



NEARBY RED DWARFS & THEIR DANCE PARTNERS: THE MULTIPLICITY OF OUR LOW-MASS NEIGHBORS

JEN WINTERS

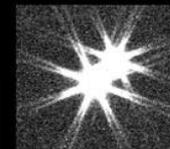
3 July 2017



HARVARD-SMITHSONIAN
CENTER FOR ASTROPHYSICS



The
MEarth
Project



RECONS
Research Consortium on Nearby Stars

MOTIVATION

- **What fraction of M dwarf systems are multiples?**
- **Where are the companions to M dwarfs found?**
- **How much mass is in multiple systems, & what is the distribution of that mass?**
- **How do M dwarf systems evolve over time?**
- **Are star systems primarily singles or multiples?**

M DWARF PROPERTIES

- **MASSES:** $0.075 - 0.64 M_{\odot}$
- **RADII:** $0.1 - 0.6 R_{\odot}$
- **LUMINOSITIES:** $0.02\% - 6\% L_{\odot}$
- **TEMPERATURES:** $2400 - 3900 \text{ K}$
- **COLORS:** $3.7 \leq (V-K) \leq 9.5$
- $8.8 < M_V < 20.0$

Tarter+ 2007
Benedict+ 2016
RECONS

PREVIOUS NEARBY M DWARF MULTIPLICITY SURVEYS

Reference	# of Targets	Technique	Notes
Skrutskie+ (1989)	55	Infrared Imaging	Mult. Fr. not given
Henry & McCarthy (1990)	27	Infrared Speckle	BD search
Henry (1991)	74	Infrared Speckle	Amazing!!!
Fishcher & Marcy (1992)	28-62	Various	Varied Sample
Tokovinin (1992)	200	Radial Velocity	K's&M's; BD search
Simons+ (1996)	63	Infrared Imaging	BD search
Delfosse+ (1999)	127	Radial Velocity	Mult. Fr. not given
Endl+ (2006)	90	Radial Velocity	Jovian search
Law+ (2006)	32	Lucky Imaging	M5 – M8
Law+ (2008)	77	Lucky Imaging	Later M's
Bergfors+ (2010)	124	Lucky Imaging	Young M0 – M6
Law+ (2010)	36	Adaptive Optics	Wide binary search
Dieterich+ (2012)	126	HST-NICMOS	BD search

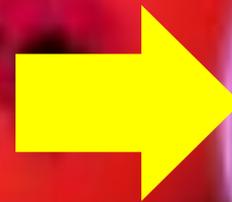
ANSWER:

7 – 42 %

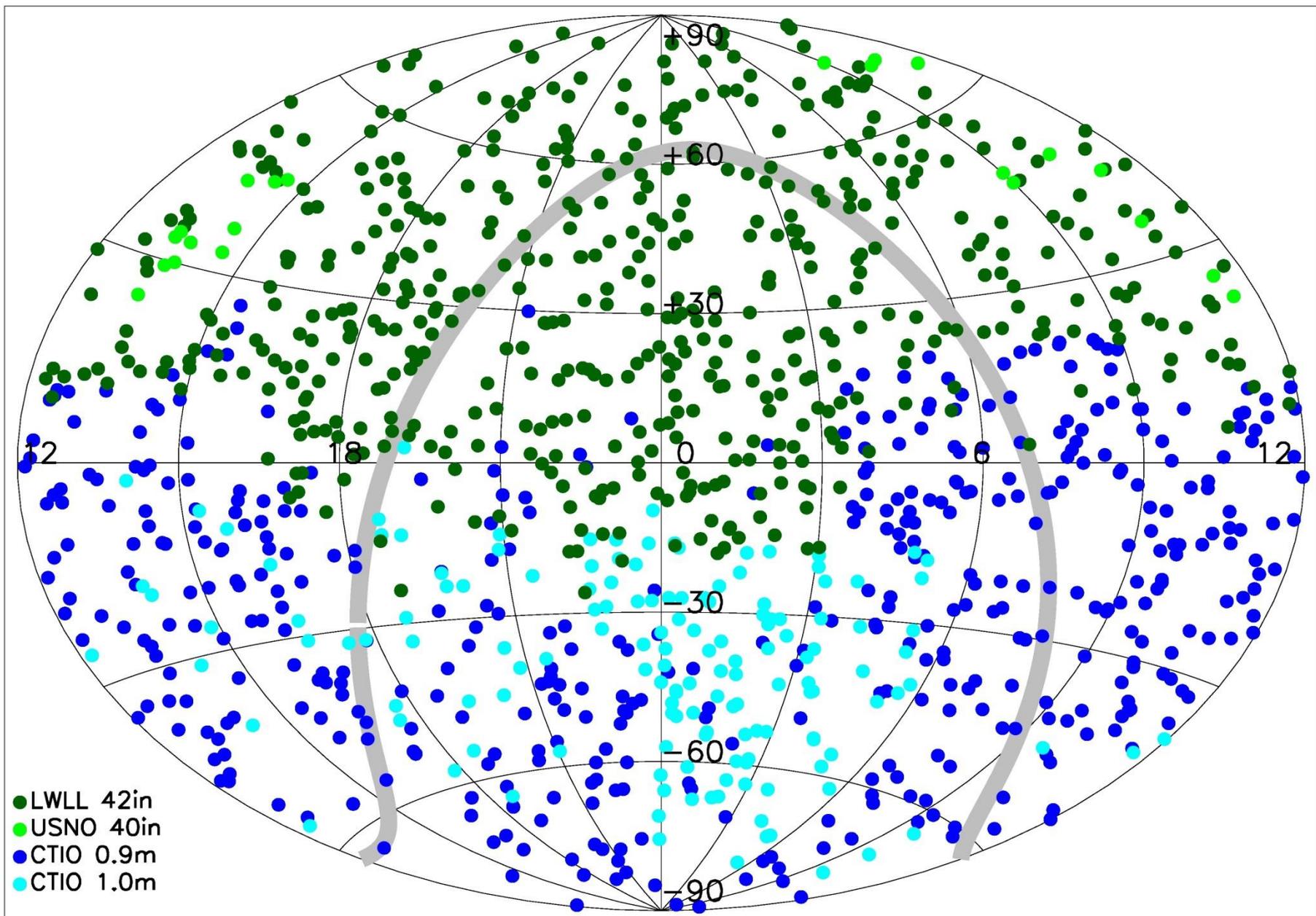
**MY GOAL:
LOOK AT > 1000
M DWARFS FOR
COMPANIONS**

DEFINING THE SAMPLE

- ALL-SKY
- NO PROPER MOTION LIMITS
- COLORS: $3.7 \leq (V-K) \leq 9.5$
- $8.8 \leq M_V \leq 20.0$
- $\pi \geq 40 \text{ mas}, \sigma_\pi \leq 10 \text{ mas}$
- NO WHITE DWARFS
- NO BROWN DWARFS



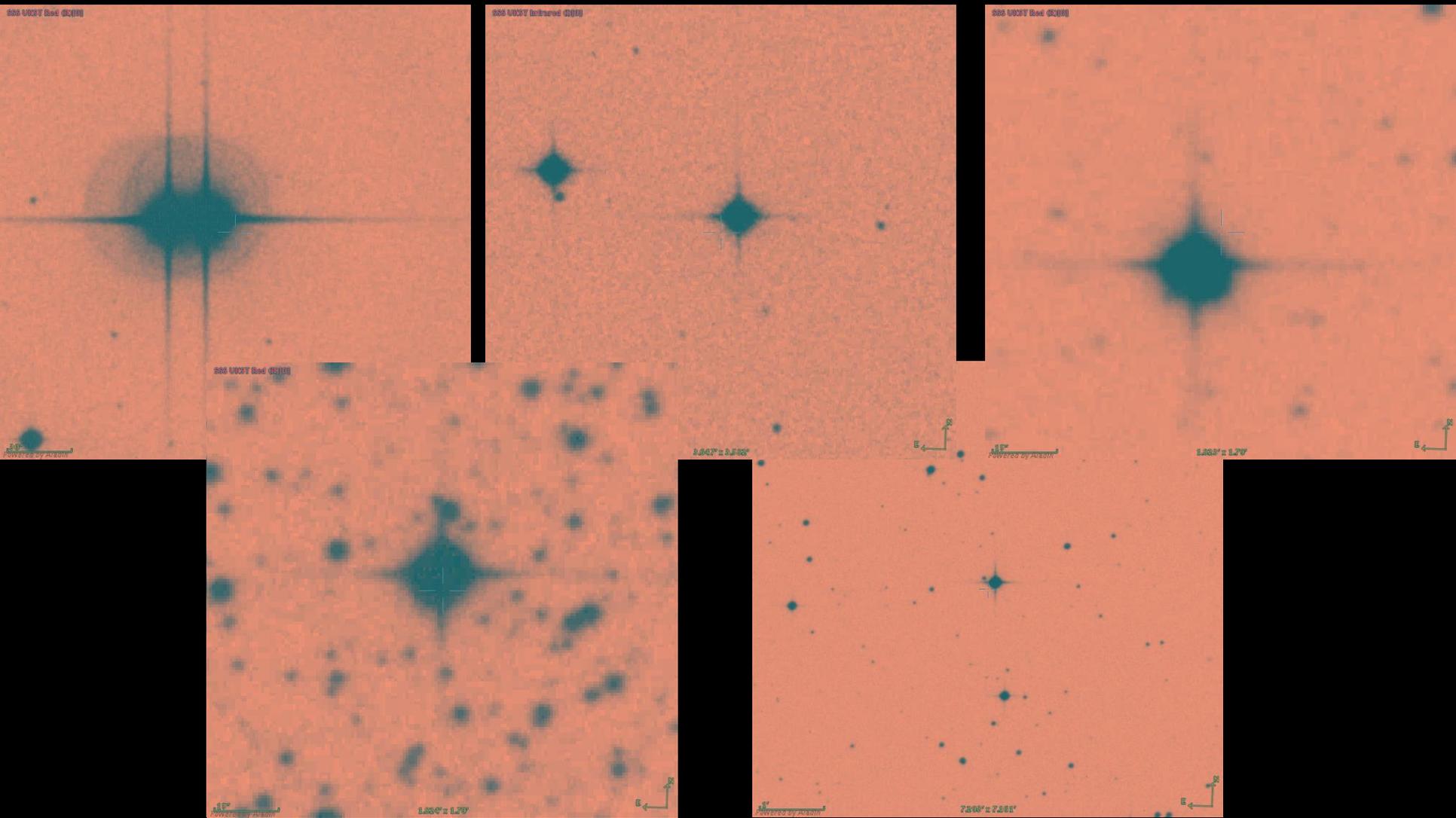
1121
Systems



SEARCH REGIONS

- *Literature Search*
- **Wide Separations: 5 – 300''**
(50 – 7500 AU)

WIDE COMPANIONS



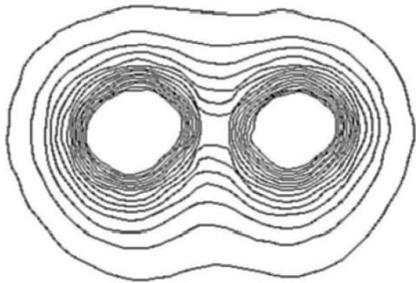
SEARCH REGIONS

- *Literature Search*
- **Wide Separations: 5 – 300''**
(50 – 7500 AU)
- **Close Separations: 2 – 5''**
(20 – 125 AU)

CLOSE COMPANIONS

Lowell Obs. +
Cerro Tololo Obs.
Separations:
2 - 5" (20 - 125 AU)

