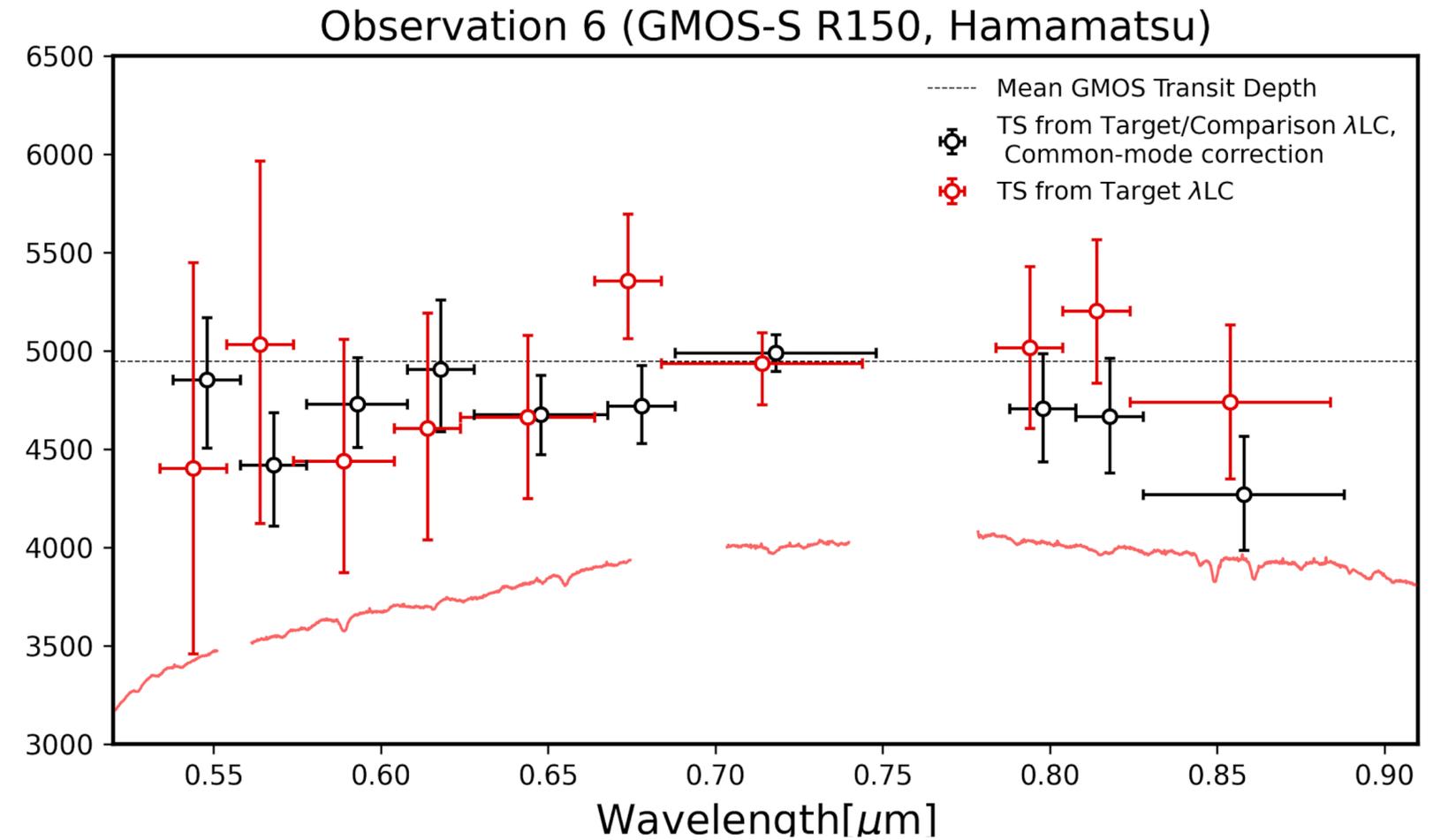
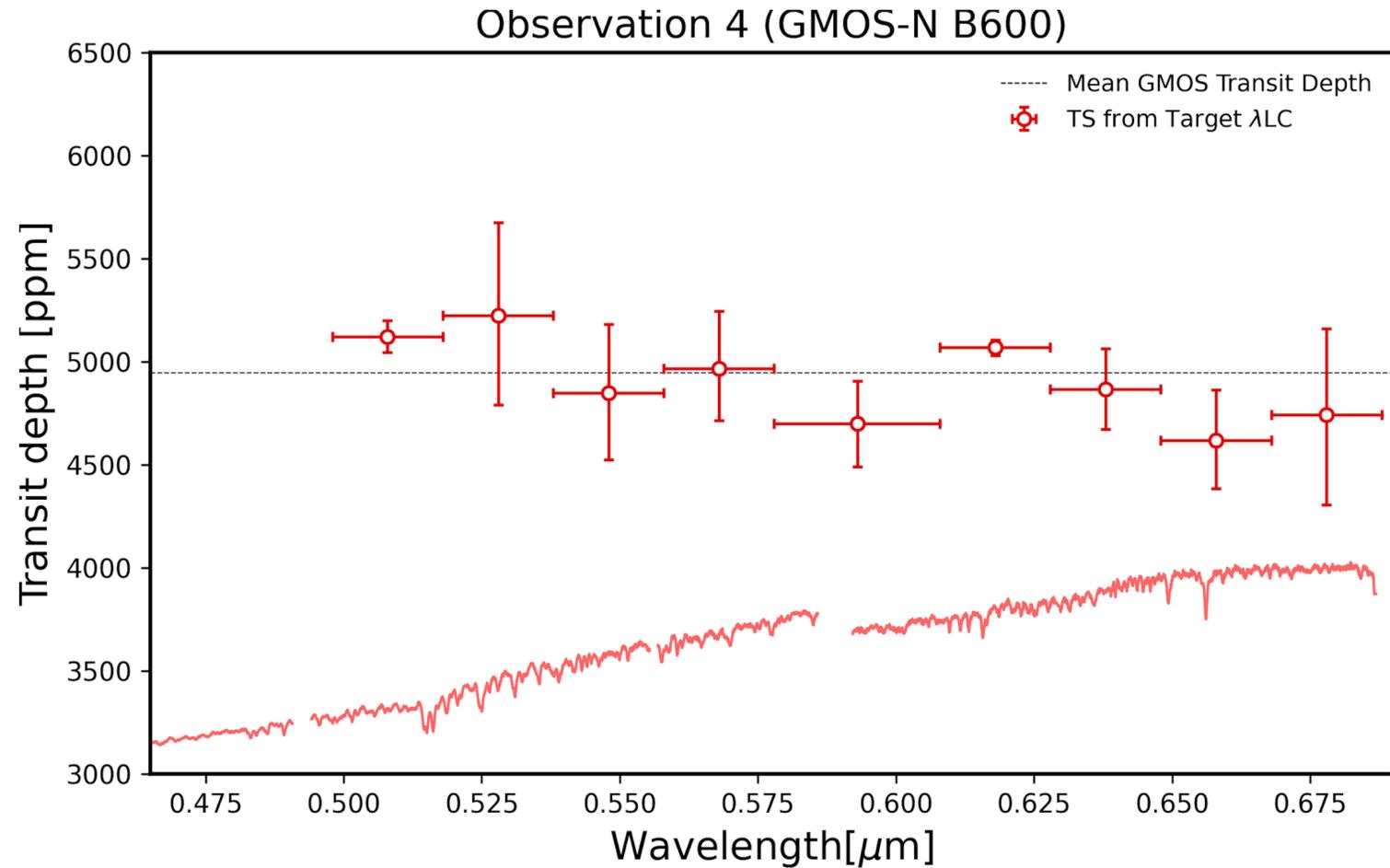
Observation 4 (GMOS-N B600)



Observation 6 (GMOS-S R150, Hamamatsu)

