4MOST – 4m Multi-Object Spectroscopic Telescope

ETC algorithms and assumptions Genoveva Micheva (AIP) 4MOST Instrument Scientist







- Provides the user with some idea of how long the exposure times would be
- Assumptions based on the Top Of the Atmosphere to Detector (TOAD) instrument simulation software.

The survey nature of 4MOST implies that...



- User cannot choose specific sky conditions
- User cannot choose airmass
- User cannot choose seeing
- User cannot choose time of observations
- User cannot choose sky positions



- User must choose targets ③
- User must provide required S/N

Some key 4MOST features



- FoV 4.2 deg^2
- Minimum distance between fibers: 15 arcsec
- 2436 fibers (LRS 2x812; HRS 812)
- LRS 370 950 nm
- HRS 392 435 nm; 516 573 nm; 610 679 nm

ETC assumptions:



- 4MOST key features ETC must use some representative fiber properties
- User cannot choose observing conditions ETC must use typical observing conditions

The ETC uses time-averaged fiber-averaged properties

Representative Sky



Representative Sky





Representative Sky



- ESO's SkyCalc models
- Default SkyCalc parameters
 - Moon/target separation 60 deg
 - Moon/Sun separation (53; 90;143 deg)
 - Moon altitude 45 deg
 - 350 1000 nm at fixed spectral resolution of 100 000

Representative Airmass:



- User provides target coordinates
- ETC converts DEC to representative airmass:
 - ± 15 degrees away from the meridian, i.e. Hour Angle =1h. \rightarrow target altitude a

$$-AM = \left\{\frac{1}{\cos(a)}\right\}$$

 – Circumpolar targets (DEC ≤ -79.6) at AM≥ 1.74 (ZD ~55 deg) → limit of the ADC





- Airmass bins [1.05, 1.1, ...,2.0]
- Each bin Dark, Grey, Bright sky









 Reference seeing values are scaled with airmass and wavelength:

•
$$S(X) = \sqrt{\left(S_{ref} \cdot \left(\frac{X}{X_{ref}}\right)^{0,6}\right)^2 + S_{Vista}^2}$$

• $S(\lambda) = S_{ref} \cdot \lambda^{-0.2}$

Reference Seeing





Based on old statistics of DIMM June 2014- Sept 2017. New statistics show smaller Median seeing of 0.7 arcsec. ETC uses a slightly pessimistic Value of 0.8 arcsec.



Mirror reflectivity



- A function of wavelength and time
- ETC assumes 86% reflectivity (as a function of $\lambda)$



Telescope vignetting



- TOAD model contains vignetting due to
 - Central obstruction
 - Spider vanes of the M2 holding cell
 - The M2 mirror is undersized
 - The Wide field corrector optics
 - The atmospheric dispersion compensator

• ETC uses field centre.

Telescope vignetting





Telescope PSF and fiber size



- ETC uses telescope PSF at field center (function of λ)
- ETC uses a fixed fibre core diameter of 85µm



Fiber aperture size on sky



- Is a function of field position for fixed fiber core size
- 1.56^{''2} 1.63^{''2} ; ETC 1.605^{''2}



Intrinsic fiber transmission efficiency

- Relatively stable for each fiber
- ETC uses the mean fibre transmission efficiency of 0.895



Residual atmospheric dispersion

- 4
- Residual between 0" (field center, zenith) and 0.25" (field edge, ZD = 55 deg)
- ETC takes (field center, zenith)





Fiber tilt-induced vignetting losses



- 2 effects:
 - Fiber entrance: focus shift, changes the PSF seen by the fiber
 - Fiber output: broadening of the light beam, which is then vignetted by the spectrographs
- ETC assumes a non-tilted 1.0 fiber
 - Minimum losses due to the broadening of the beam
 - But Non-optimal due to slight de-focus



Detector characteristics



- Detectors have 4 quadrants, each with a RON, DC, Gain.
- ETC uses the mean of all quadrants for all 3.
 - E.g. DC = 2.288 e-/h
 - E.g. RON = 2.5 e-
 - E.g. Gain = 1.06875 e-/ADU

Charge transfer efficiency



 ETC uses position close to center of slit (~0.99)



Other effects/assumptions



- Fiber-to-target alignment accuracy of 0.1" (between fiber center and target position).
- ETC uses a single Focal Ratio Degradation model (based on lab measurements).
- Spectrograph image quality
- Material transmission efficiency

S/N calculation



- On-detector binning only in dispersion direction
- No binning in cross-dispersion direction (pipeline limitations)
- Per-pixel S/N converted to per-Å S/N using:

$$-SNR_{spec} = \sqrt{N_{pix}} \cdot SNR_{pixel}$$

 $N_{\text{pix}}(\lambda, LRS/HRS, red/green/blue)$ = # pixels that make up 1 Å





- ETC does not apply any reddening internally
- User templates must have reddening applied already!





- User provides total Pogson magnitudes of target
- AB or Vega
- Available filters (next slide)
- ETC internally calculates the magnitude inside the aperture of the fiber

Filters



- GAIA DR2r: G_{BP} , G_{RP} , G
- DECam (CTIO): g,r,i,z
- VISTA: Z, Y, J, H, Ks
- Standard filters:
 - Bessel U
 - Johnson B,V
 - Cousins R,I
- All filter transmission is photon-counting curves!





- Point source ETC assumes Moffat with β =2.5
- Extended source ETC assumes Sersic profile, user provides R_{eff} and n

SED shape: Templates



- ESO standard templates:
 - MARCS
 - Kinney- Calzetti
 - HII region
 - Kuruzc
 - Pickles
 - QSO
 - PNe
- User-defined spectrum

Templates

Target model could be a power law spectrum:

• $F(\lambda) = F_0 * \left(\frac{\lambda}{\lambda_c}\right)^p$

where

 λ – wavelength

 F_0 – continuum flux level @ λ_c

p – powerlaw index

•
$$F_0 = 10^{-(0.4*M_{obs}+ZP)}$$

Where

 M_{obs} – apparent magnitude ZP – zero point of the observing band



Scaling of Input Source Templates



- User provides total magnitude m_{Vega}(k) (or AB)
- Filter k (with transmission T_{λ})
- Source template F^{temp}
- $F_k^{temp} = \frac{\int \lambda * T_\lambda F^{temp} d\lambda}{\int \lambda * T_\lambda d\lambda}$
- $f = 10^{-0.4(m_{Vega}(k) 2.5 \log_{10} ZP_k^{Vega})}$
- Scaling factor $\xi = \frac{f}{F_{\mu}^{temp}}$
- $F_k^{scaled} = \xi F^{temp}$

Take-away message



- The ETC takes a number of shortcuts
- The virtual fiber used in the ETC does not exist
- But averaged over the survey lifetime, the ETC predictions will give correct results!