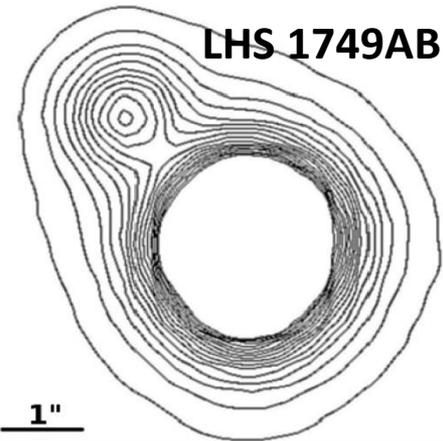
LHS 225AB



$\Delta I = 0.1$
 $\rho = 2.5''$

$\Delta I = 1.0$
 $\rho = 1.6''$

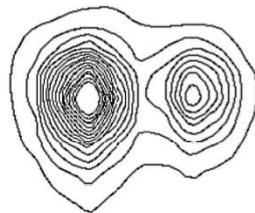
LHS 1749AB



$\Delta V = 3.0$
 $\rho = 2.9''$

$\Delta I = 0$
 $\rho = 1.1''$

SCR 0644-4223AB



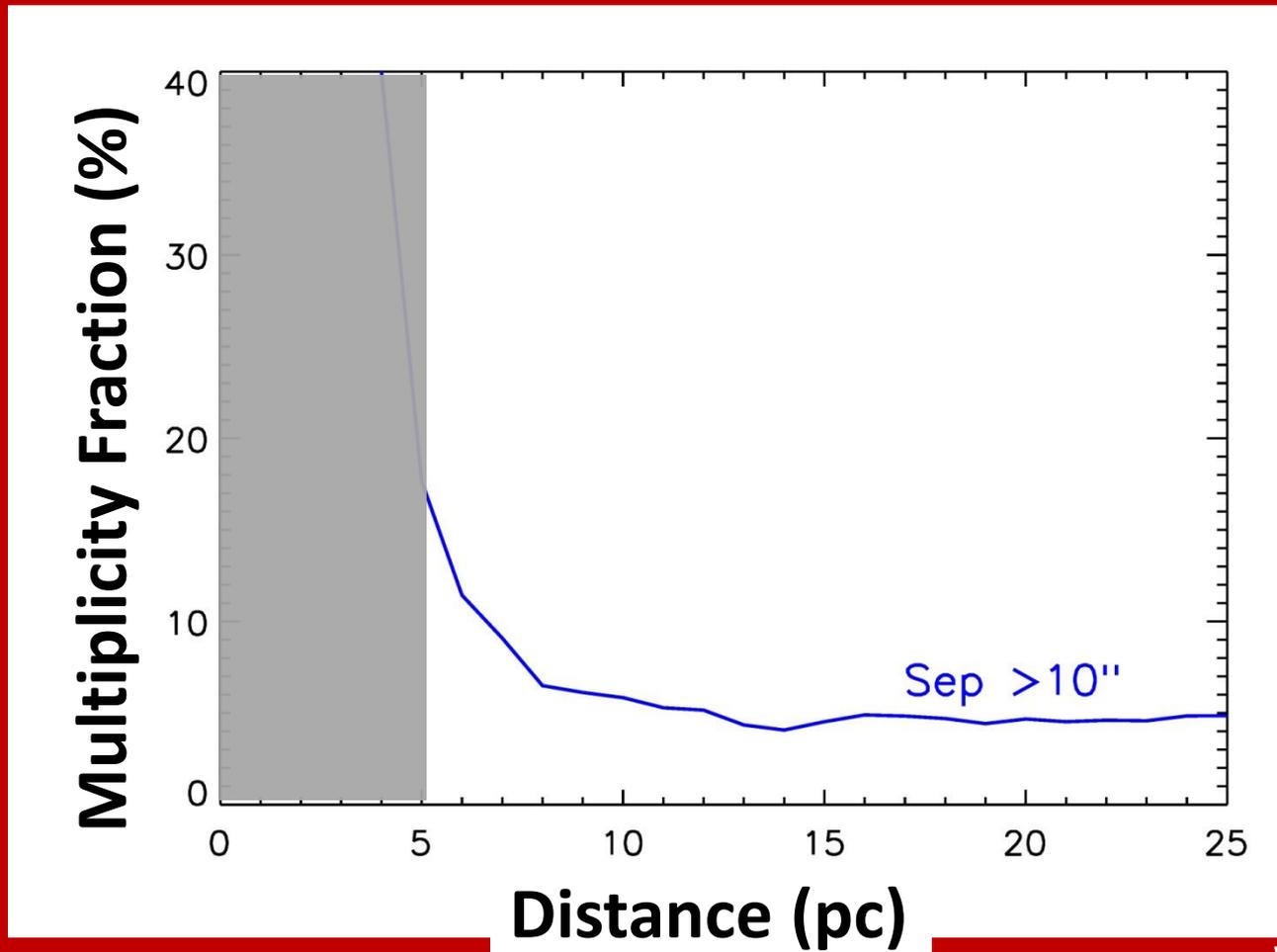
SCR 1856-4704AB



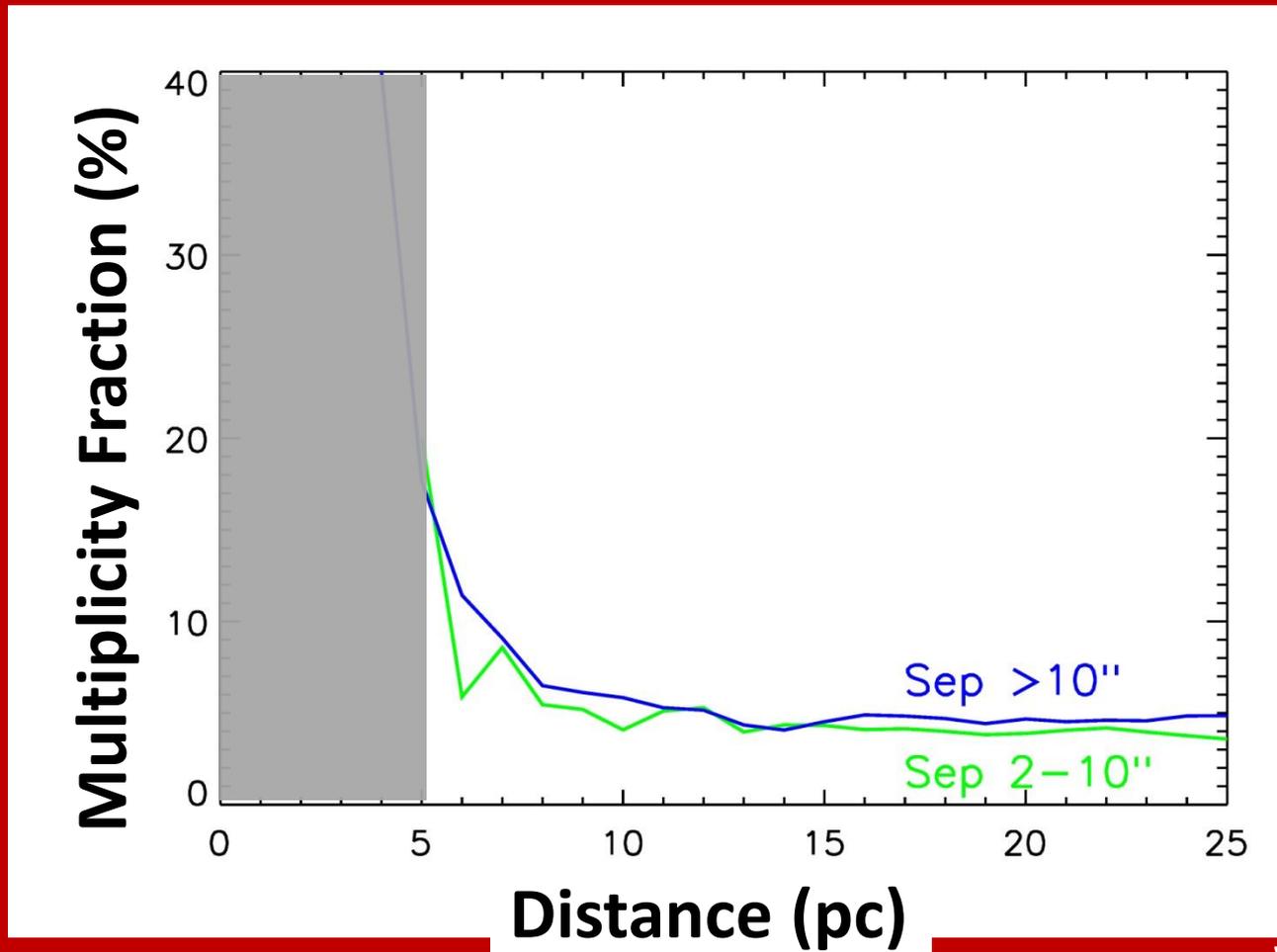
SEARCH REGIONS

- *Literature Search*
- **Wide Separations: 5 – 300''**
(50 – 7500 AU)
- **Close Separations: 2 – 5''**
(20 – 125 AU)
- **Even Closer Separations: < 2''**
(< 50 AU)

