

BLACK HOLES IN THE UNIVERSE

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THE IDEA OF BLACK HOLE

First proposed by John Michell in 1783

(Philosophical Transactions of the Royal Society)

IN THE CONTEXT OF CORPUSCULAR THEORY OF LIGHT:

“...suppose the particles of light to be attracted
in the same manner as all other bodies...”

THERE SHOULD EXIST BLACK HOLES:

“...there should exist in nature bodies from which light could not arrive at us...”

CAN BE DETECTED BY THE MOTION OF COMPANION STARS:

“...we might still perhaps from the motions of these revolving bodies infer the
existence of the central objects with some degree of probability...”

SUPERMASSIVE BLACK HOLES

(in the context of Newtonian physics)

BLACK HOLES MAY EXIST:

“...tous ces **corps devenus invisibles**...”

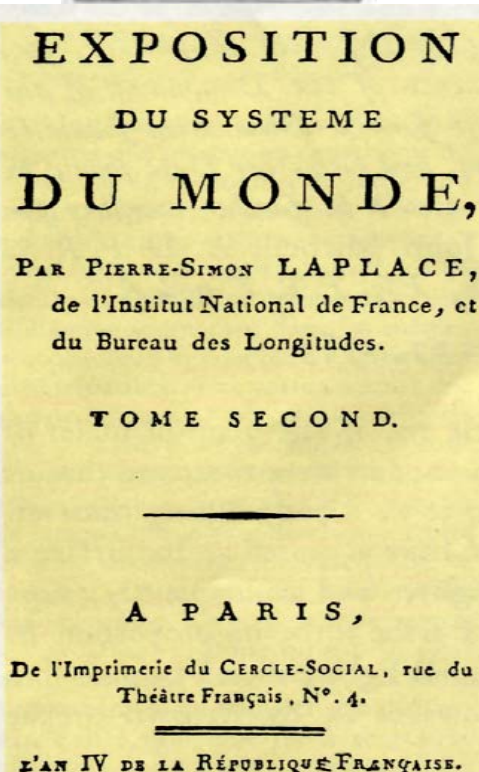
IN VERY LARGE NUMBERS:

“...Il existe donc dans les espaces celestes, des corps obscurs aussi considerables, et peut etre en aussi grand nombre, que les etoiles.”

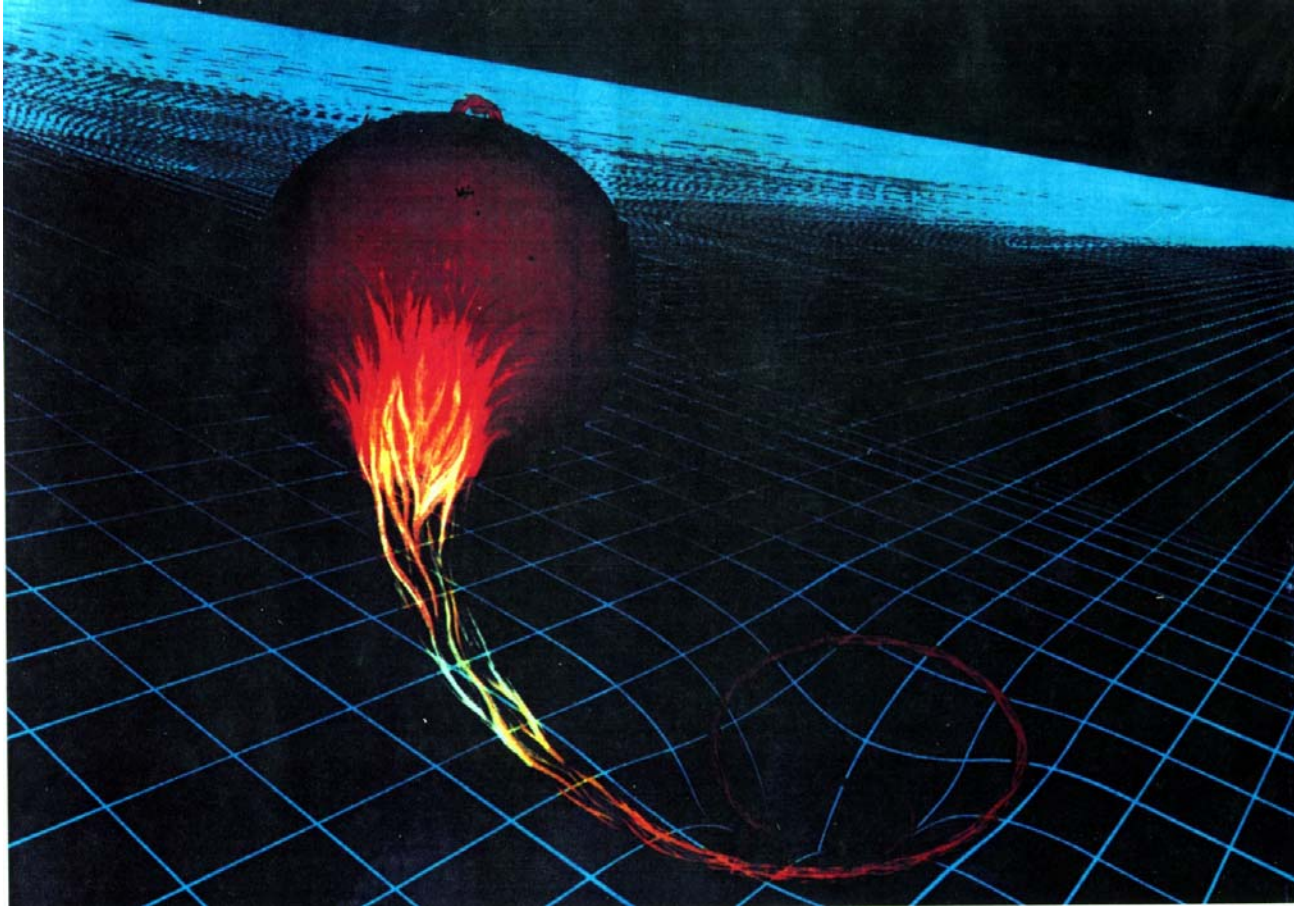
THE LARGEST OBJECTS IN THE UNIVERSE:

“...ne lesserait en vertu de son attraction, parvenir aucun de ses rayons jusqu'a nous; il est donc possible que les plus grands corps lumineux de l'univers, soient par cela meme, invisibles.”

