

# WINGS survey of local clusters of galaxies

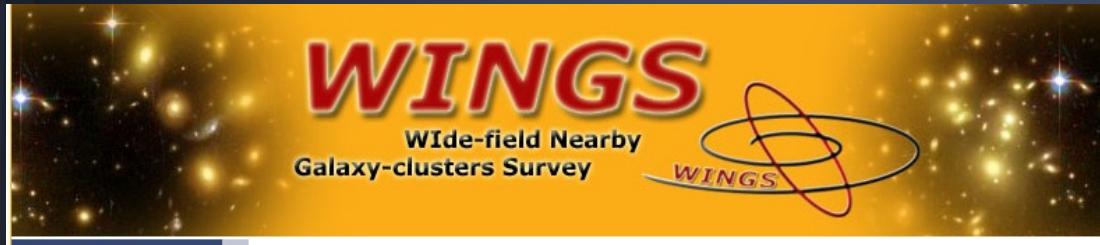
## The database

A. Moretti for the WINGS team  
Padova University/INAF-OaPD

Fasano  
Poggianti  
Bettoni  
D'Onofrio  
Omizzolo  
Gullieuszik  
Valentinuzzi  
Bindoni

Couch  
Dressler  
Moles  
Kjaergaard  
Fritz  
Cava  
Varela

# The survey



**WINGS**  
WIde-field Nearby  
Galaxy-clusters Survey

**Wings Project**

WINGS (WIde-field Nearby Galaxy-cluster Survey: Fasano et al. 2002; Fasano et al. 2006) is an all-sky ( $|b|>20$ ) survey of a complete, X-ray selected sample of galaxy clusters in the redshift range 0.04–0.07. The goal of the WINGS project is the systematic study of the local cosmic variance of the cluster population and of the properties of cluster galaxies as a function of cluster properties and local environment. This data collection allows the definition of a local, 'zero-point' reference against which to gauge the cosmic evolution when compared to more distant clusters. The core of the WINGS project is the optical (B,V) imaging survey. It provides photometric data for huge samples of galaxies (~550,000) and stars (~190,000) in the inner field ( $34''\times34''$ ) of 77 nearby galaxy clusters, as well as structural and morphological information for a subsample (~50,000) of relatively bright galaxies. Spectroscopic information for ~6,000 galaxies in 48 WINGS clusters is provided by a follow-up multi-fiber, medium-resolution survey, while additional photometric information comes from follow-up imaging surveys in the NIR (J,K; 28 WINGS clusters) and U-band (18 WINGS clusters). When compared to other literature sky surveys it is clear that WINGS is the only one providing complete and homogeneous data for a huge sample of galaxies in the field of nearby galaxy clusters.

Abell Richness

Redshift

N

Bautz–Morgan class

$\log L_{X}(0.1-2.4 \text{ keV})$

Home Page

Cluster Sample

Optical

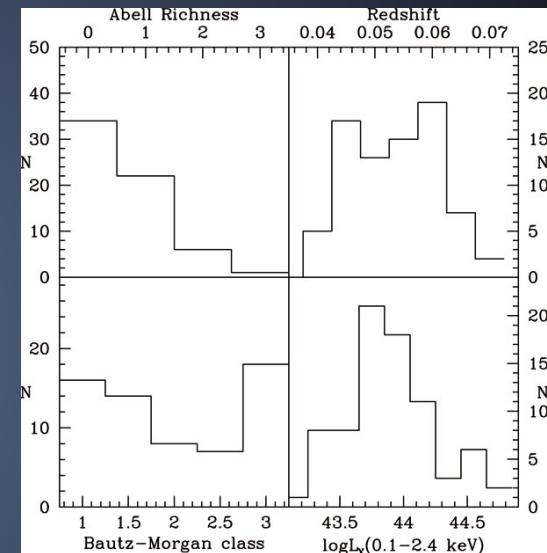
Near Infrared

Spectroscopy

People

Get Data

Contact



76 clusters  
(36 north and 42 south)

$0.04 < z < 0.07$

Fasano et al., 2006

# WINGS data (raw)

## Photometric

B,V wide field (34'x 34')  
(INT, WFI) [WINGS-OPT, 76  
clusters]

J, K wide field photometry  
(UKIRT) [WINGS-NIR, 28  
clusters]

U wide field photometry  
(BOK, INT, LBT) [WINGS-U, 9  
& 11 clusters]

V,B,u' wide wide field (VST)  
[OMEGA-WINGS, 59 clusters]

## Spectroscopic

WYFFOS (north) & 2dF  
(south) spectra for ~6000  
objects [WINGS-SPE, 48  
clusters]

# WINGS science

## B & V photometry

- Fasano et al., 2006  
Ramella et al., 2007  
Fritz et al., 2007  
D'Onofrio et al., 2008  
Cava et al., 2009  
Varela et al., 2009  
Valentinuzzi et al., 2009  
Poggianti et al., 2009  
Fasano et al., 2010
- Valentinuzzi et al., 2010  
Fritz et al., 2011  
Vulcani et al., 2011a  
Vulcani et al., 2011b  
D'Onofrio et al., 2011  
Bettoni et al., 2011  
Valentinuzzi et al., 2011  
Fasano et al., 2012  
Vulcani et al., 2012