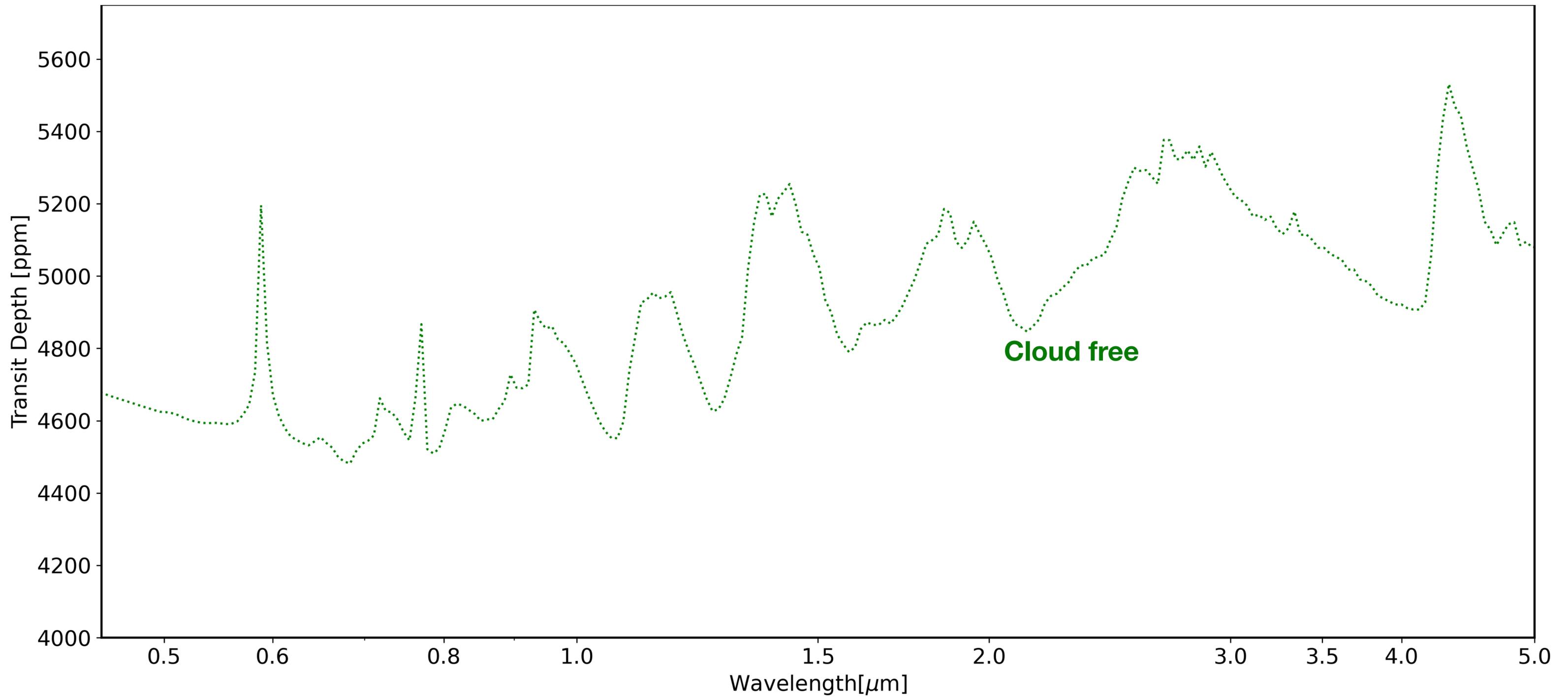


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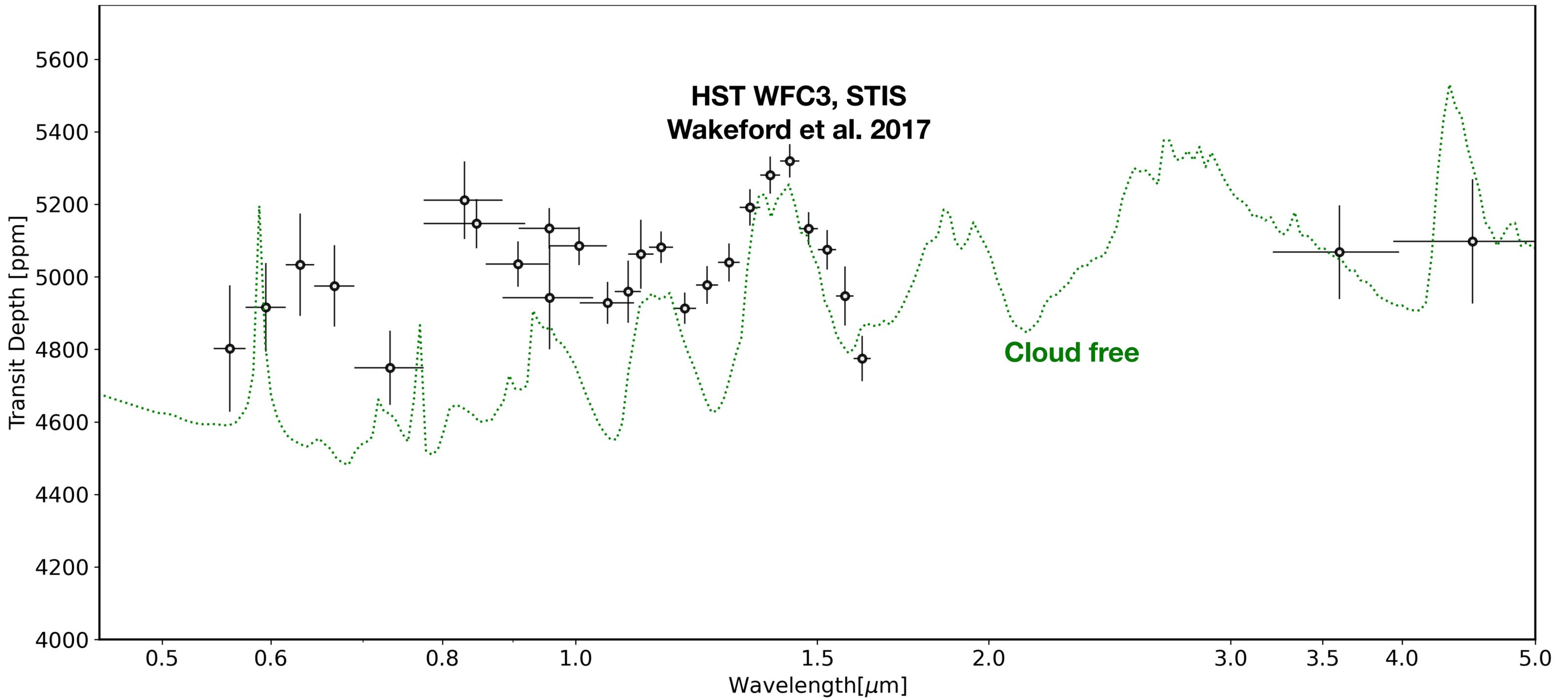


Propagating uncertainties from common mode correction within Bayesian framework of GPs

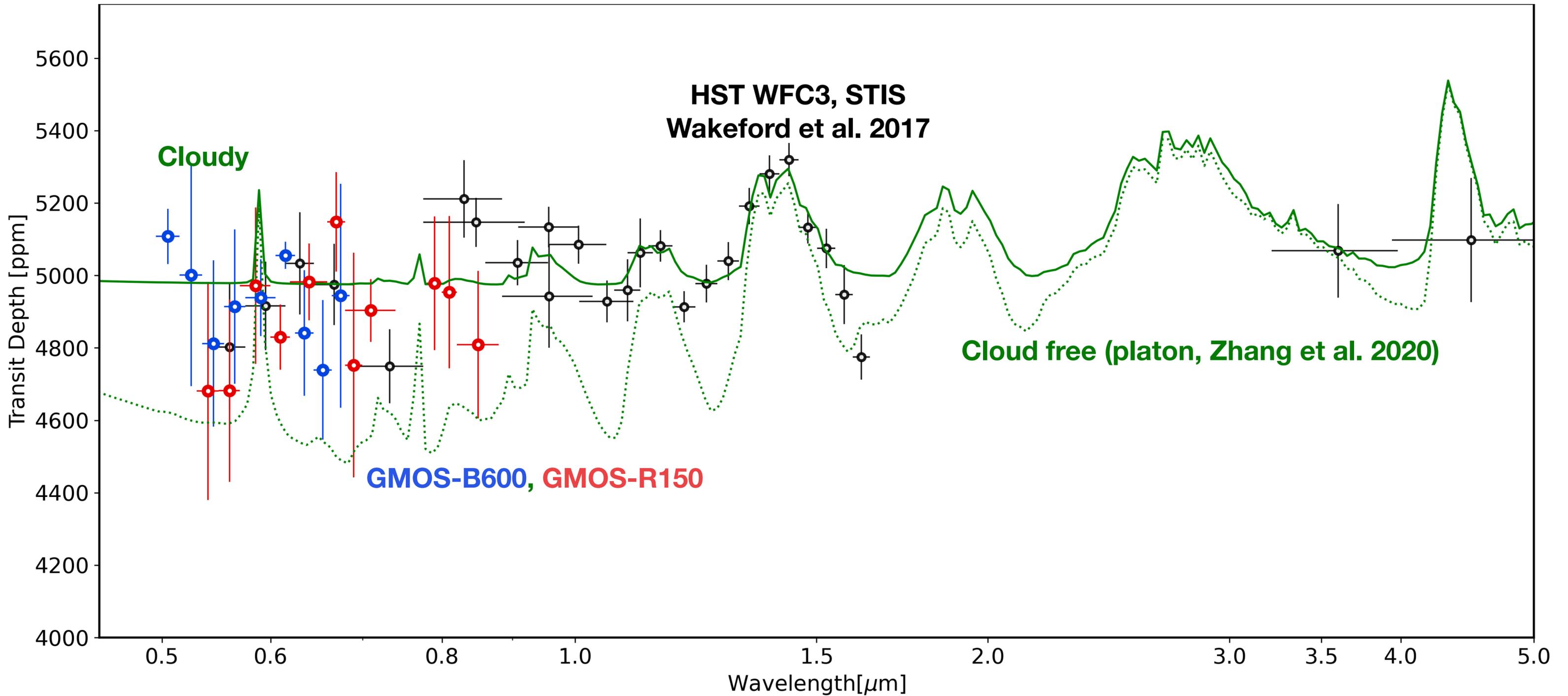
Gemini/GMOS transmission spectrum of the warm Neptune HAT-P-26b