Idea remained silent due to the ondulatory theory of light



RELATIVISTIC GRAVITY



LIGHT IS DEFLECTED AND λ MODIFIED IRRESPECTIVE
OF THE CORPUSCULAR OR WAVE THEORY OF LIGHT

RELATIVISTIC THEORY OF BLACK HOLES

Einstein



1916

Schwarzschild



1916

Eddington



1924

Chandrasekhar



1930

Oppenheimer



1939

Theory of
gravitation:
black holes
should exist

Horizon:
 $R_s = 2 G M / c^2$
 $\sim 3 M / M_\odot \text{ km}$

White dwarfs:
 $R_{wd} = R_T$

Limit of mass
for white dwarfs:
 $M_{wd} < 1.4 M_\odot \Rightarrow$

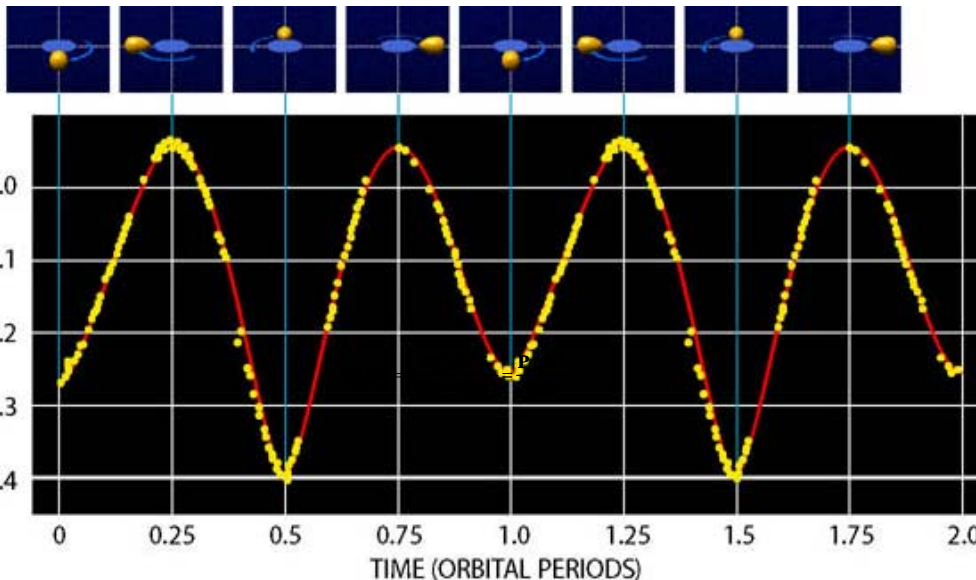
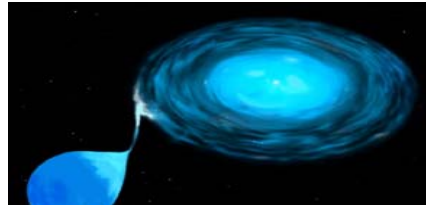
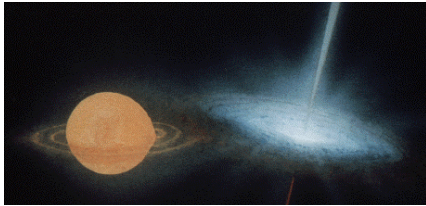
Neutron stars:
 $M > 1.4 M_\odot$
 $R \sim 20 \text{ km}$
Black holes:
 $M > 3 M_\odot$

Black holes have mass, angular momentum, charge (?), but no material surface

OBSERVATION OF STELLAR-MASS BHs

DISCOVERED AS X-RAY SOURCES

High Mass and Low Mass X-Ray Binary Systems



Fonction de masse: $f_x(M) = \frac{M_n^3 \sin^3 i}{(M_n + M_x)^2} = \frac{P_{orb} K}{2\pi G}$

Minimum de masse de l'objet compact

$K = (V_{max} - V_{min}) / 2 ; \quad f < M_n$

(Giacconi et al. 1962)

IN BINARY STELLAR SYSTEMS:

as predicted by Michell (1783)