V band used for detection/photometry

Big haloes separately analyzed

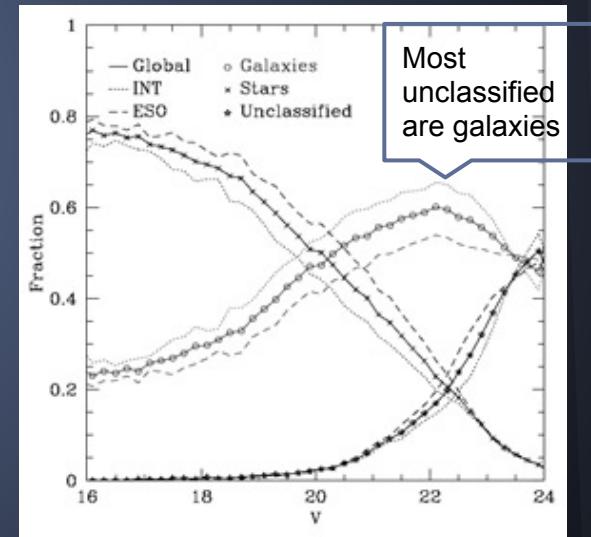
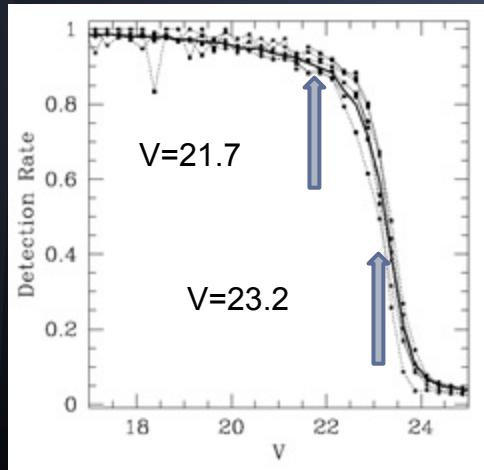
Detection limit at  $m_V \sim 25.7$  mag/arcsec<sup>2</sup> (0.6 deeper than SDSS)

Rejection of detections in only 1 band

Classification using sextractor stellarity

(Gx  $\leq 0.2$ ; ~400000)

Simulations on each field to compute completeness

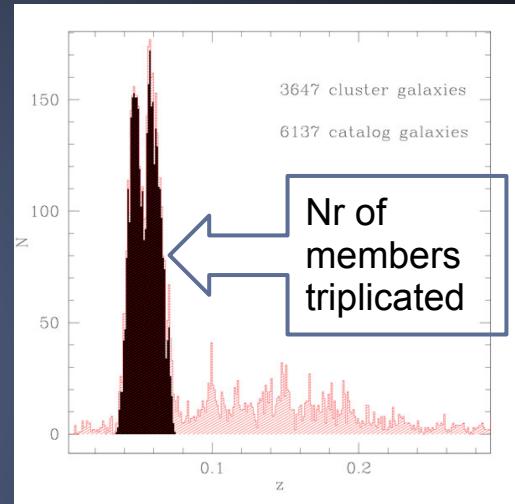
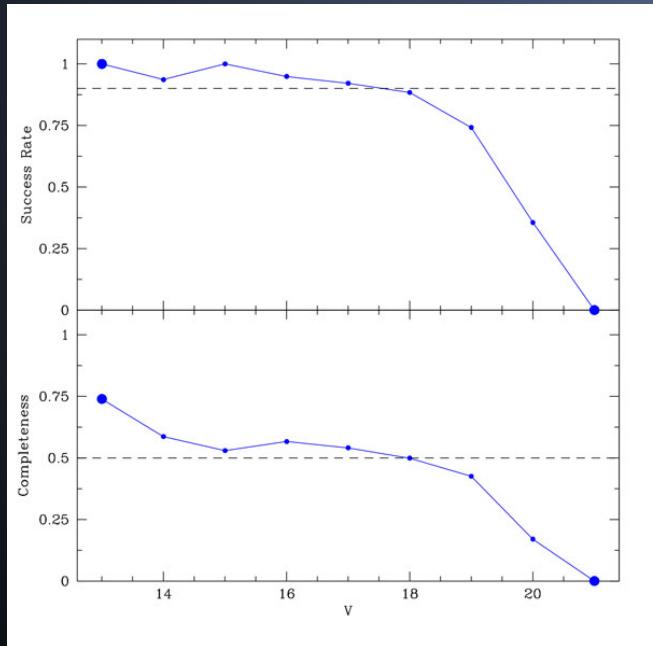


# WINGS science Spectroscopy

48 clusters (26 N, 22 S) - ~6000 galaxies  
(30% of which already in literature)

WYFFOS@WHT 1.6" 3800-7000 Å  
2dF@AAT 2" 3600-8000 Å

- Fasano et al., 2006
- Ramella et al., 2007
- Fritz et al., 2007
- D'Onofrio et al., 2008
- Cava et al., 2009
- Varela et al., 2009
- Valentinuzzi et al., 2009
- Poggianti et al., 2009
- Fasano et al., 2010
- Valentinuzzi et al., 2010
- Fritz et al., 2011
- Vulcani et al., 2011a
- Vulcani et al., 2011b
- D'Onofrio et al., 2011
- Bettoni et al., 2011
- Valentinuzzi et al., 2011
- Fasano et al., 2012
- Vulcani et al., 2012



Nr of successful redshift measurement /  
number of spectra  
Nr of spectra / Nr of photometric objects with  
same cuts

SPECTROSCOPIC completeness (to be used  
when studying magnitude limited properties  
(21 clusters have completeness > 50%)

# WINGS science

## J & K photometry

- Fasano et al., 2006  
Ramella et al., 2007  
Fritz et al., 2007  
D'Onofrio et al., 2008  
Cava et al., 2009  
Varela et al., 2009  
**Valentinuzzi et al., 2009**  
Poggianti et al., 2009  
Fasano et al., 2010
- Valentinuzzi et al., 2010  
Fritz et al., 2011  
Vulcani et al., 2011a  
Vulcani et al., 2011b  
D'Onofrio et al., 2011  
Bettoni et al., 2011  
Valentinuzzi et al., 2011  
Fasano et al., 2012  
Vulcani et al., 2012

28 clusters (with good sampling of cluster properties)