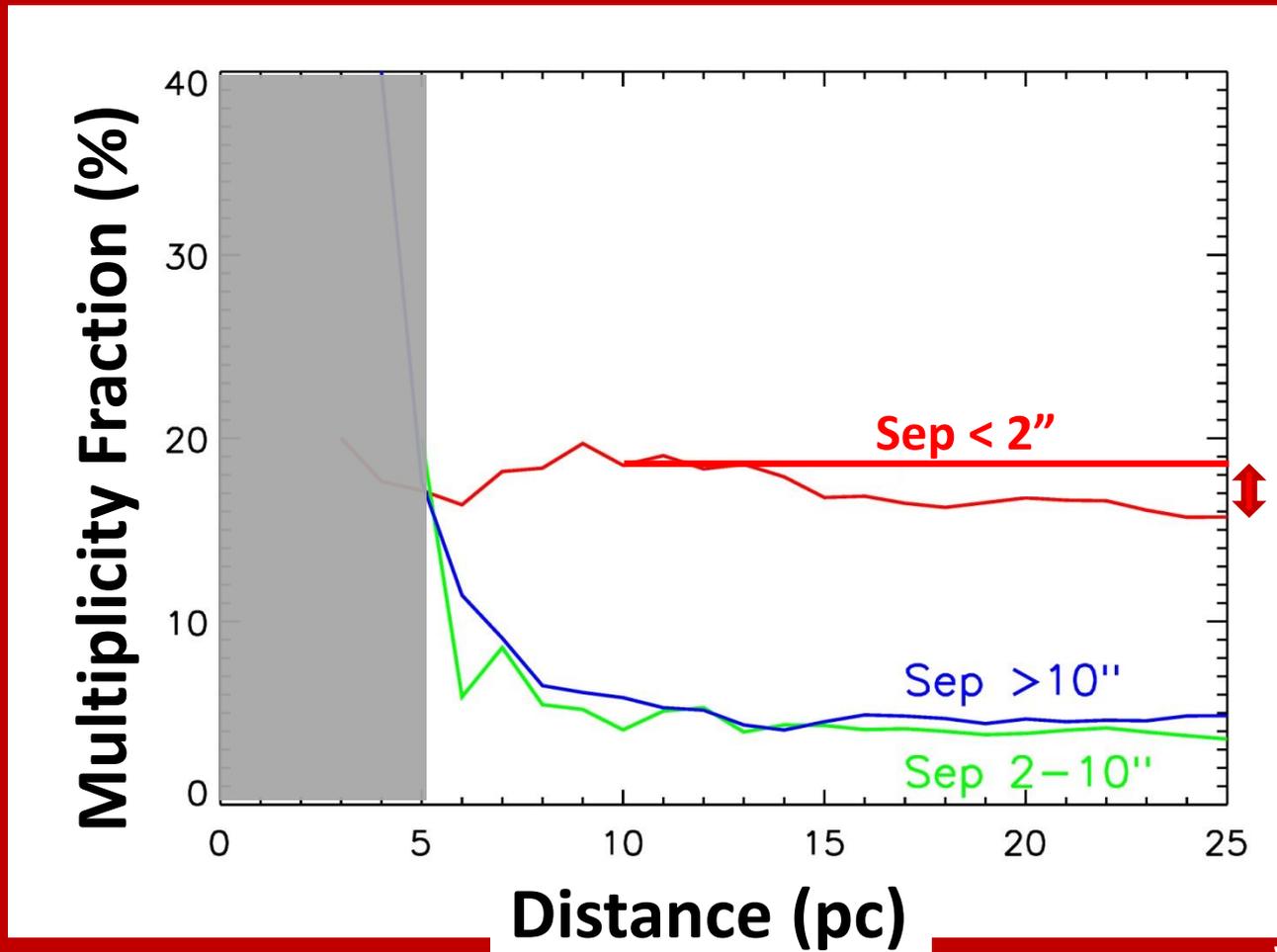
CORRECTING THE MULTIPLICITY FRACTION



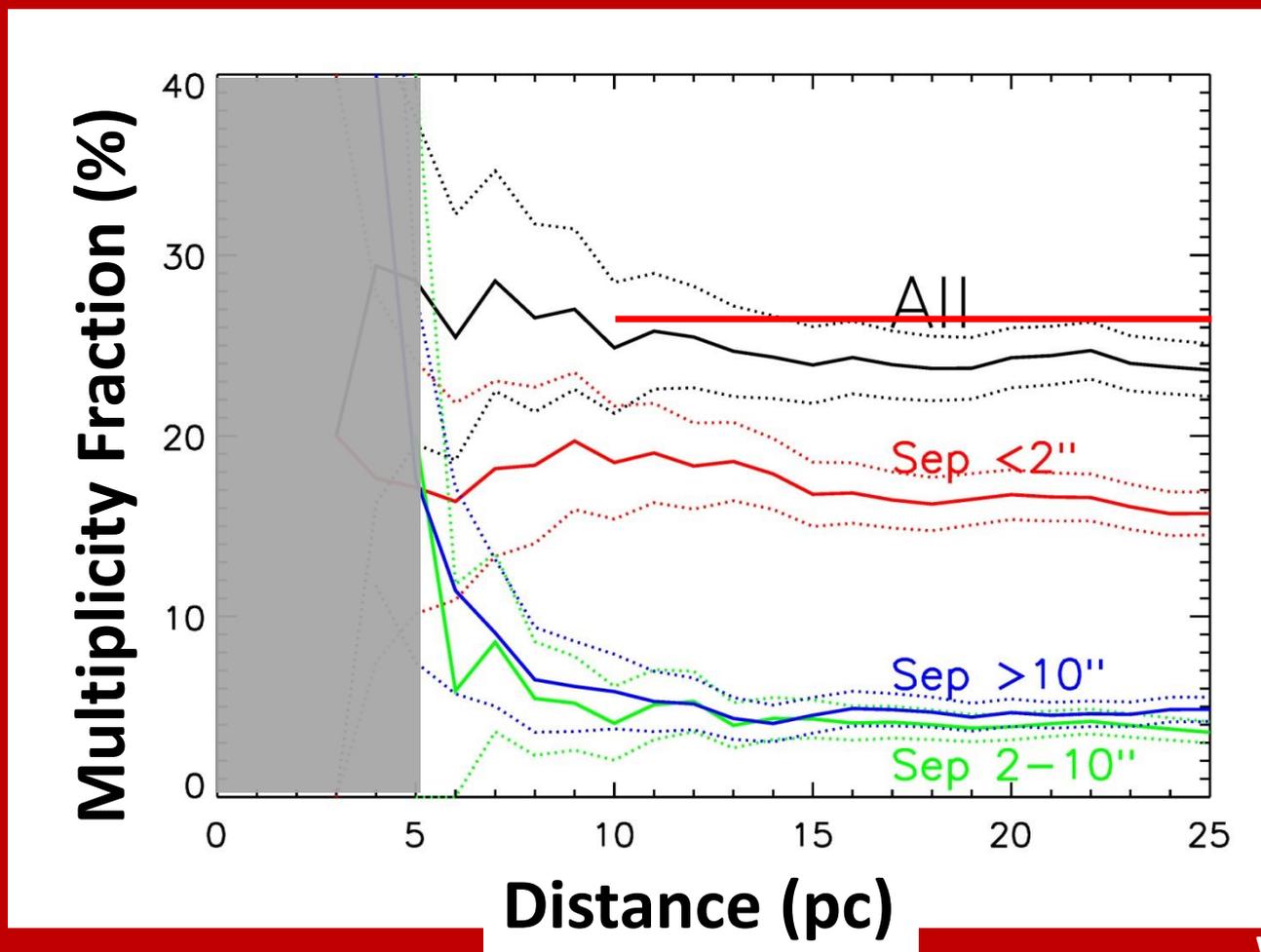
CORRECTING THE MULTIPLICITY FRACTION



CORRECTING THE MULTIPLICITY FRACTION



CORRECTING THE MULTIPLICITY FRACTION



M DWARF MULTIPLICITY

26% \pm 1.3%

M DWARF MULTIPLICITY

26% \pm 1.3%



Singles

856

(77%

M DWARF MULTIPLICITY

26% \pm 1.3%

Singles:Binaries

856 :223

(77% :20%



M DWARF MULTIPLICITY

26% \pm 1.3%

Singles:Binaries:Triples

856 :223 :39

(77% :20% :3%



Winters, in prep

M DWARF MULTIPLICITY

26% \pm 1.3%

Singles:Binaries:Triples:Quads

856 :223 :39 :1

(77% :20% :3% :0.1%



Winters, in prep

M DWARF MULTIPLICITY

26% \pm 1.3%

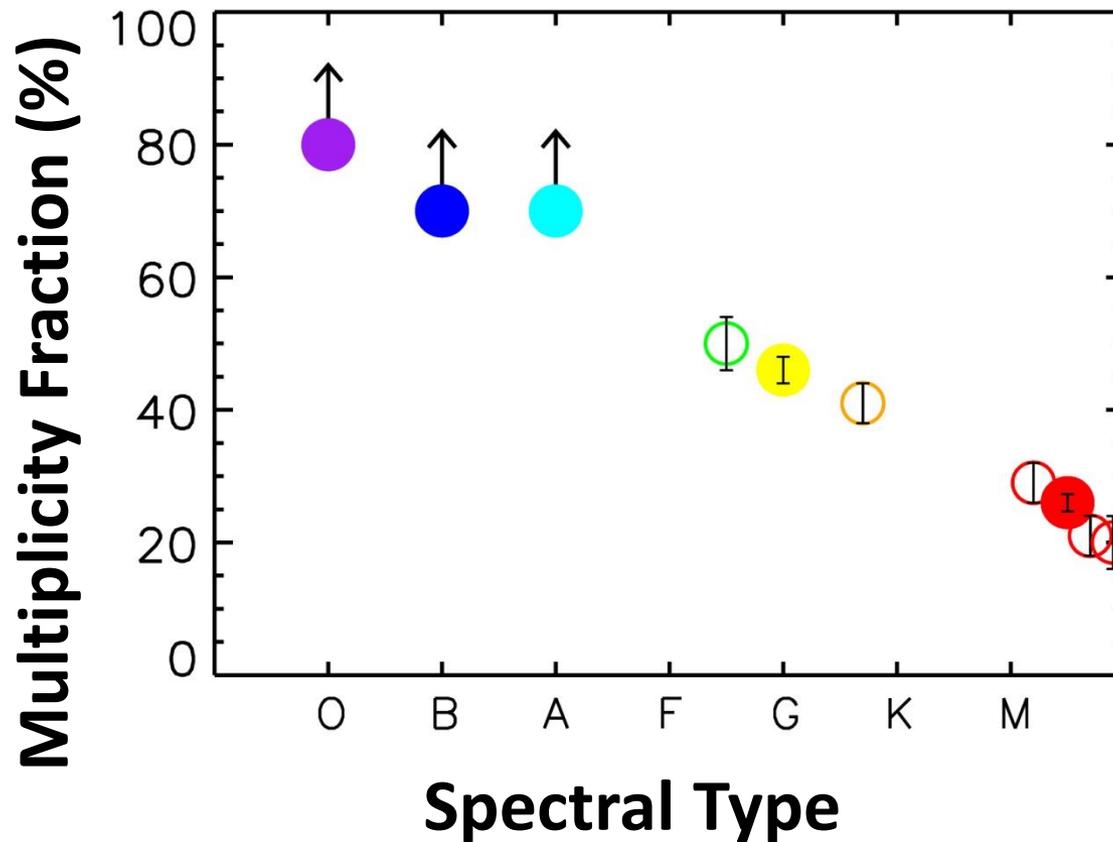
Singles:Binaries:Triples:Quads:Quins

856 :223 :39 :1 :2
(77% :20% :3% :0.1% :0.2%)



Winters, in prep

MULTIPLICITY IN CONTEXT

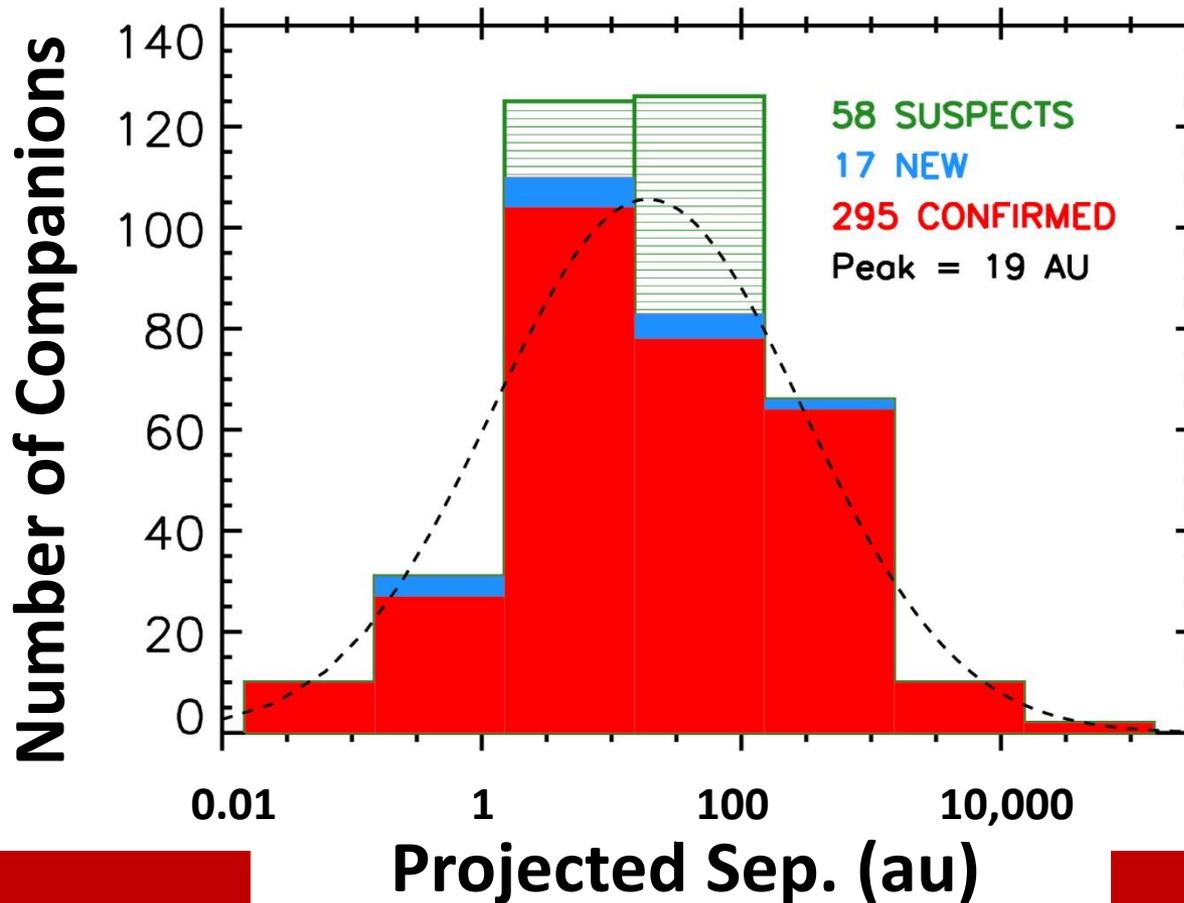


ARE MOST SYSTEMS SINGLE OR MULTIPLE?