$M > 3 M_{\odot} \Rightarrow \text{BLACK HOLE}$

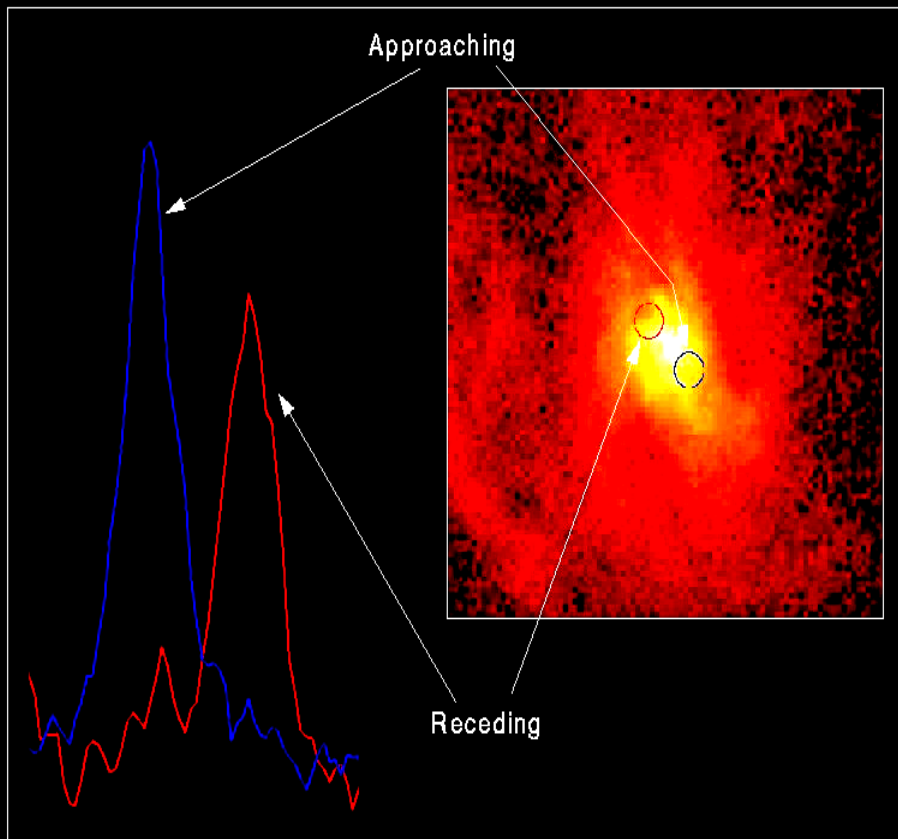
- 20 BHs known in binaries and other 20 additional candidates
- Estimated population in the Galaxy $\sim 3 \times 10^8 \Rightarrow$
- Assuming $\sim 10 M_{\odot}$ this form of dark mass is $\sim 4\%$ of total baryonic mass of the Galaxy
- Outweighs the supermassive black hole at Galactic Centre by a factor of 10^3

SUPERMASSIVE BLACK HOLES IN THE UNIVERSE

Kinematics of the H_α line with HST

$$\Rightarrow M_{\text{BH}} \sim 10^8 M_\odot \text{ in M 87}$$

Spectrum of Gas Disk in Active Galaxy M87



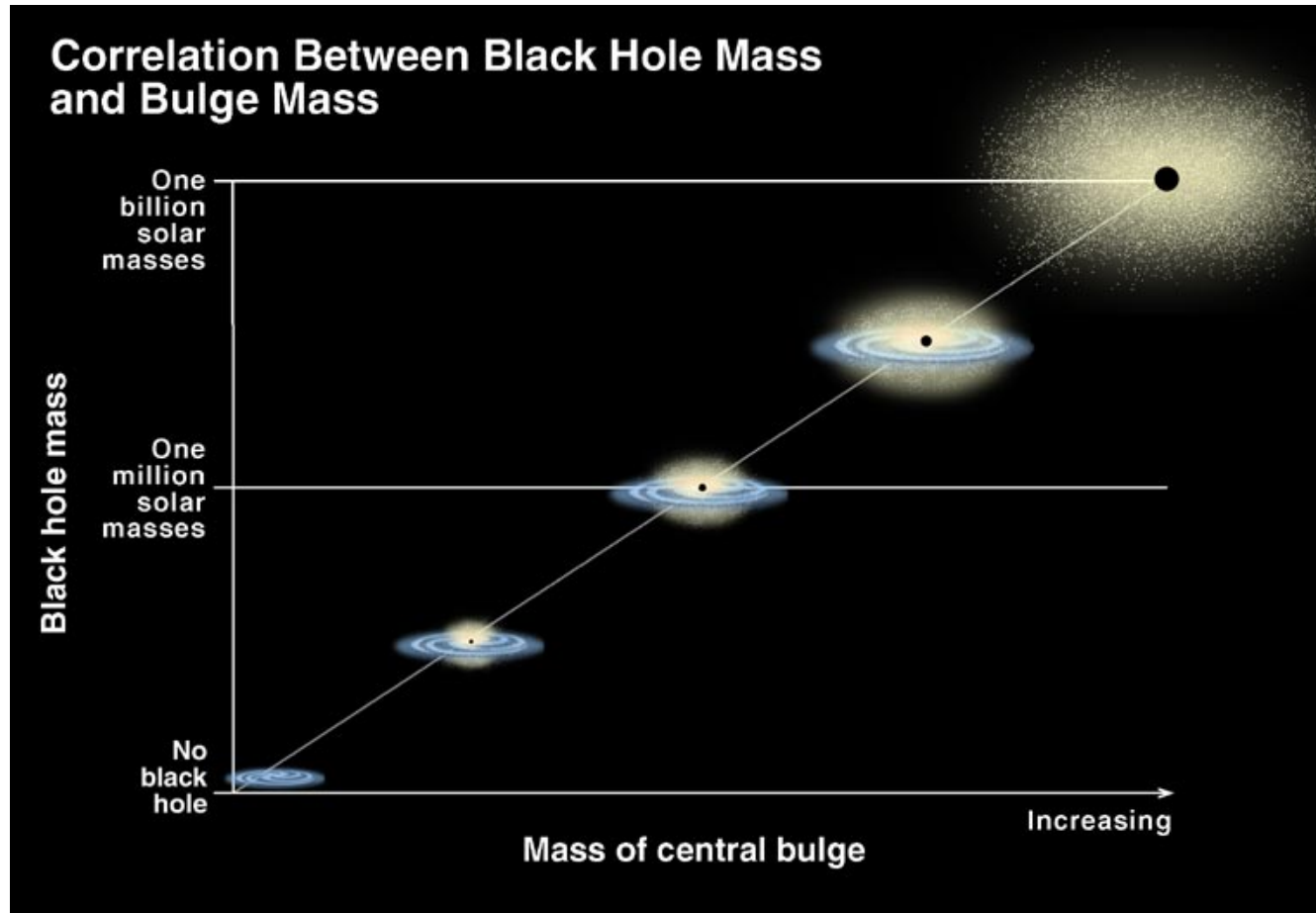
Hubble Space Telescope • Faint Object Spectrograph