pixel resolution ~ 0.2 Kpc

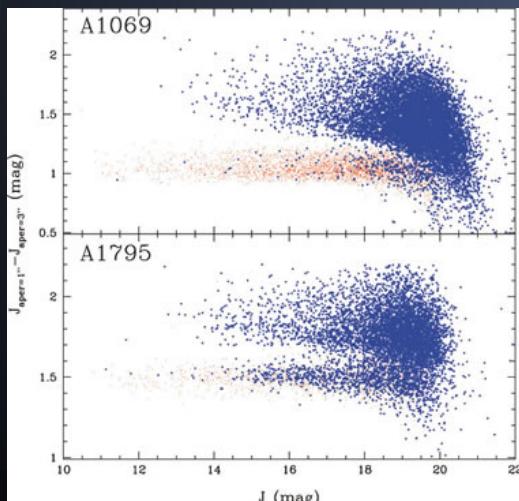
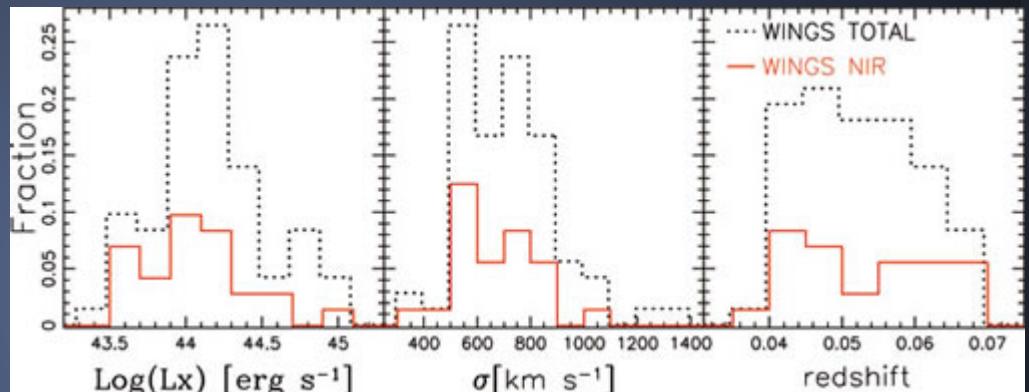
~750000 extended sources

90% detection completeness @

J=20.5, K=19.4

90% classification compl. @

J=19.5, K=18.5



Interactive cleaning of galaxy sample (to remove stars)

# WINGS science

## SED fitting

SED fit on ~5300 galaxies  
(12 ages, 3 metallicities)  
70% with chi-sq <3.0

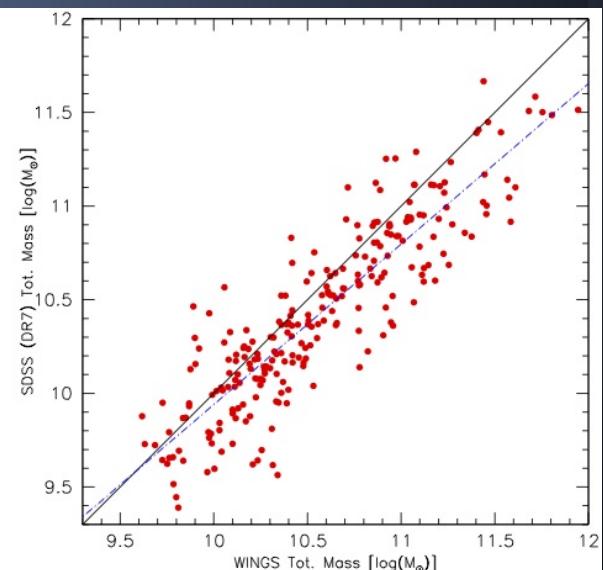
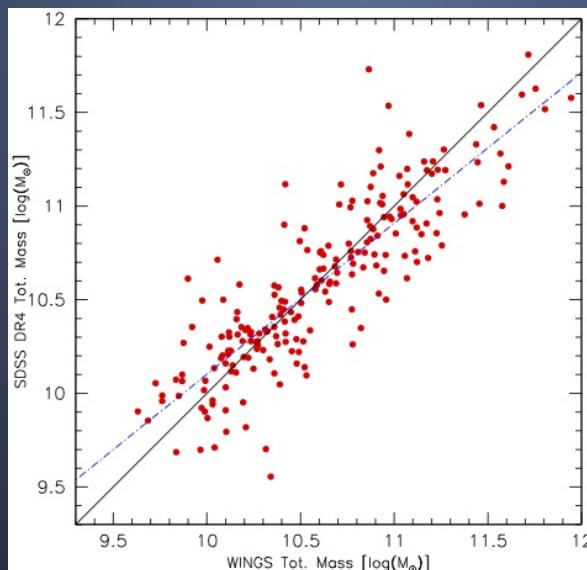
->Stellar masses

->Star Formation  
Histories

->Ages (MW & LW)

Fasano et al., 2006  
Ramella et al., 2007  
Fritz et al., 2007  
D'Onofrio et al., 2008  
Cava et al., 2009  
Varela et al., 2009  
Valentinuzzi et al., 2009  
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Vulcani et al., 2001b  
D'Onofrio et al., 2011  
Bettoni et al., 2011  
Valentinuzzi et al., 2011  
Fasano et al., 2012  
Vulcani et al., 2012



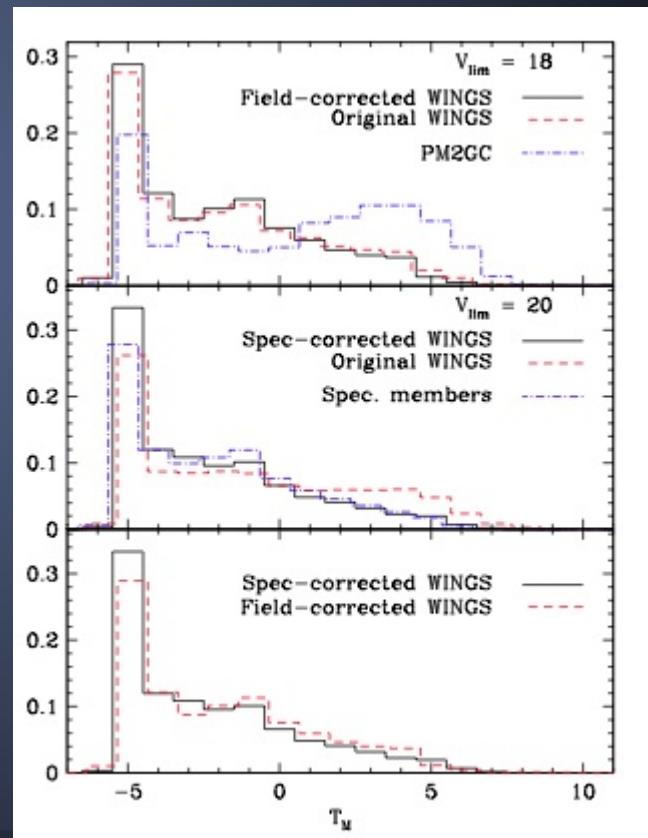
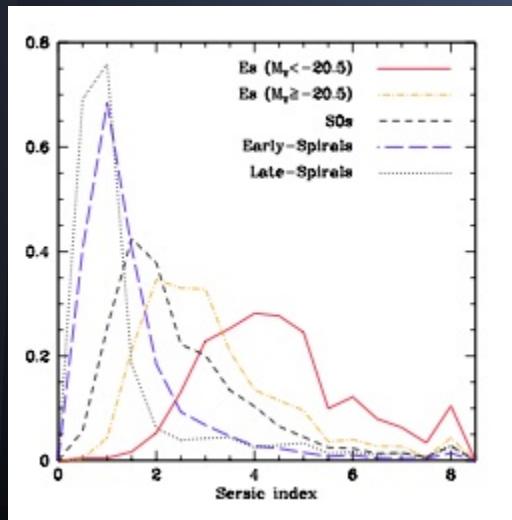
# WINGS science

