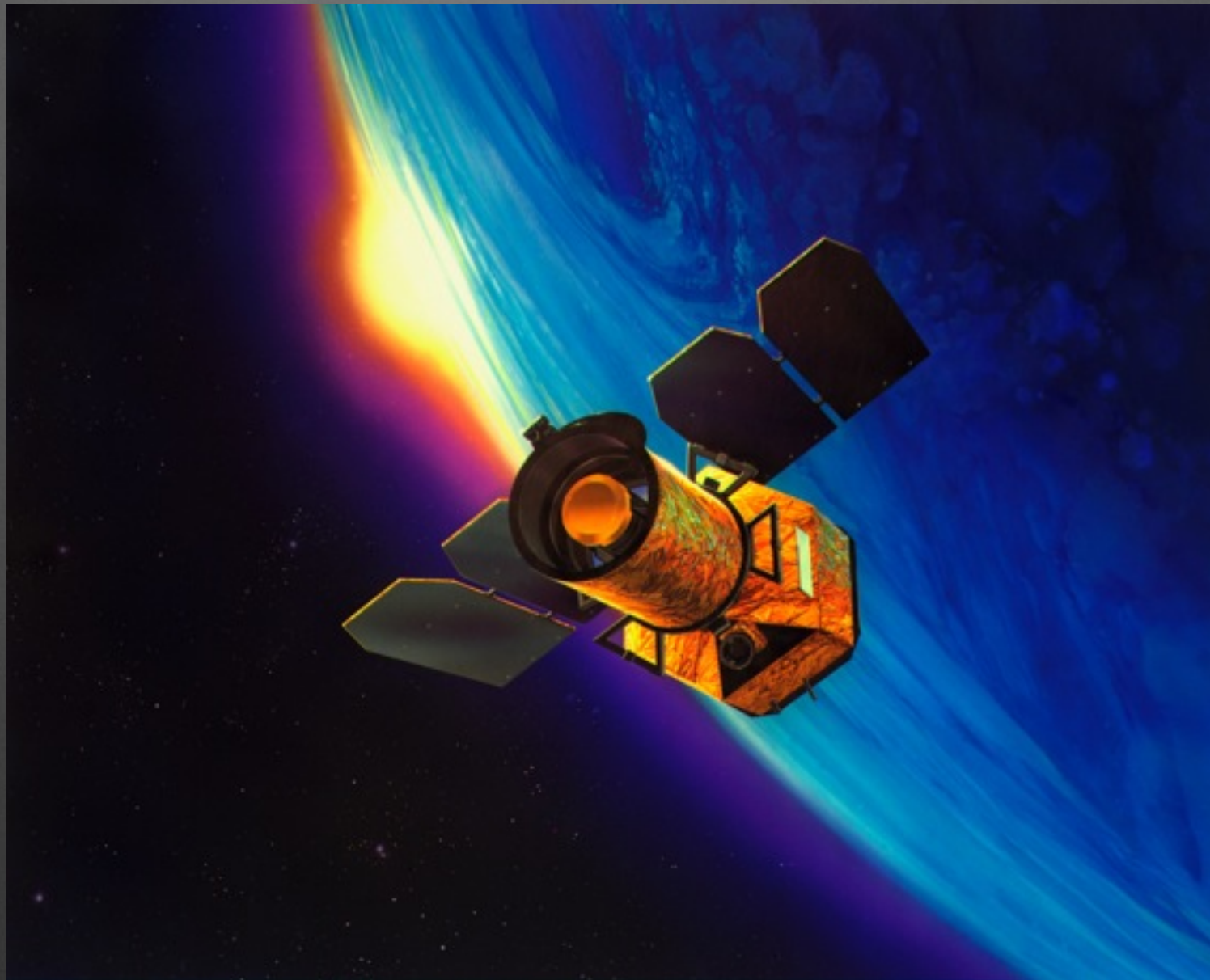


RR Lyrae in Gaia



Giuliano Giuffrida

RASPUTIN

Resolved And unresolved Stellar PopUlaTIoNs

ASDC-INAF

Giuseppe Bono (Univ. Tor Vergata Rome)

Vittorio Braga (Univ. Tor Vergata Rome)

Massimo Dall'Ora (INAF-OACN)

Francesco Verrecchia (ASDC-INAF)

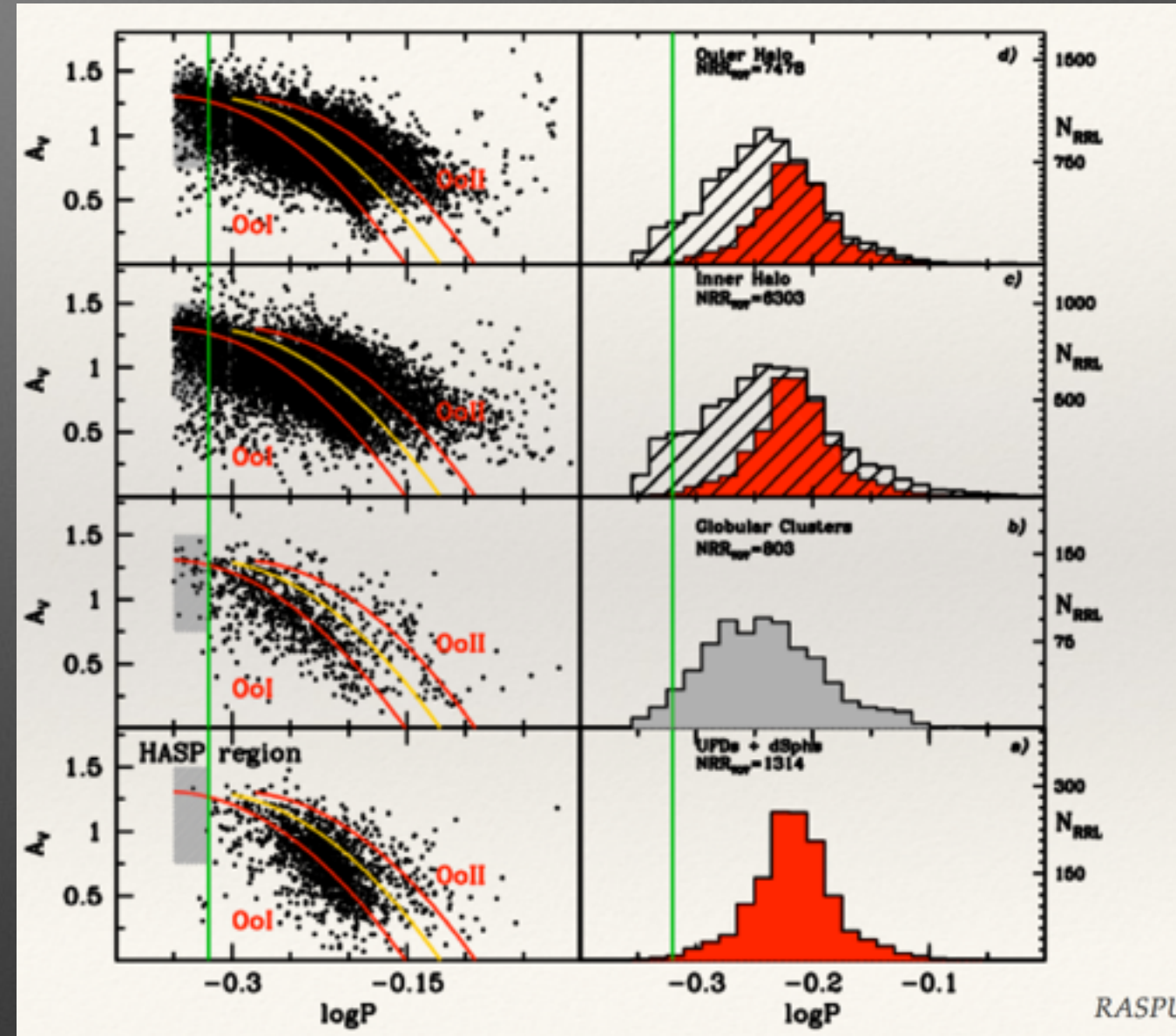
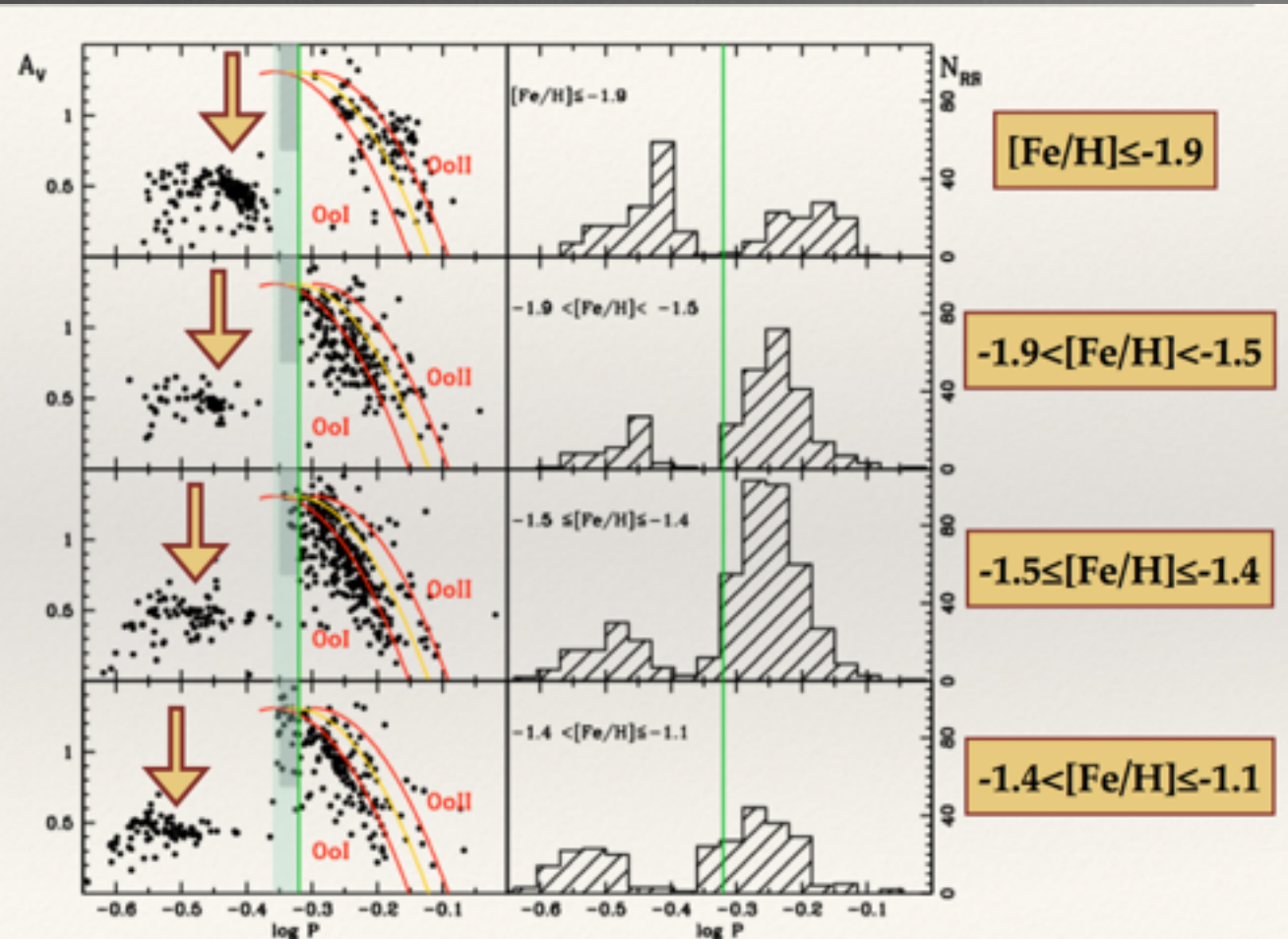
Giacinto Iannicola (INAF-OAR)

Ivan Ferraro (INAF-OAR)

Roberto Buonanno (INAF-OACTe)