H_2O masers with VLBA

$$\Rightarrow M_{\text{BH}} \sim 10^7 M_\odot \text{ in NGC 4258}$$



MASSIVE GALAXIES HOST SUPERMASSIVE BLACK HOLES

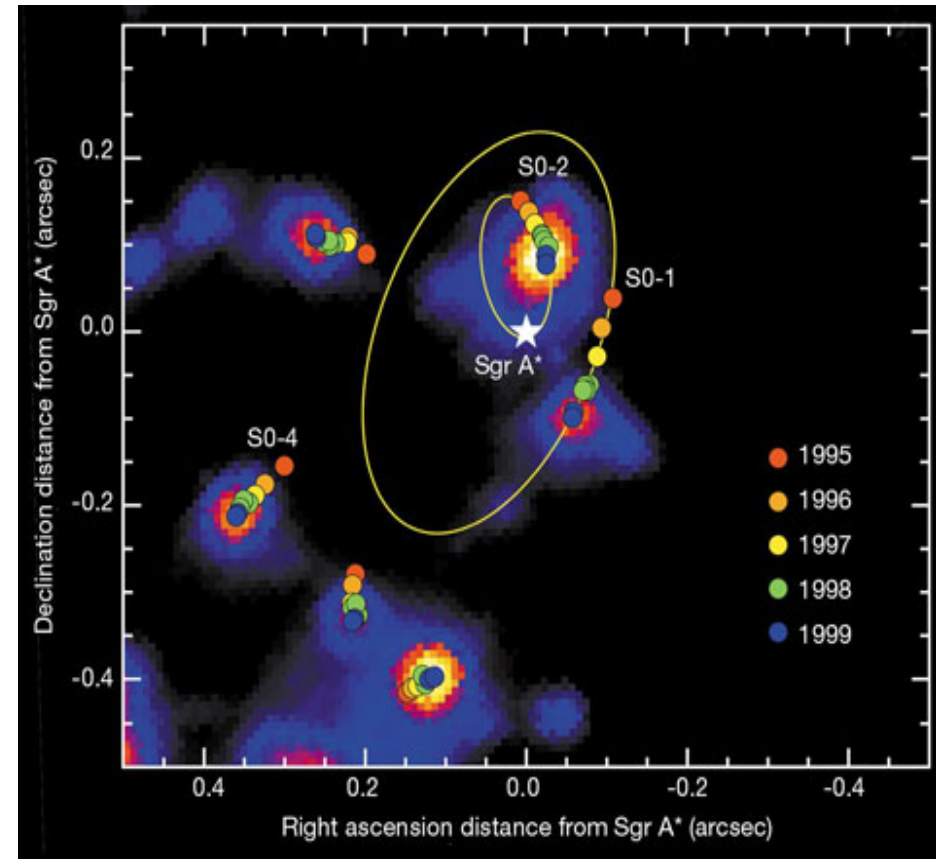
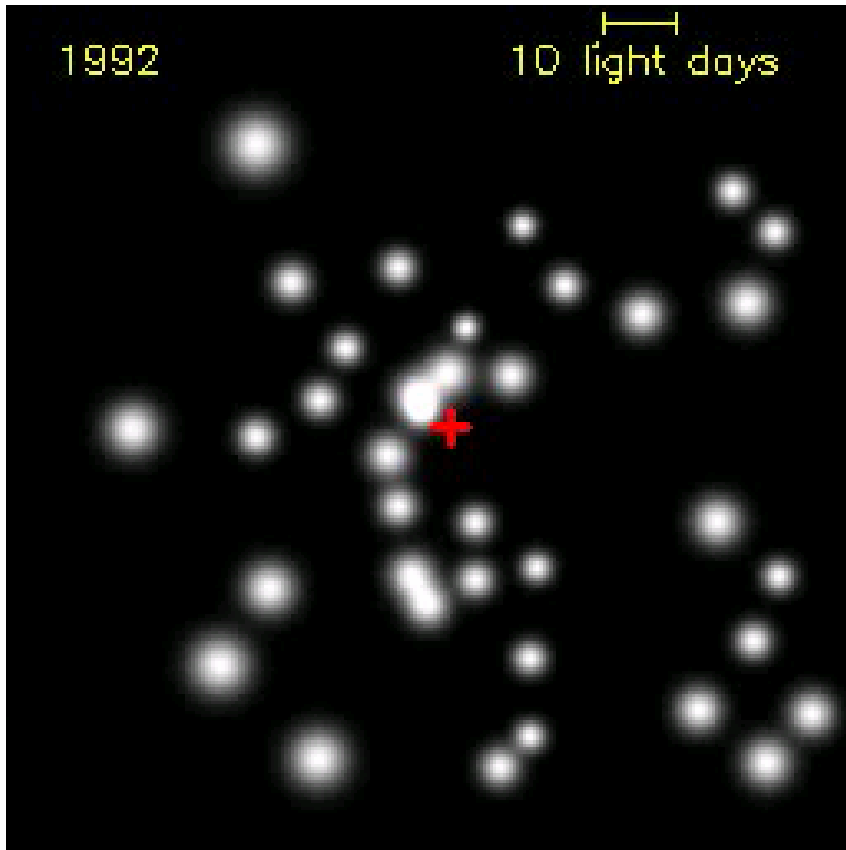


- HOW ARE SUPERMASSIVE BLACK HOLES FORMED ?
- ARE THERE BHs OF INTERMEDIATE MASS (10^2 - $10^4 M_{\odot}$) ?