- Fasano et al., 2006  
Ramella et al., 2007  
Fritz et al., 2007  
D'Onofrio et al., 2008  
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Varela et al., 2009  
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Poggianti et al., 2009  
Fasano et al., 2010
- Valentinuzzi et al., 2010  
Fritz et al., 2011  
Vulcani et al., 2011a  
Vulcani et al., 2011b  
D'Onofrio et al., 2011  
Bettoni et al., 2011  
Valentinuzzi et al., 2011  
Fasano et al., 2012  
Vulcani et al., 2012

MORPHOT: automatic classification of morphological types based on  
21 diagnostics & ML + NN technique

Validation using ~ 1000 visually class. galaxies  
[particularly efficient in separating E and S0,  
unlike other automatic tools]

-> M type for ~ 40000 WINGS galaxies



# WINGS WiP

Moretti et al. [LF]

Moretti et al. [Database]

Gullieuszik et al. [OmegaWINGS]

Omizzolo et al. [U,B]

Fasano et al. [Morph/Density]

Hanssen et al. [Lick indices]

Marziani et al. [AGN]

Bindoni et al. [Surf. photom.]

# OmegaWINGS

Moretti et al. [LF]  
Moretti et al. [Database]  
Gullieuszik et al. [OmegaWINGS]  
Omizzolo et al. [U,B]  
Fasano et al. [Morph/Density]  
Hanssen et al. [Lick indices]  
Marziani et al. [AGN]  
Bindoni et al. [Surf. photom.]

## Motivations:

study outskirts of clusters and the infalling regions, where clusters accrete new galaxies.  
morphologies, sizes, structural parameter in different vs. distance from the center

## Data:

**u', B- and V- band OmegaCAM imaging for all WINGS clusters  
observable from Paranal (59/76)**

2 observing programs in GTO time (PI Poggianti [BV] and D'Onofrio [U]). 60+60 hrs

Observations started 1 year ago (ESO P88) [18 clusters already observed]

The two programs will be likely completed in ~1year (end of ESO P92)

## Data-reduction:

ESO/MVM or "Alambic" + simple Python scripts.

Completed automated pipeline. From raw telescope frames to stacked mosaics

# OmegaWINGS

Moretti et al. [LF]  
Moretti et al. [Database]  
Gullieuszik et al. [OmegaWINGS]  
Omizzolo et al. [U,B]  
Fasano et al. [Morph/Density]  
Hanssen et al. [Lick indices]  
Marziani et al. [AGN]  
Bindoni et al. [Surf. photom.]

## Photometric calibration:

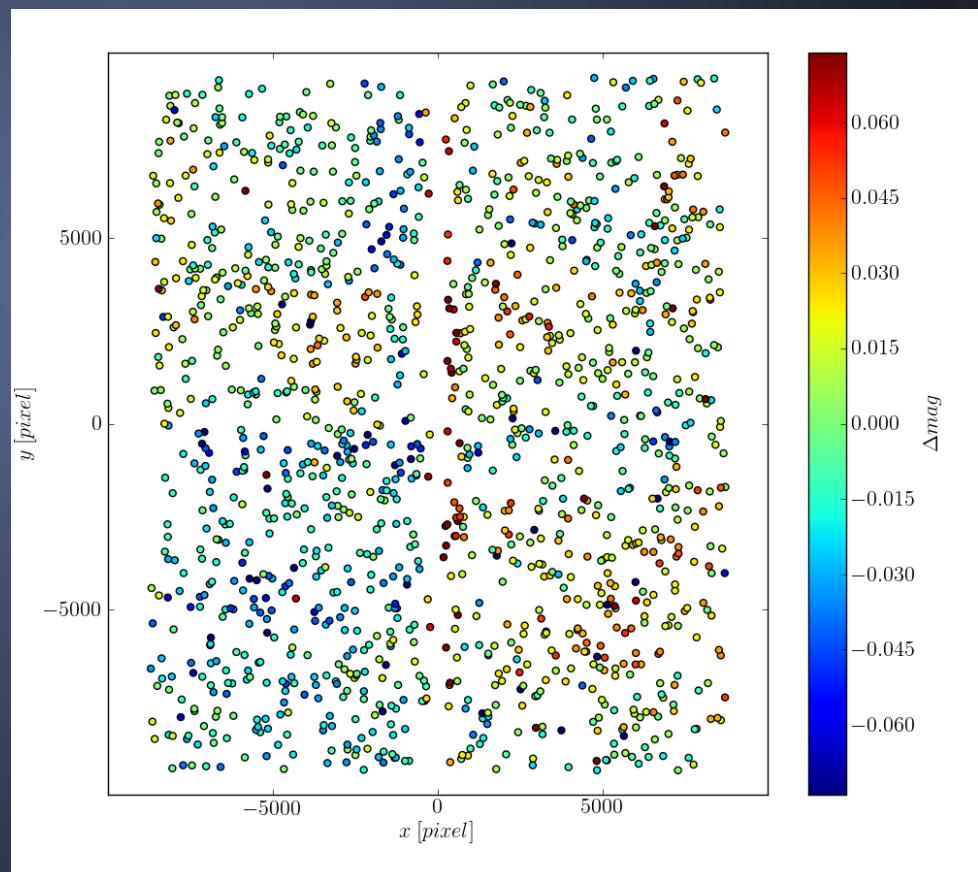
**CCDs gain variation:** calculated from background value