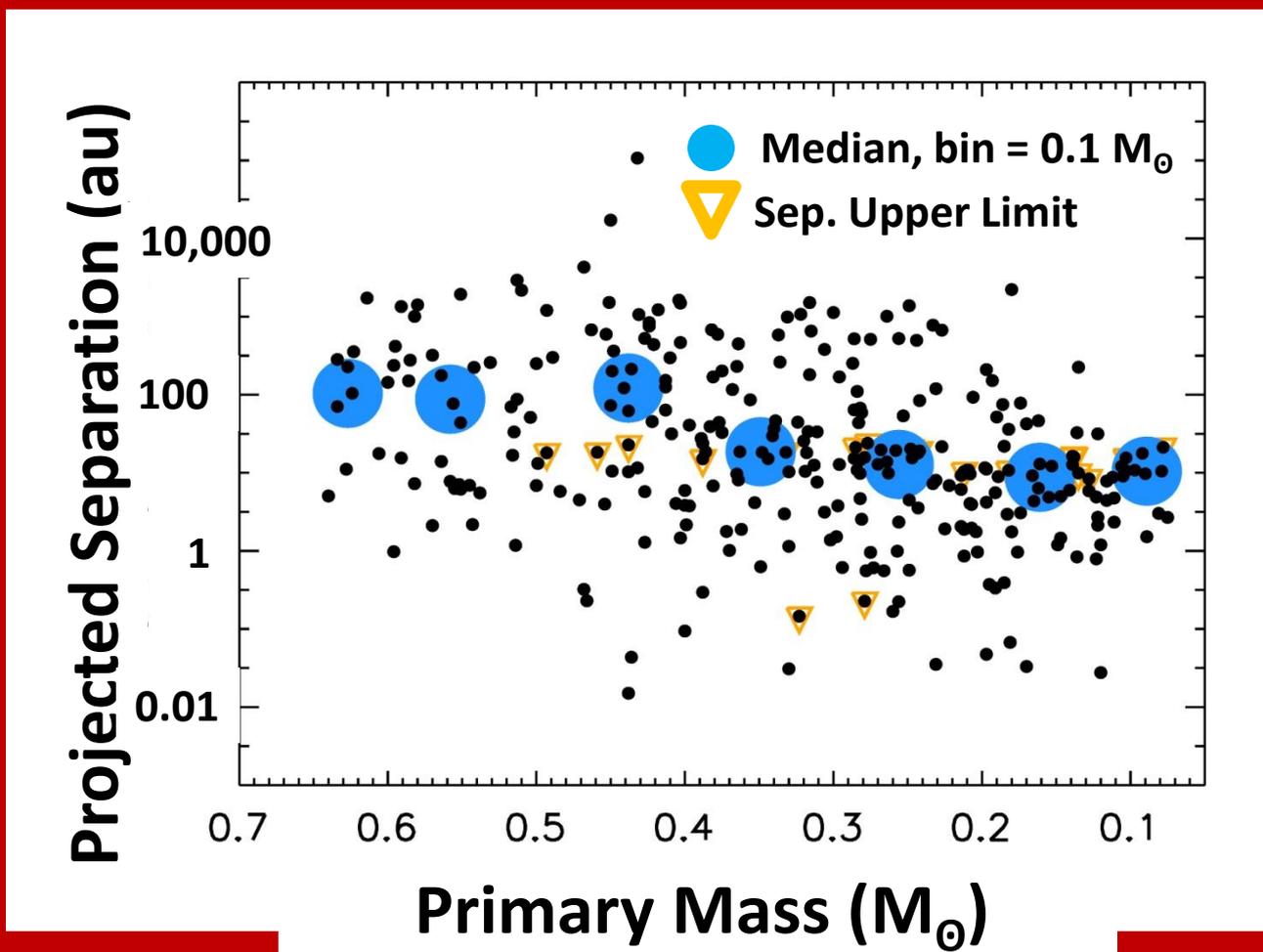
SPECTRAL TYPE	% OF ALL SYSTEMS	MULTIPLICITY FREQUENCY	REFERENCE
O	<0.003	>80	Mason+ 2009
B	0.1	70	Raghavan+ 2010
A			
F			
G			
K	10.2	40?	
M	75	26	This work

**Mult. Fr. ~30%
FOR ALL SYSTEMS**

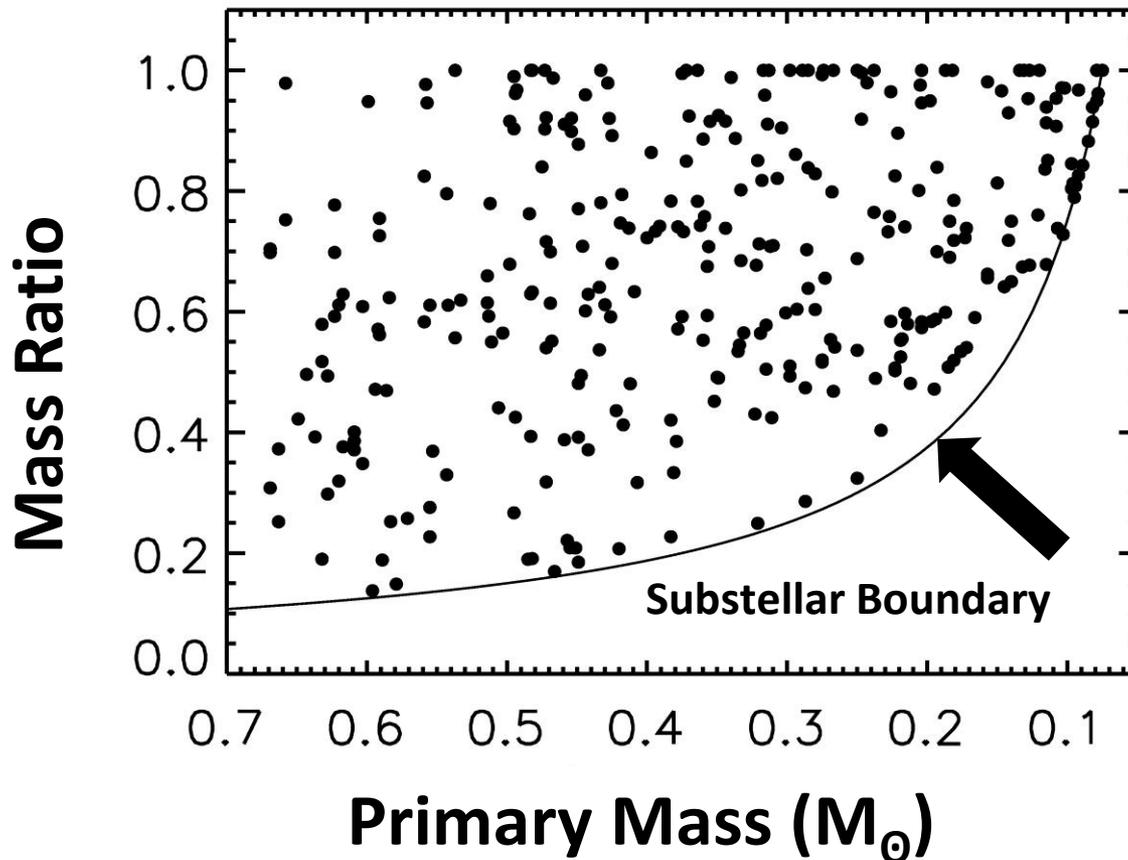
COMPANIONS TO M DWARFS ARE FOUND AT SOLAR SYSTEM SCALES



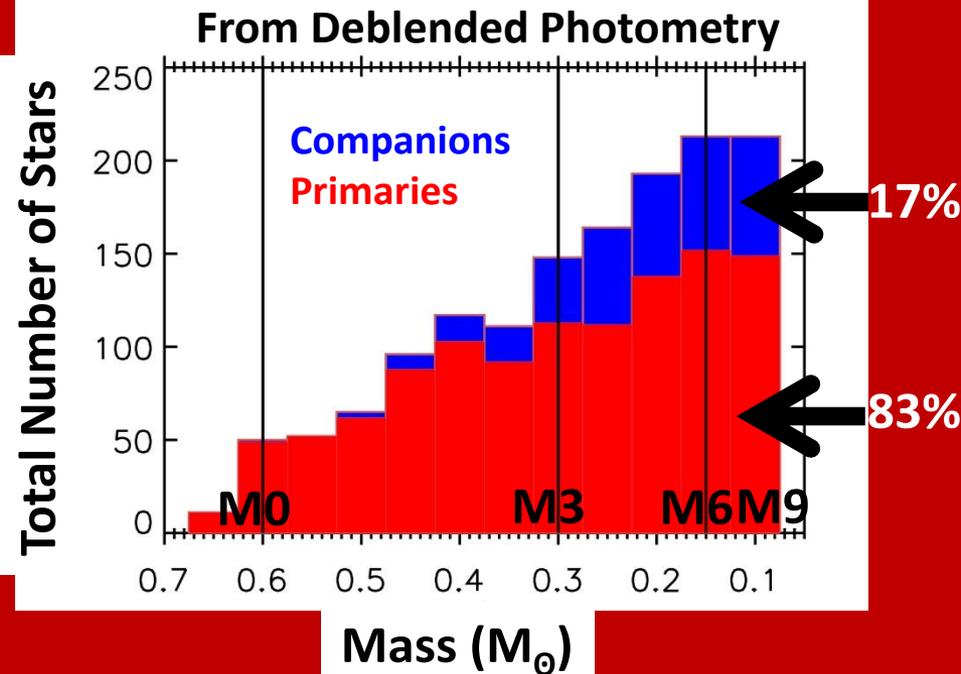
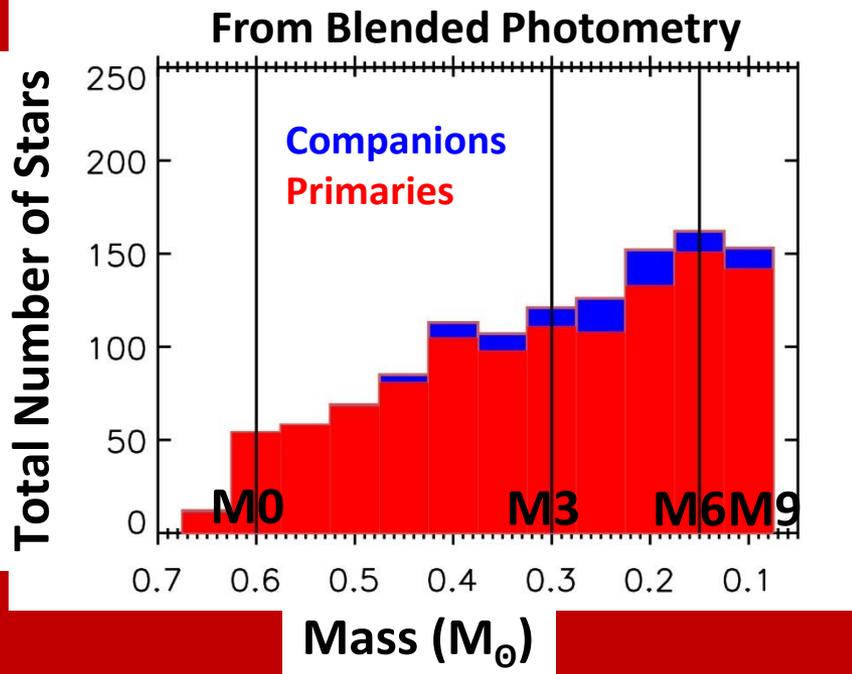
IS COMPANION SEPARATION A FUNCTION OF PRIMARY MASS?