BLACK HOLE IN THE GALACTIC CENTRE

Eckart & Genzel (IR with VLT-ESO)

Ghez, Morris et al. (IR with Keck)



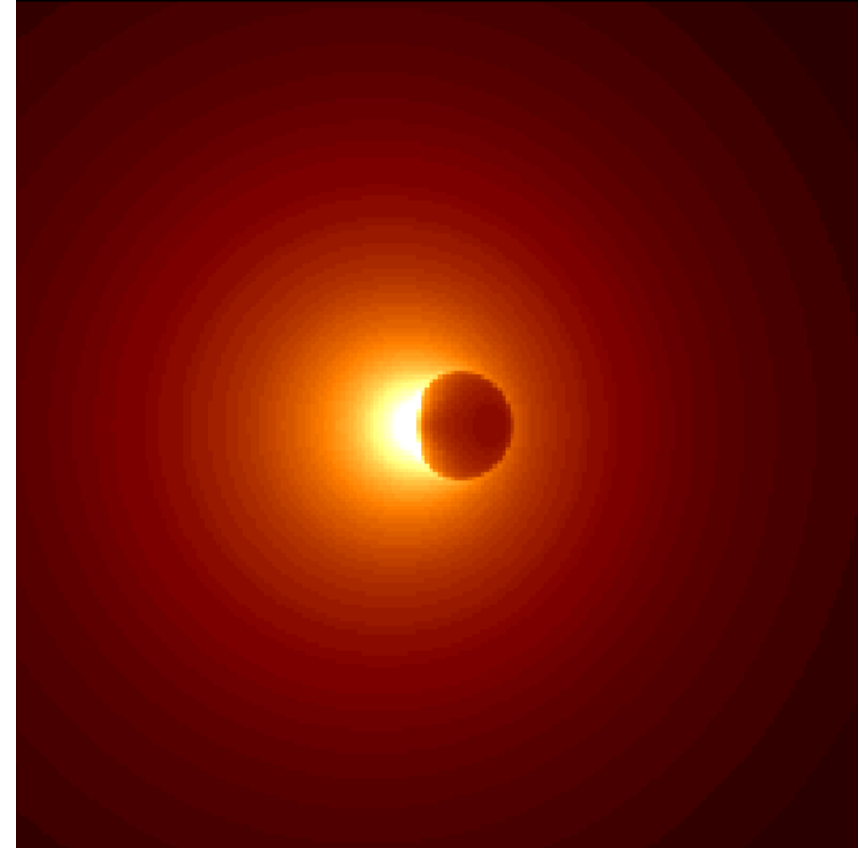
- Must be a single object with $M_{\text{BH}} \sim 3 \times 10^6 M_{\odot}$
- Black Hole in Andromeda $M_{\text{BH}} \sim 10^8 M_{\odot}$ (Bender et al. 2005)
- How stars $< 10^7$ yr old can exist in such environments ?
- But this is an indirect evidence of black hole

CAN WE OBTAIN A DIRECT IMAGE OF Sgr A* ?

“HORIZON”: DEFINING CONCEPT OF BLACK HOLE

J.P. Luminet

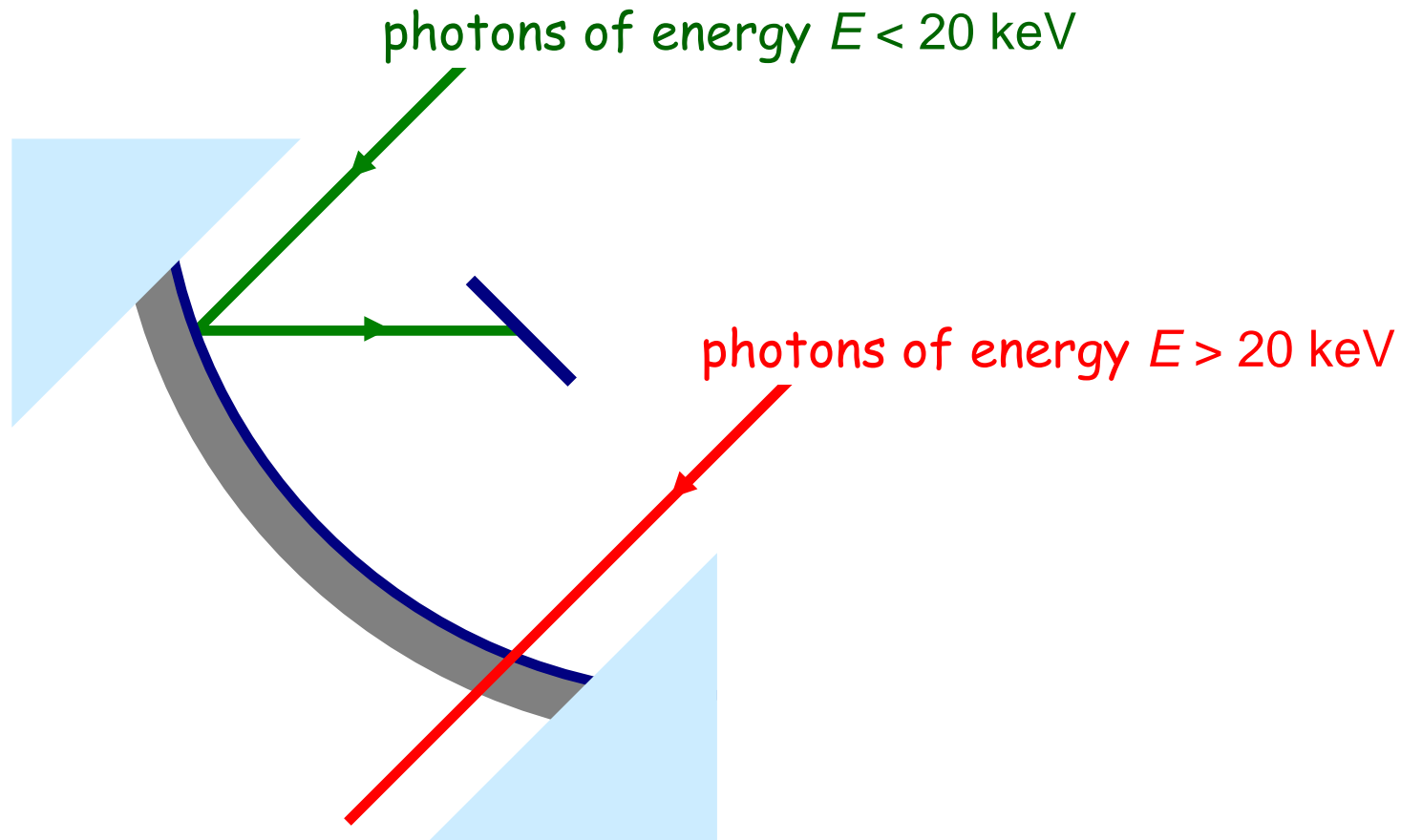
- Dark circle caused by radiation from sources behind that are being swallowed by the event horizon.
- Bright ring due to rays deflected by BH
- Shadow is off-centre due to flung of photons in the direction of BHs' rotation