**Illumination correction:** based on dithered observations of SA107  
[up to 0.2 mag]

**ZP and Colour terms:**  
by comparison with previous WINGS photometry

## Astrometric calibration:

SDSS DR8 (when available) or 2MASS



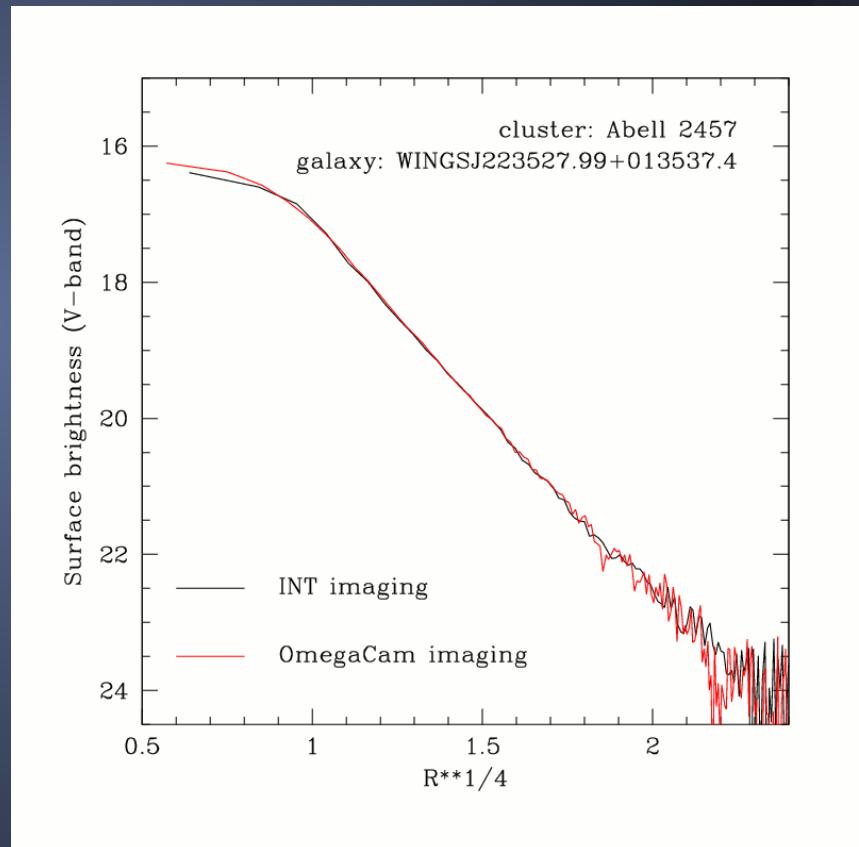
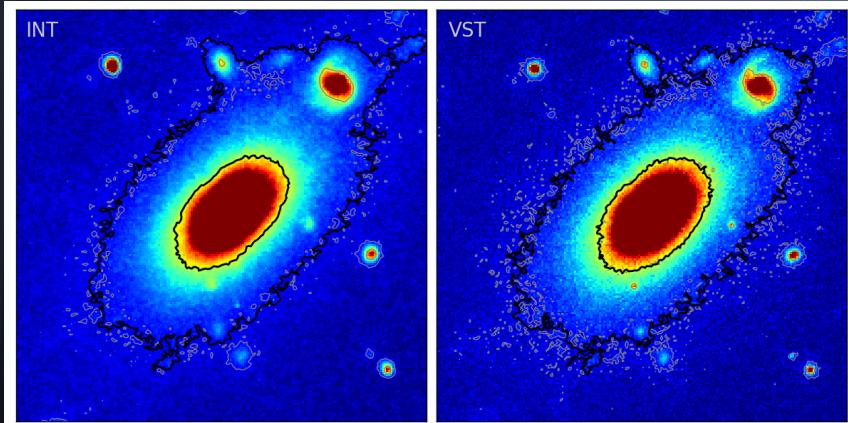
SDSS DR8 -> Johnson (Jordi et al., 2005) [Entire field]

# OmegaWINGS

Moretti et al. [LF]  
Moretti et al. [Database]  
Gullieuszik et al. [OmegaWINGS]  
Omizzolo et al. [U,B]  
Fasano et al. [Morph/Density]  
Hanssen et al. [Lick indices]  
Marziani et al. [AGN]  
Bindoni et al. [Surf. photom.]

## Sky subtraction

2step combination of  
dithered images  
+ source masking



# WINGS data (catalogs)

- █ CDS
- █ VO [NEW]
- █ In publ.

Optical photometry (~759000) and surface photometry (~42000)

NIR photometry (~960000 [k] ~630000[j]) and surface photometry (~72000 [k])

SFH (~5300)

Morphologies (~40000)

EW (~4400)

Local densities (~66000)

Lick indices (~4500)

# The VO tools and WINGS

thanks to IA2 team in Trieste

34'x34' B, V images for 76 clusters available (Aladin, TOPCAT)

Mosaic

Fully calibrated (astro+photo)

Ready to download and use

Surface photometry (V, K)  
catalogs

Equivalent widths catalogs

Local densities catalogs  
[Unpublished yet]



# The VO tools and WINGS: how to

SIA query

Registry: euro-  
vo

Keyword:  
WINGS

SIA parameters:  
Optional

Send table to  
Aladin

Simple Image Access (SIA) Query

Available SIA Services

Registry: <http://registry.euro-vo.org/services/RegistrySearch>

Keywords: wings

Match Fields:  Short Name  Title  Subjects  ID  Publisher  Description

Accept Resource Lists

Short Name	Title	Subjects	Identifier	Publisher	Contact
WINGSOptima	WINGS Optical wide--field images	imaging survey optical	ivo://ia2.inaf.it/hosted/wings/siap/opt	IA2	Alessia Moretti <alessia.mor...