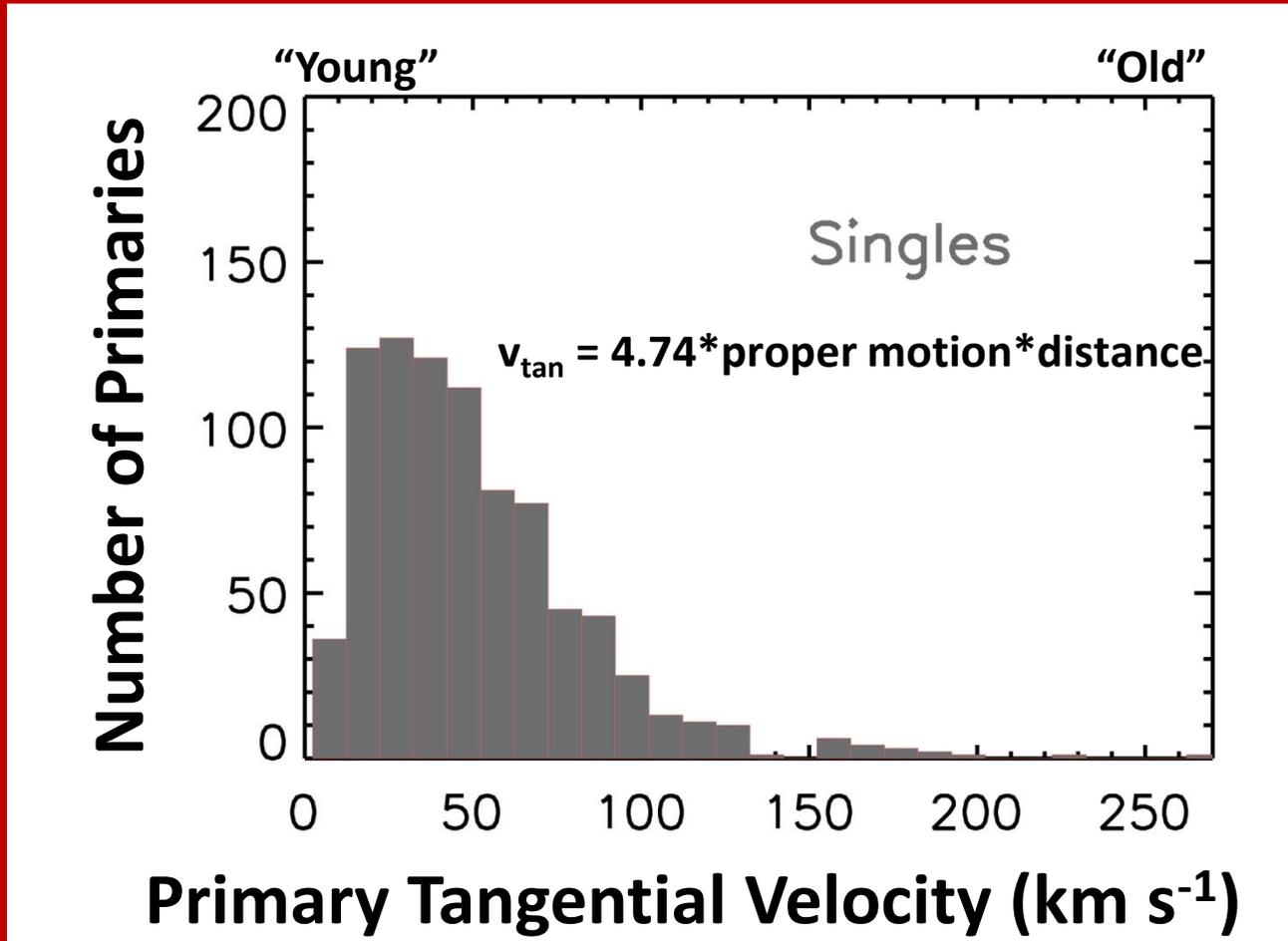
NO PREFERRED MASS RATIO



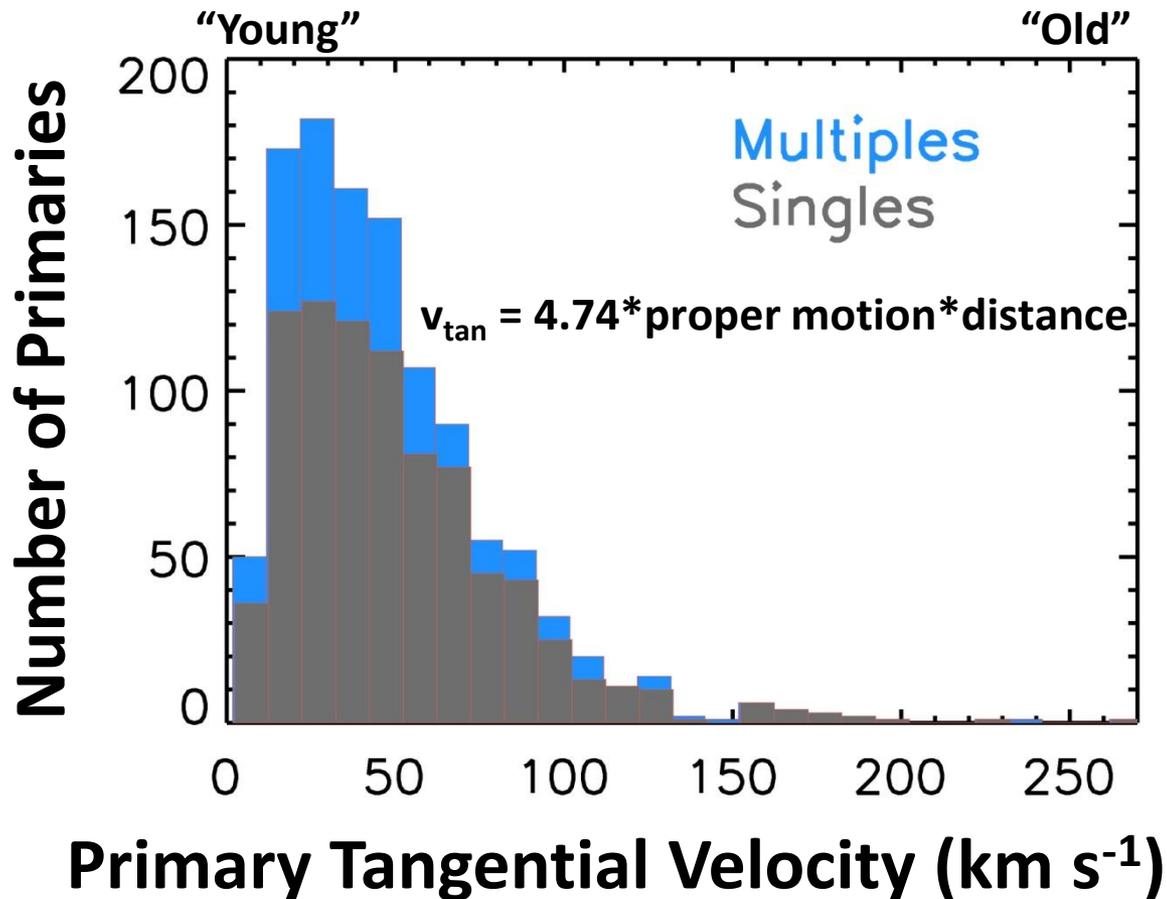
MASS FUNCTION



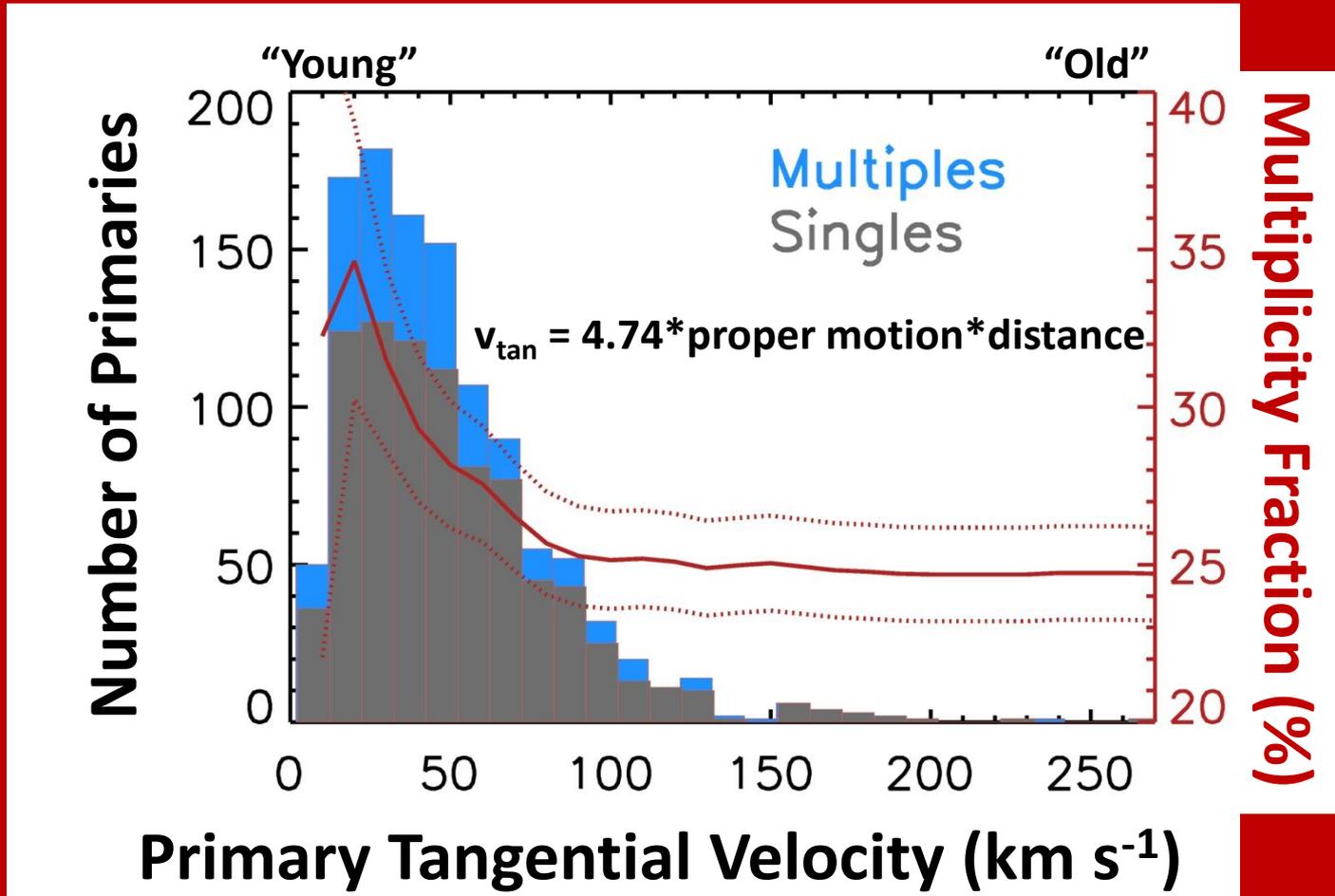
M DWARF MULTIPLICITY AS A FUNCTION OF AGE