- $3 \times 10^6 M_{\odot}$ confined in a region enclosed by the orbit of the Earth
- $D = 30 \mu\text{arcsec}$ to be imaged with VLBI at sub-millimeter or X-rays

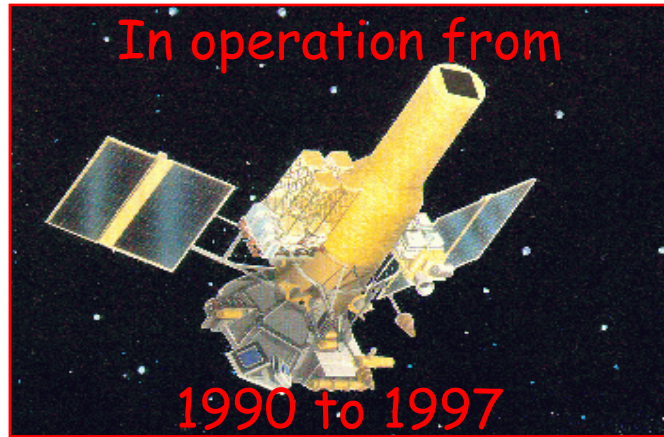
Is Sgr A* an annihilation source of e^+e^- at 511 keV ?

OBSERVATIONAL LIMITATIONS IN HIGH ENERGY ASTROPHYSICS



- The imaging at $E > 20$ keV is done with coded mask techniques which can provide positions for the sources with resolutions of a few arcmin
- Focusing photons with energy $E > 20$ keV, a challenge for the future