AccessURL Description Version

SIA Parameters

SIA URL:

Object Name: A1831

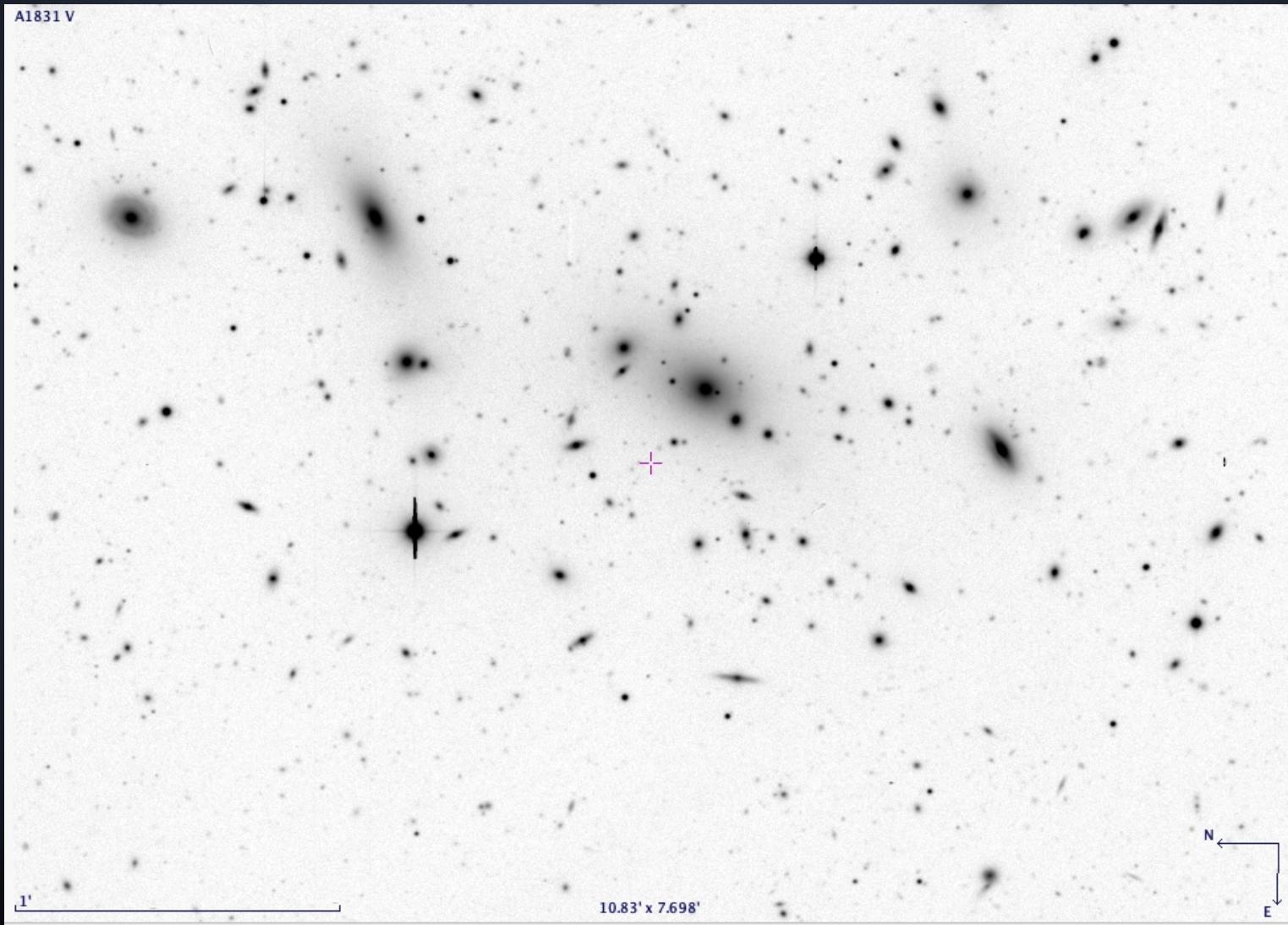
RA: 209.7925 degrees (J2000)  Accept Sky Positions

Dec: 27.99111 degrees (J2000)

Angular Size: 1 degrees

Image Format: image/fits

# The VO tools and WINGS: how to



# The VO tools and WINGS: how to

Cone search

Registry: euro-vo

Keyword:  
WINGS

cone  
parameters:  
Optional

Send table to  
Aladin

Screenshot of the Cone Search interface showing the results for a query with Registry: <http://registry.euro-vo.org/services/RegistrySearch>, Keywords: wings, and Match Fields: Short Name, Title, Subjects, ID, Publisher.

The results table shows the following data:

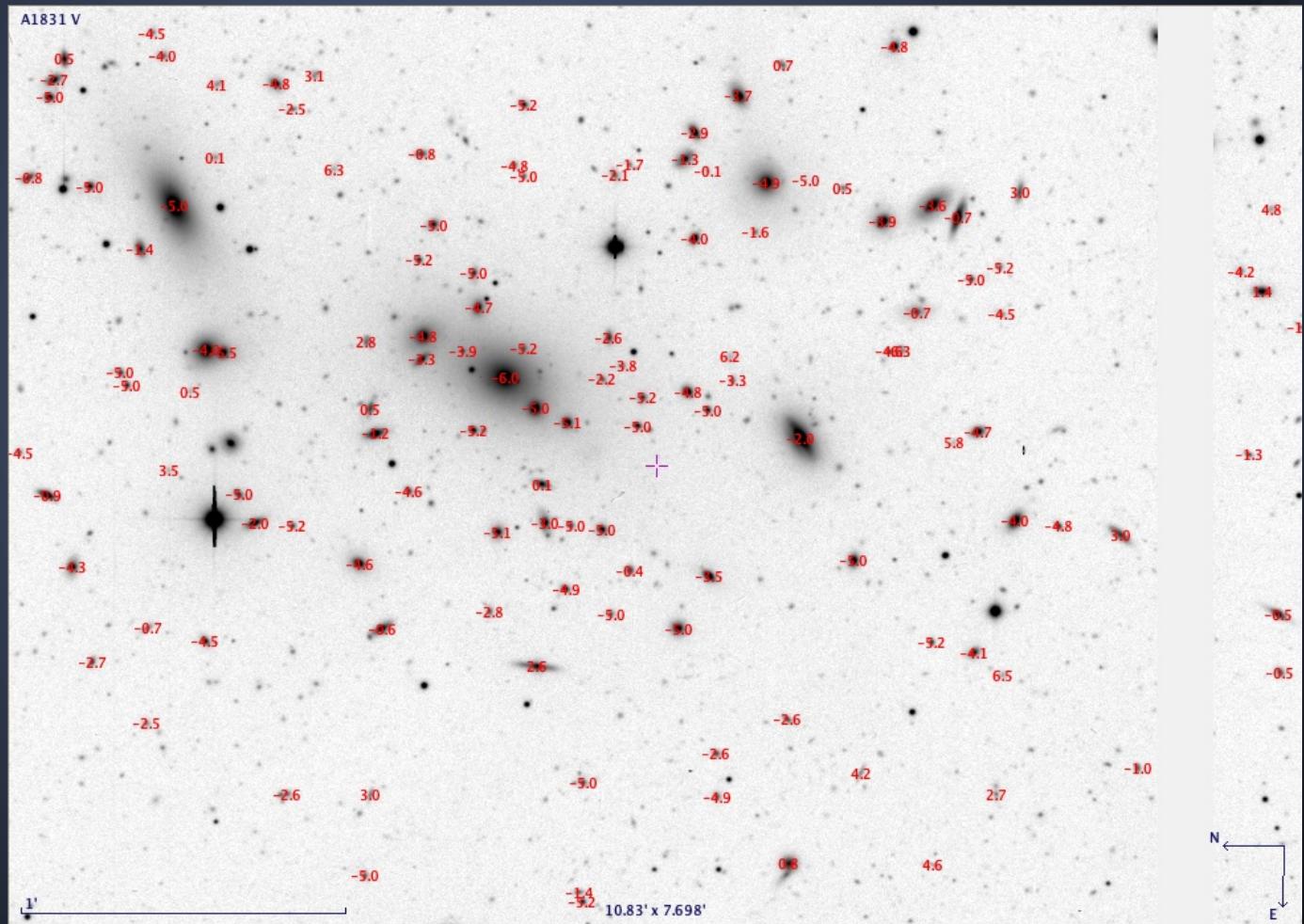
Short Name	Title	Subjects	Identifier
J/A+A/470/39	Substructures in WINGS clusters (Ramella+, 2007)	Clusters_of_galaxies	ivo
J/A+A/495/707	WINGS spectroscopy of 48 galaxy clusters (Cava+, 2009)	Clusters_of_galaxies, Velocities, Redshifts	ivo
J/A+A/497/667	WINGS: Deep optical phot. of 77 nearby clusters (Varela+, 2009)	Clusters_of_galaxies, Photometry, Galaxies	ivo
J/A+A/501/851	WINGS JK photometry of 28 galaxy clusters (Valentinuzzi+, 2009)	Clusters_of_galaxies, Photometry, Galaxies	ivo
J/A+A/526/A45	WINGS-SPE II catalog (Fritz+, 2011)	Clusters_of_galaxies, Photometry, Galaxies, Photometry:wide-band	ivo
J/ApJS/173/85	HI 21cm forbidden-velocity wings (Kang+, 2007)	Interstellar_Medium	ivo
J/MNRAS/420/926	Morphology of galaxies in WINGS clusters (Fasano+, 2012)	Clusters_of_galaxies, Galaxies	ivo
WINGSEqwidth	WINGS equivalent widths	optical lines equivalent width catalog	ivo
WINGSGasphotk	WINGS K surface photometry	surface brightness galaxies catalog	ivo
WINGSGasphotv	WINGS V surface photometry	surface brightness galaxies catalog	ivo
WINGSLocDens19.5	WINGS local densities (Mv<-19.5)	galaxies local density catalog	ivo

The bottom section shows the Cone Parameters:

- Cone URL: [Input field]
- Object Name: [Input field]
- RA: [Input field] degrees (J2000)  Accept Sky Positions
- Dec: [Input field] degrees (J2000)
- Radius: [Input field] degrees

