Convolutional neural networks, trained

on simulations, can discover strong

lenses in survey image data.

An extended catalog of galaxy-galaxy strong gravitational lenses discovered in DES using convolutional neural networks

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INTRO

- Strong lensing is driving much discovery in astronomy, including galaxy evolution, cosmology and dark matter¹.
- Thousands of lenses are discoverable in current imaging; tens of thousands to come in







RESULTS

- Total of 511 high-quality strong lens candidates discovered.
- 84 high-redshift (lens at z > 0.8) discovered (see Jacobs et al 2019a)⁴.

Deployment in the cloud

This methodology was tested both on local HPC resources and in the Microsoft Azure Cloud (the author received sponsored credit from Microsoft⁵).

The trained TensorFlow model was deployed onto a GPU-enabled virtual machine. This has the advantage of bringing the code closer to the data – the model can be deployed on an instance in a region as close to the data as possible.

- LSST and Euclid².
- Deep Learning techniques from computer vision can make lens discovery in large image sets efficient.

METHOD

- Simulate galaxy-galaxy strong lenses with the seeing and other characteristics of Dark Energy Survey (DES) co-add images.
- Train CNNs, using simulated lenses, non-lenses and real galaxies.
- 3. Score ~8 million postage stamps,
 chosen to match the colors of
 simulated lenses in survey
 photometry³ (Fig. 2).



Figure 1. Top: Four of 84 high-redshift (*z* > 0.8) lens candidates from Jacobs et al 2019a[4]. **Middle**: Four bright 5 arcs detected in the DES search of Jacobs et al 2019b. **Bottom**: Images used in training the neural networks. *Left column*: Simulated lenses and lensed sources. *Second from left*: Simulated lens galaxies without a lensed source. *Third from left*: Real ellptical galaxies and simulated lensed sources. *Right column*: Real field galaxies, used as negative examples.

DISCUSSION

Discovery of candidates involved manual inspection of ~20,000 images. False positives are still an issue.
Improvements in simulations (better PSF matching, a greater variety of lens colors, greater use of real sources) likely to bring gains in accuracy.



This reduced data transfer times, a significant bottleneck, by a factor of 5, without requiring GPU access co-located with the survey data.

Both the training process and the trained model can be deployed on demand in a container in the cloud (using AzureML or the equivalent from another provider). This way, a **trained model can be easily shared** with the astronomical community as as web service.

5 Disclosure: Jacobs was awarded a Cloud Research Software Fellowship by Microsoft Asia in May 2019 which includes \$3000 of Azure credits and \$2000 travel support.

References

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g-i

Figure 2. *g-i* vs *g-r* colors of simulated lenses and non-lenses, with sources from the DES catalog shown in grey. The box indicates the color-cut used to select a catalog of 7.9 million sources for scoring by CNNs.

Poster template: Mike Morrison





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