GRANAT SURVEY OF THE GALACTIC CENTRE

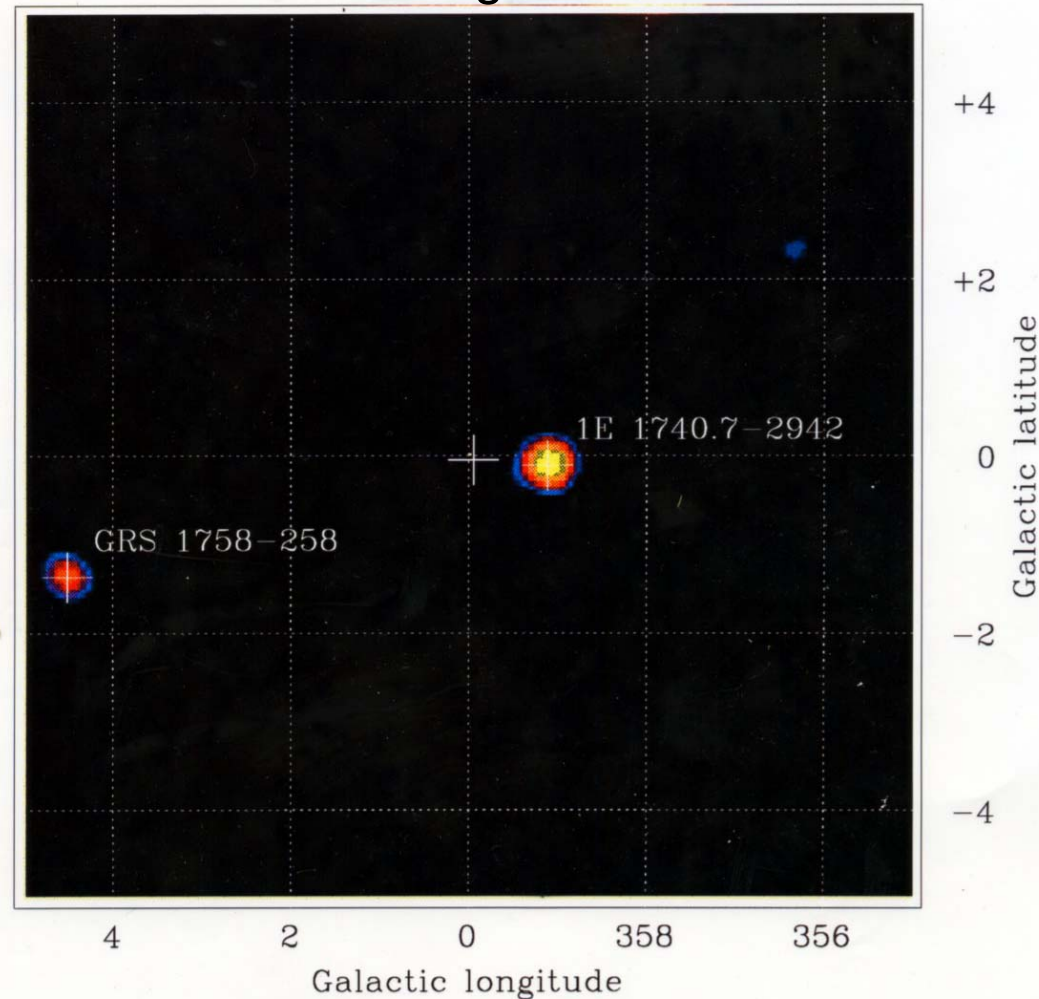


with a simultaneous
multiwavelength approach

VERY LARGE ARRAY



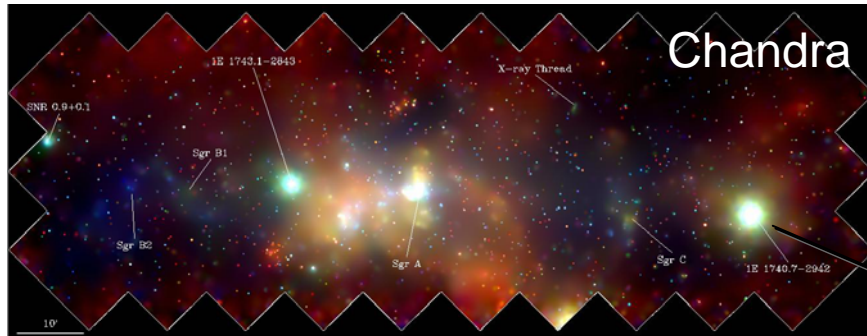
75-150 keV image for 1990-1997



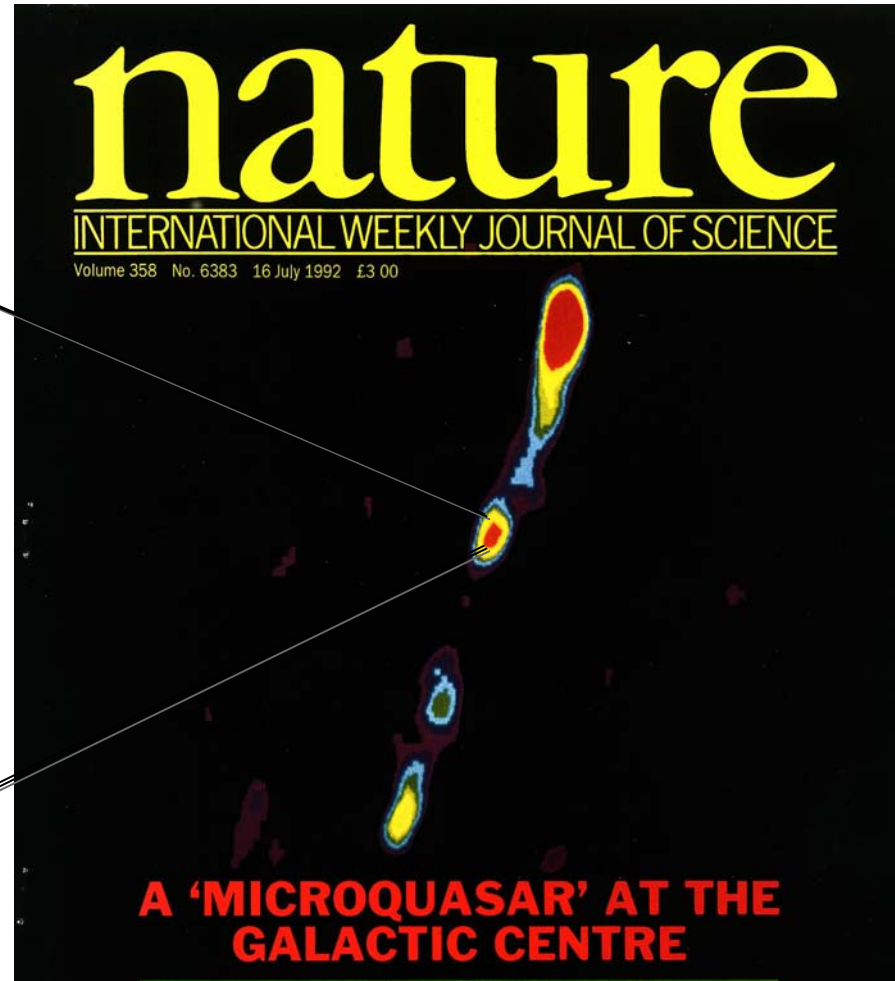
In 1990-97 Sgr A* was silent (Goldwurm et al.) \Rightarrow ADAF (Narayan et al.)
In 2002-5 Chandra & XMM detected flares at energies <10 keV