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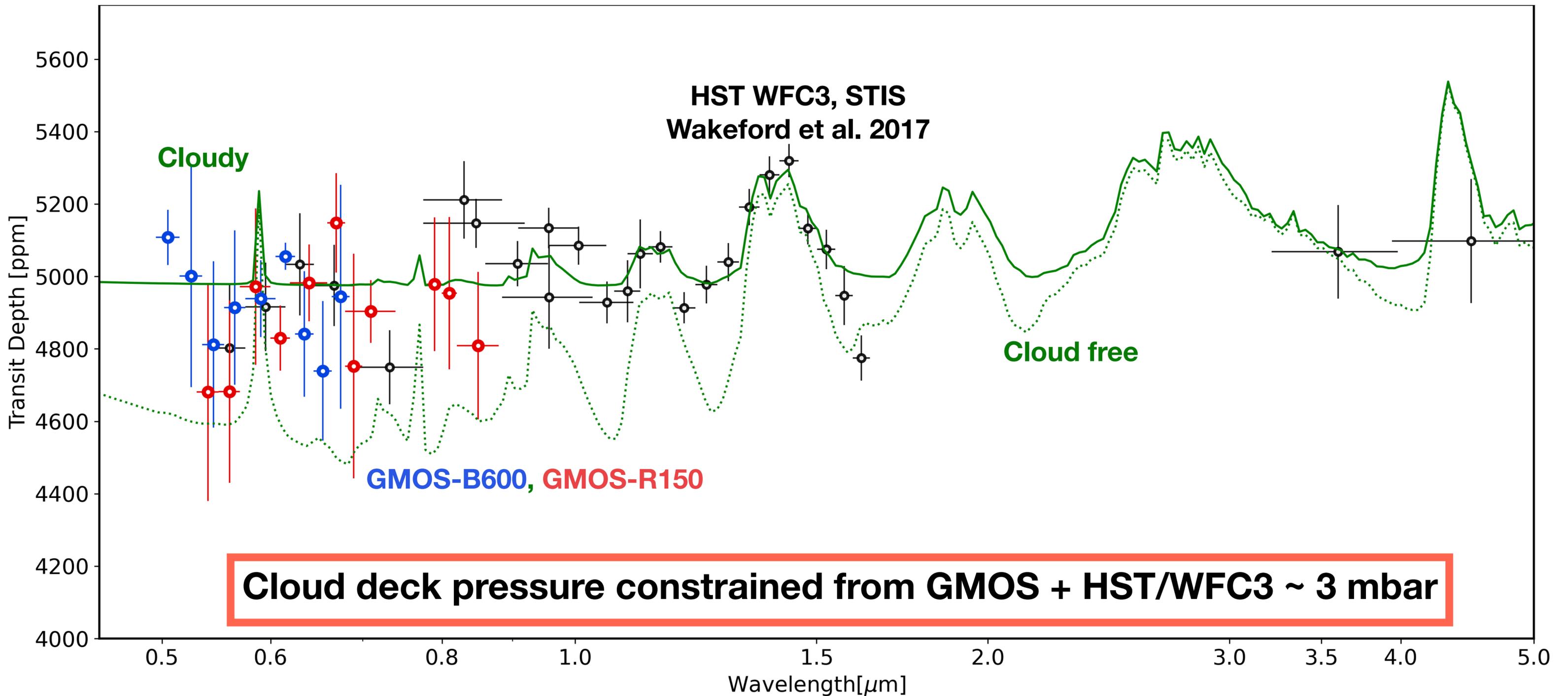


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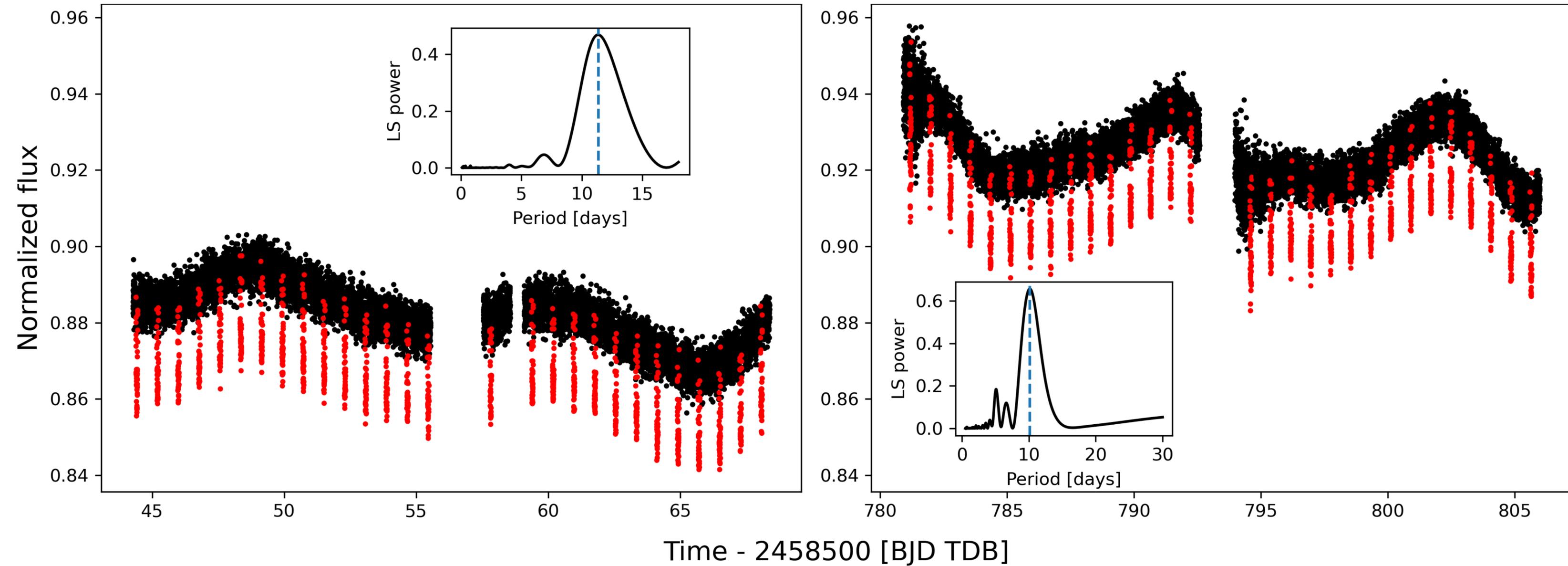
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Constraining the cloud deck pressure level



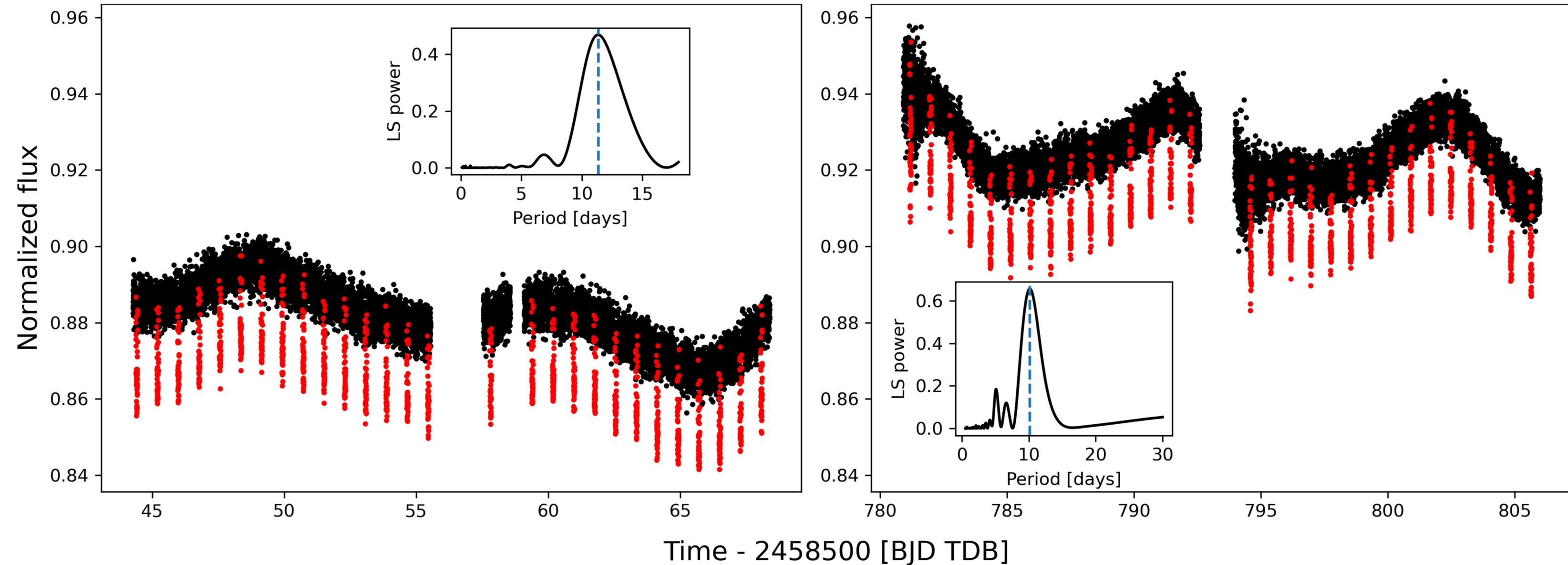
Measuring Accurate transit depth necessary for active host stars

