

Introduction to Bibliometrics and Tools for Organizing References

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Astronomy
evaluate?



Possible measures

- ▶ Number of talks
- ▶ Invitations to conferences
- ▶ Students, graduations
- ▶ Press releases
- ▶ Research grants
- ▶ Number of papers

————→ **bibliometrics**



Bibliometrics

Set of **methods** used for **publication** and **citation analysis**
in order to explore **impact** in respective field.

Bibliometrics

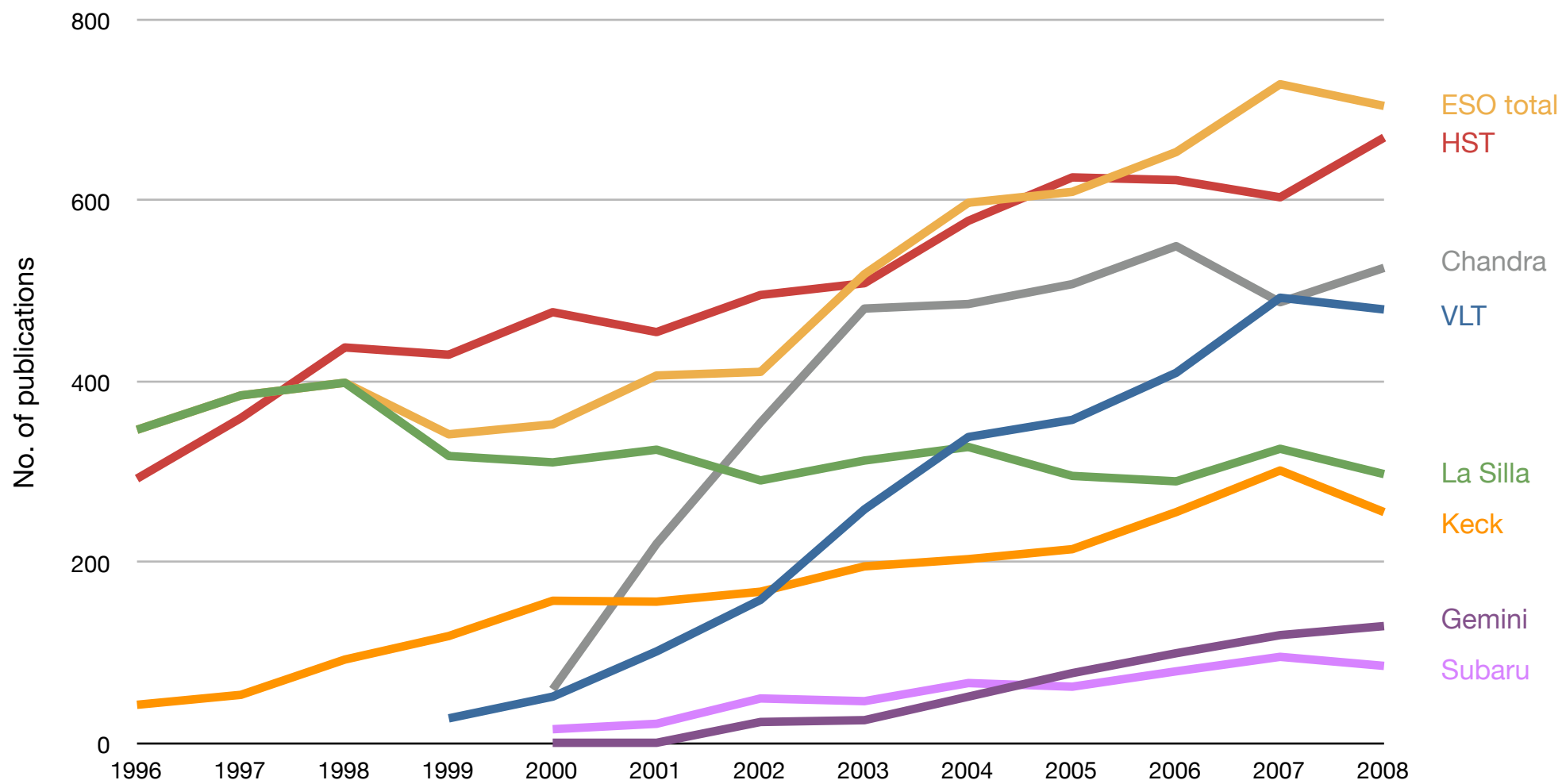
Ingredients:

- ▶ “Objects” for evaluation (researchers, observatories)
- ▶ Appropriate measures (methods)
- ▶ Pool of information (publication & citation databases)

Common Methods of Evaluation

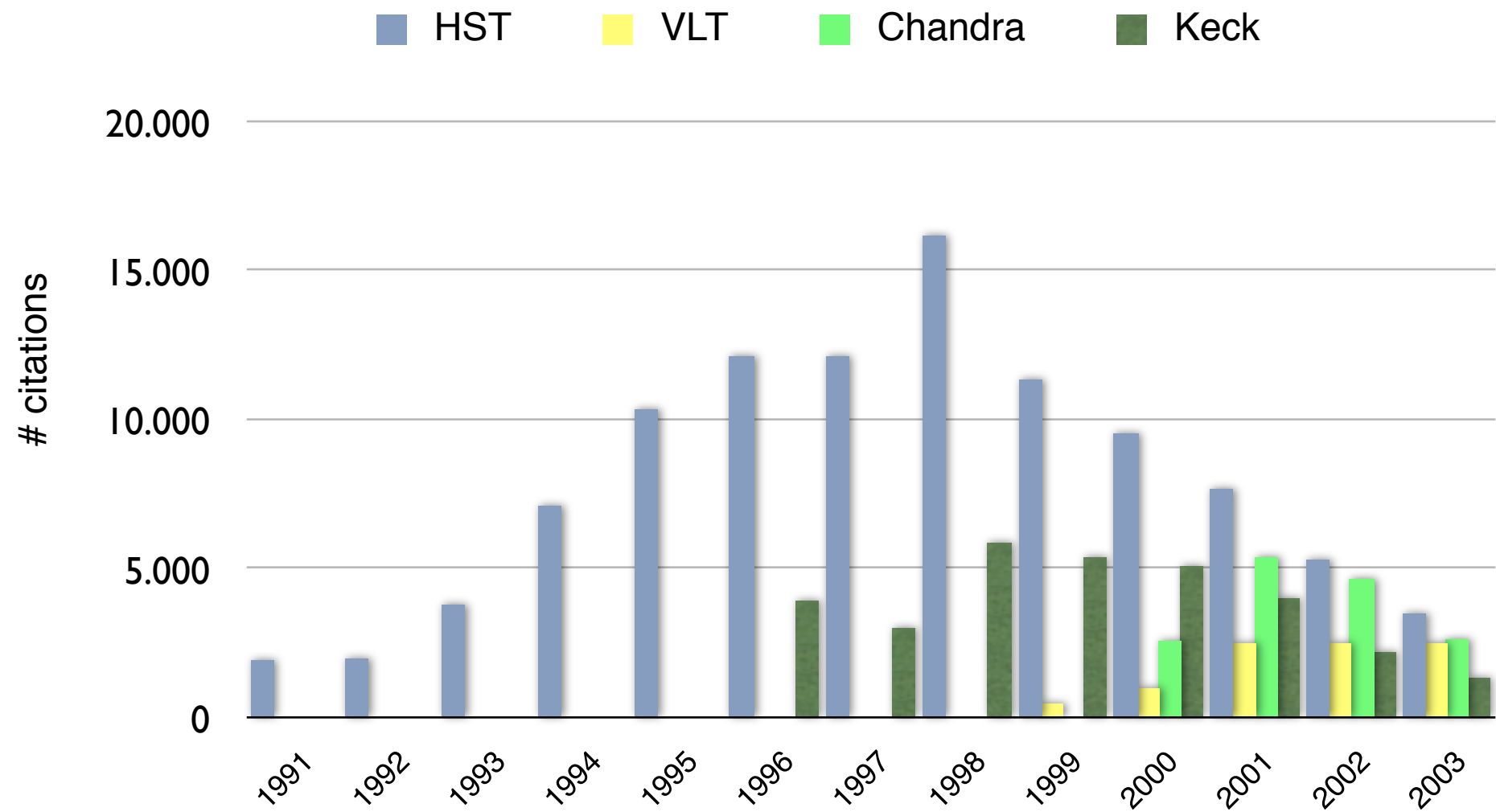
	Good	Bad
# Publications	productivity	no impact