HIGH ENERGY SOURCES IN THE GALACTIC CENTRE

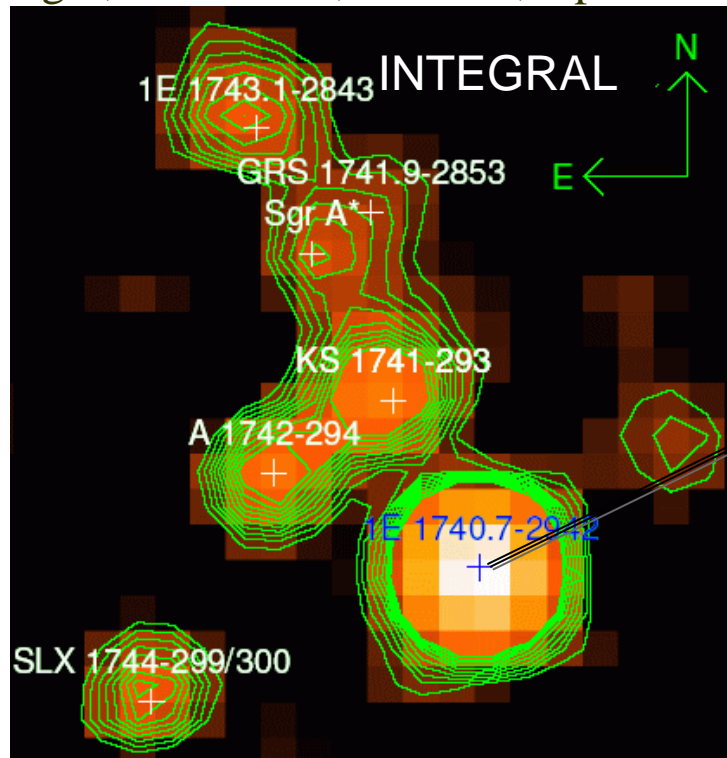
Wang et al. ApJ 2002



M., Rodriguez, et al, 1992



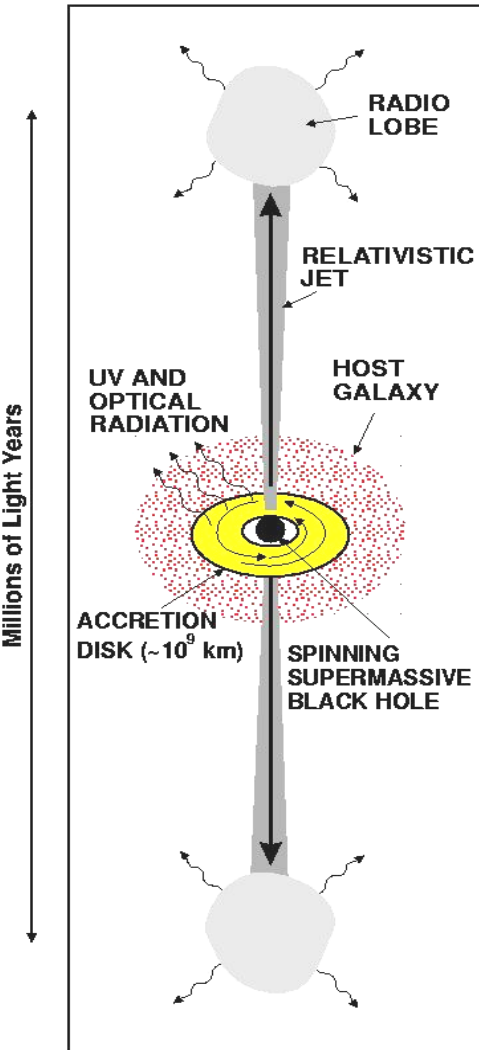
Belanger, Goldwurm, Goldoni, ApJ 2003



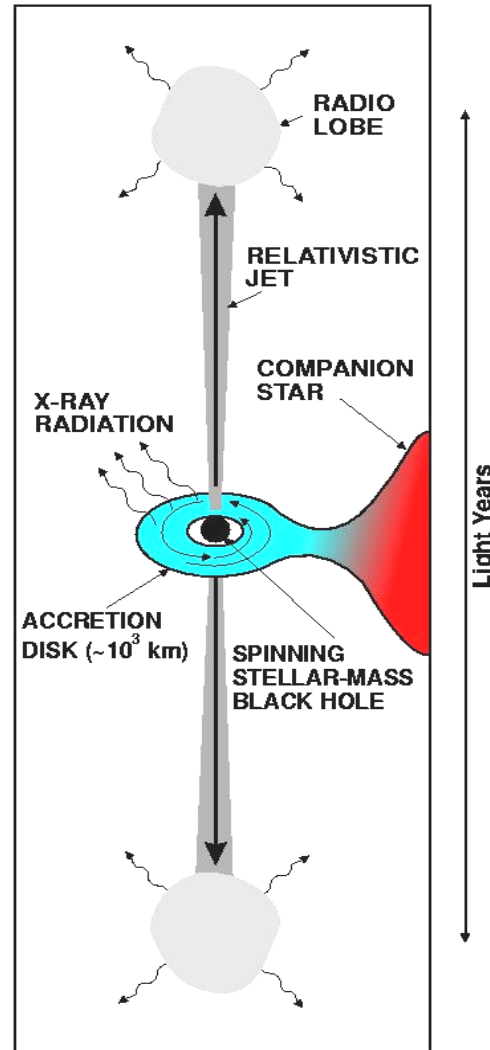
Five international workshops

QUASAR-MICROQUASAR ANALOGY

QUASAR



MICROQUASAR



M. & Rodriguez; Nature 1998

The scales of length and time are proportional to M_{BH}

$$R_{sh} = 2GM_{BH}/c^2 ; \Delta T \propto M_{BH}$$

Unique system of equations:

The maximum color temperature of the accretion disk is:

$$T_{col} \propto (M/10M_{\odot})^{-1/4}$$

(Shakura & Sunyaev, 1976)

Waited era of space astronomy

For a given accretion rate:

$$L_{Bol} \propto M_{BH} ; l_{jet} \propto M_{BH} ;$$

$$\phi \propto M_{BH}^{-1} ; B \propto M_{BH}^{-1/2}$$

(Sams, Eckart, Sunyaev, 96; Rees 04)

APPARENT SUPERLUMINAL MOTIONS IN μ QSOs AS IN QSOs ?