WASP-19b observed by
TESS



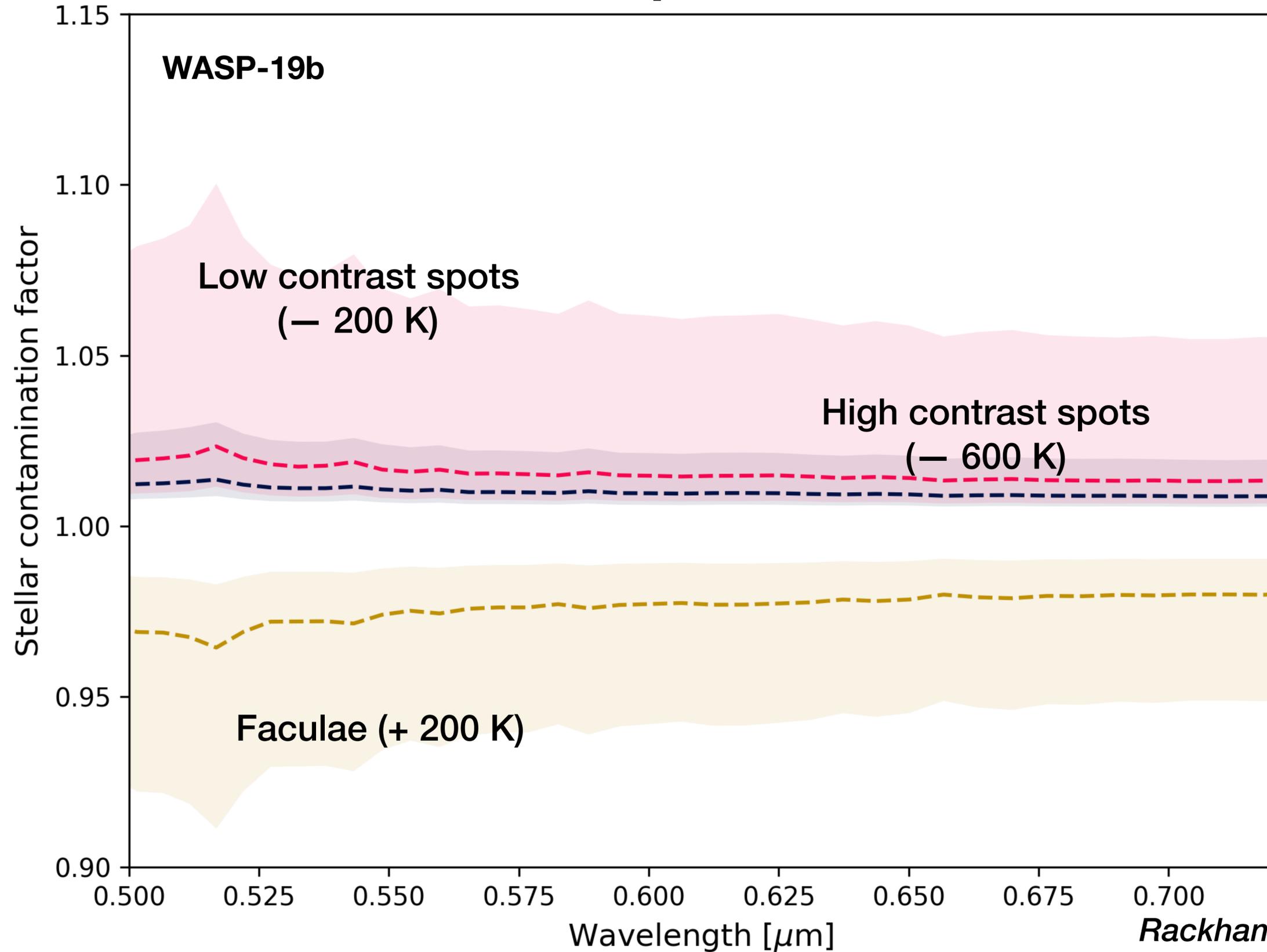
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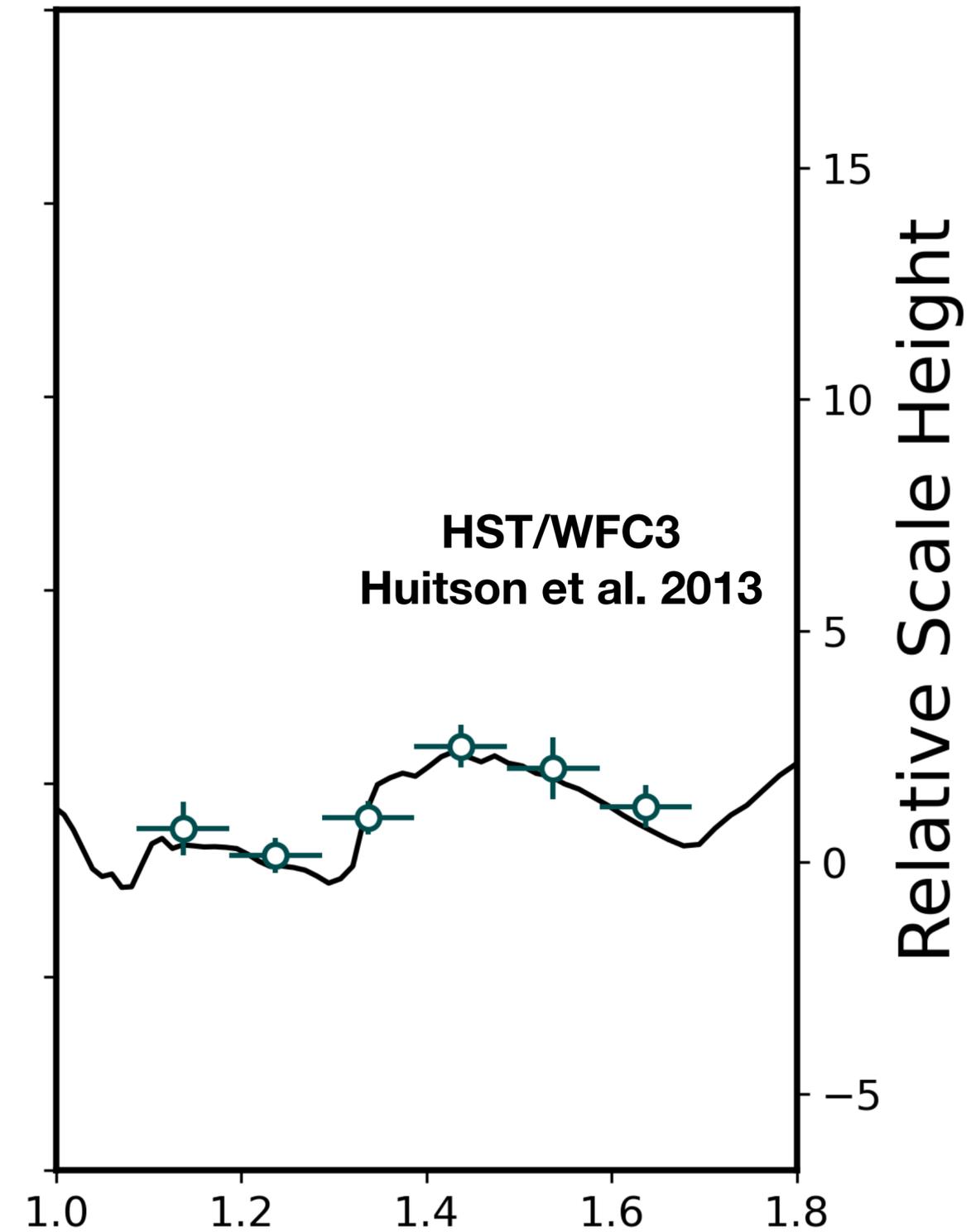
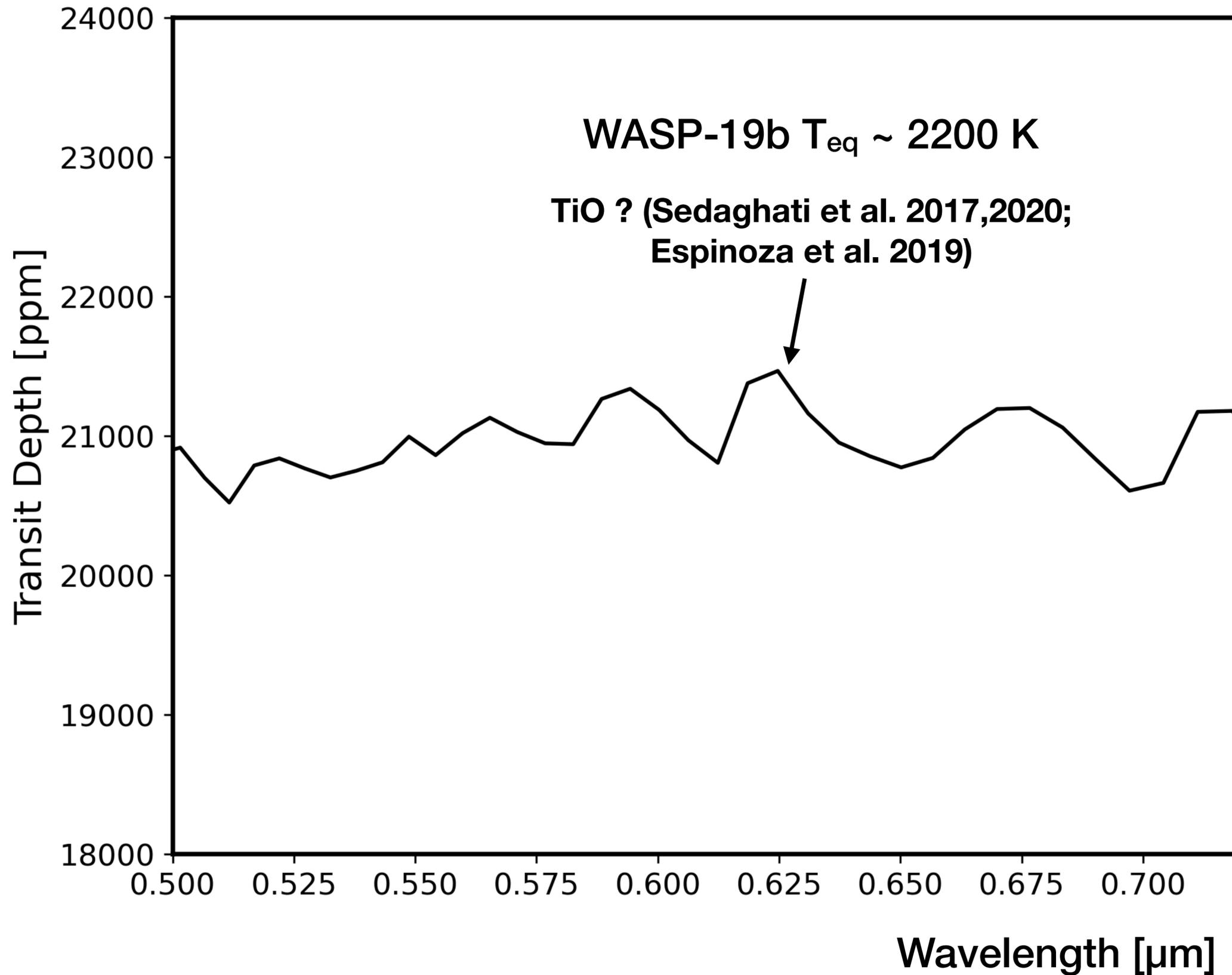
How does this change planet's spectrum?