Fiorentino

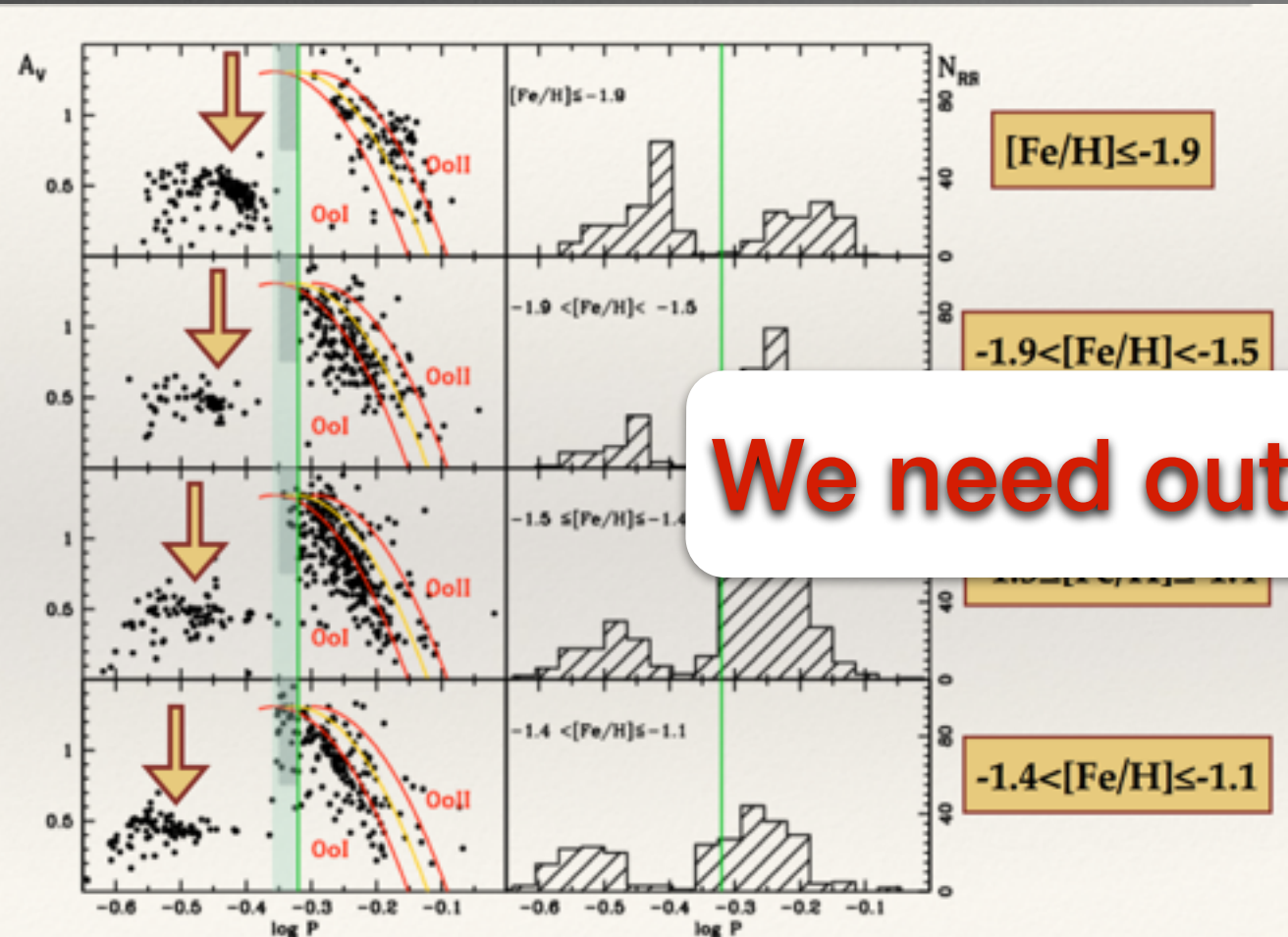
Monday's talk



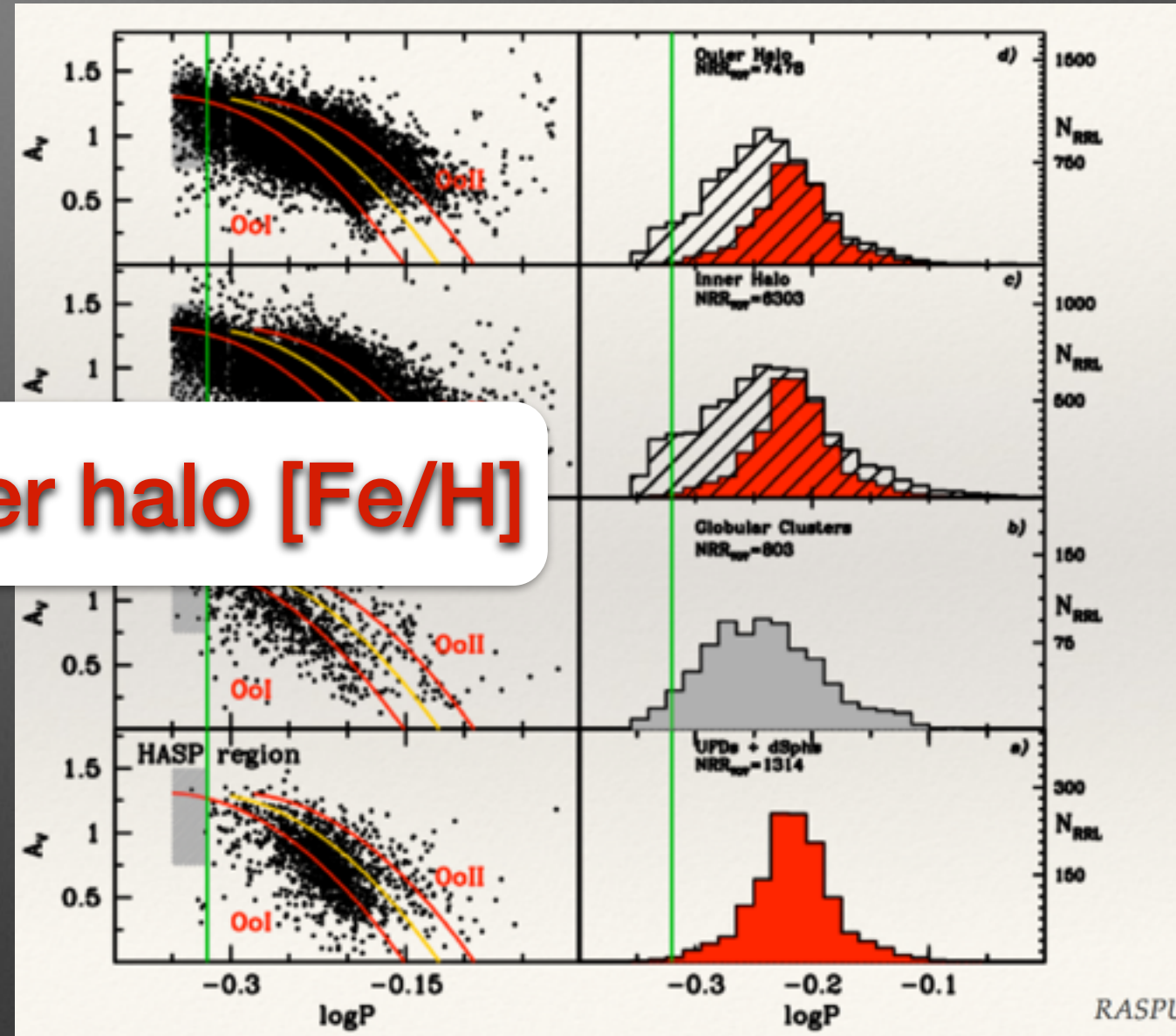
+ Gilmore Tuesday talk and others

Fiorentino

Monday's talk



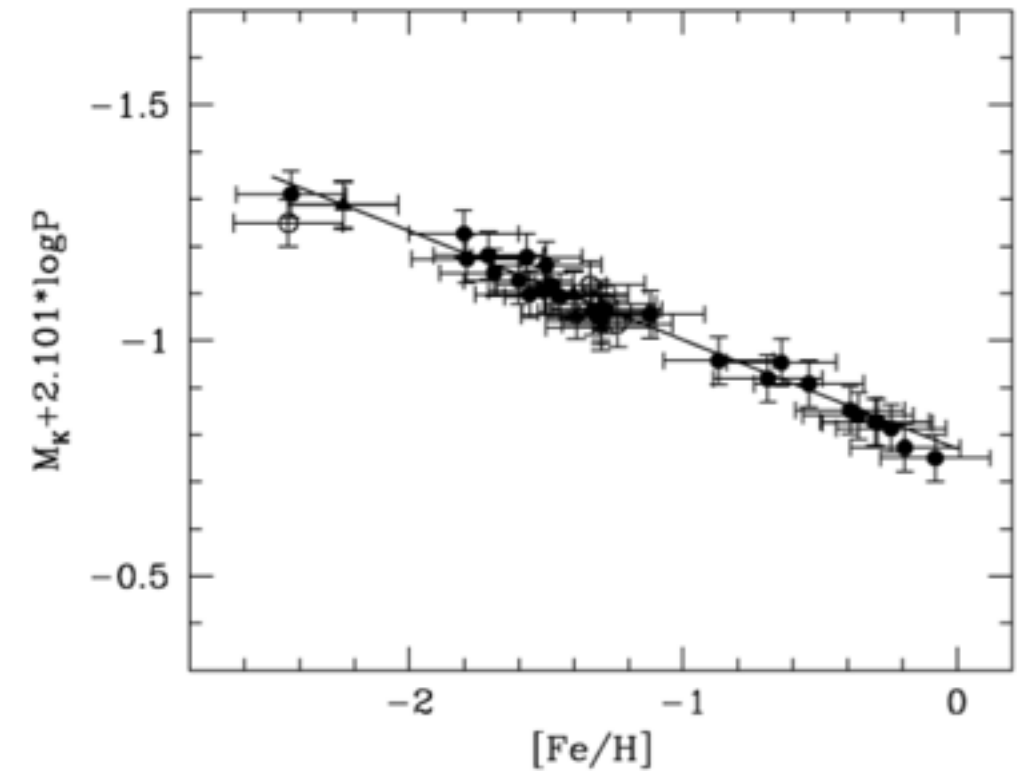
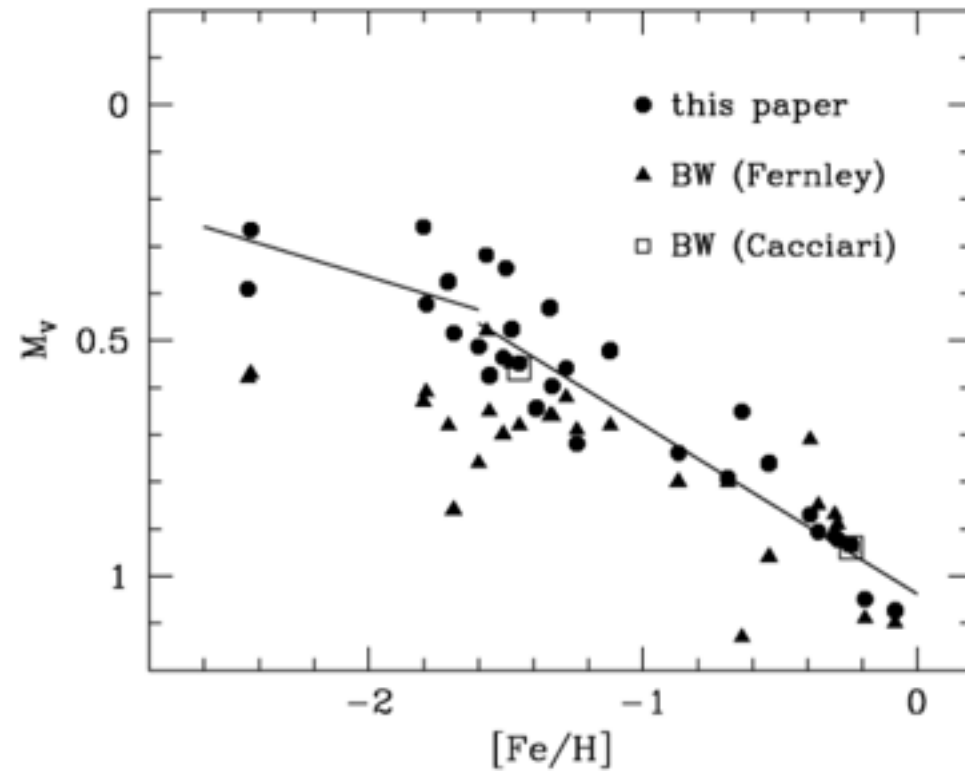
We need outer halo $[\text{Fe}/\text{H}]$



+ Gilmore Tuesday talk and others

RR Lyrae

Period-Luminosity-Metallicity



Bono et al 2003

$$MK = -0.770 + 0.231 * [Fe/H] - 2.101 * \log(Pf)$$

$$MV = 0.718 + 0.177(\pm 0.020) * [Fe/H] \quad ([Fe/H] \leq -1.6)$$

$$MV = 1.038 + 0.359(\pm 0.020) * [Fe/H] \quad ([Fe/H] > -1.6)$$

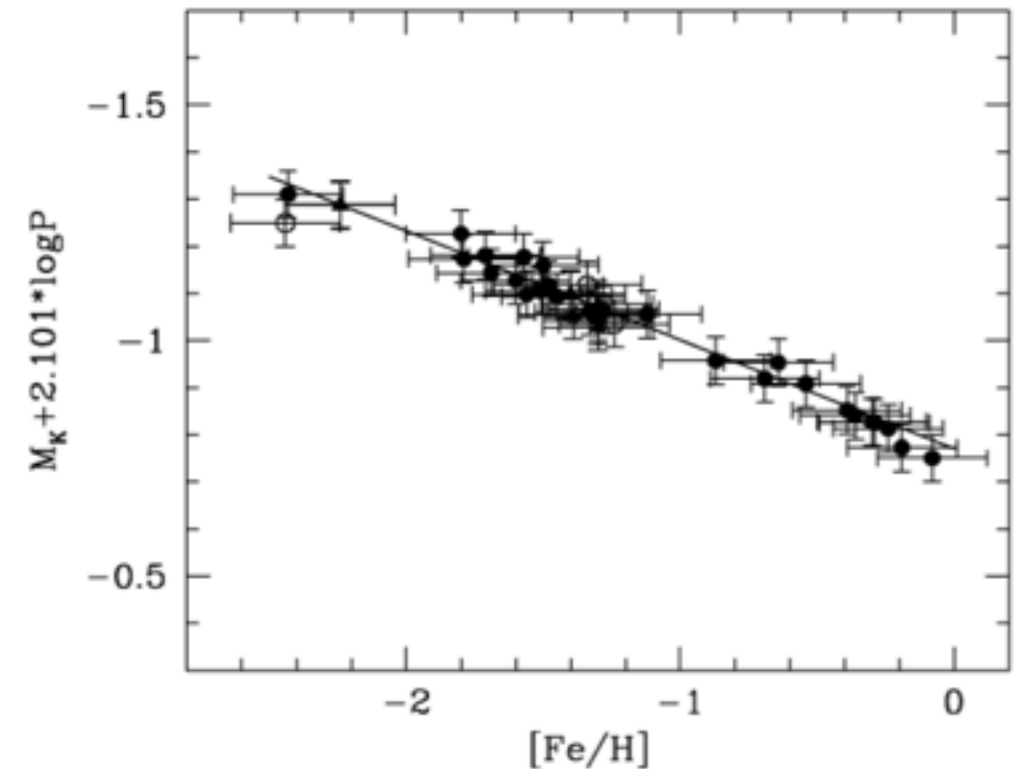
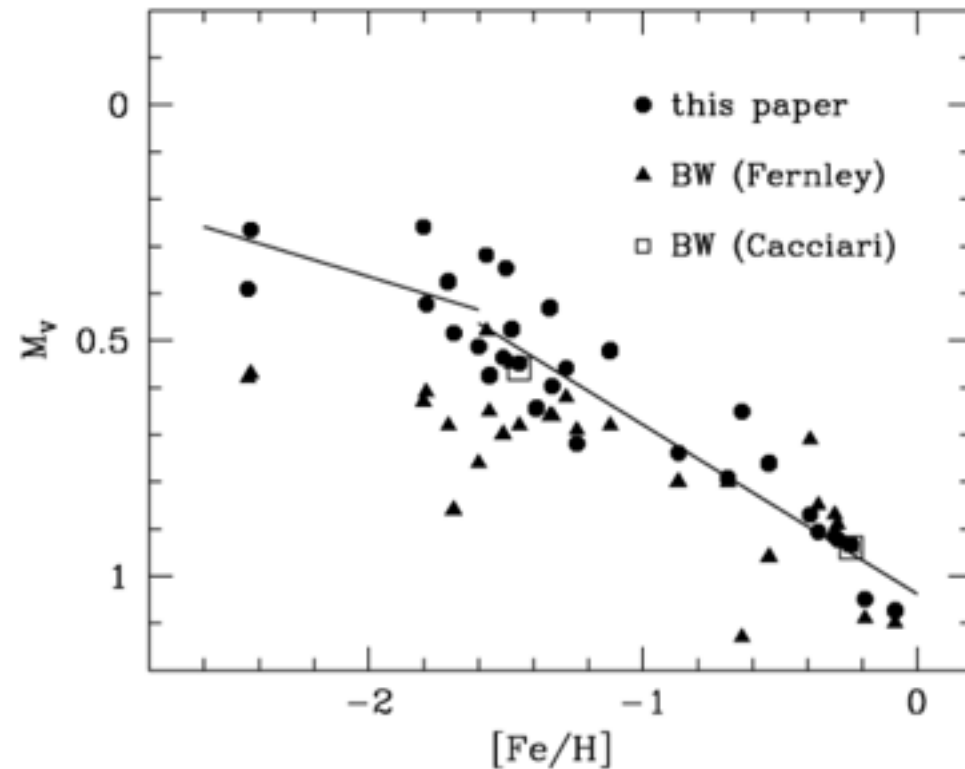
Dambis et al 2013

$$MK = -0.769(\pm 0.088) + 0.088(\pm 0.026) * [Fe/H] - 2.33 * \log(Pf)$$

$$MV = 1.094(\pm 0.091) + 0.232(\pm 0.020) * [Fe/H]$$

RR Lyrae

Period-Luminosity-Metallicity



Bono et al 2003

Dambis et al 2013

$$M_K = -0.770 + 0.231 * [Fe/H] - 2.101 * \log(Pf)$$

$$M_K = -0.769(\pm 0.088) + 0.088(\pm 0.026) * [Fe/H] - 2.33 * \log(Pf)$$

$$M_V = 0.718 + 0.177(\pm 0.020) * [Fe/H] \quad ([Fe/H] \leq -1.6)$$

$$M_V = 1.094(\pm 0.091) + 0.232(\pm 0.020) * [Fe/H]$$

$$M_V = 1.038 + 0.359(\pm 0.020) * [Fe/H] \quad ([Fe/H] > -1.6)$$

UV?

Outline

- **Galex data**
- **Period-Color-[Fe/H] on FUV-NIR**
- **[Fe/H] on Halo**

Galex

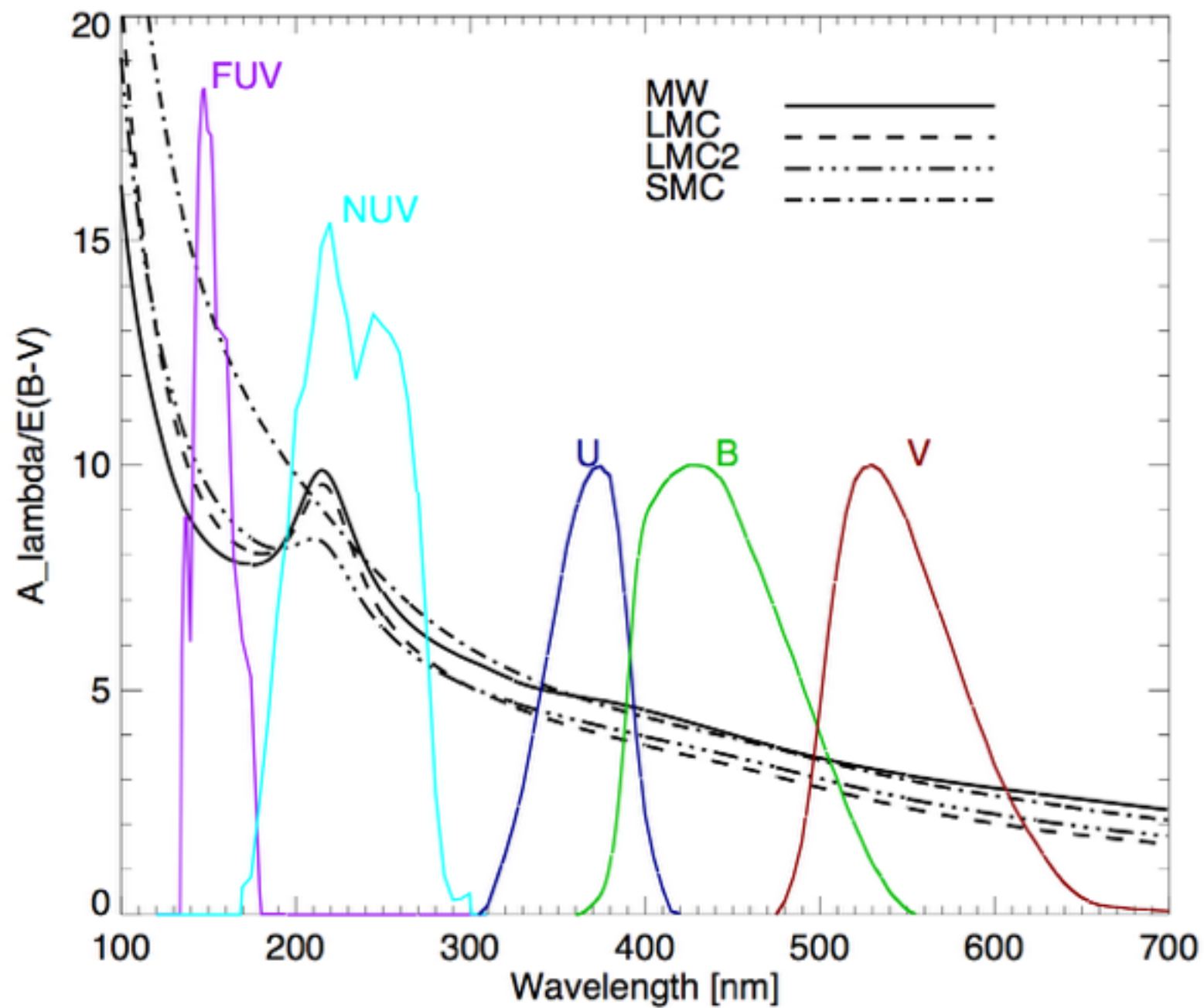
Launched in 2003

2 Filters FUV (135-175 nm)
 NUV (175-275 nm)

Field of view: 1.2 d (circular)

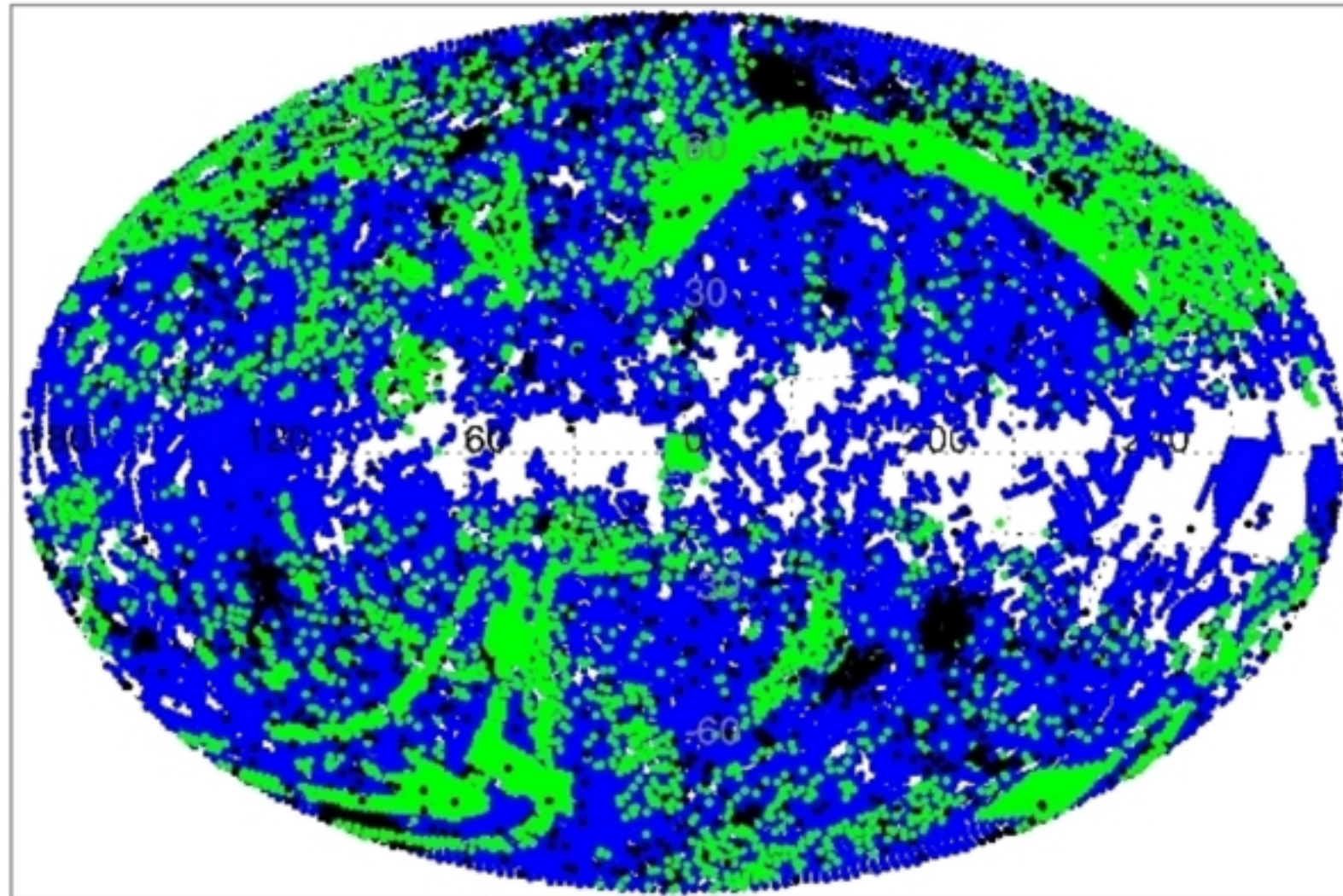
Angular resolution: 4.5'' (FUV)
 6.5'' (NUV)





http://dolomiti.pha.jhu.edu/papers/Bianchi_GALEX_SF.pdf

GALEX GR6/7 NUV



All Sky Survey (AIS)

Medium Imaging Survey (MIS)

Bianchi 2014

<http://galex.stsci.edu/GR6/>

145,755 fits
598 GB

762,339,557
Source measurements

Barbara A. MIKULSKI ARCHIVE FOR SPACE TELESCOPES

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GALEX GR6/GR7 Home About GALEX Getting Started GalexView CASJobs Scan Mode

GALEX GR6/7 Data Release

[GR6 Summary](#) [GR7 Summary](#)
[GR6 Pipeline Description](#) [GR7 Pipeline Description](#)

MORE GR6 and NEW GR7 data now available, click for summary.