Publications of major observatories by year



Common Methods of Evaluation

	Good	Bad
# Publications	productivity	no impact
# Citations	impact	delayed



Common Methods of Evaluation

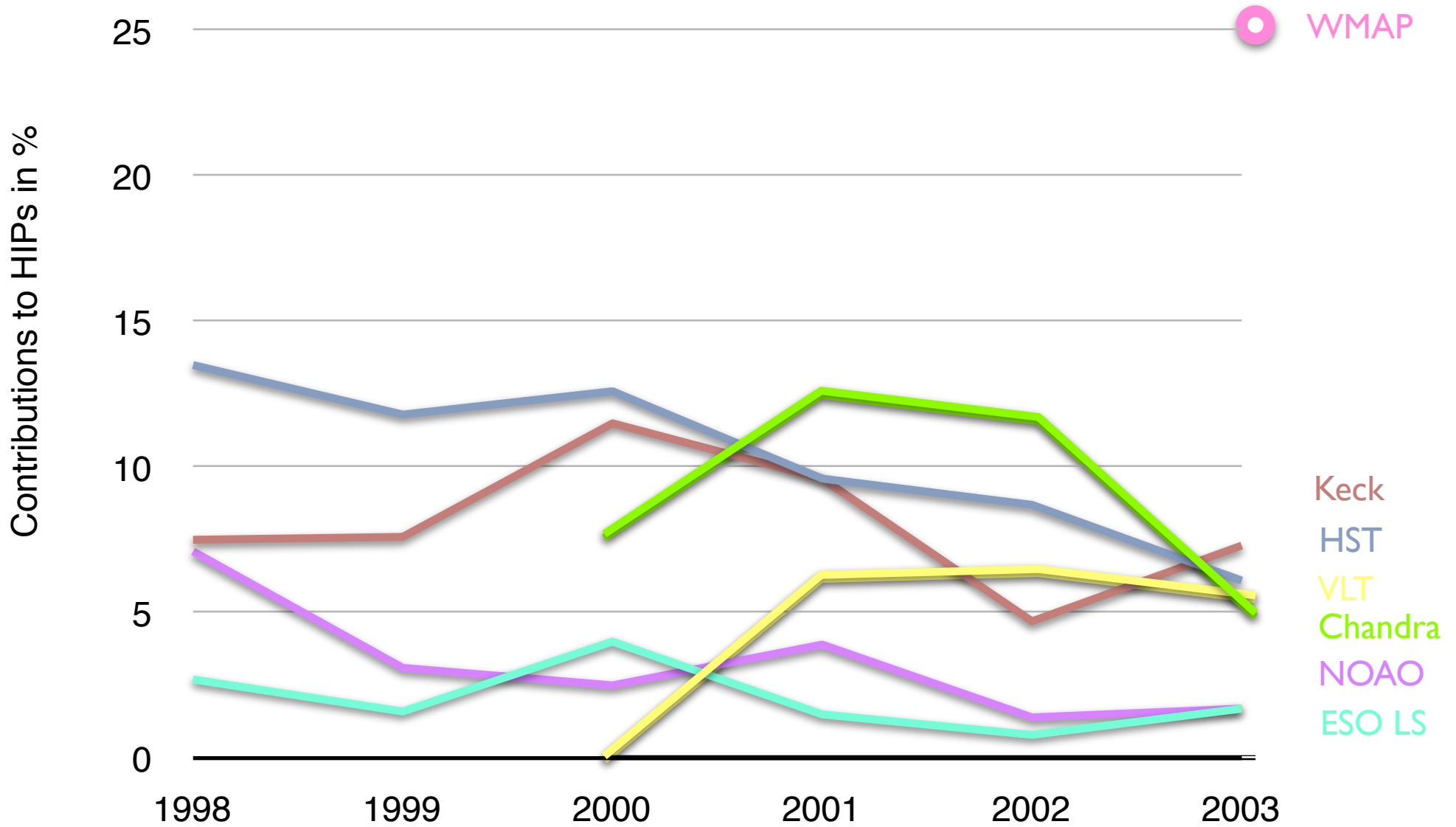
	Good	Bad
# Publications	productivity	no impact
# Citations	impact	delayed
mean / median cites per paper	allows comparison of different ages	rewards low productivity