Contamination of transmission spectrum due to stellar spots/faculae

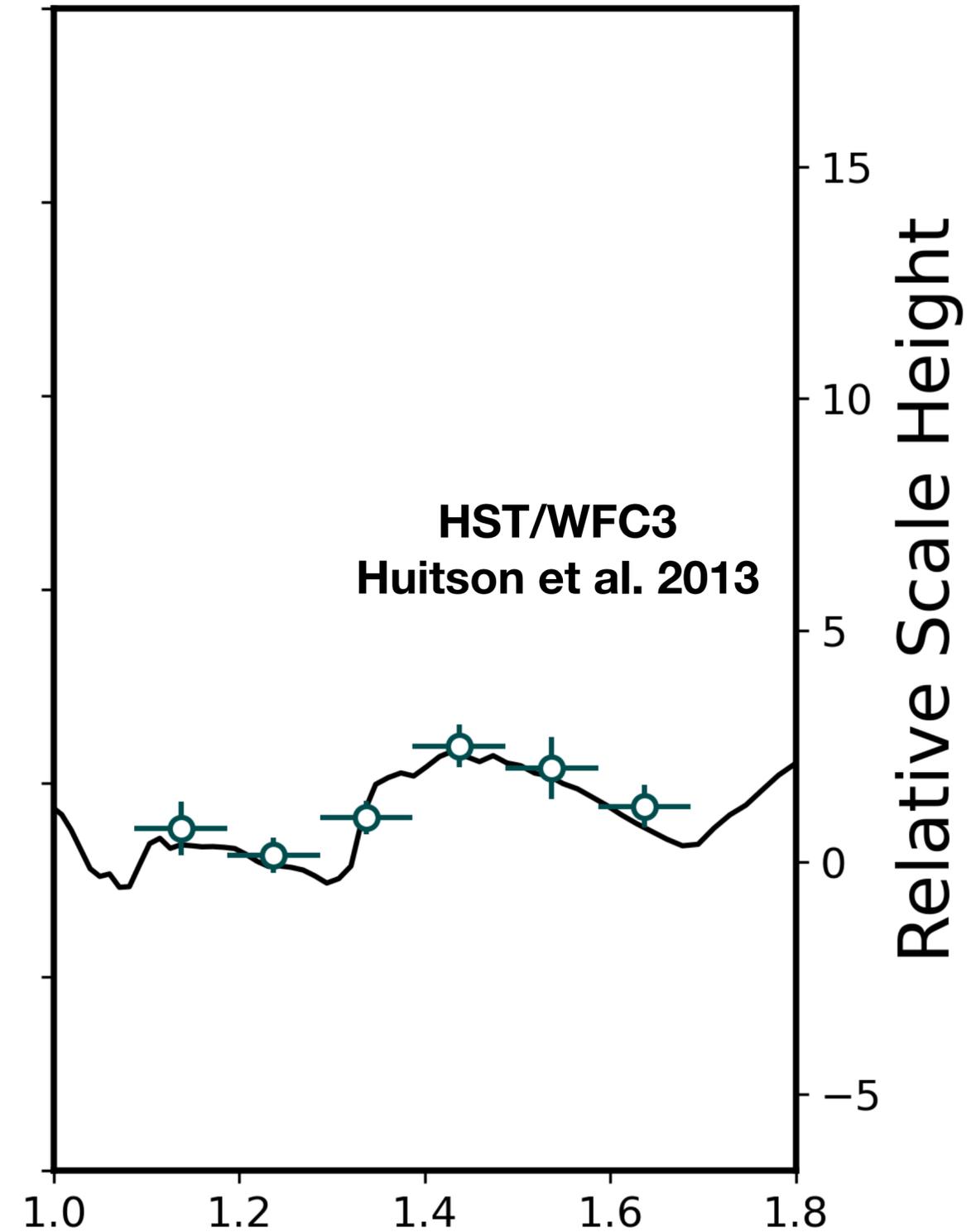
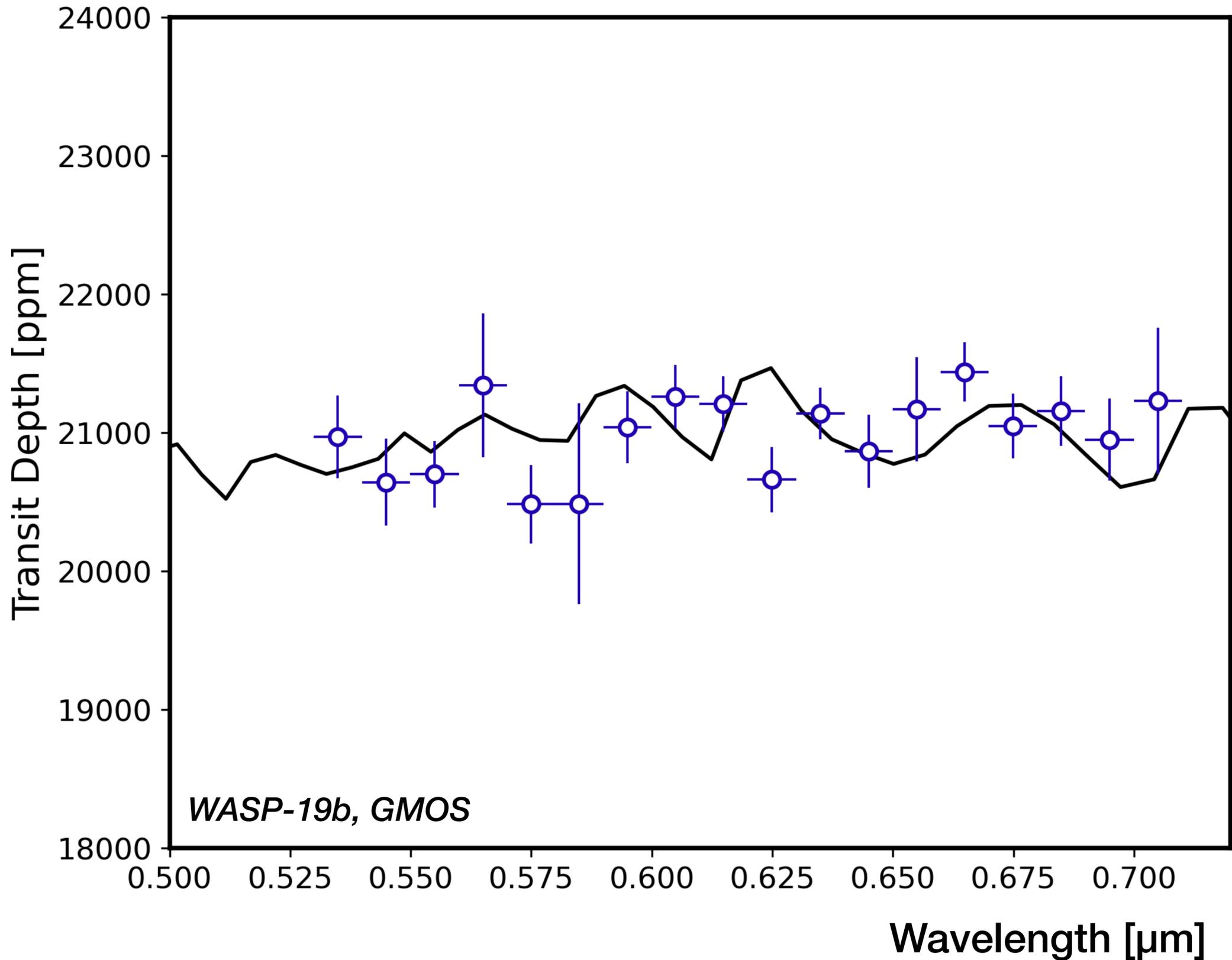