	AIS	DIS	MIS	NGS	GII	CAI	SPECTRA	TOTAL
GR1	3074	14	112	52	-	-	7	3259
GR2/GR3	15721	165	1017	296	288	20	41	17548
GR4/GR5	28269	292	2161	458	788	38	174	32180
GR6/GR7	31285	720	6964	716	2112	87	311	45195

Welcome to the GALEX Public Archive

- [The GALEX Mission](#)
- [GALEX Surveys](#)
- [Public GALEX Releases](#)
- [Guest Investigator Program](#)
- [Close-out Deliverables \(GR7\)](#)
- [Scan Mode Data](#)

The GALEX Mission

Welcome to the MAST public access site for browsing and retrieving GALEX Release 7 data products. The Galaxy Evolution Explorer (GALEX) satellite is a NASA mission led by the [California Institute of Technology](#) to investigate how star formation in galaxies evolved from the early Universe up to the present. GALEX uses microchannel plate detectors to obtain direct images

NEWS

Feb 27, 2013:
[MORE GR6 and NEW GR7 data now available, click here to see a summary](#)

Oct 07, 2011:
[GR6 Spectra are now available, click here to find out more](#)

Oct 07, 2011:
[GR6 addendum released, comprised of 12 DIS, 389 MIS and 66 GII tiles, click here to find out more](#)

Mar 16, 2011:
[Catalogs of GALEX UV unique sources and of UV-optical matched sources](#)

Dec 21, 2010:
[GALEX GR6 cross matched against SDSS DR7, available thru casjobs only. Click here for a tutorial](#)

RSS 2.0

Globular Cluster M2

Halo RR Lyrae - Public catalogs

Position

[Fe/H]

Variability Parameters

Spectroscopic/Photometric

ASAS3

Catalina

Linear

Catalina
SDSS

Dambis

Loneos

Quest

ROTSE

HR

Mateu

NSVS

Wallerstein et al 2010
Hansen et al 2011
Lambert et al 1996
Clementini et al 1995
Kolenberg et al 2010

Pojmanski et al 2005 (ASAS3)

Drake et al 2013-2014 (Catalina)

Dambis et al 2013 (Dambis)

Mateu et al 2012 (Quest-low)

Sesar et al 2013 (Linear)

Miceli et al 2008 (Loneos)

Vivas et al 2004-2006 (Quest)

Akerlof et al 2000 (ROTSE)

Hoffman et al 2009 (NSVS)

Homogeneous Reddening and Distance

Schlafly
&
Finkbeiner
(2011)

Position Variability Parameters			[Fe/H] Spectroscopic/Photometric	
ASAS3	Catalina	Linear	Catalina SDSS	Dambis
Loneos	Quest	ROTSE	HR	Mateu
NSVS				

Wallerstein et al 2010
Hansen et al 2011
Lambert et al 1996
Clementini et al 1996
Kolenberg et al 2010



PL relation

FU

$$M_j = -2.13557 \cdot \text{Log}(P) - 0.79116$$

$$M_h = -2.40348 \cdot \text{Log}(P) - 1.0591$$

$$M_k = -2.43557 \cdot \text{Log}(P) - 1.11740$$

$$M_{w1} = -2.38 \cdot \text{Log}(P/0.55) - 0.495$$

Braga et al 2014 (submitted)

Marconi et al (in preparation)