Common Methods of Evaluation

	Good	Bad
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'High-Impact Papers'	shows trends	favors 'hot topics'

High Impact Papers 2005

from Juan Madrid (STScl)



The World's Top Ten Telescopes [2006]

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Published online 6 February 2009 | Nature | doi:10.1038/news.2009.81

News

The world's top ten telescopes revealed

The best observatories ranked by their scientific impact.

[Eric Hand](#)

It doesn't take a big mirror to have a big impact. The Sloan Digital Sky Survey, a project conducted with a modest 2.5-metre-wide telescope in New Mexico, performed the most highly cited science in 2006, according to a new analysis of the top ten 'high impact' astronomical observatories¹.

"It measures how hot the science of the telescope is," says Juan Madrid of McMaster University in Hamilton, Canada, of the top-ten table he has released for most years since 1998. "In a way it measures how good the time-allocation committee is and how good the telescope is. I will also say it measures how good the




SDSS image of Messier 51, the Whirlpool Galaxy.

Sloan Digital Sky Survey

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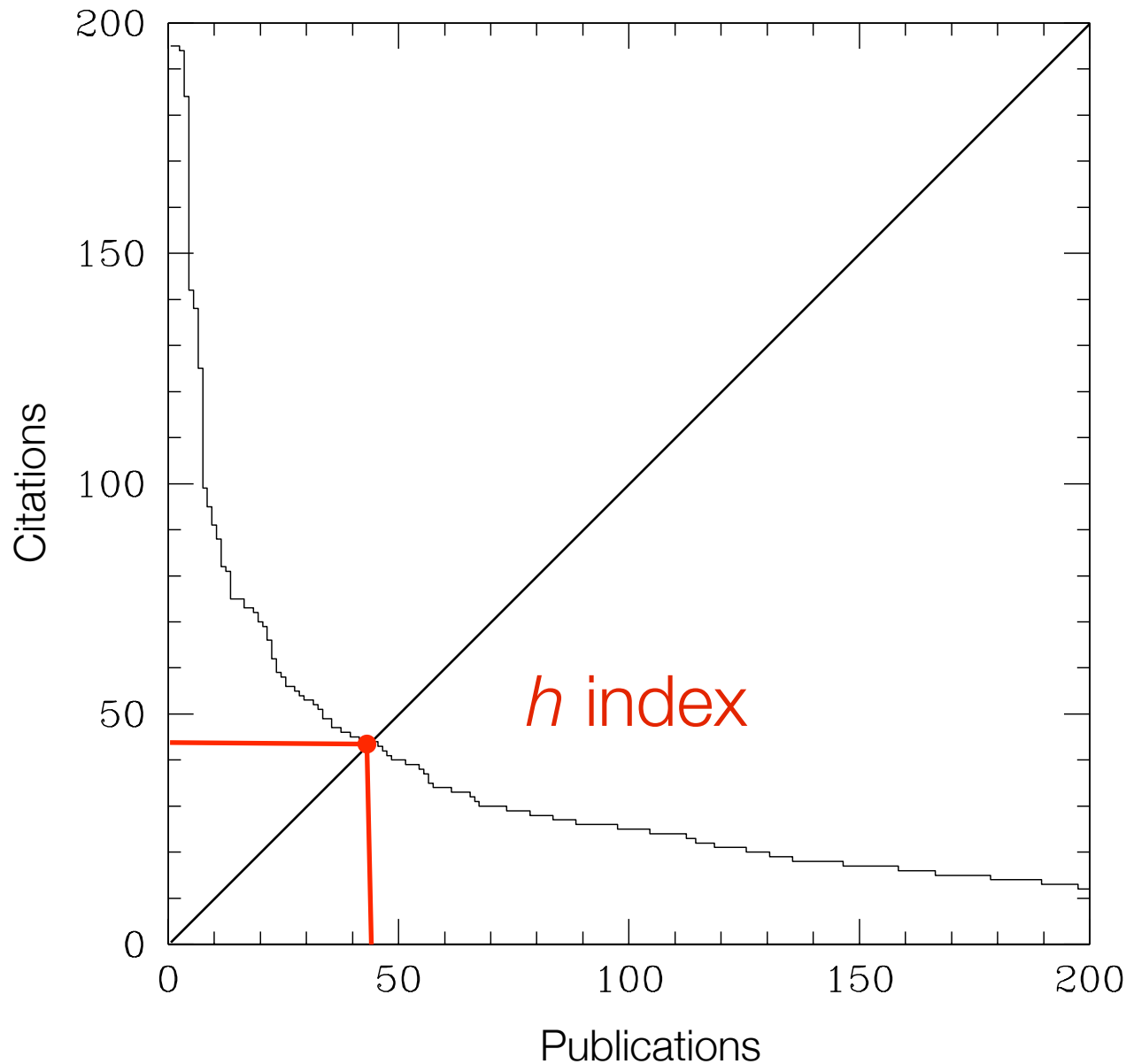
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Common Methods of Evaluation