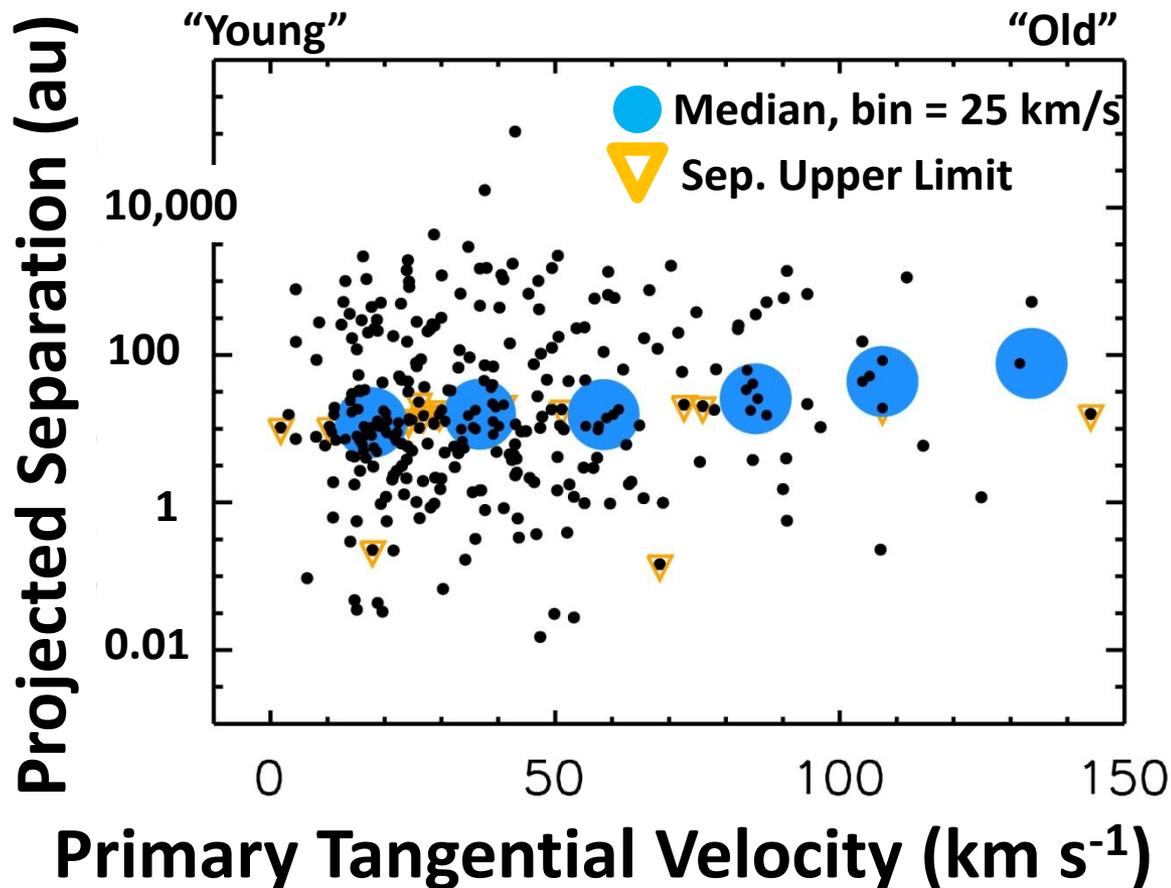
M DWARF MULTIPLICITY AS A FUNCTION OF AGE



M DWARF MULTIPLICITY AS A FUNCTION OF AGE



IS COMPANION SEPARATION A FUNCTION OF AGE?



RESULTS

- **26%** of M dwarf systems are multiples.
- Companions to M dwarfs are found at **separations on Solar System scales: 1-100 au.**
- Mass function for all M dwarfs **increases** to the end of the main sequence.
- M dwarf multiplicity & age *may* be related.
- *At least* **17% of M dwarf mass** is found in companions.
- Most stellar systems are **single.**

BIG PICTURE

- The Universe makes LOTS of little stars.
- Most are alone.
- Are they born that way, or do they become separated?

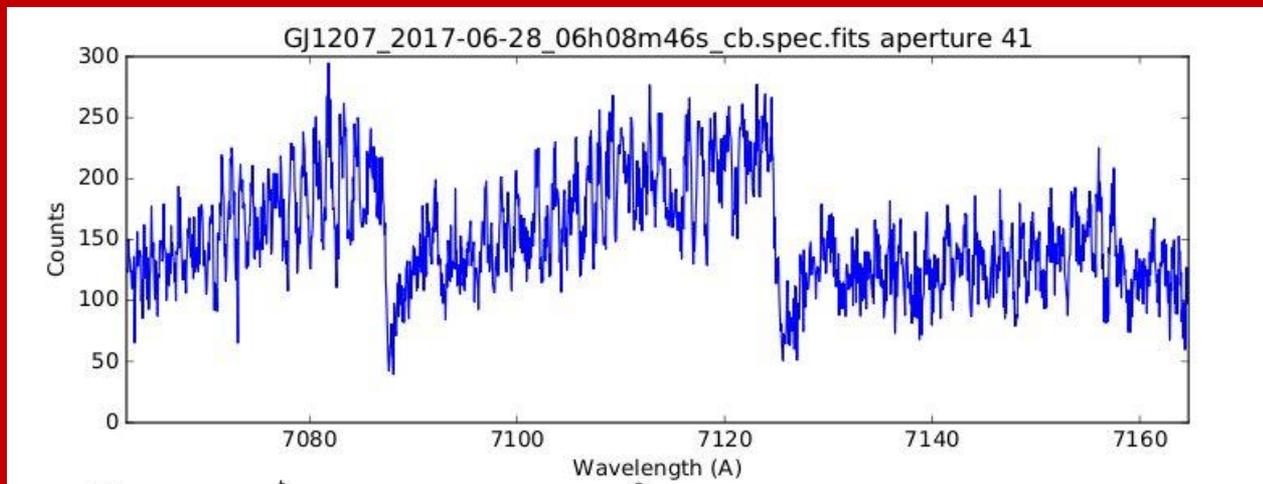


Image credit:

<http://www.geeksofdoom.com/2014/04/26/geek-gear-the-hitchhikers-guide-to-the-galaxy-constellation-42-marvin-on-life>

FUTURE

- **NEW** high-resolution spectroscopic project targeting the **mid-M dwarfs ($0.1 - 0.3 M_{\odot}$)** within **15 pc**
 - **TRES** ($R \approx 44,000$) on FLWO 1.5m on Mt. Hopkins in AZ
 - **CHIRON** ($R \approx 80,000$) on CTIO 1.5m in Chile
 - ID **SBs**, measure **RVs**, ***vsinis***, **H α** equivalent widths



RESULTS

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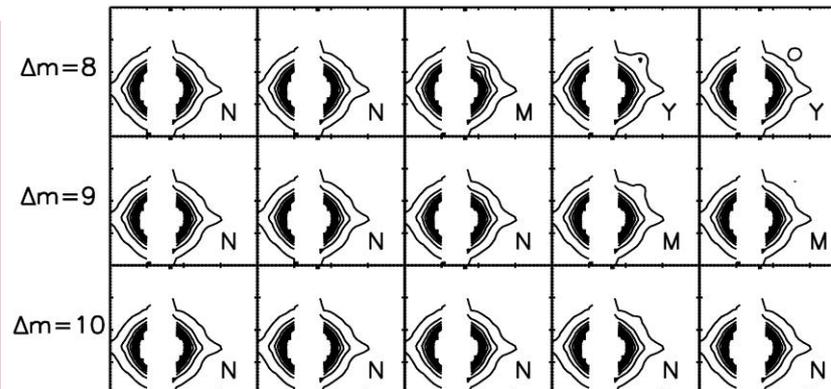
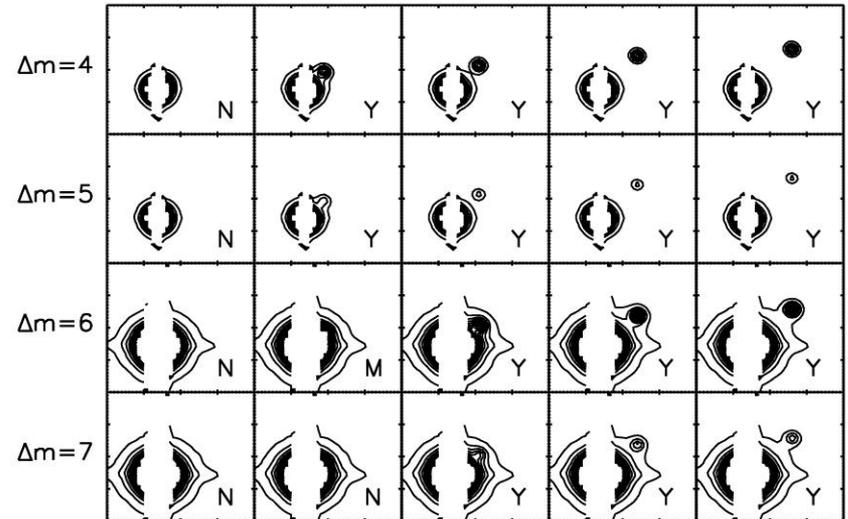
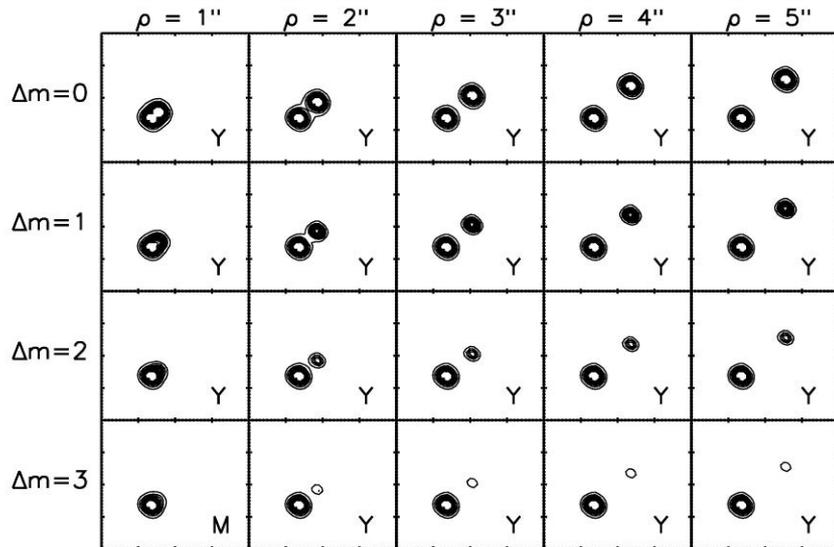
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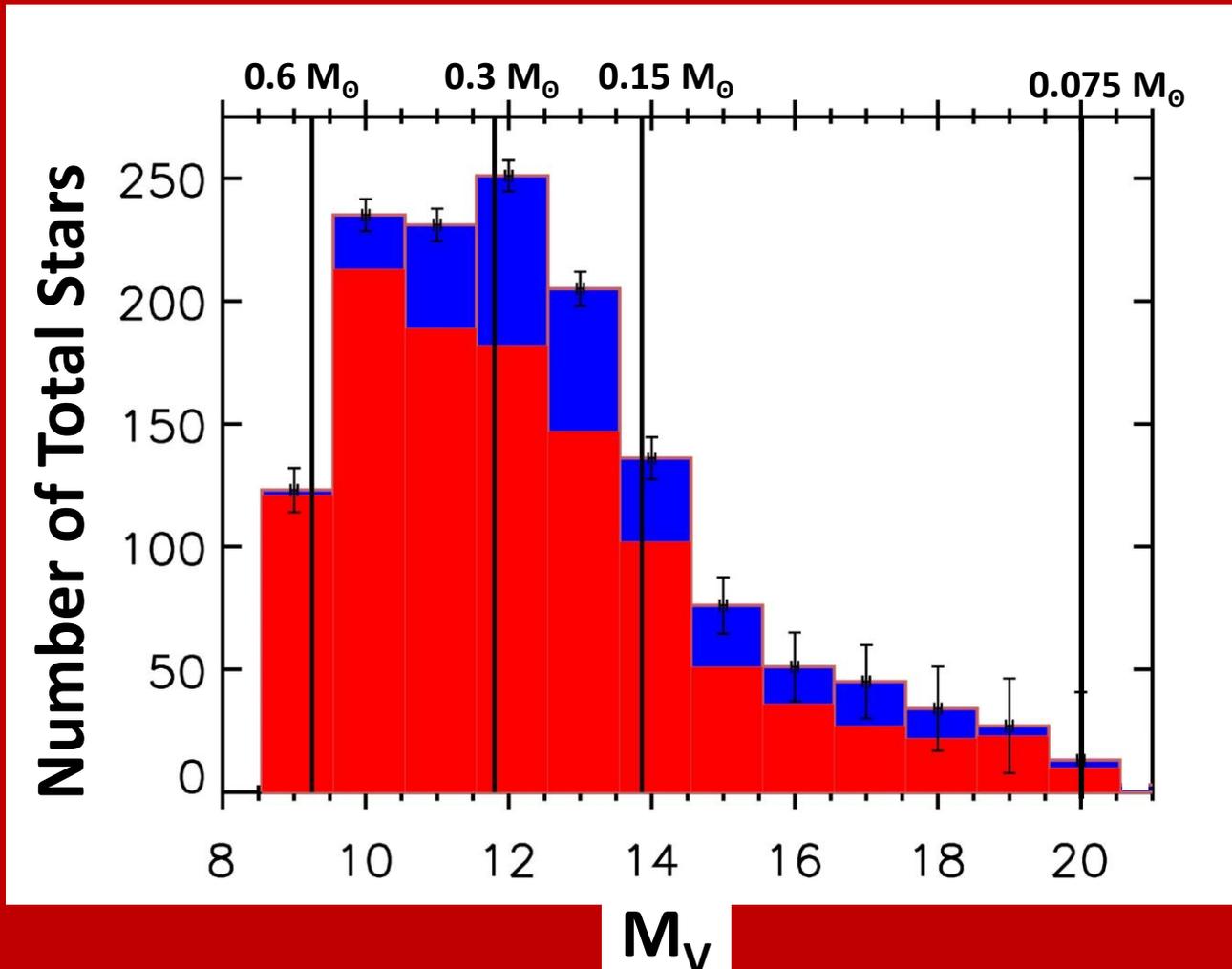
BIASES & SELECTION EFFECTS