Klein et al 2014

FO

$$M_j = -2.34836 \cdot \text{Log}(P) - 1.24632$$

$$M_h = -2.57585 \cdot \text{Log}(P) - 1.49593$$

$$M_k = -2.59197 \cdot \text{Log}(P) - 1.54985$$

$$M_{w1} = -1.64 \cdot \text{Log}(P/0.32) - 0.231$$



A_v

+

Distance



RR Lyrae

```

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'id'	Objectid:5411ba2d20a33cbb63b0c6b0e7	'name1'	'BS'	'name2'	'A'
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WISE

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Galex

[illegible]

The diagram illustrates the relationship between variability parameters and spectroscopic/photometric parameters for white dwarf classification. It is divided into two main sections: 'Position' and 'Spectroscopic/Photometric'.

Position Section:

- ASAS3:** A grid of 100 small images showing variability patterns.
- Catalina:** A grid of 100 small images showing variability patterns.
- Linear:** A grid of 100 small images showing variability patterns.
- Loneos:** A grid of 100 small images showing variability patterns.
- Quest:** A grid of 100 small images showing variability patterns.
- ROTSE:** A grid of 100 small images showing variability patterns.
- NSVS:** A grid of 100 small images showing variability patterns.

Spectroscopic/Photometric Section:

- Catalina SDSS:** A grid of 100 small images showing variability patterns.
- Dambris:** A grid of 100 small images showing variability patterns.
- HR:** A grid of 100 small images showing variability patterns.
- Mateu:** A grid of 100 small images showing variability patterns.

A red arrow points from the 'NSVS' section to a box containing the following text:

Wallerstein et al 2010
Hansen et al 2011
Lambert et al 1996
Clementini et al 1996
Kolenberg et al 2010

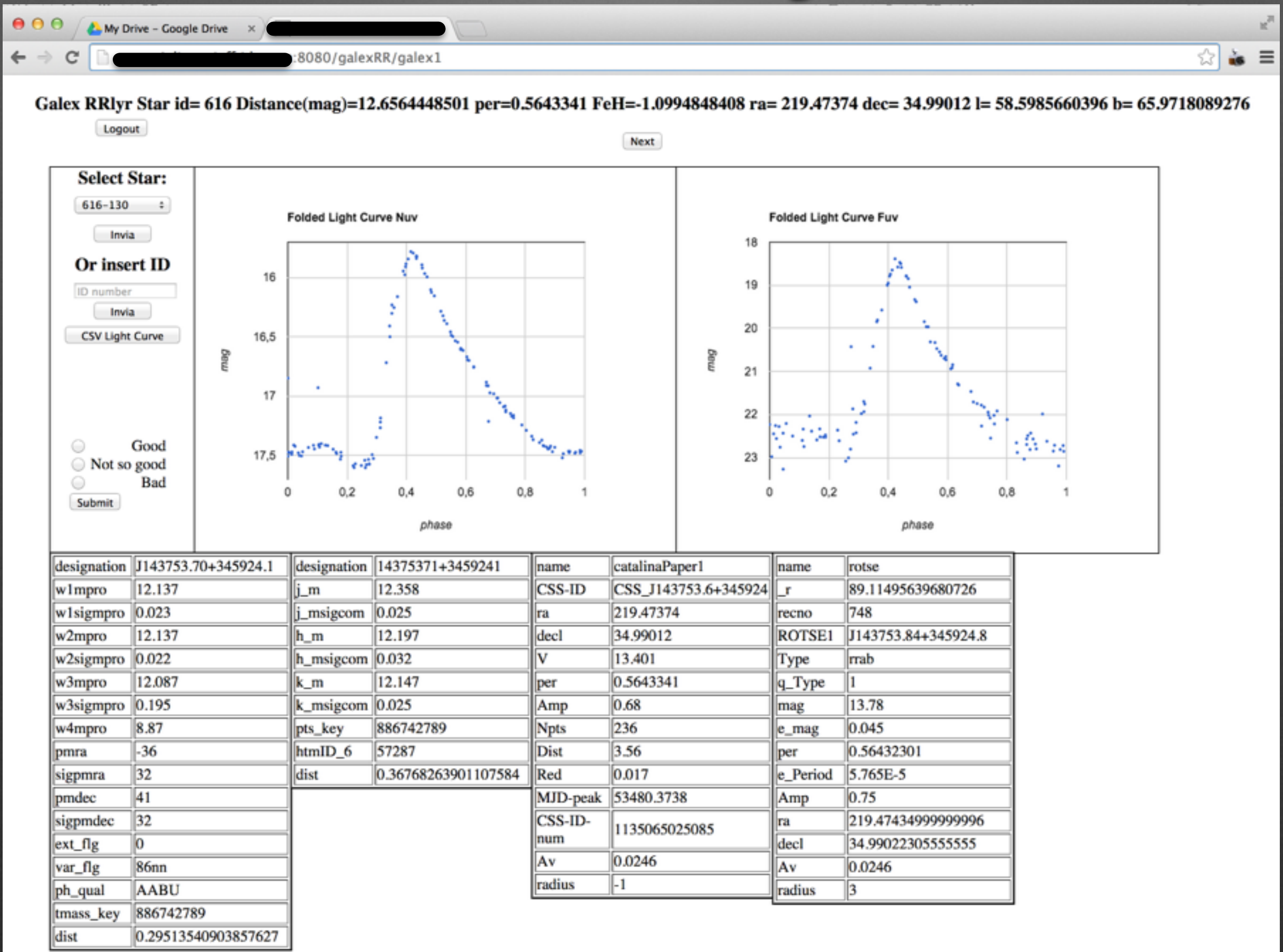
Wallerstein et al 2010
Hanson et al 2011
Lambert et al 1996
Clementini et al 1995
Kolenberg et al 2010

RR Lyrae Catalog

RR Surveys + 2MASS + WISE + Galex Obs

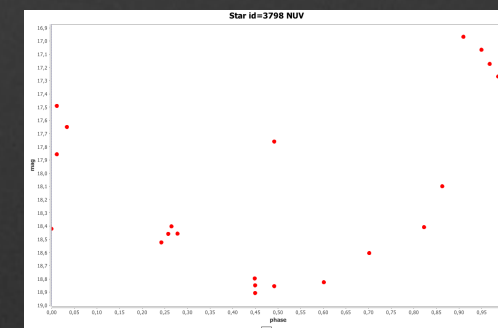
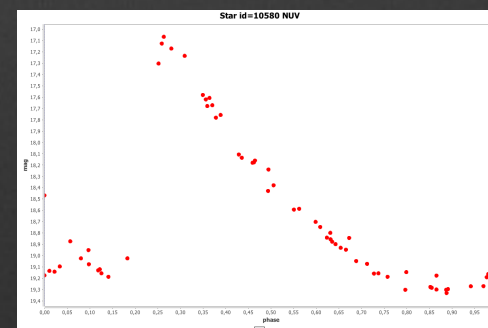
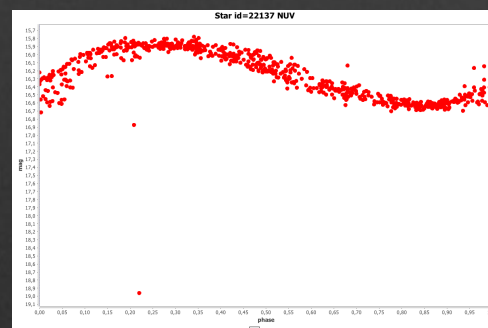
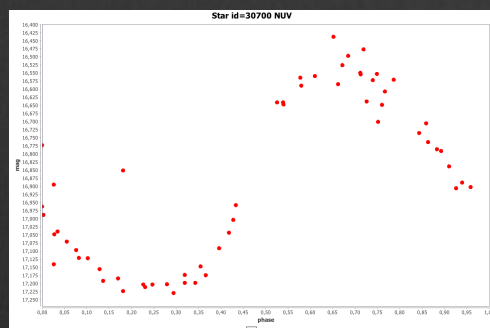
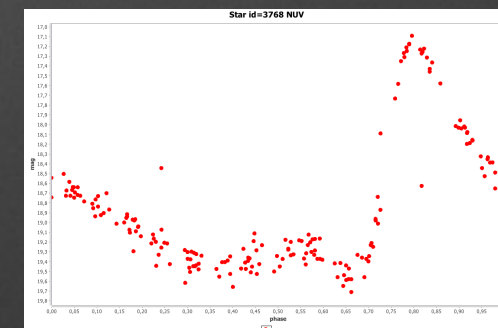
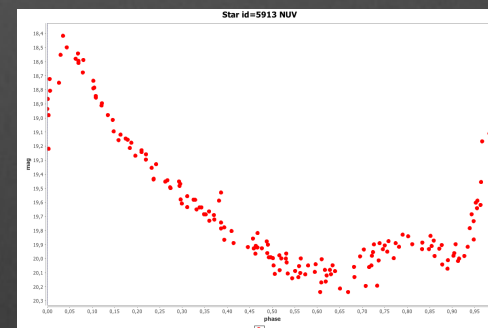
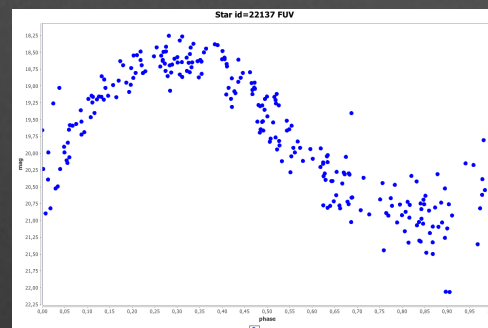
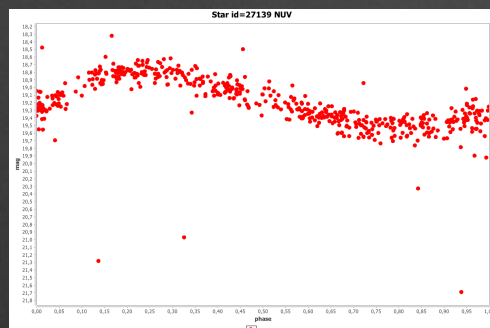
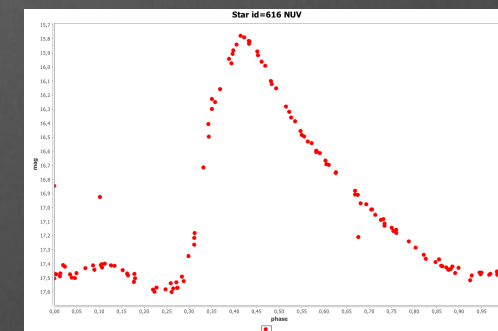
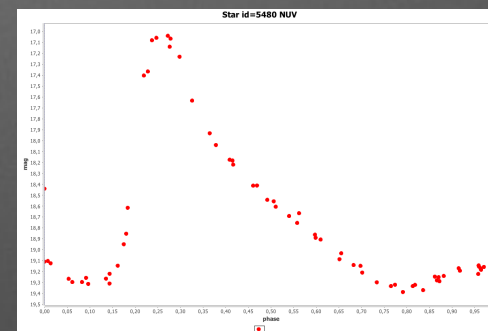
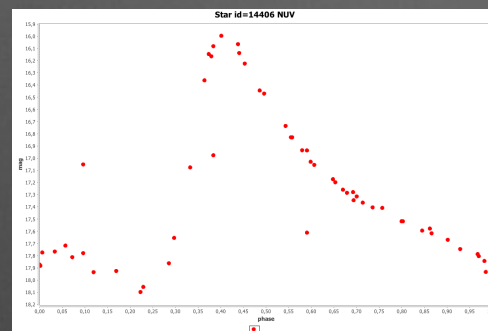
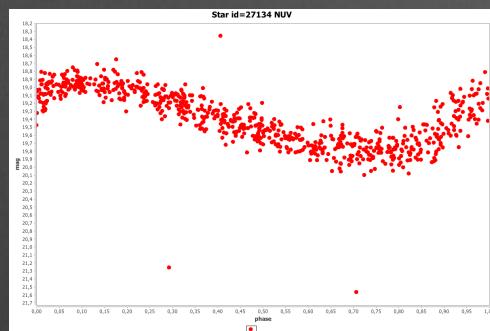