	Good	Bad
# Publications	productivity	no impact
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mean / median cites per paper	allows comparison of different ages	rewards low productivity
'High-Impact Papers'	shows trends	favors 'hot topics'
<i>h</i> -index	productivity + impact	determined by years of operation

h -index



h papers with
at least
 h citations each

(Hirsch 2005)

h -index from the ADS

no.

citations

34	<input type="checkbox"/> 1991MNRAS.249..742M McHardy, I. M.; Abraham, R. G.; Crawford, C. S.; Ulrich, M.-H.; Mock, P. C.; Vanderspeck, R. K.	37.000	04/1991	A	F G	PKS1413 + 135 - A BL Lac object in a
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36	<input type="checkbox"/> 2002ApJ...570...54C Chen, Hsiao-Wen; McCarthy, P. J.; Marzke, R. O.; Wilson, J.; Carlberg, R. G.; Firth, A. E.; Persson, S. E.; Sabbey, C. N.; Lewis, J. R.; McMahon, R. G.; and 9 coauthors	35.000	05/2002	A	E F	The Las Campanas Infrared Survey. III. Photometric Catalogs

h -index / m -parameter

- h will be different for different disciplines and for different sub-fields within a discipline (e.g., planetary vs. UV astronomy)
- does not reflect number of years of activity
- m -parameter: $m = h / t$

Problems

Caveats of citation analysis:

- ▶ **Incompleteness**: listings often incomplete
 - ▶ **Incorrectness**: incorrect citing, multiple journal abbreviations
 - ▶ **Citing behavior**: cite well-known authors, friends citing friends
 - ▶ **Multi-author papers**: self-cites (introduce normalized citation counts)
-
- * Retrieving citations: **don't rely on only one source**
 - * Beware comparisons
 - * Bad style: citing without reading original

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Title: The morphological identification of the rapidly evolving population of faint galaxies

Authors: [Glazebrook, Karl](#); [Ellis, Richard](#); [Santiago, Basilio](#); [Griffiths, Richard](#)

Affiliation: AA(Institute of Astronomy, Madingley Road, Cambridge CB3 0HA, UK), AB(Institute of Astronomy, Madingley Road, Cambridge CB3 0HA, UK), AC(Institute of Astronomy, Madingley Road, Cambridge CB3 0HA, UK), AD(Department of Astronomy, Johns Hopkins University, 3400 North Charles St, Baltimore, MD 21218, USA)

Publication: Monthly Notices of the Royal Astronomical Society, Volume 275, Issue 2, pp. L19-L22. ([MNRAS Homepage](#))