- **Close Companions (corrected)**
- **Malmquist Bias (\uparrow MR)**
- **Parallax Error Limits (\downarrow MR)**
- **Faint Companions (\downarrow MR)**
- **Missing Faint Primaries (\uparrow MR)**

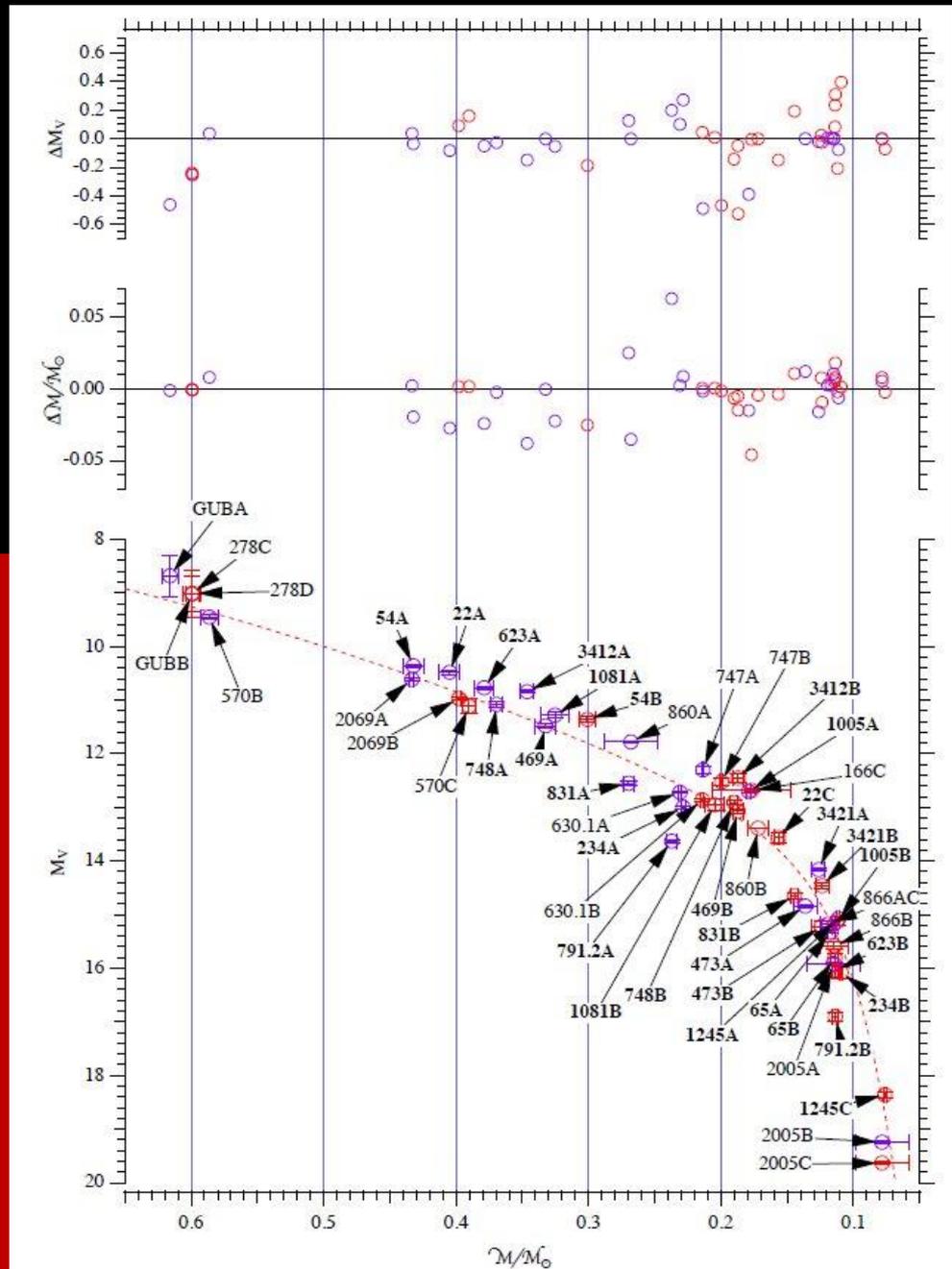
DETECTION LIMITS



LUMINOSITY FUNCTION – Primaries + Companions



MASS – LUMINOSITY RELATION

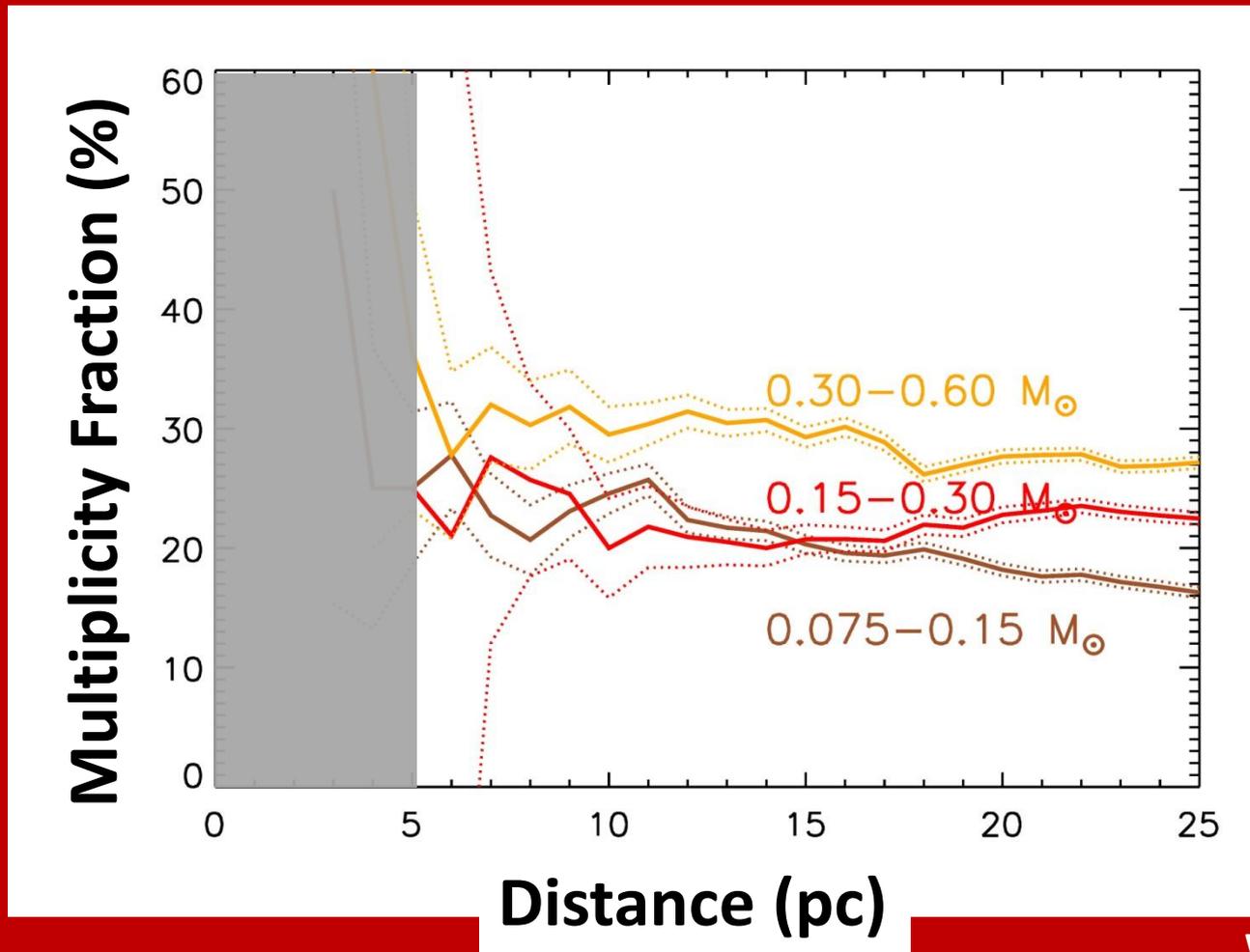


from Benedict+ 2016

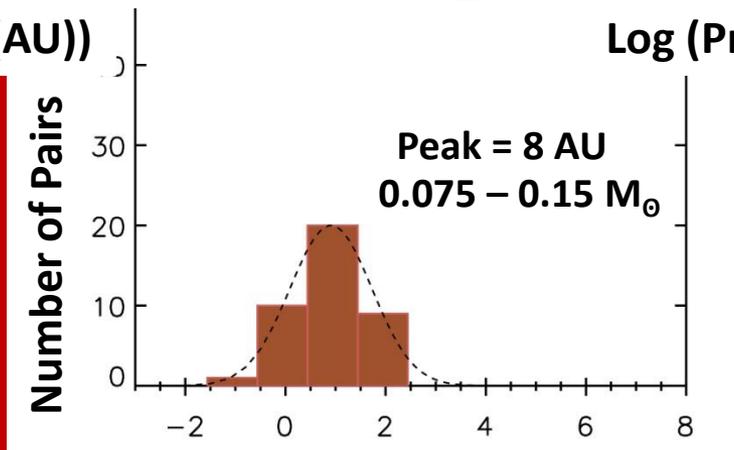
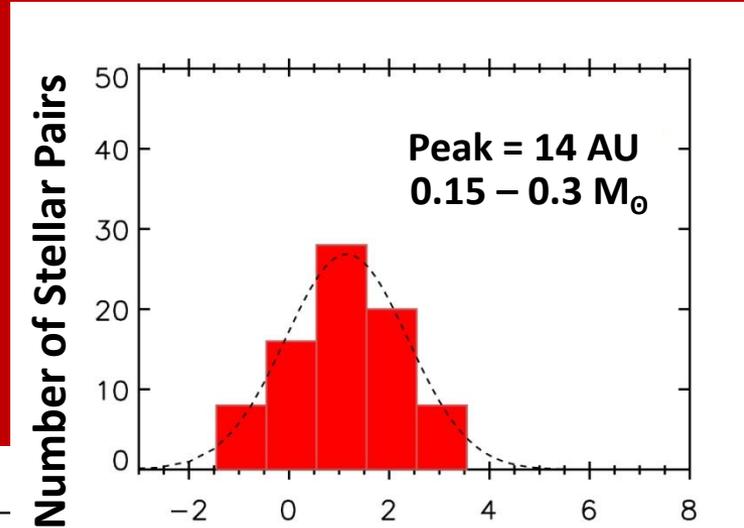
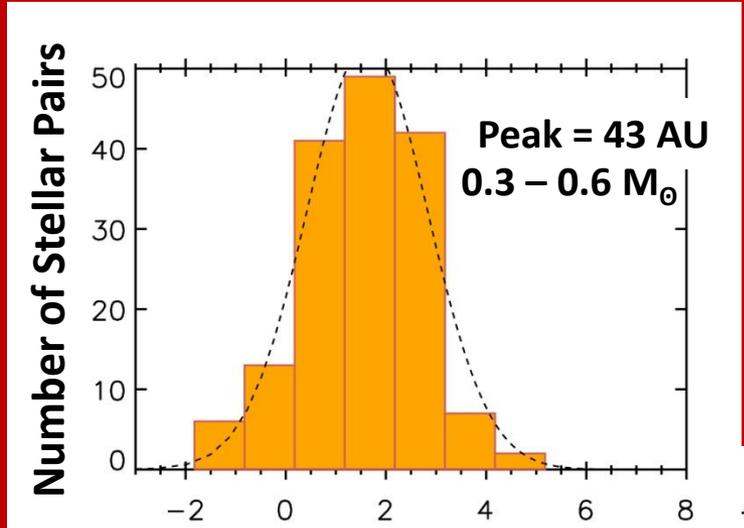
PARALLAX SOURCES