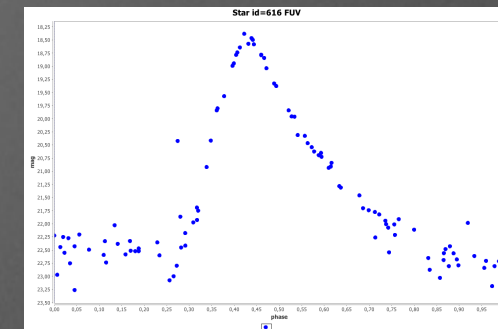
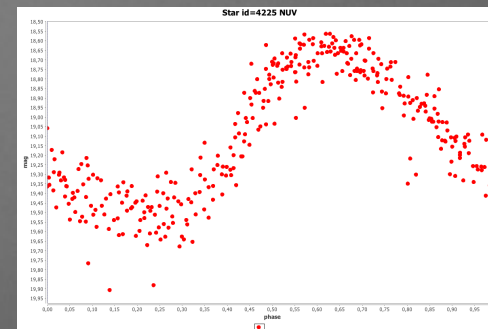
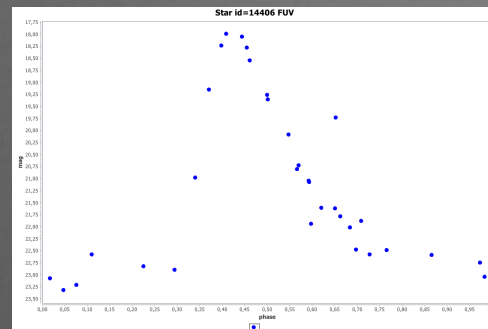
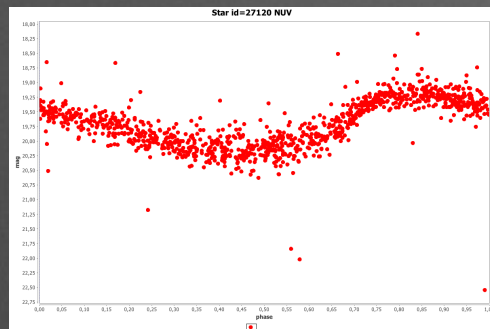
- ~30,000 RR Lyrae
- 1,502 with more than 10 Galex obs
- 19,759 with Kmag from 2MASS
- 1,977 with [Fe/H] (34 HR, 1415 Catalina/SDSS, 366 Dambis, 162 Mateu)

Galex RR Lyrae

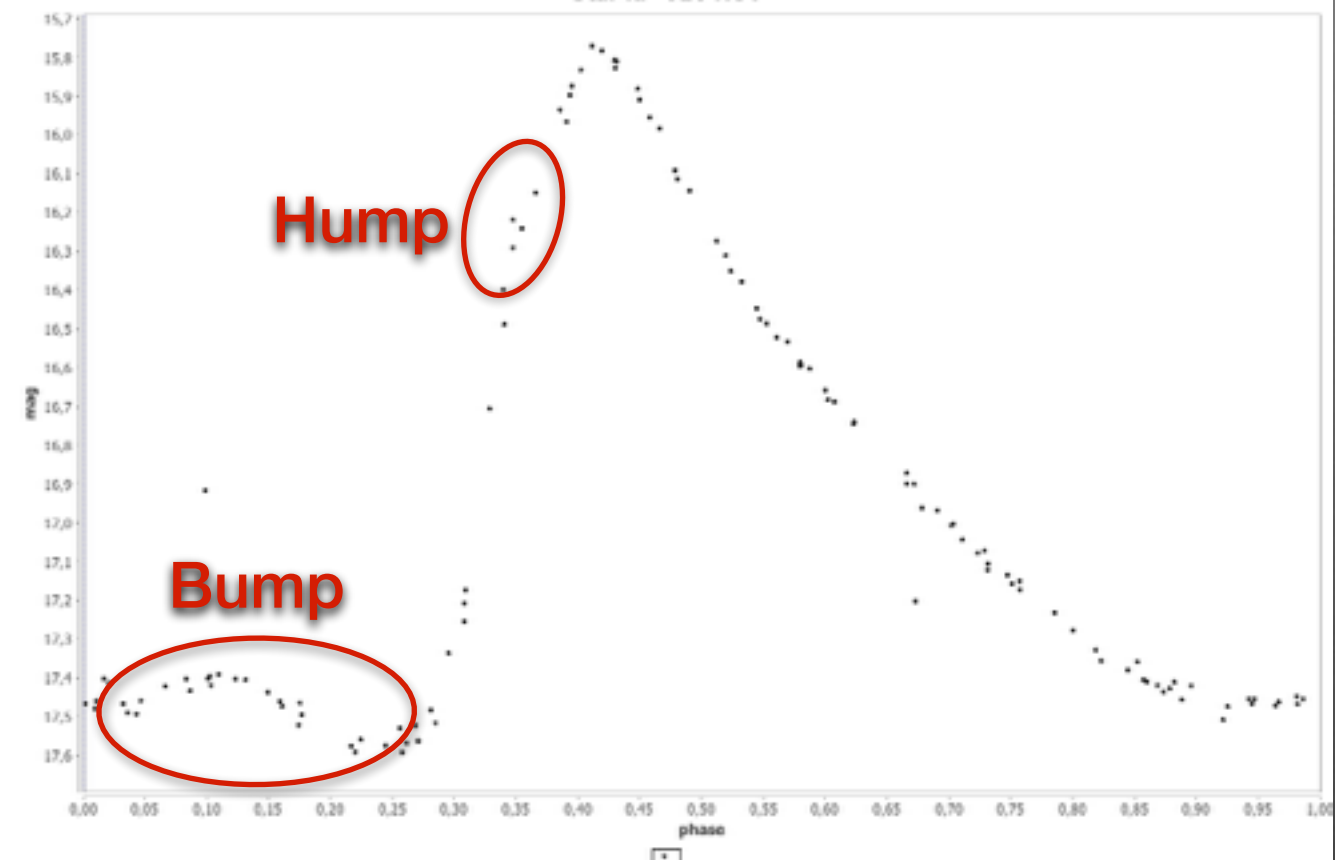


Galex RR Lyrae

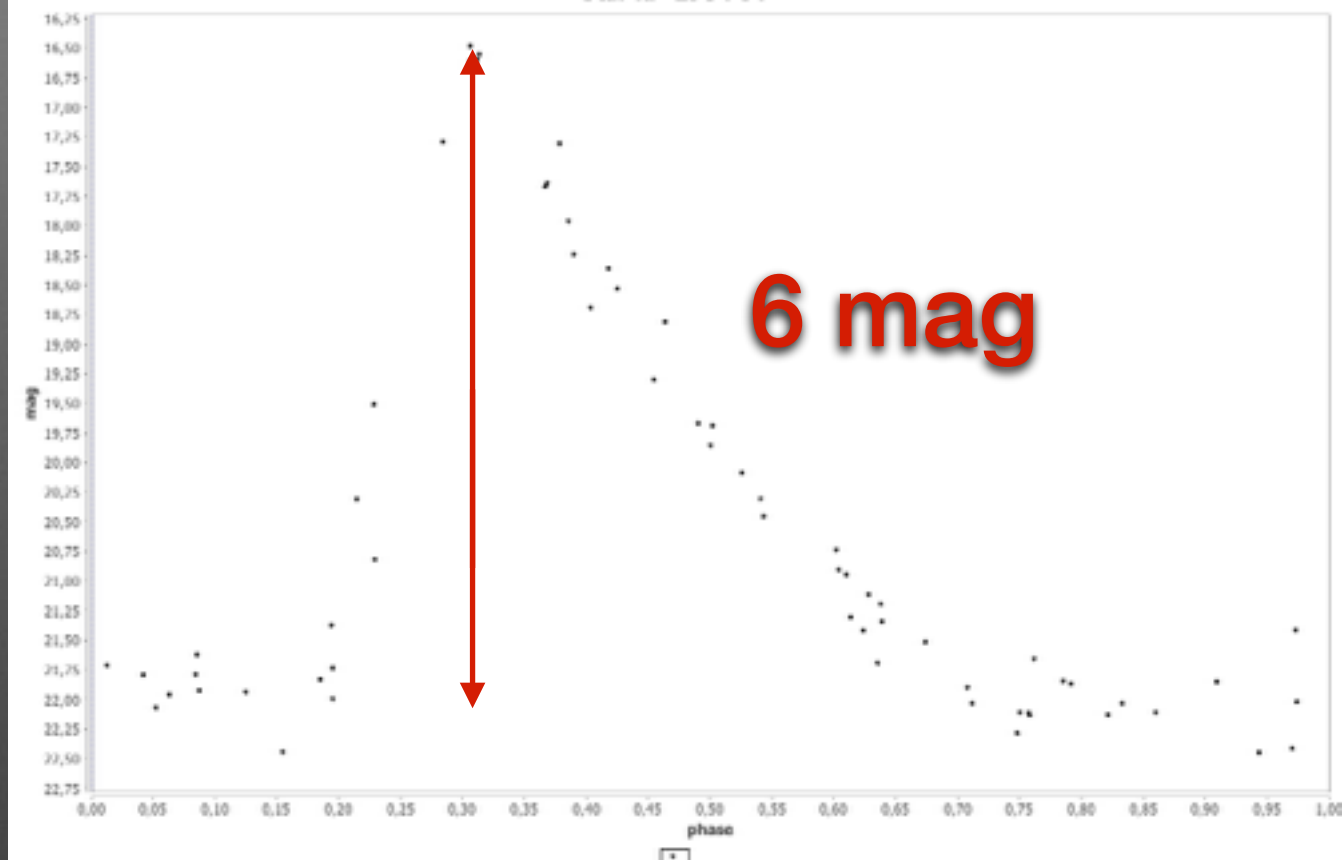
181 (**NUV**) + ~30 (**FUV**) light curves



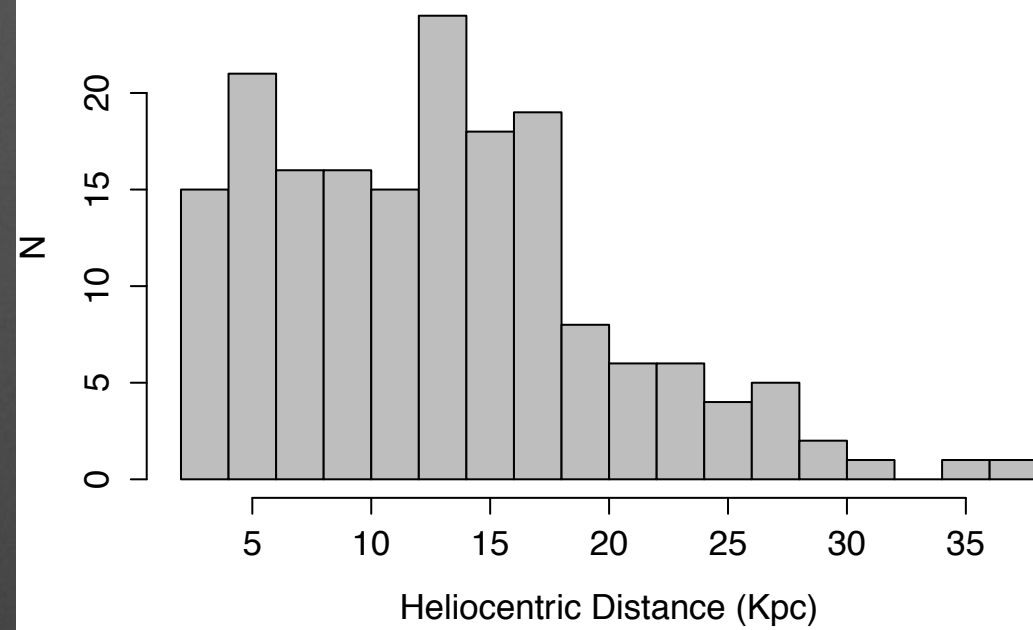
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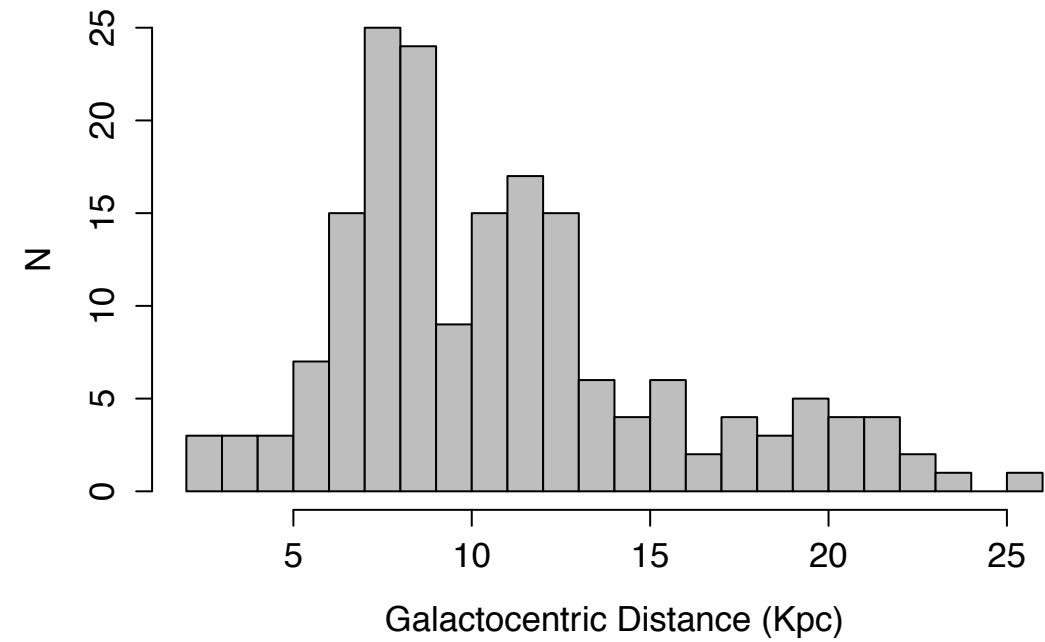
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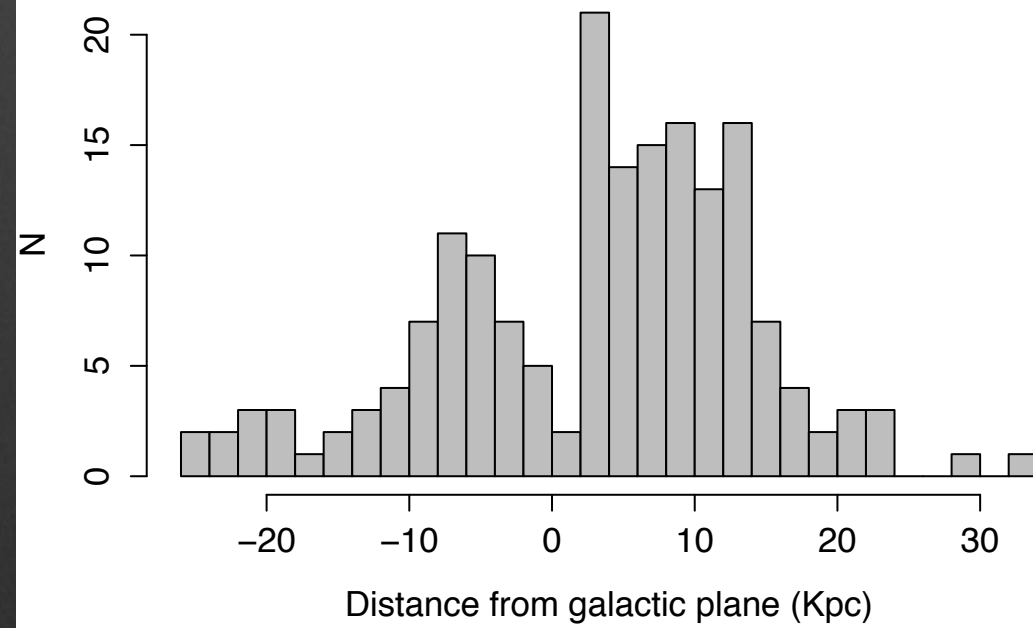
Galex Stars (181)



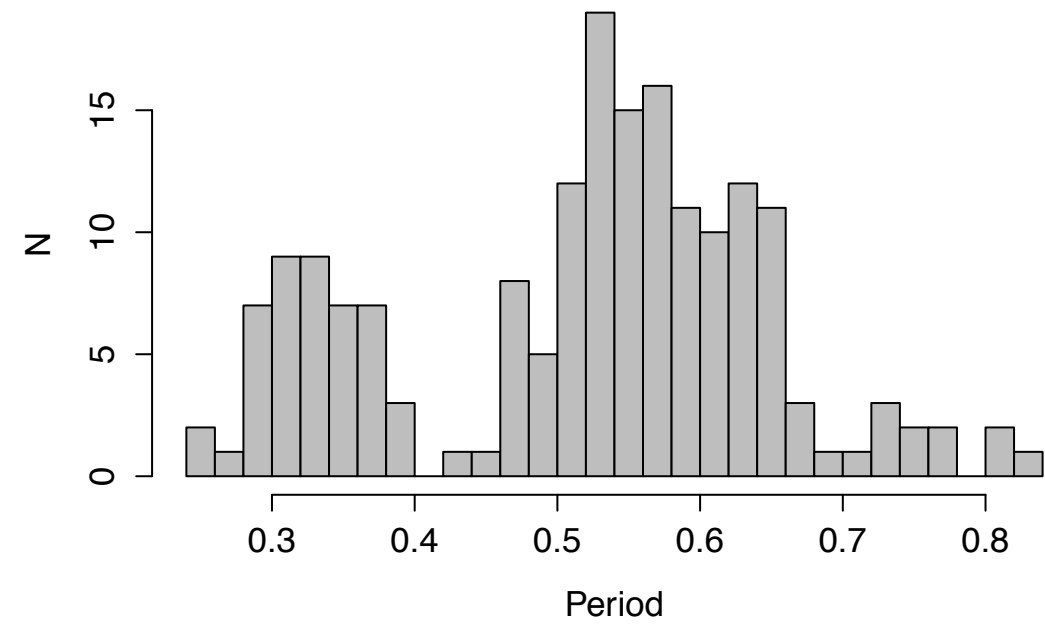
Galex Stars (181)



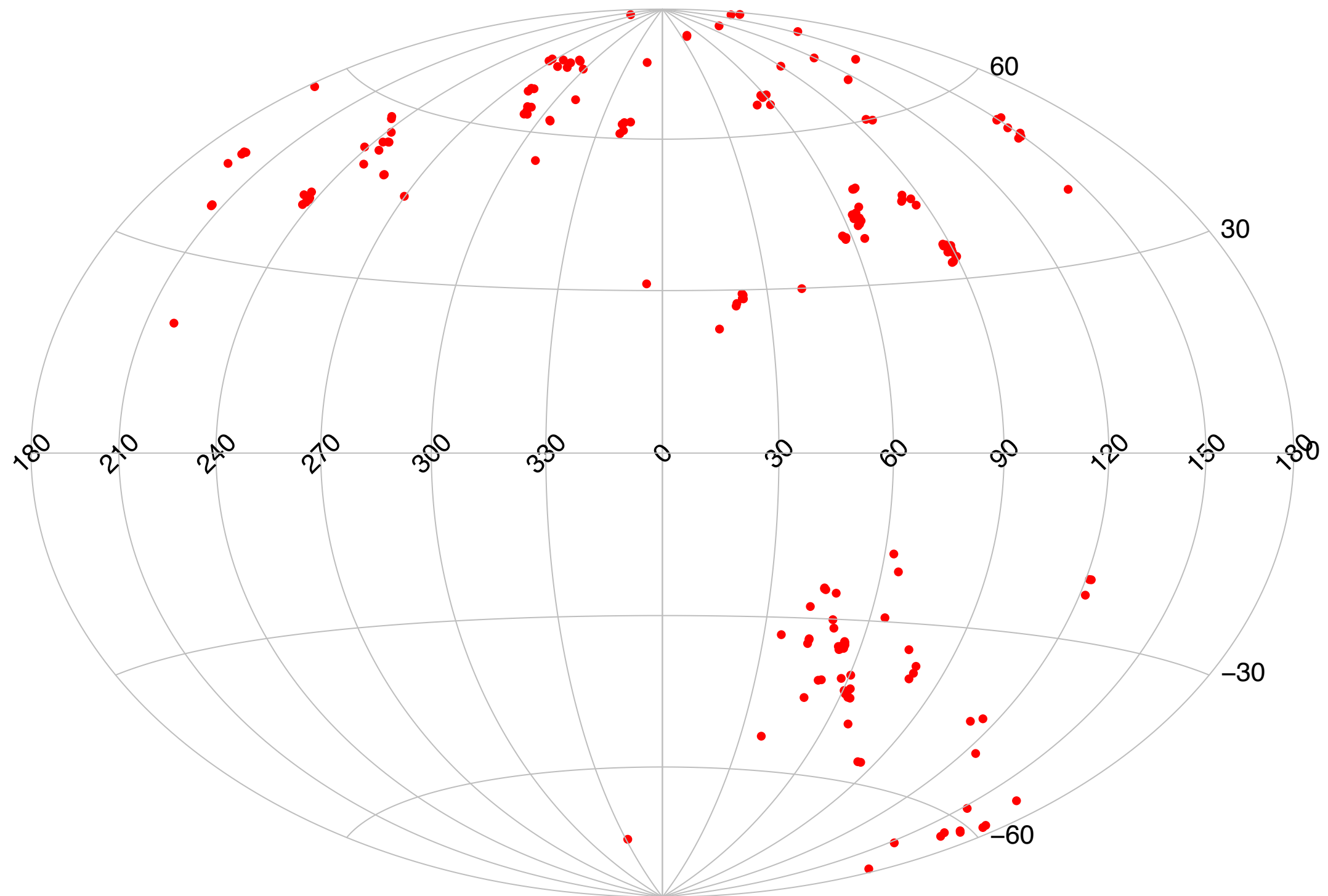
Galex Stars (181)



Galex Stars (181)



Galex RR Lyr - sky distribution



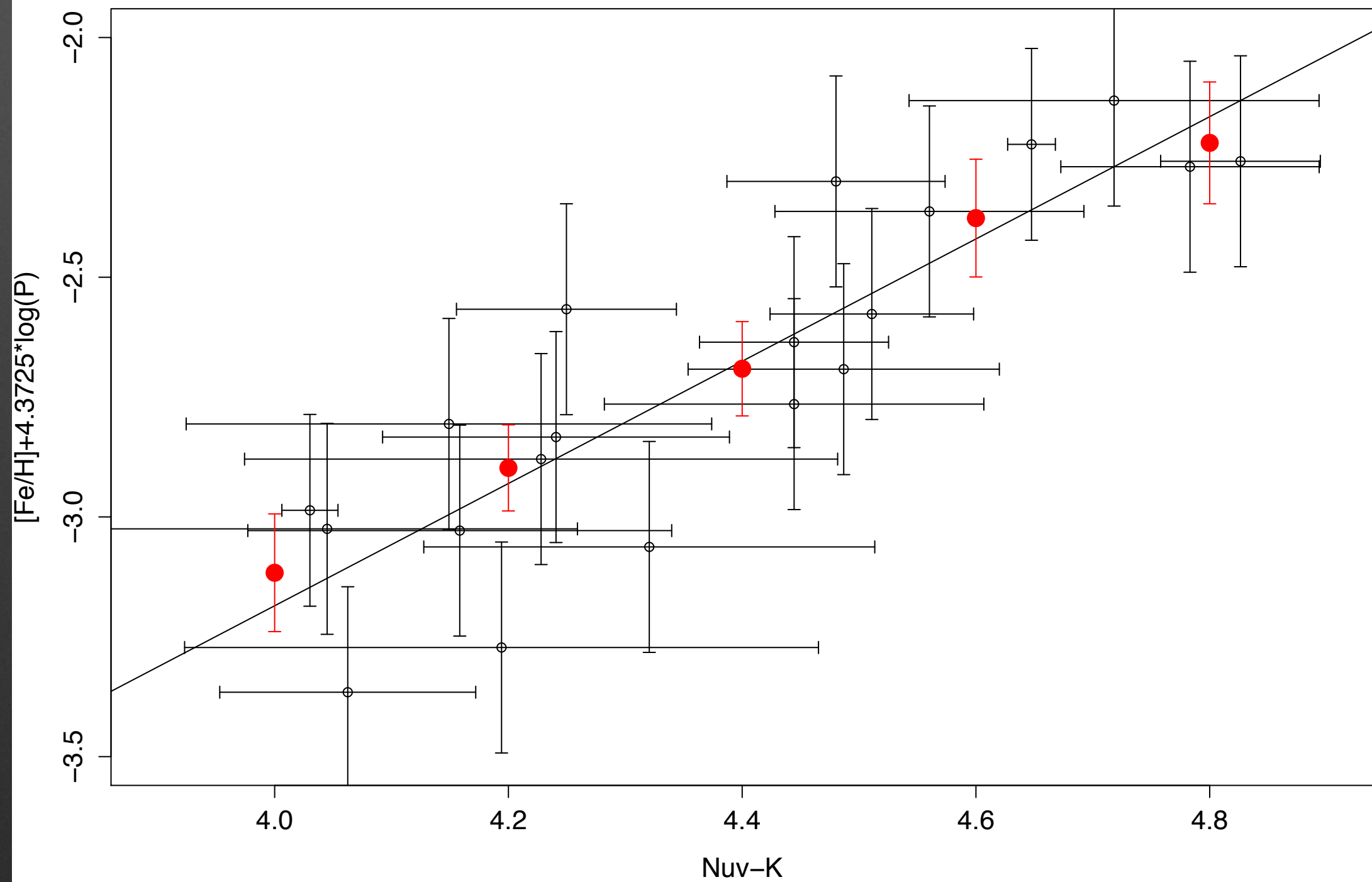
Outline

- Galex data
- Period-Color-[Fe/H] on FUV-NIR
- [Fe/H] on Halo

Galex RR Lyrae - [Fe/H]

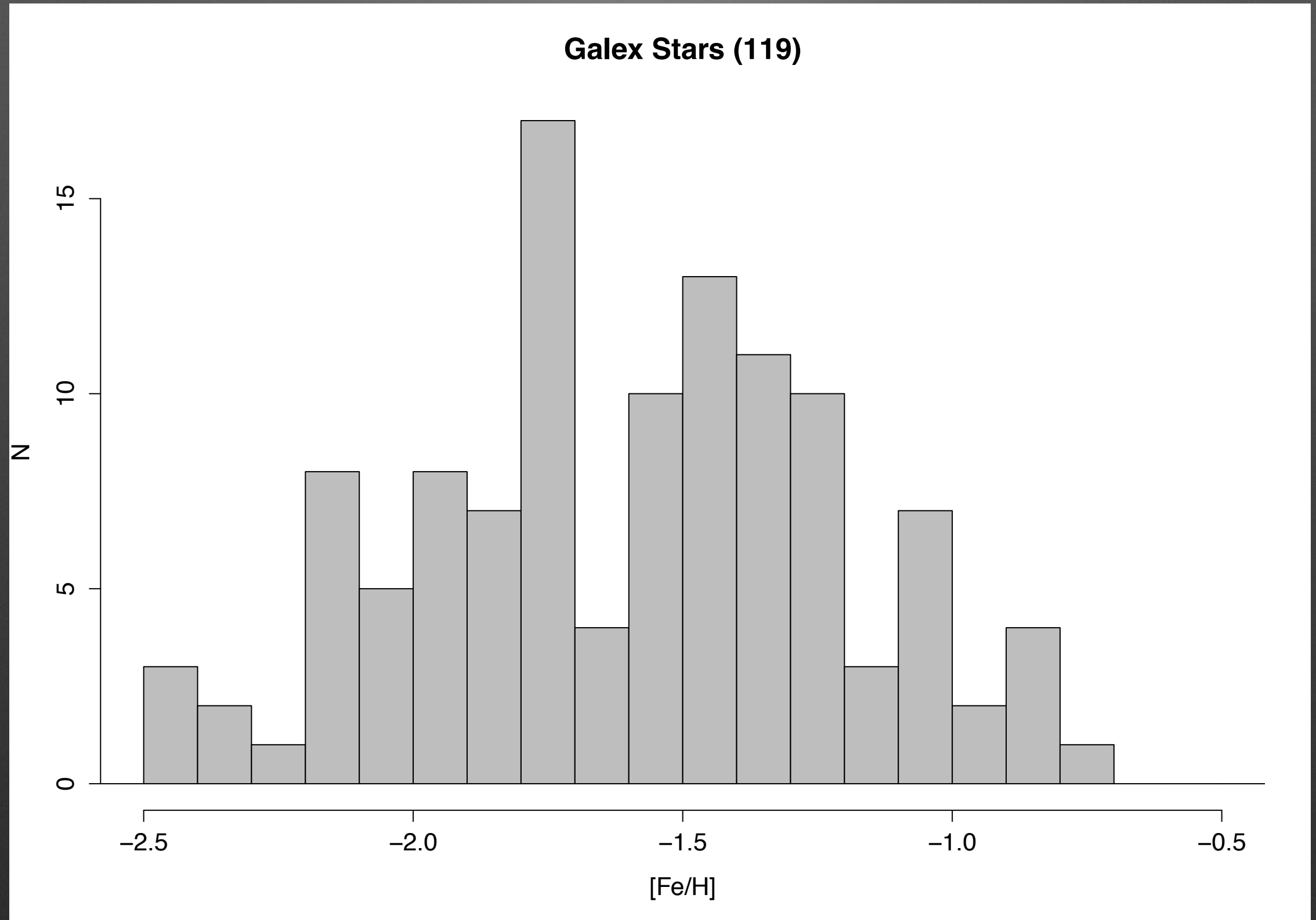
- 28 stars with [Fe/H] from Catalina/SDSS (Drake et al 2013)
10 out of 28 stars without 2MASS K (95% confidence upper limit)