Buttons: OK

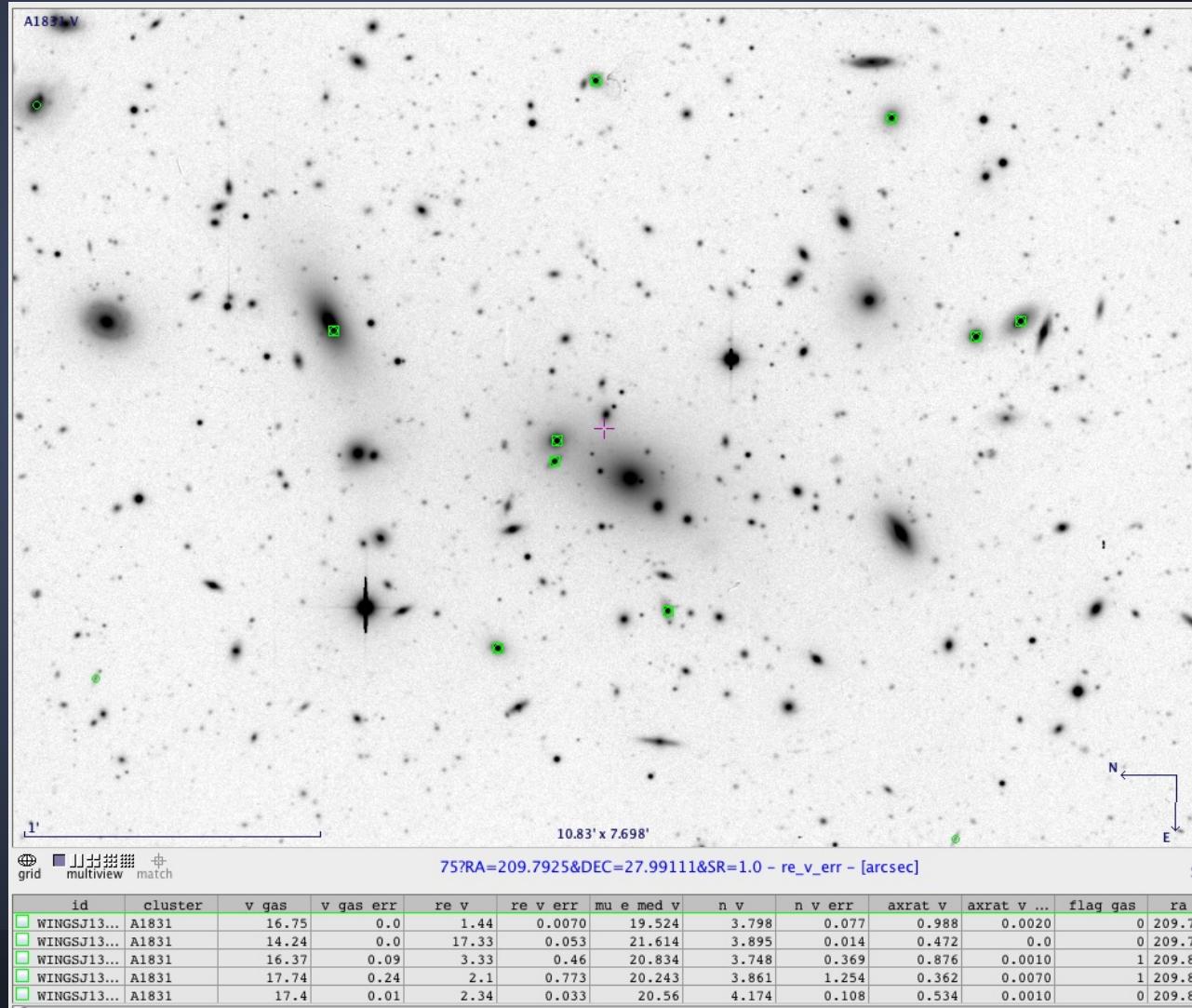
# The VO tools and WINGS: how to



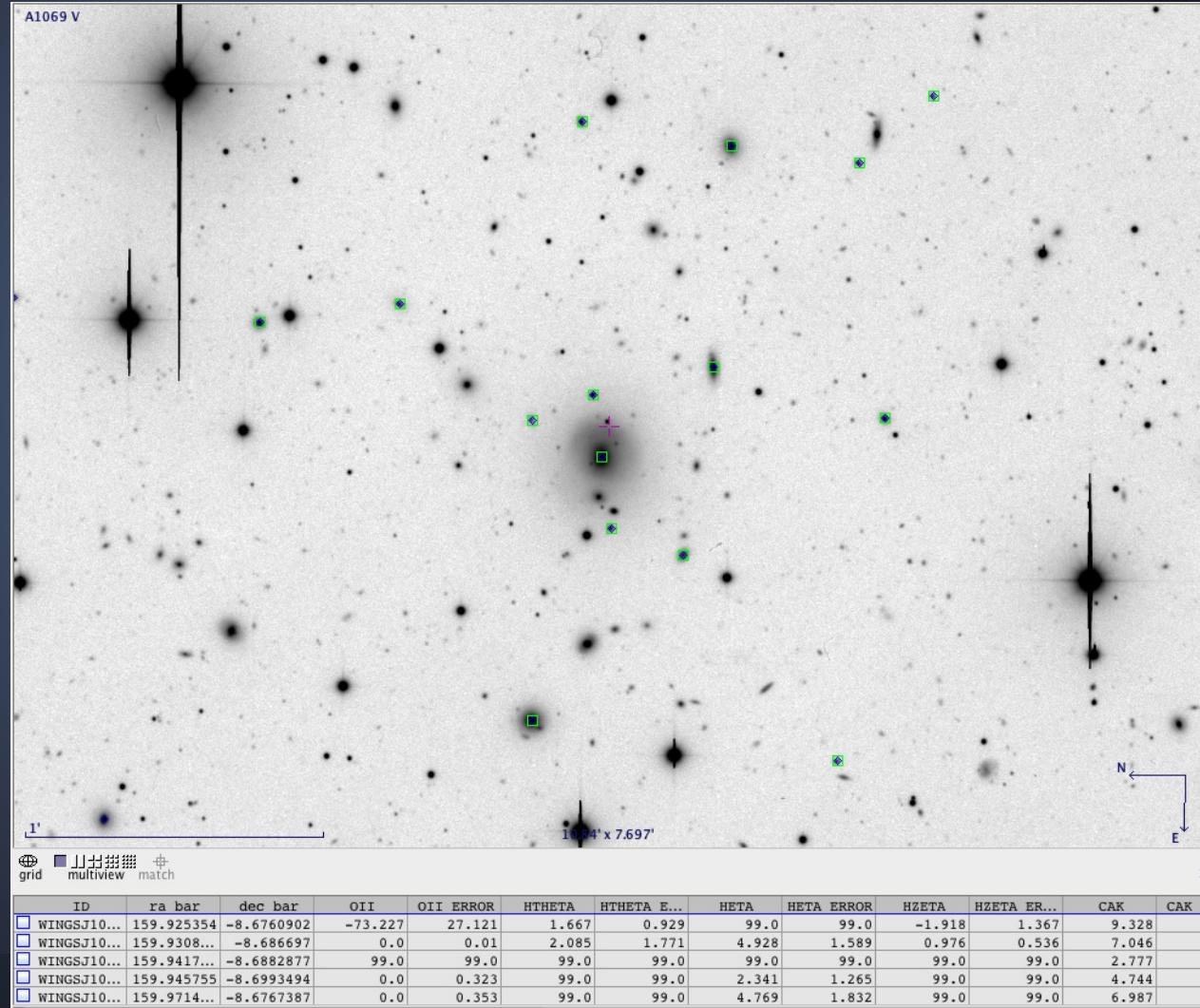
# The VO tools and WINGS: how to

Gasphot catalog (V, K)

- > Surf. brightness
- > Sersic index
- > Axial ratio



# The VO tools and WINGS: how to



# Concluding remarks: USE WINGS!

Cone Search

Available Cone Services

Registry: <http://registry.euro-vo.org/services/RegistrySearch>

Keywords: wings

Match Fields:  Short Name  Title  Subjects  ID  Publisher  Description

Accept Resource Lists

Cancel Query  Submit Query

Short Name	Title	Subjects	Identifier
J/A+A/470/39	Substructures in WINGS clusters (Ramella+, 2007)	Clusters_of_galaxies	ivo
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AccessURL      Description      Version

Cone Parameters

Cone URL:

Object Name:   Resolve

RA:  degrees (J2000)  Accept Sky Positions

Dec:  degrees (J2000)

Radius:  degrees

OK