*Rackham et al. 2018,
Stellar spot parameters from Espinoza et al. 2019*

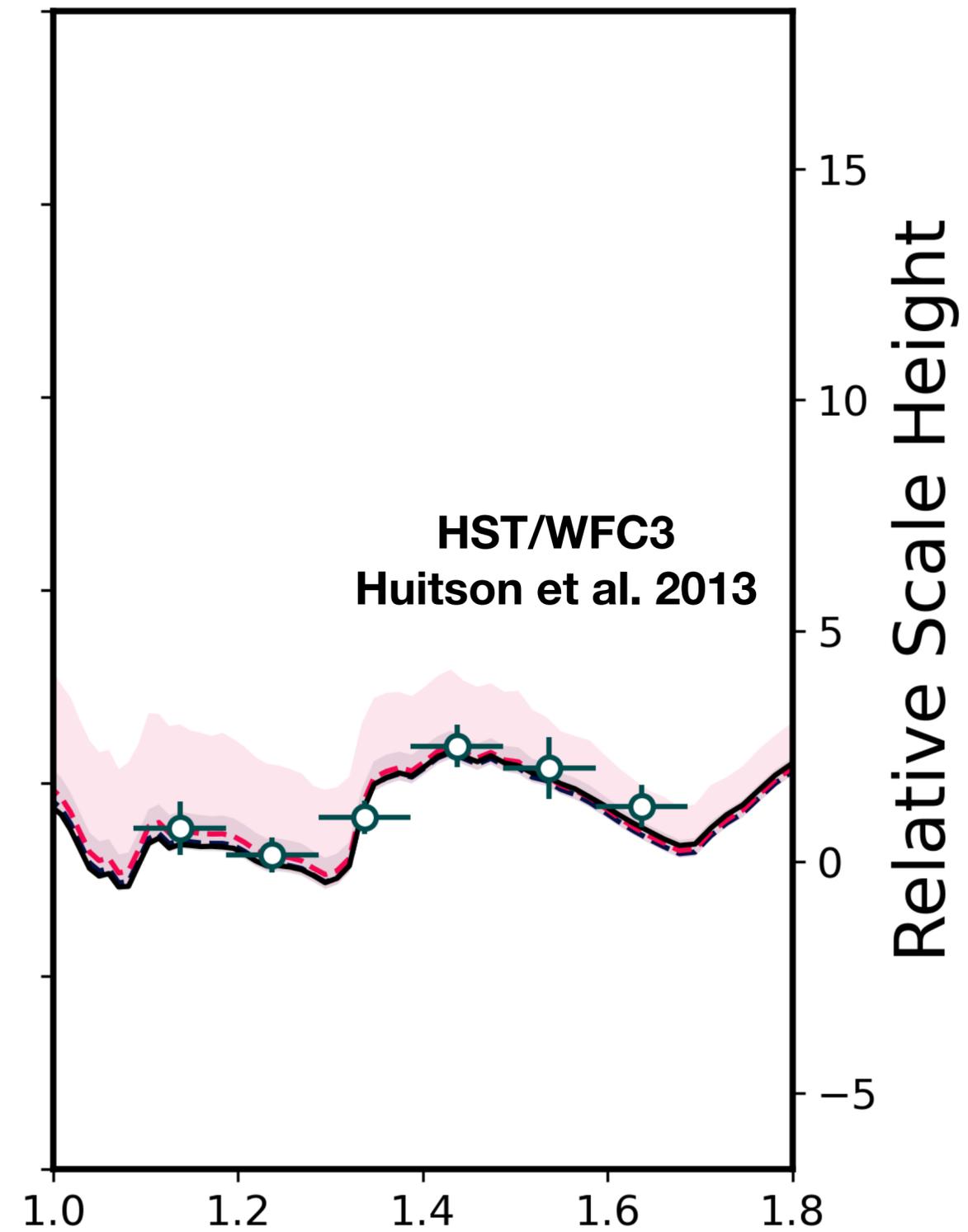
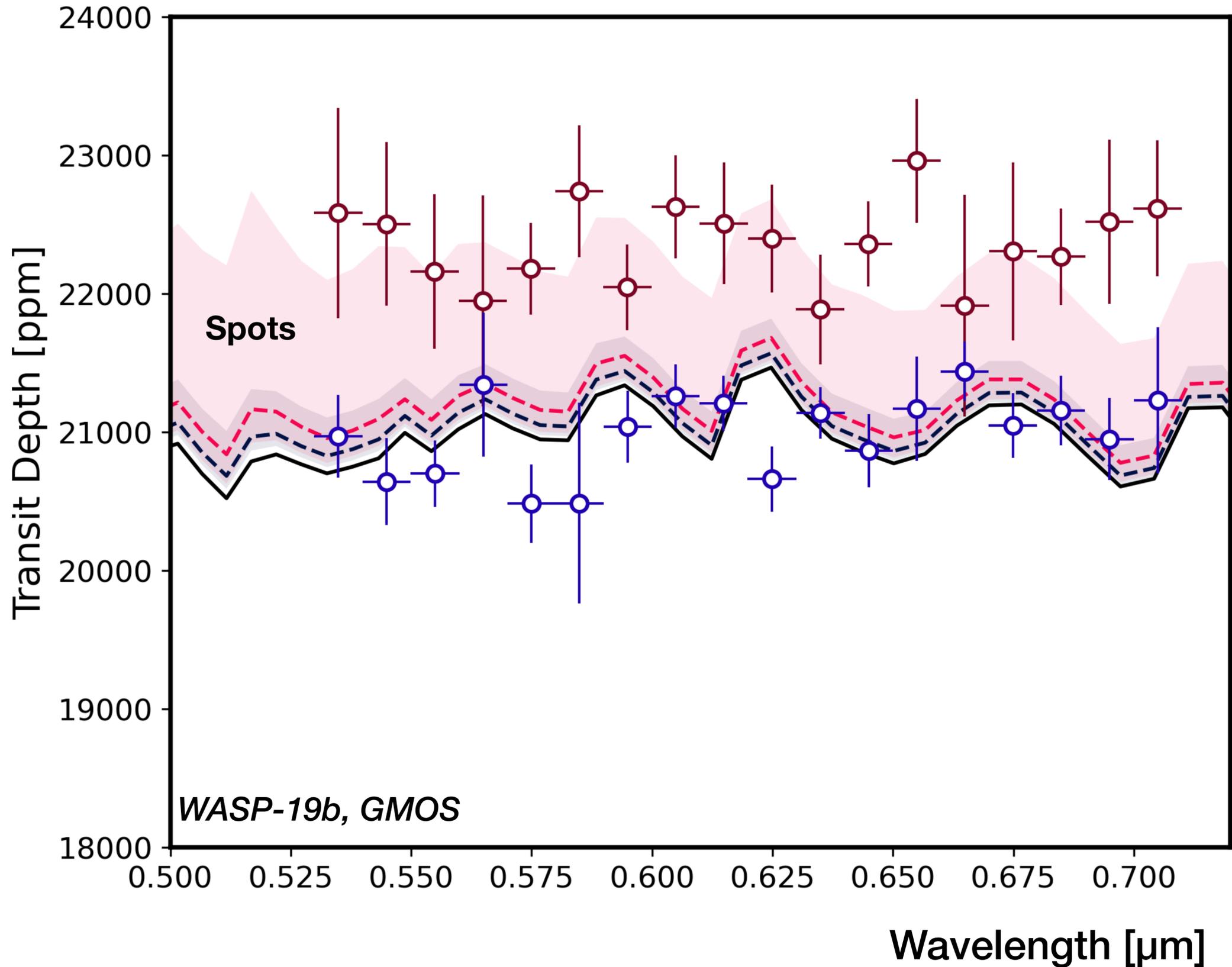
WASP-19b observed by Gemini/GMOS over multiple epochs