- 2 stars with [Fe/H] from Dambis (Dambis et al 2013)

$$[\text{Fe}/\text{H}] = -8.3 + 1.3 \cdot (\text{Nuv-K}) - 4.4 \cdot \log(P)$$



Galex RR Lyr - [Fe/H]

119 new [Fe/H]

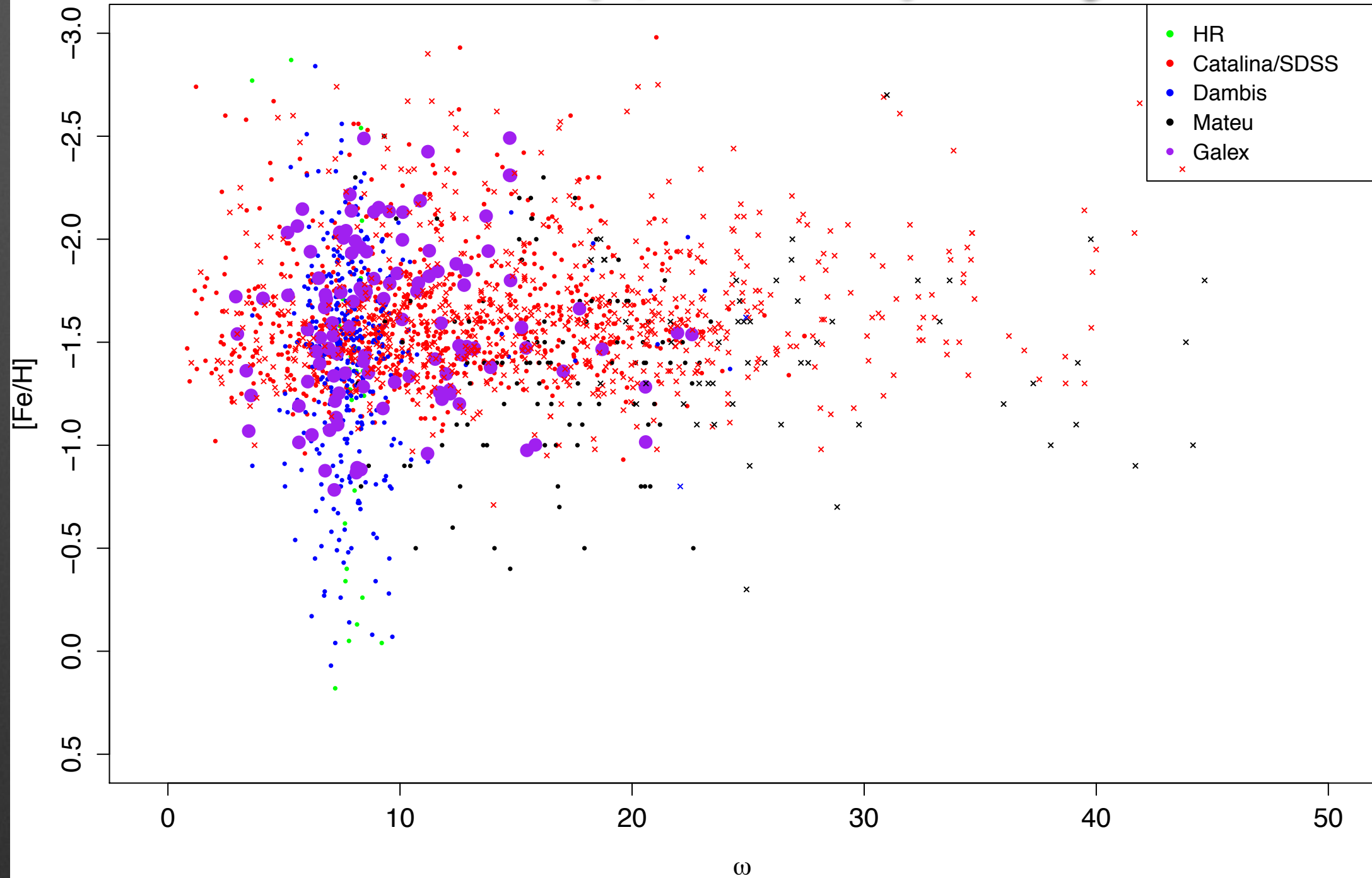


Outline

- Galex data
- Period-Color-[Fe/H] on FUV-NIR
- [Fe/H] on Halo

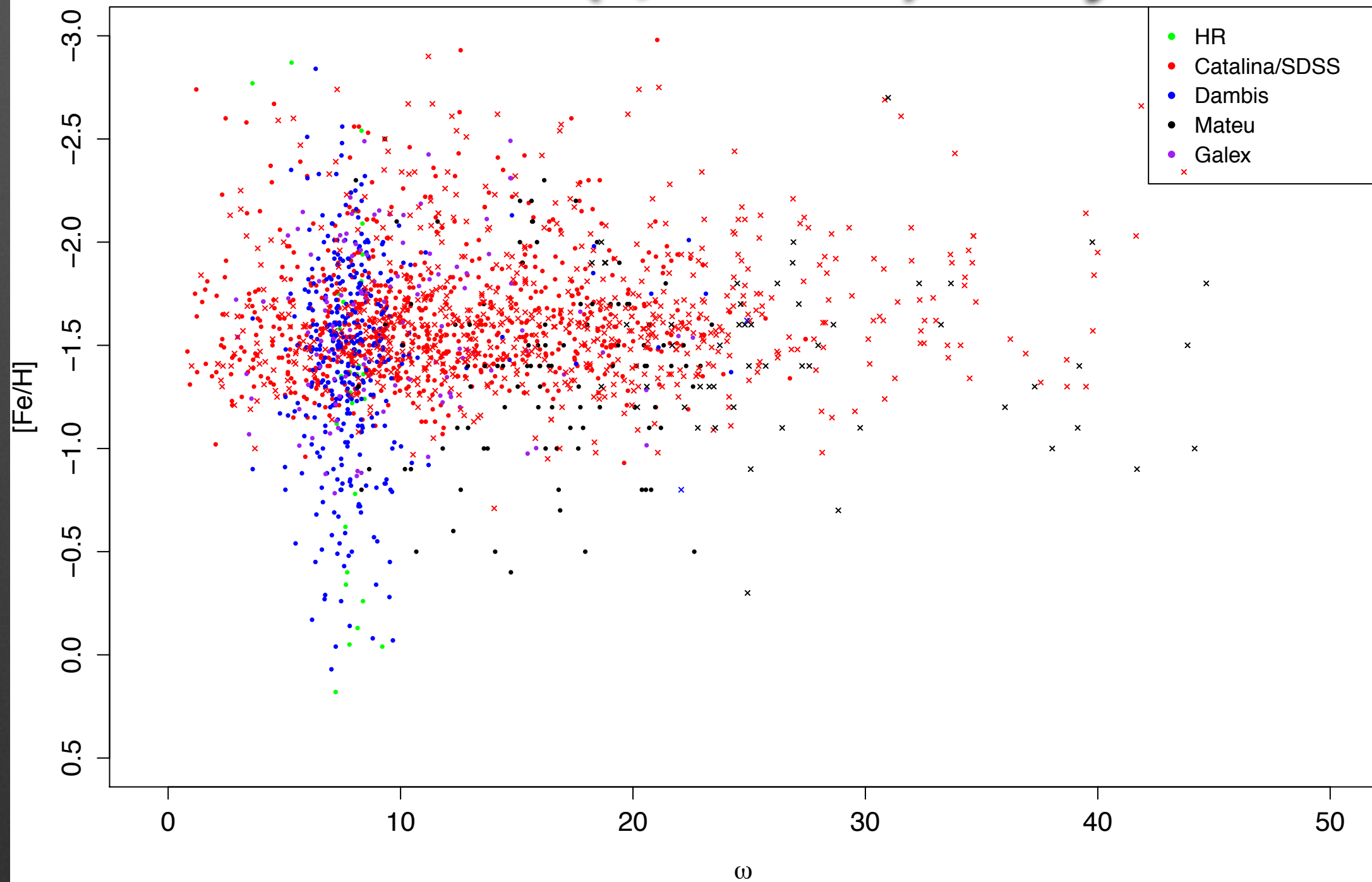
Galactocentric distance (projected onto the galactic plane)

2096 (1,977+119) RR Lyrae



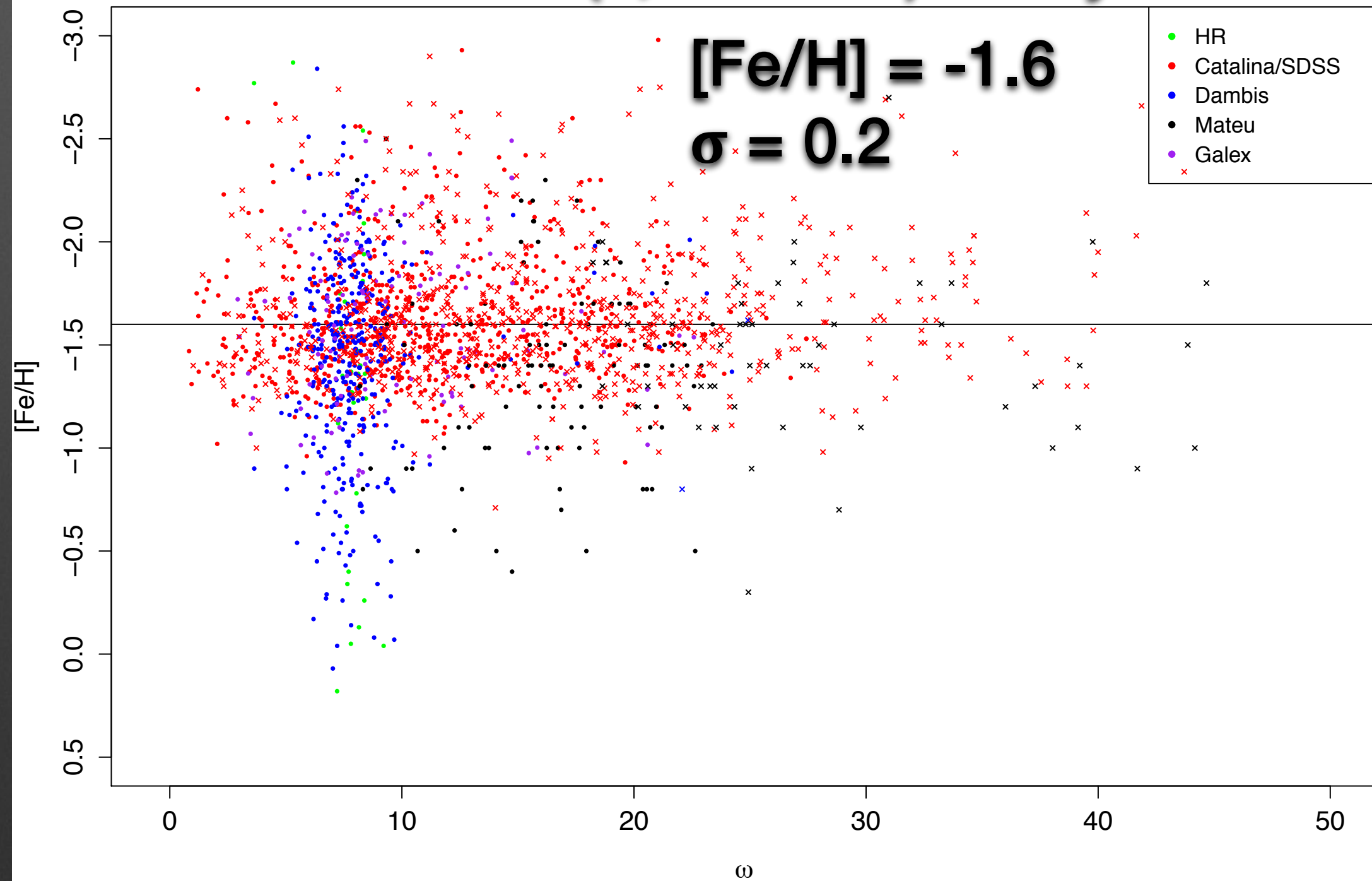
Galactocentric distance (projected onto the galactic plane)

2096 (1,977+119) RR Lyrae

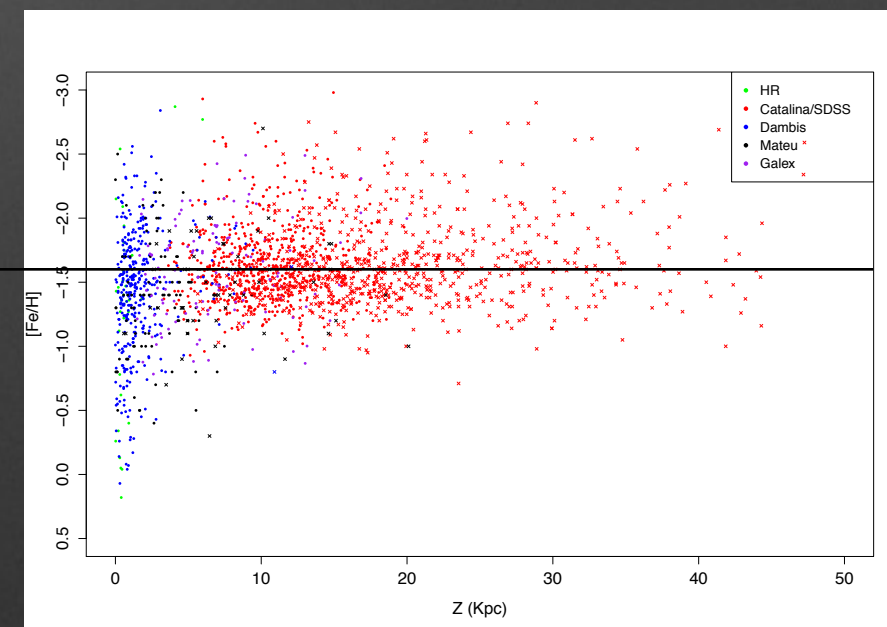
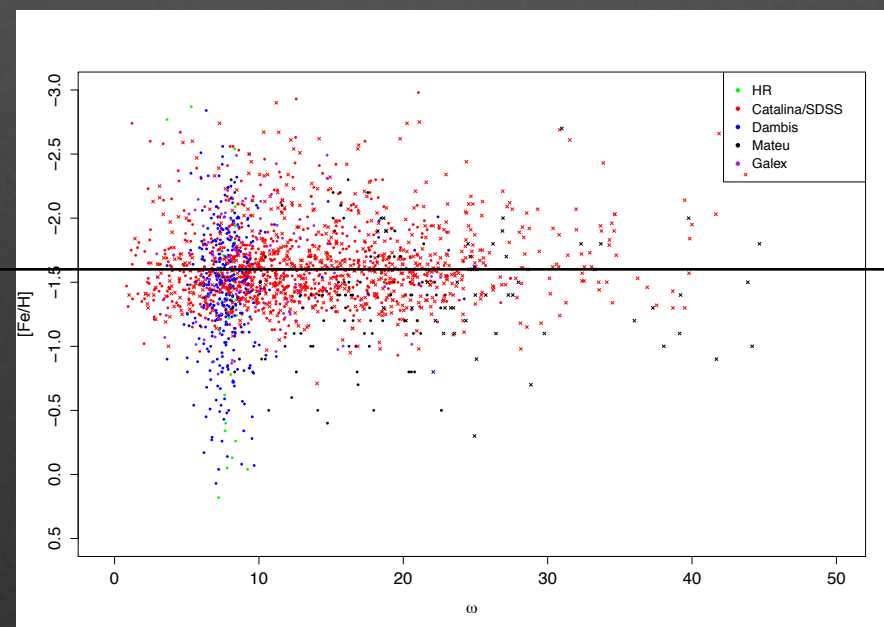
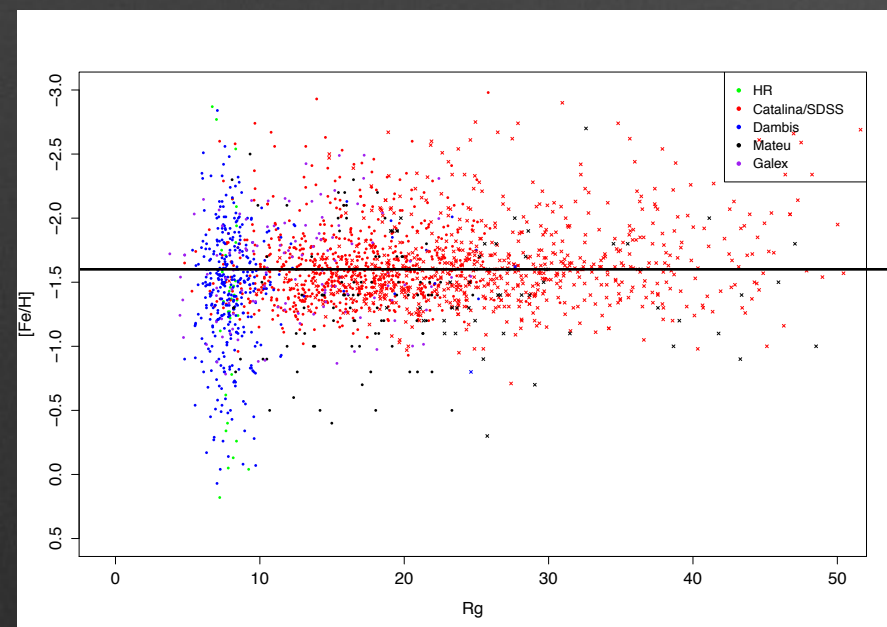
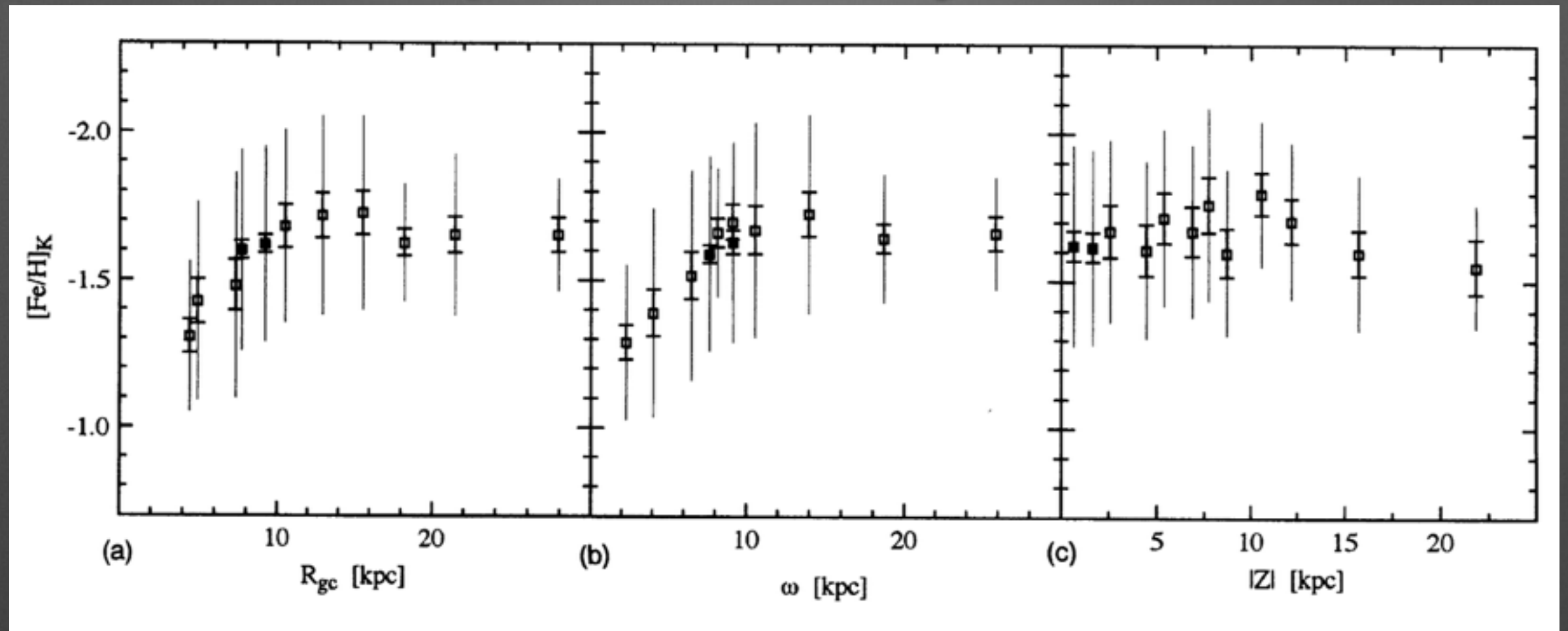


Galactocentric distance (projected onto the galactic plane)

2096 (1,977+119) RR Lyrae



Comparison with Layden 1994



Conclusions

- We increase by a factor of 6 the Layden+Suntzeff sample of RR Lyrae with either spectroscopic or photometric $[Fe/H]$ estimation
- The mean metallicity is relatively constant (-1.6 ± 0.2 supporting Layden+Suntzeff)
- Evidence of a metal poor tail
- No evidence of a metal rich tail but the solar circle
- No evidence of clear separation between inner/outer halo

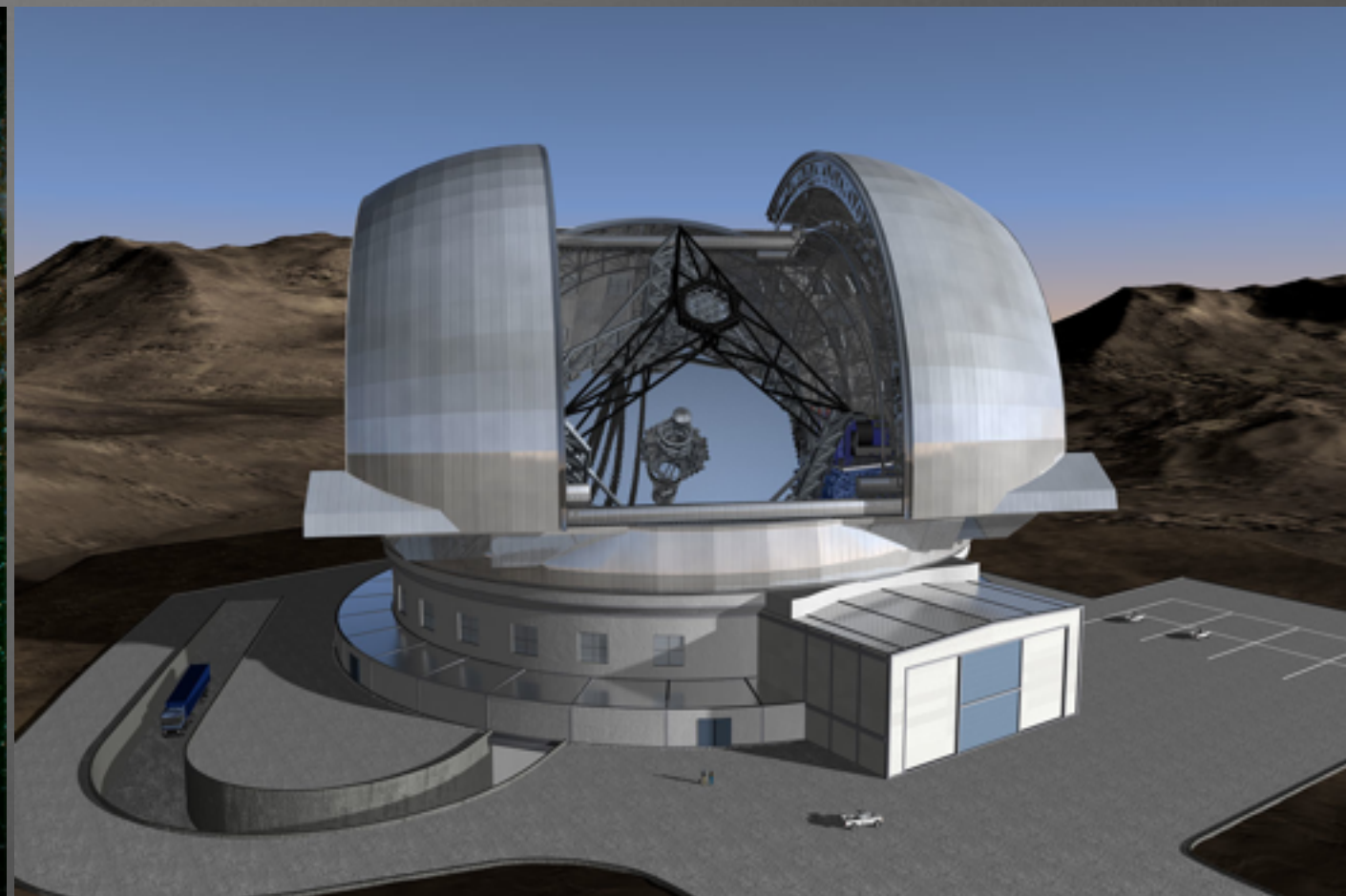
