

## 4MOST - StePS

## Angela Iovino <br> INAF-OABrera

## on behalf of StePS collaboration

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Sanchez-B1azquez, A. Vazdekis, B. Vulcani, A. Zanella, M. Annunziatel1a, F. Belfiore, L. Cassarà,
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## StePS - Stellar Population Survey

- Mapping galaxy evolution over the past 7 Gyrs

StePS Science Goal in a sentence: detail the processes that shape galaxy evolution in the past $\sim 7$ Gyrs (half the life-time of Universe) and that produce galaxy properties as observed at $\mathbf{z} \sim 0$

StePS Ingredients: high $\mathrm{S} / \mathrm{N}$, high resolution spectra with wide wavelength coverage of $\mathrm{IAB}_{\mathrm{AB}}<20.5$ selected galaxies in the range $0.3<\mathrm{z}<0.7$

## StePS Ingredients

## $I_{A B<20.5} \& 0.3<z<0.7$ : a redshift range still Iargely unexplored

StePS fits nicely in the redshift niche between SDSS and LEGA-C


Stellar Mass compleness


| Redshift | $\log \left(m / M_{0}\right)$ |
| :---: | :---: |
| 0.3 | 10.4 |
| 0.5 | 11.0 |
| 0.7 | 11.3 |

## StePS - Stellar Population Survey

- Mapping galaxy evolution over the past 7 gyrs

StePS Ingredients: high spectral quality, high resolution spectra with wide wavelength coverage of $\mid A B<20.5$ selected galaxies in the range $0.3<z<0.7$

StePS Products: - age of the stellar component

- star-formation activity time-scale
- metal abundances in stars and gas
- presence/absence of AGN activity
- galaxy stellar and dynamical mass
- presence of gas inflows and outflows


## Environment information

## StePS Ingredients

## High S/N + High resolution

Together enable good estimates of key spectral indices that are stellar age and metallicity indicators

VIPERS (---) vs SDSS (一) R200 vs R2000



## StePS Ingredients

## High S/N + High resolution + Wide $\lambda$ coverage

Coverage of main spectral indices and abs/em lines as a function of redshift in the range of interest


## StePS ideal instruments

## WEAVE@WHT and/or 4MOST@VISTA

a new window of opportunity for some years to come


Next big step forward will be MSE: 10 mt telescope, 10 years timescale!


## StePS ideal instruments



|  | WEAVE @ WHT | 4MOST @ VISTA |
| :--- | :--- | :--- |
| Telescope size | 4mt class | 4mt class |
| FoV | 3 sq degs | 4 sq degs |
| R @ Low resolution mode | 5000 | 6000 |
| Lambda range | $3600-9900$ AA | $3700-9500$ AA |
| Multiplexing | 1000 | 1600 |
| Fibers on sky aperture | $1 " .3$ | $1 " .45$ |

High S/N + High resolution + Wide $\lambda$ coverage
25k spectra @ 7h texp - 0.3<z<0.7 - IAB<20.5 targetting well known extragalactic fields

Realistic end-to-end simulations using COSMOS field HST data computed in-fiber fluxes and S/N

Assuming $1^{\prime \prime} .3$ arcsec seeing



High S/N + High resolution + Wide $\lambda$ coverage
25k spectra @ 7h texp - 0.3<z<0.7 - IAB<20.5 $\rightarrow$ median ( $\mathrm{S} / \mathrm{N}$ ) ${ }^{\sim} 10$ in I-band


A S/N value of ~10 in I-band enables good estimates of major spectral features across the full wavelength range covered

## High S/N + High resolution + Wide $\lambda$ coverage

At texp~7 hours, only ~ 10\% of targets will have a median (S/N)>20 in I-band



The most massive, redder and lower redshift galaxies: far from being a representative subsample.

## Take a step forward:

Higher S/N + High resolution + Wide $\boldsymbol{\lambda}$ coverage
3.5k spectra @ 30h texp $-0.3<z<0.7-I_{\mathrm{AB}}<20.5 \rightarrow$ median $(\mathrm{S} / \mathrm{N}) \sim 30$ in I-band

Trading sample size for much higher S/N A LEGA-C like survey at $0.3<z<0.7$

## Take a step forward:

Higher S/N + High resolution + Wide $\boldsymbol{\lambda}$ coverage

## 3.5k spectra @ 30h texp - 0.3<z<0.7 - $\operatorname{IAB}<20.5 \rightarrow$ median $(S / N) \sim 30$ in l-band



Trading sample size for much higher S/N
Piggybacking on WAVES-DEEP footprint -

- Take advantage of planned repeated passes AND
- Get precise information on galaxy position within the cosmic web


## Take a step forward:

## Higher S/N + High resolution + Wide $\boldsymbol{\lambda}$ coverage

3.5k spectra @ 30h texp -0.3<z<0.7 - IAB<20.5 $\rightarrow$ median (S/N)~30 in I-band



Gallazzi et al. 2005

## Take a step forward:

Higher S/N + High resolution + Wide $\lambda$ coverage
3.5k spectra @ 30h texp -0.3<z<0.7 - IAB<20.5 $\rightarrow$ median (S/N)~30 in I-band



## Take a step forward:

## Strategy

- observe repeatedly subset of bright targets (~200 per FoV), embedded in the footprint of WAVES-Deep, where high number of passes are expected, totalling ~30h exp for ~3.5K galaxies


## Total Fibre hours budget

- 100K fiber hours - only $2.5 \%$ of total available for public surveys


## A gain for all



- 4MOST-StePS is a powerful enhancement for a full science exploitation of WAVES-Deep - we will explore the connections between observed physical properties (mass, SFR, stellar age, metallicity) and environment - down to galaxy pair scales
- 4MOST-StePS data will provide robust physical information, given their superior $\mathrm{S} / \mathrm{N}$, down to lower stellar masses and higher redshifts, thus complementing WEAVE-StePS results
- WEAVE-StePS larger statistic will help characterizing global trends in galaxy evolution mechanisms


## Take home messages

* 4MOST-StePS is a low-cost survey with a high scientific return
* 4MOST-StePS will provide an unbiased empirical description of the evolutionary path of massive galaxies in the still unexplored redshift range $0.3<z<0.7$
* The important synergies with the science case of WAVES-DEEP is a win-win opportunity that should be seized



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