SAMPLE ORIGIN	NORTH	SOUTH	Notes
YPC + HIP	472	269	Within 25.0 pc from Compendia
RECONS - Published	26	247	Already in Literature
RECONS - Unpublished	5	27	This Work + Coming Soon
From Literature (1995 – 2012)	55	20	Additions from Others
TOTAL	558	563	1121

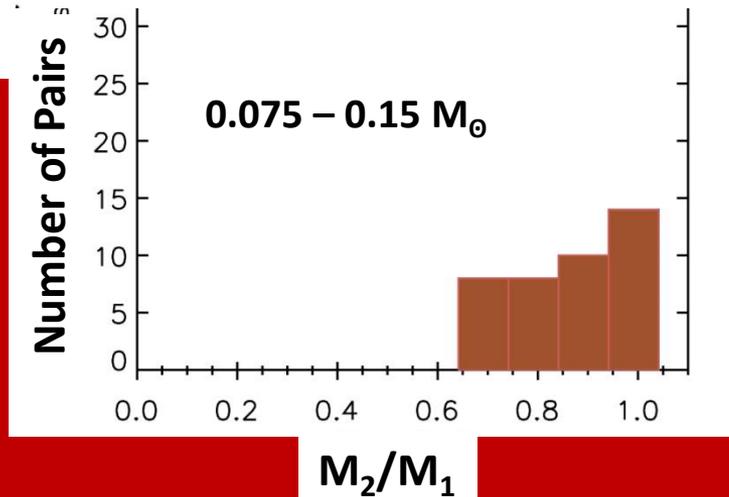
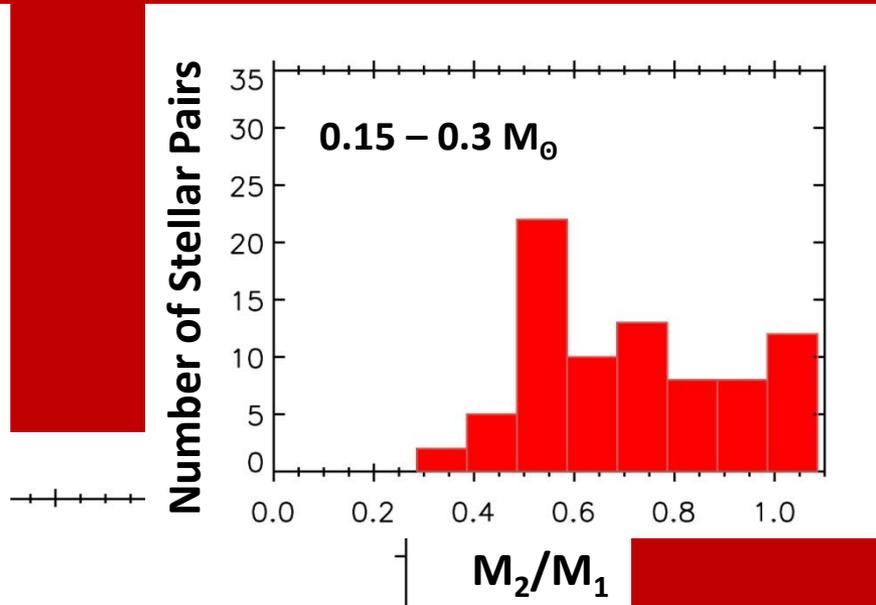
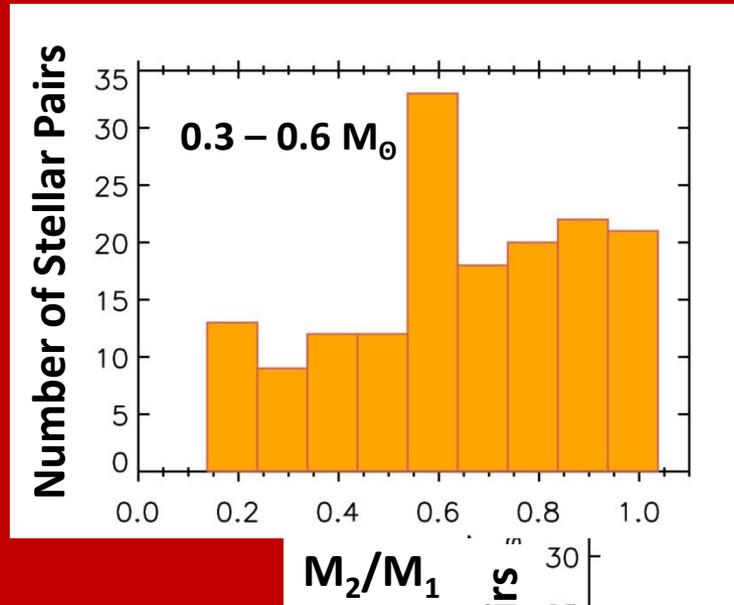
M DWARF MULTIPLICITY AS A FUNCTION OF MASS



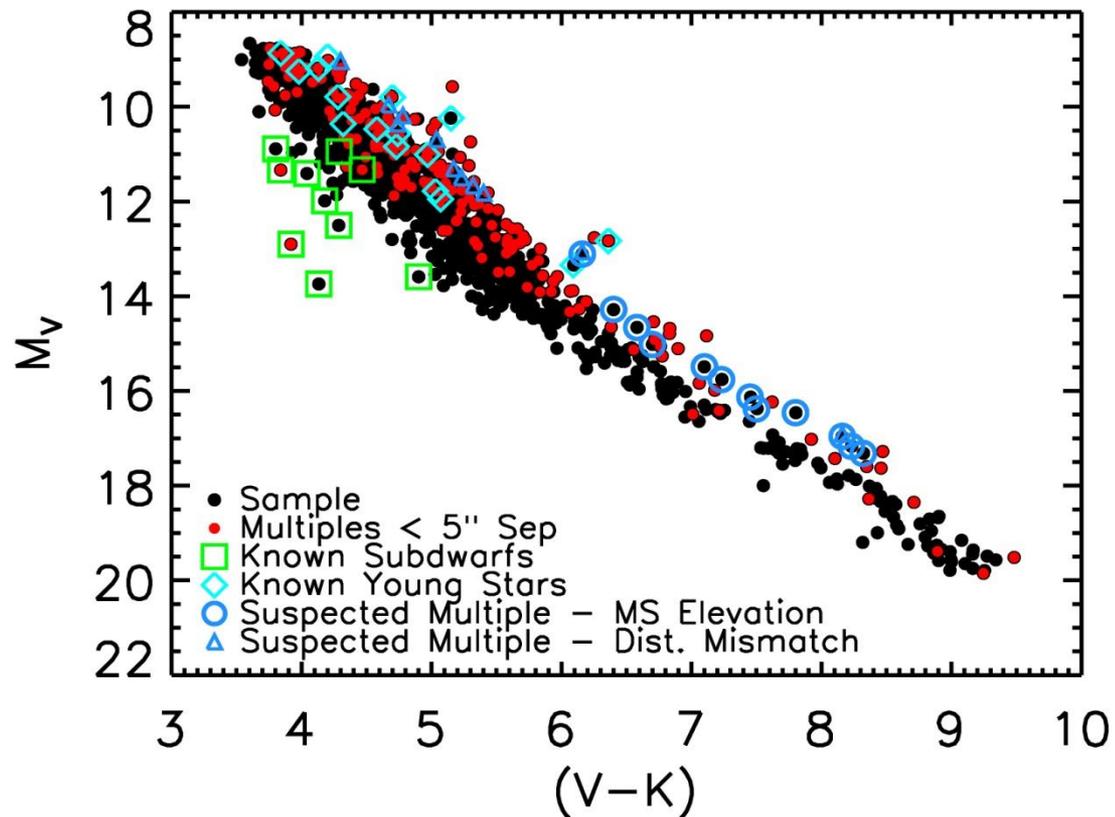
SEPARATION BY MASS SUBSET



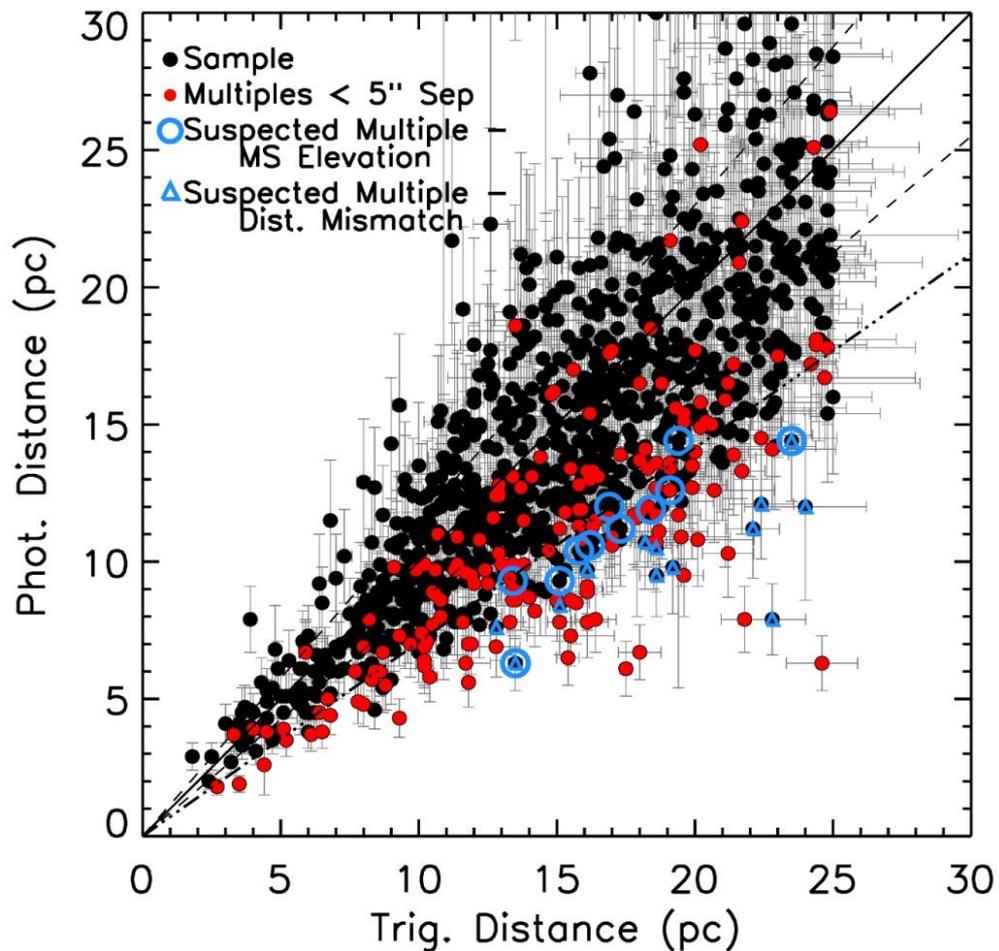
MASS RATIO BY MASS SUBSET



H-R DIAGRAM ELEVATION



DISTANCE MISMATCH



HOW IS THE MASS IN MULTIPLES DISTRIBUTED?