Next steps

- Accurate estimation of NUV and K mean mag
- Complete visual inspection of all Galex candidates
- A new spin of [FE/H] distribution on Halo RR Lyrae

E-ELT + Gaia

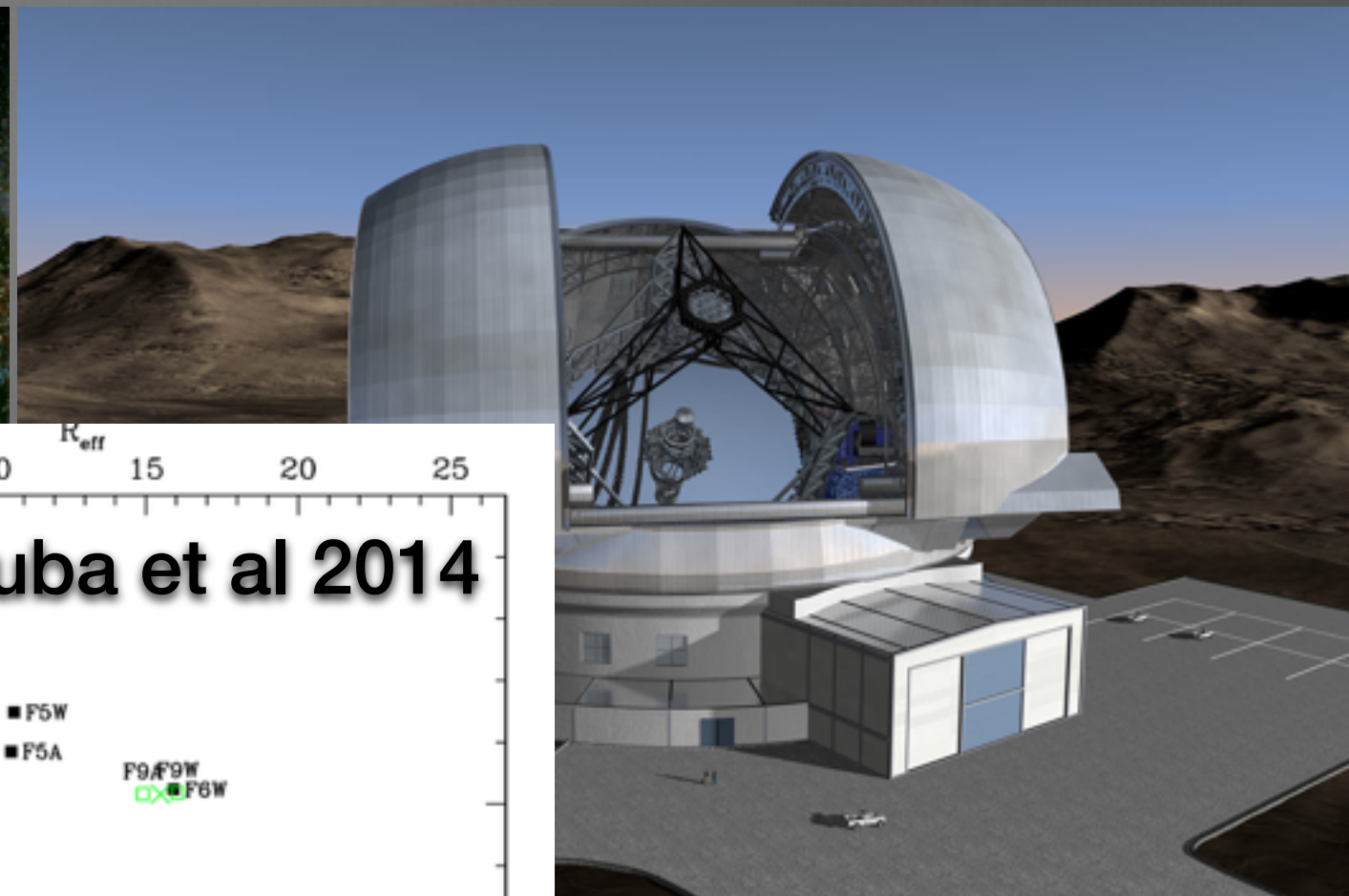
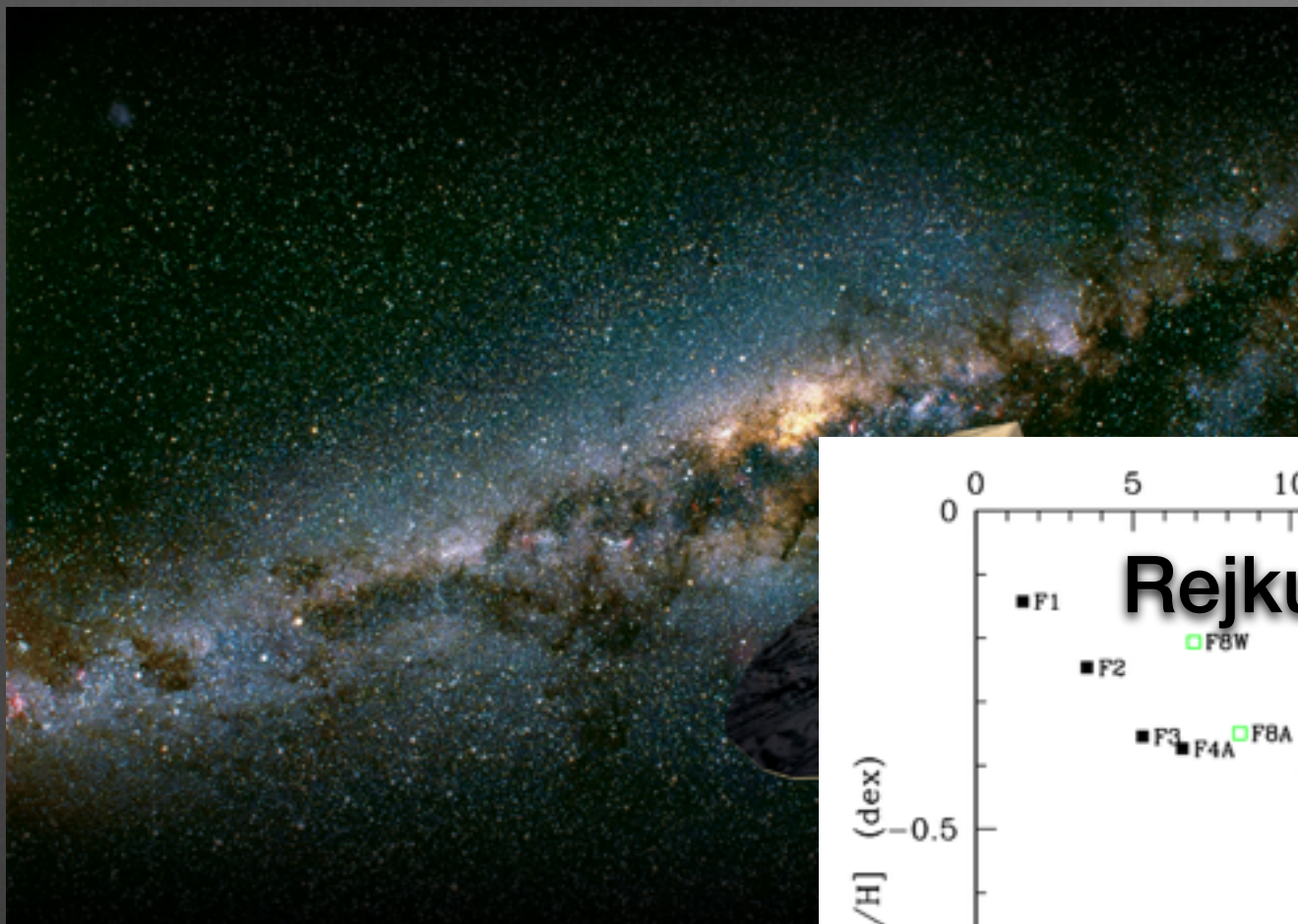


Trigonometric Parallax
+
Gaia

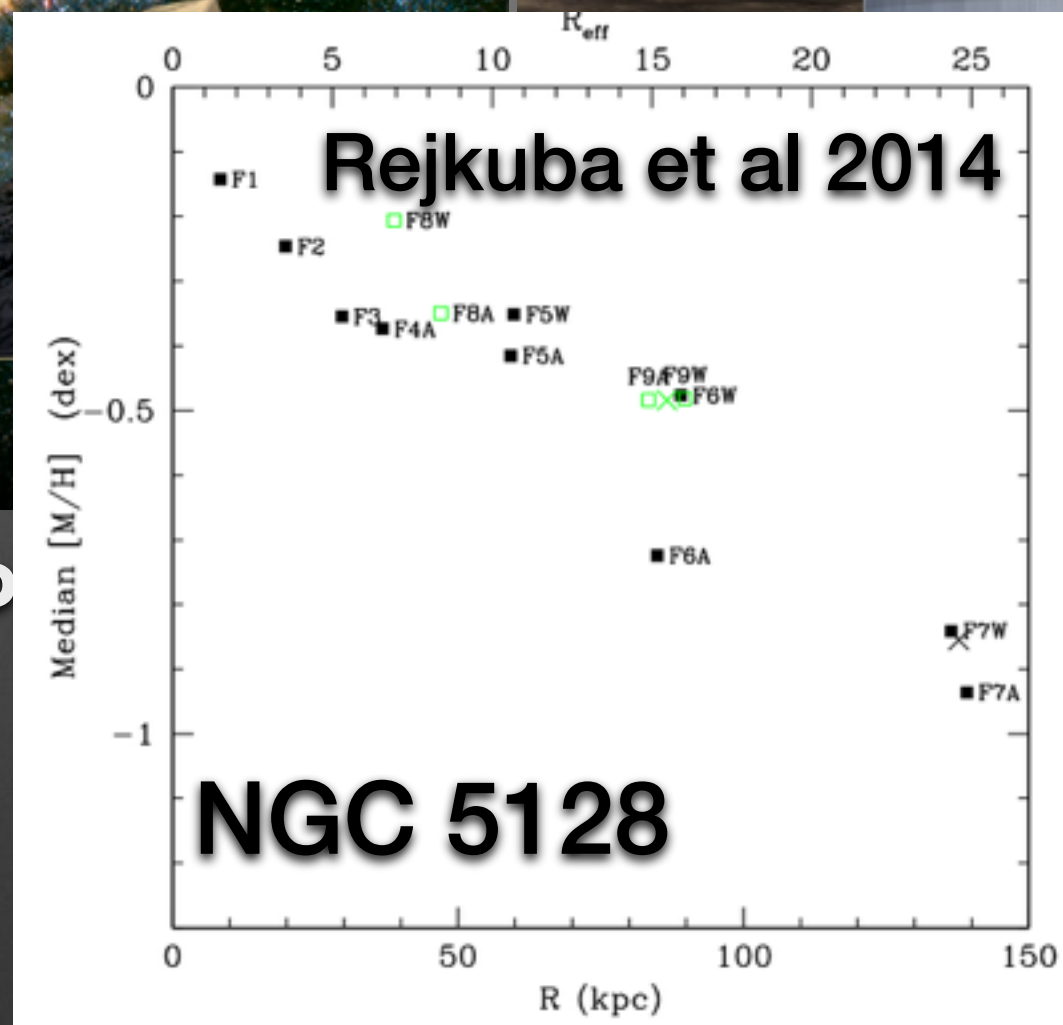


MOSAIC@E-ELT
Virgo Cluster

E-ELT + Gaia



Trigonometric P
+
Gaia



AIC@E-ELT
go Cluster

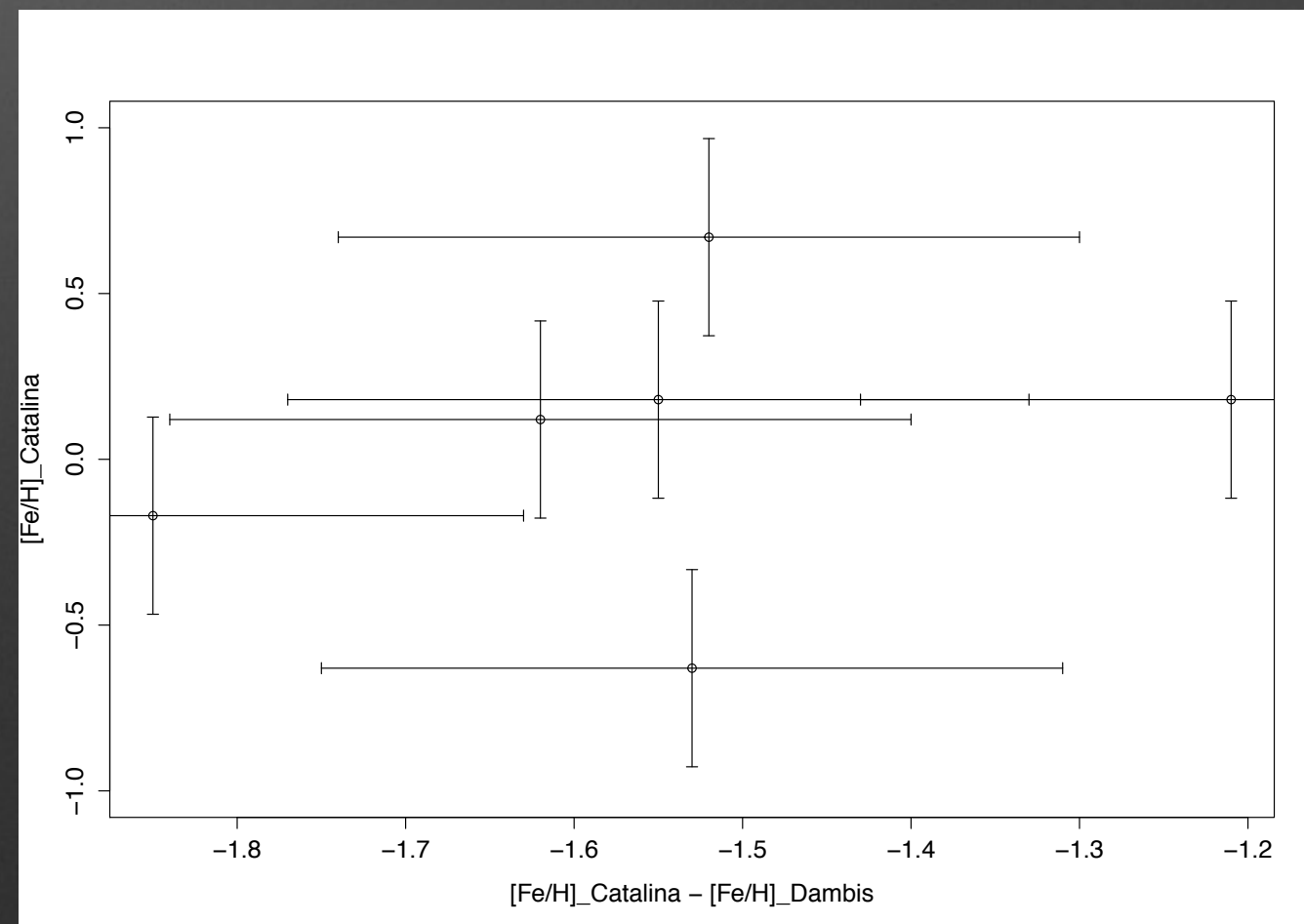
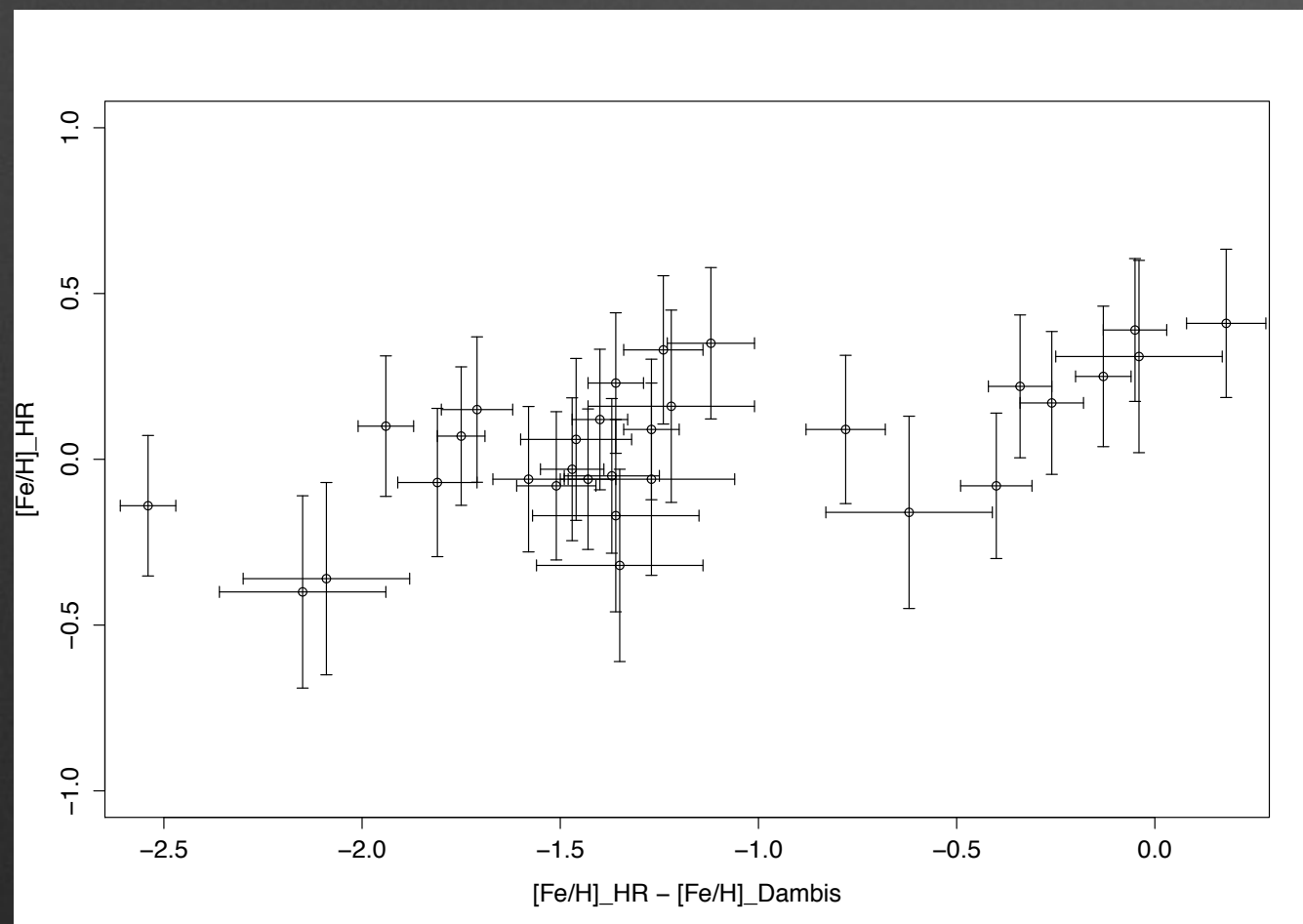
[Fe/H] comparison

- HR-Dambis : 31 match
- Catalina-Dambis 6 match
- Catalina-Mateu 1 match

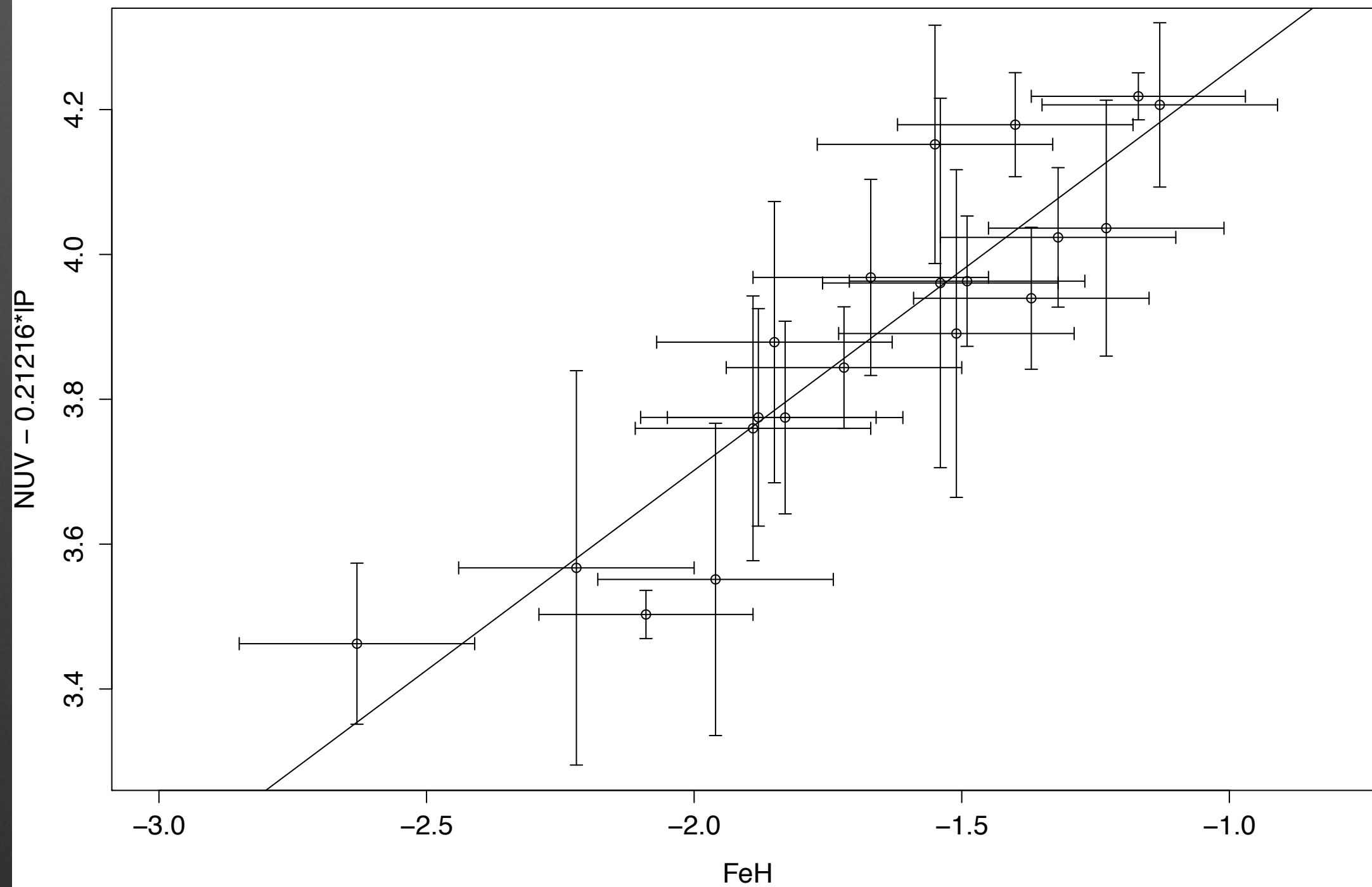
$$d[\text{Fe}/\text{H}] = -0.02 \quad \sigma=0.13$$

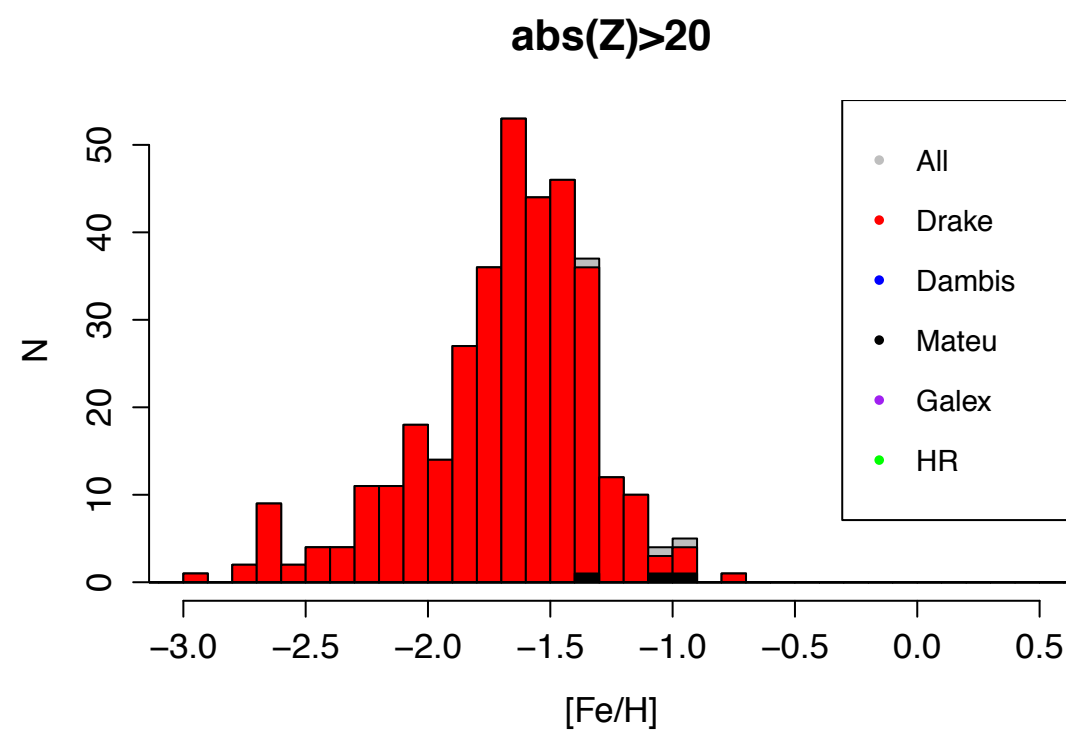
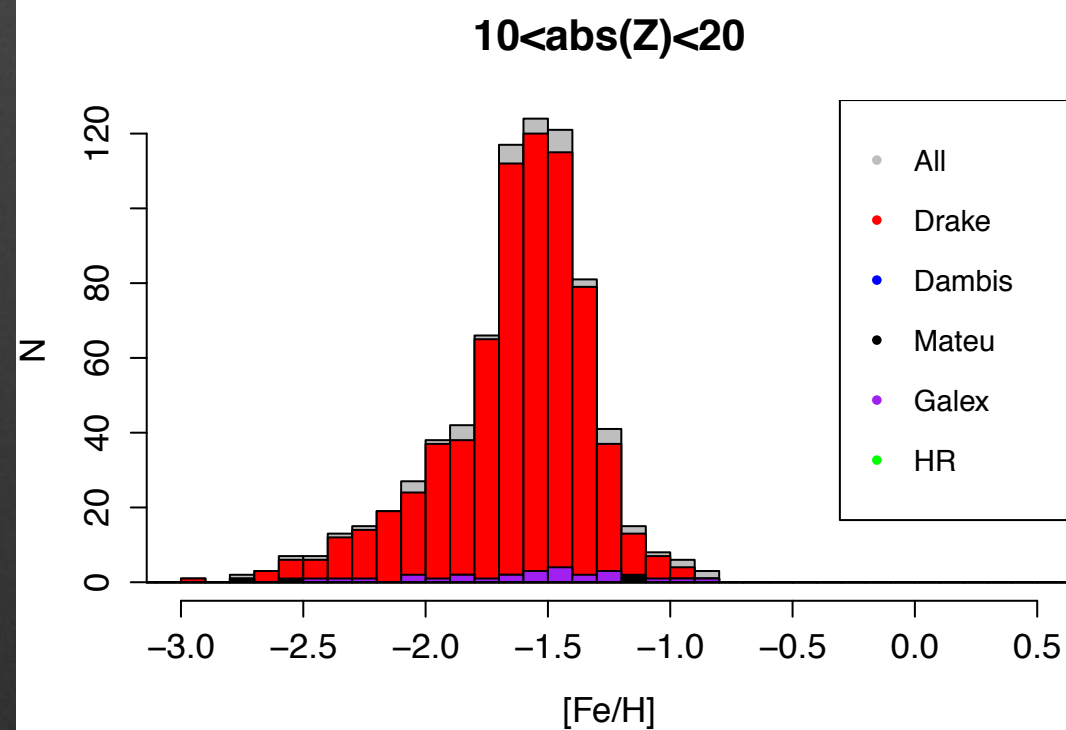
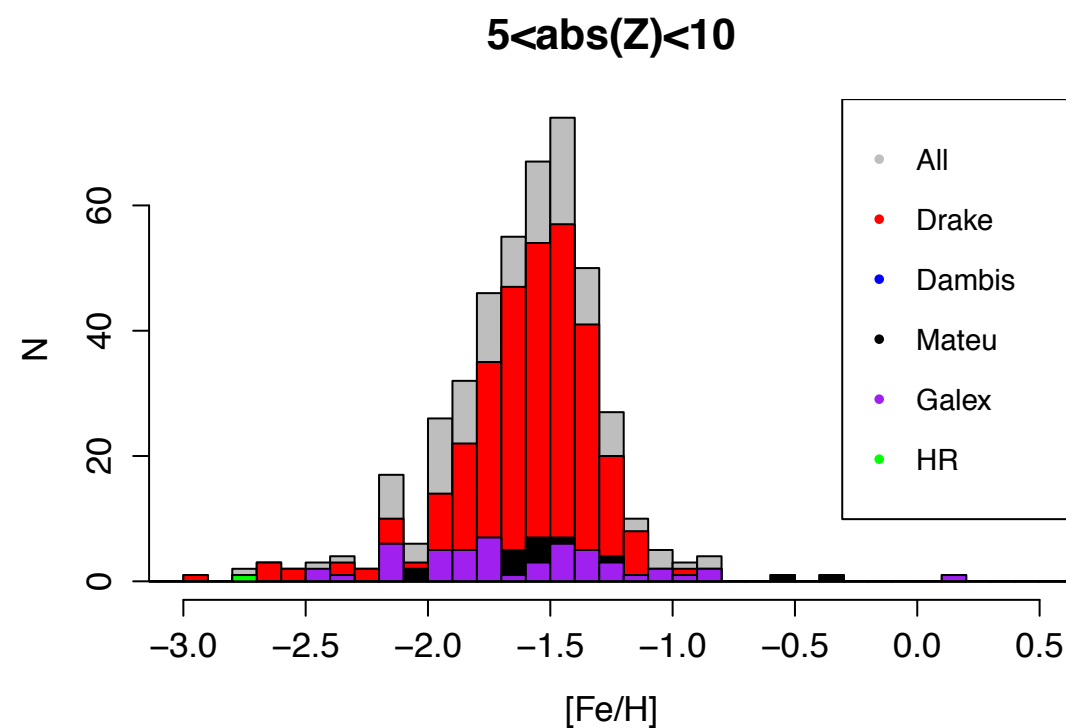
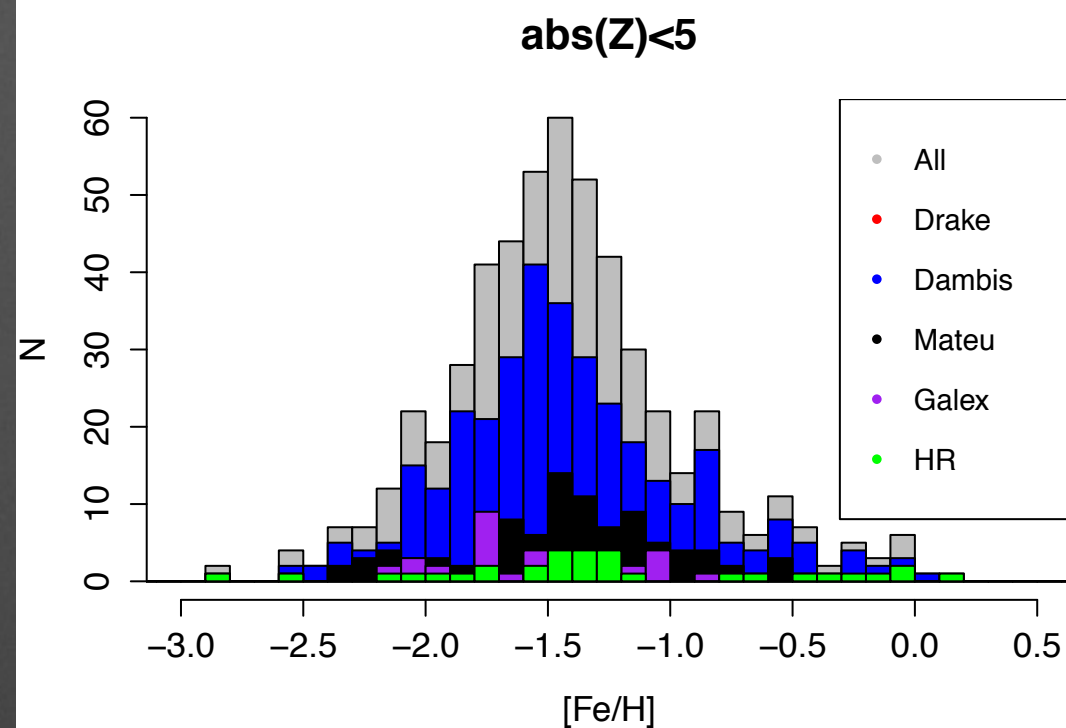
$$d[\text{Fe}/\text{H}] = 0.06 \quad \sigma=0.39$$

$$d[\text{Fe}/\text{H}] = 0.03$$

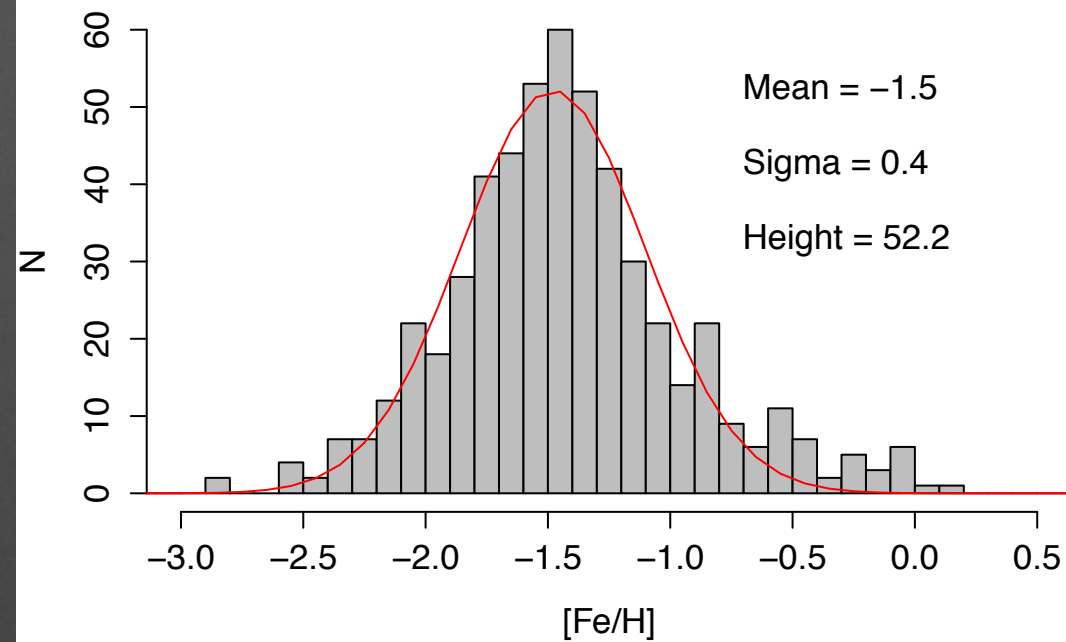


$$\text{NUV} = 0.55212 \cdot [\text{Fe}/\text{H}] + 0.21216 \cdot \log(P) + 4.806$$

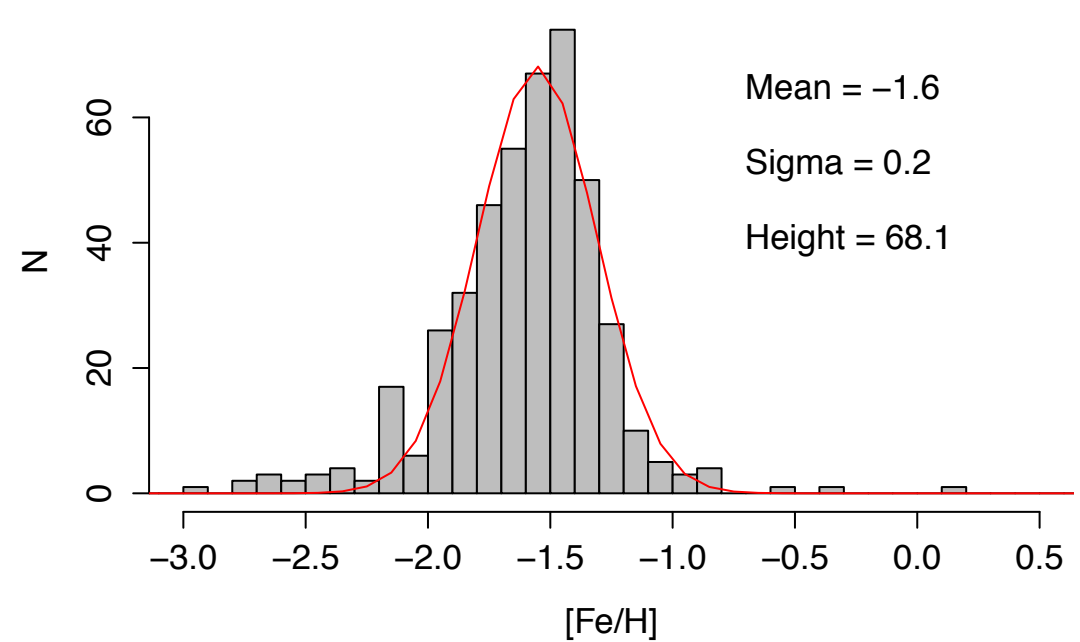




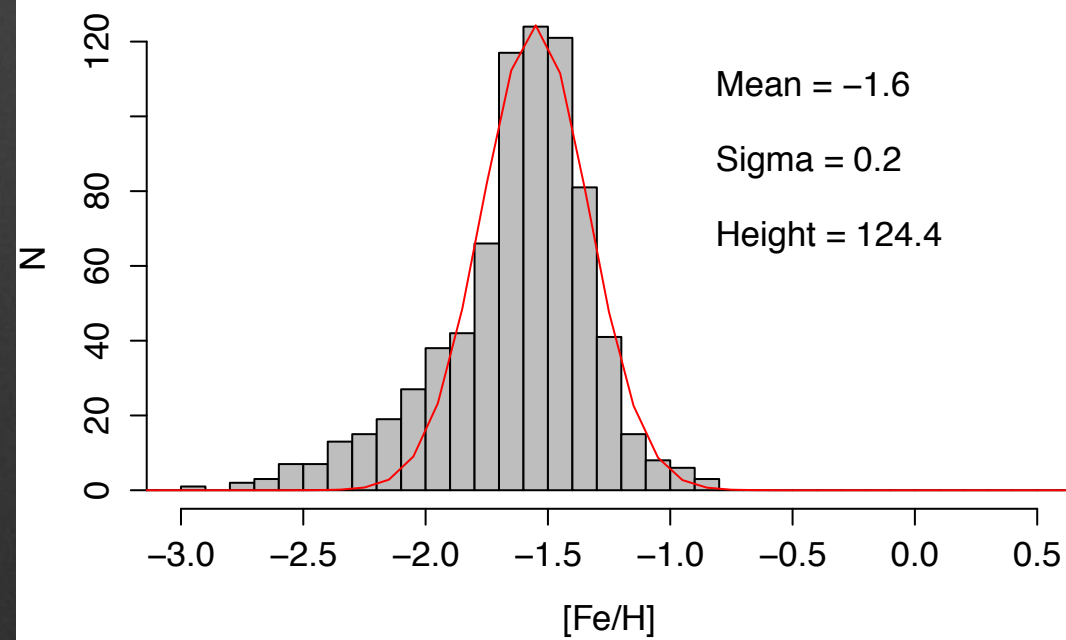
abs(Z)<5



5<abs(Z)<10



10<abs(Z)<20



abs(Z)>20