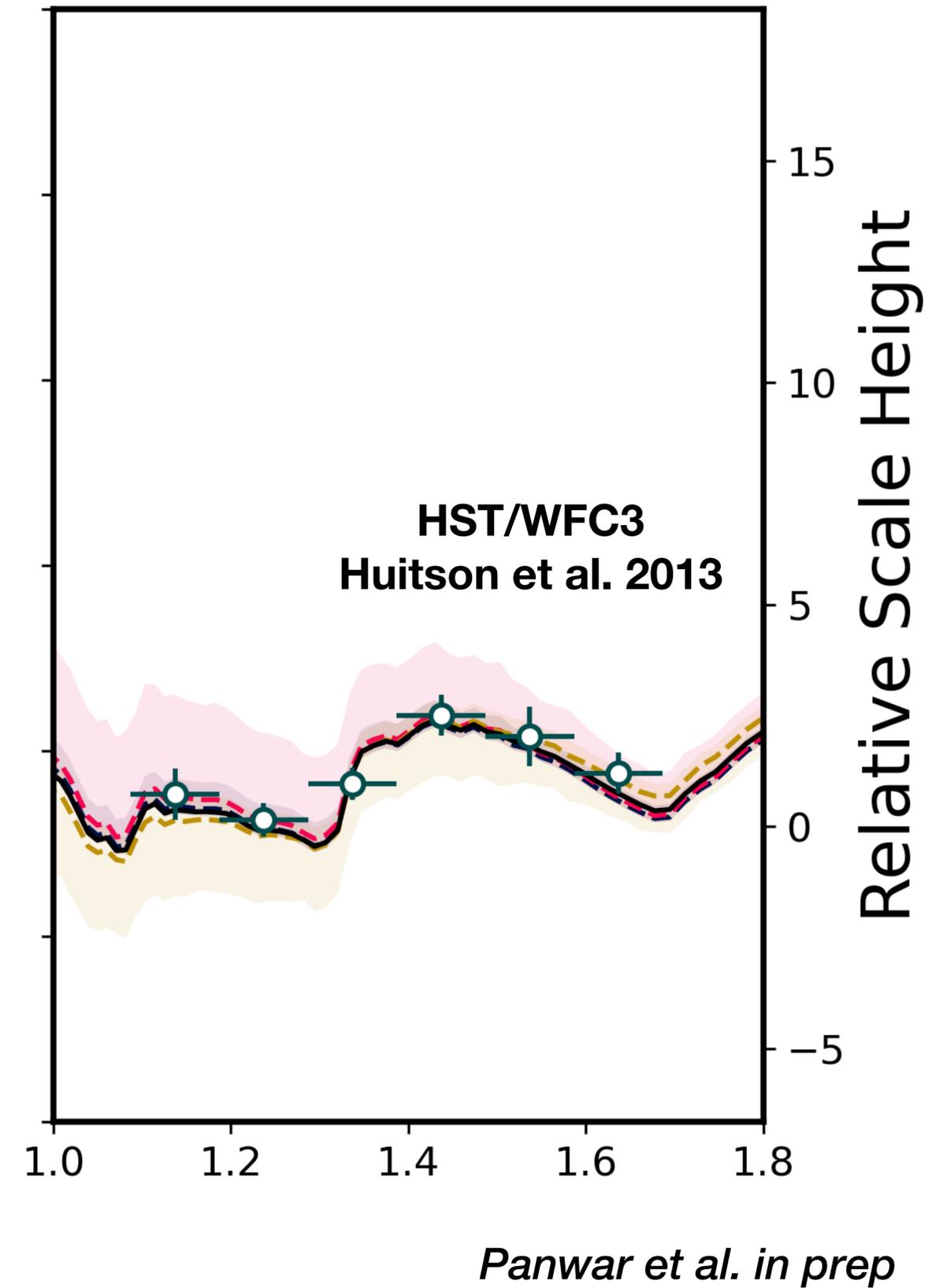
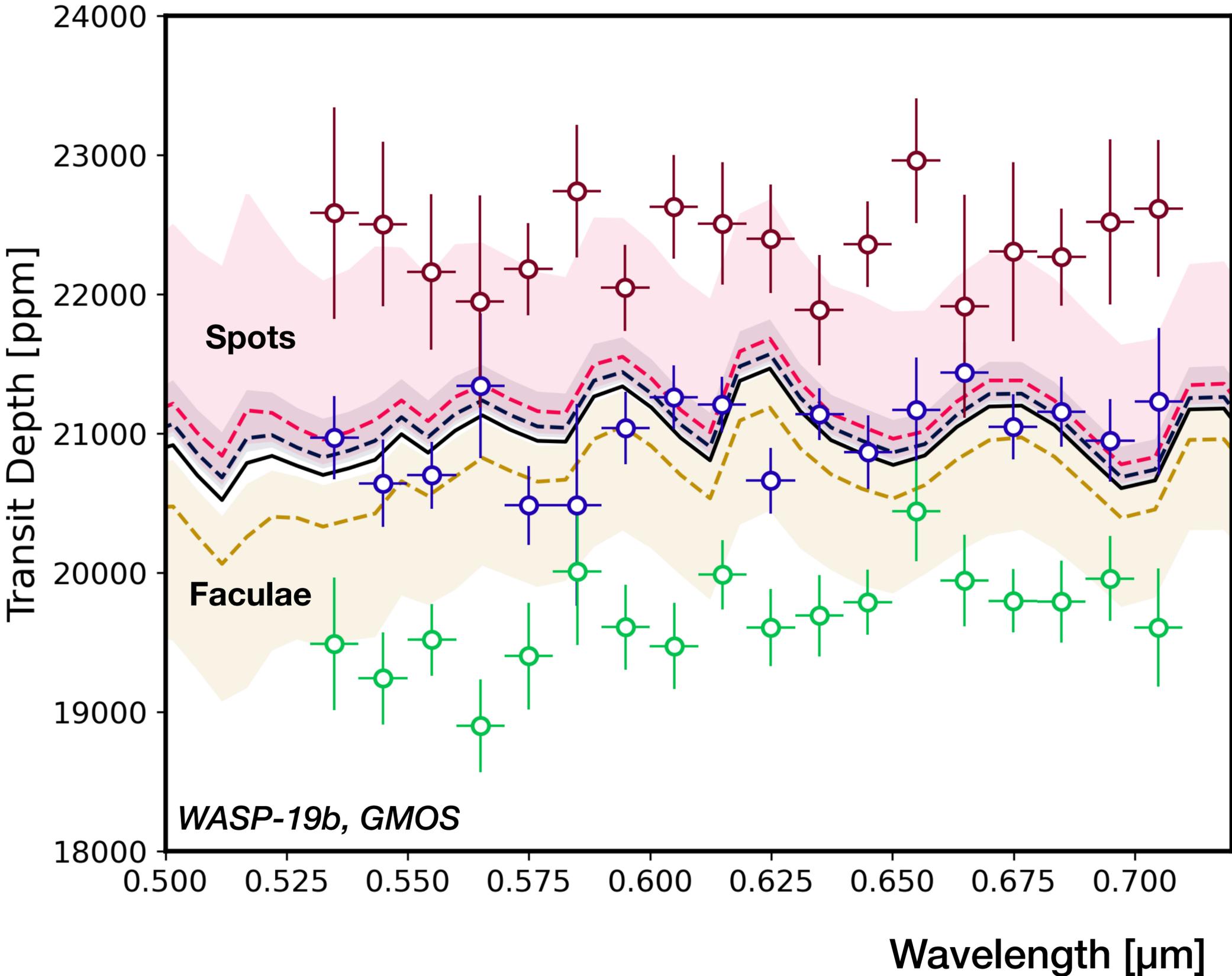
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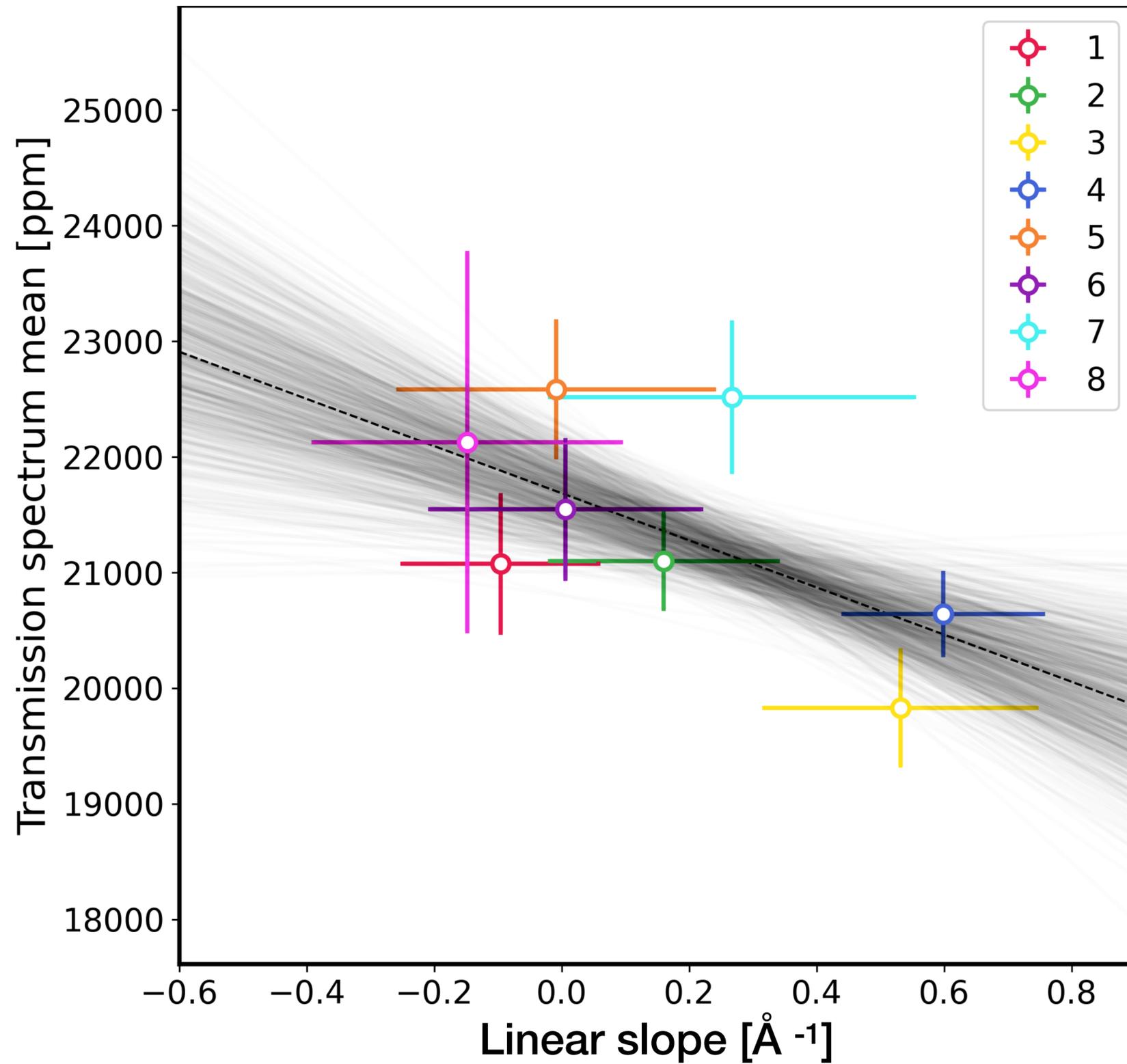
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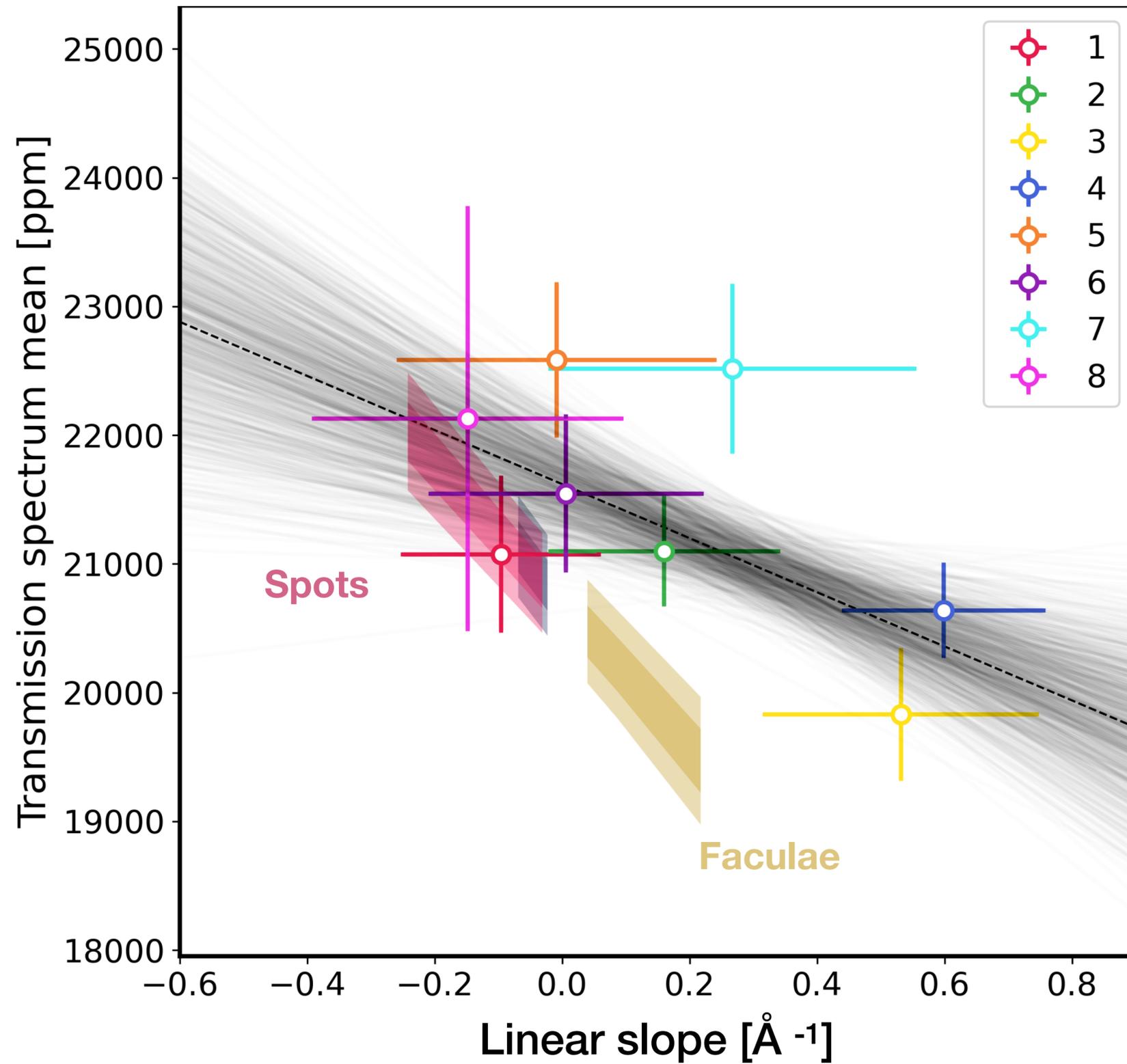
WASP-19b observed by Gemini/GMOS over multiple epochs



Stellar variability: an obstacle to combining multi-epoch spectra



Stellar variability: an obstacle to combining multi-epoch spectra



Summary and Conclusions

- We develop a **new method** to extract ground-based transmission spectra that does **not rely on comparison stars**.
- The new method is more **accurate** and more **precise**; it allows to derive **wavelength dependent absolute transit depths**.
- The new method enables **ground-based** atmospheric follow-up of **bright targets** with no suitable comparison stars nearby.
- Contamination due to **stellar variability** raises concerns on reliably combining transmission spectra over multiple epochs.