Publication Date: 07/1995

Origin: [MNRAS](#); KNUDSEN

MNRAS Keywords: SURVEYS, GALAXIES: EVOLUTION, GALAXIES: STRUCTURE, GALAXIES: PECULIAR, COSMOLOGY

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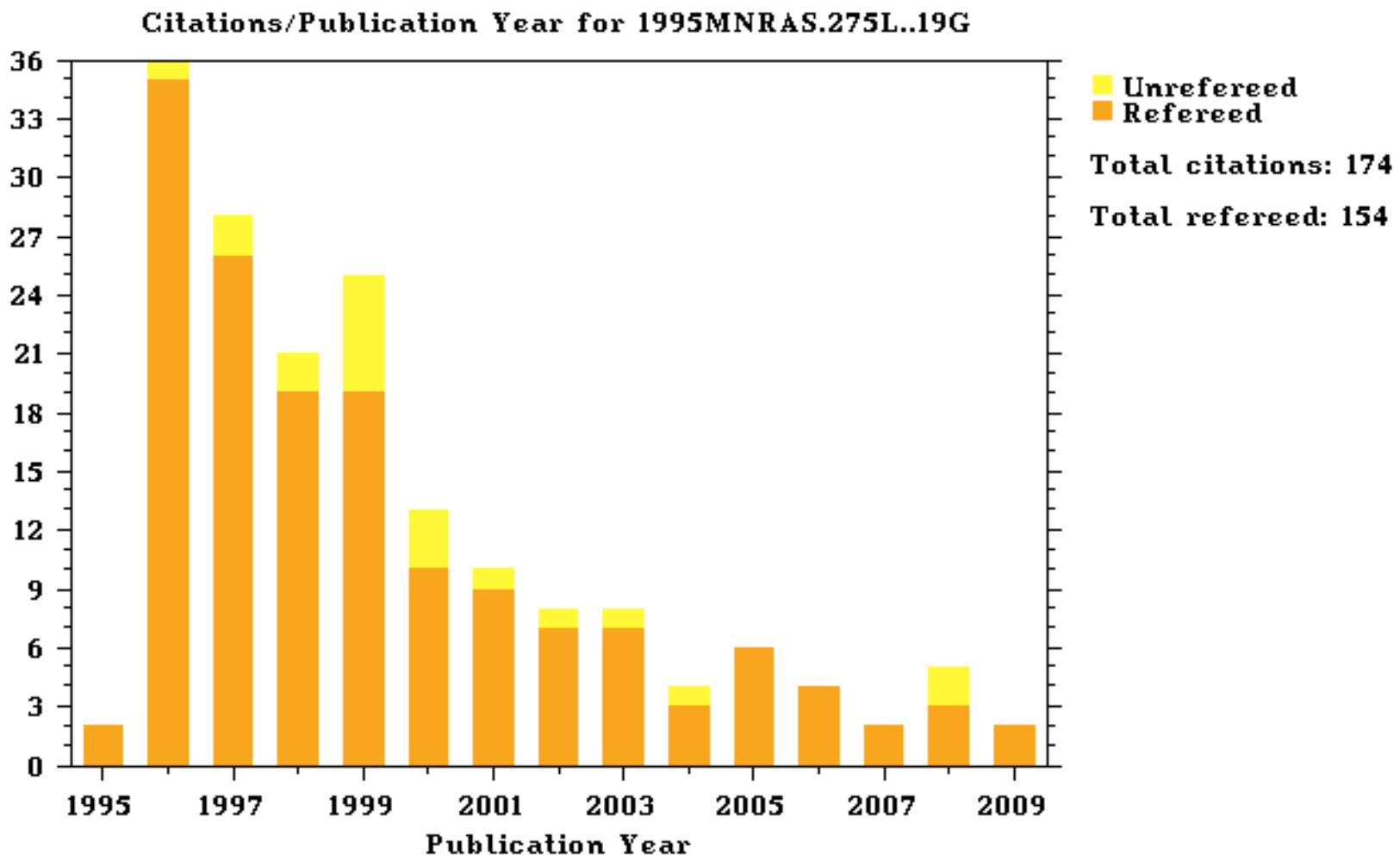
Abstract

The excess numbers of blue galaxies at faint magnitudes are a long-standing cosmological puzzle. We present new number-magnitude galactic morphology from the first deep fields of the Cycle 4 Hubble Space Telescope Medium Deep Survey project. From a sample of

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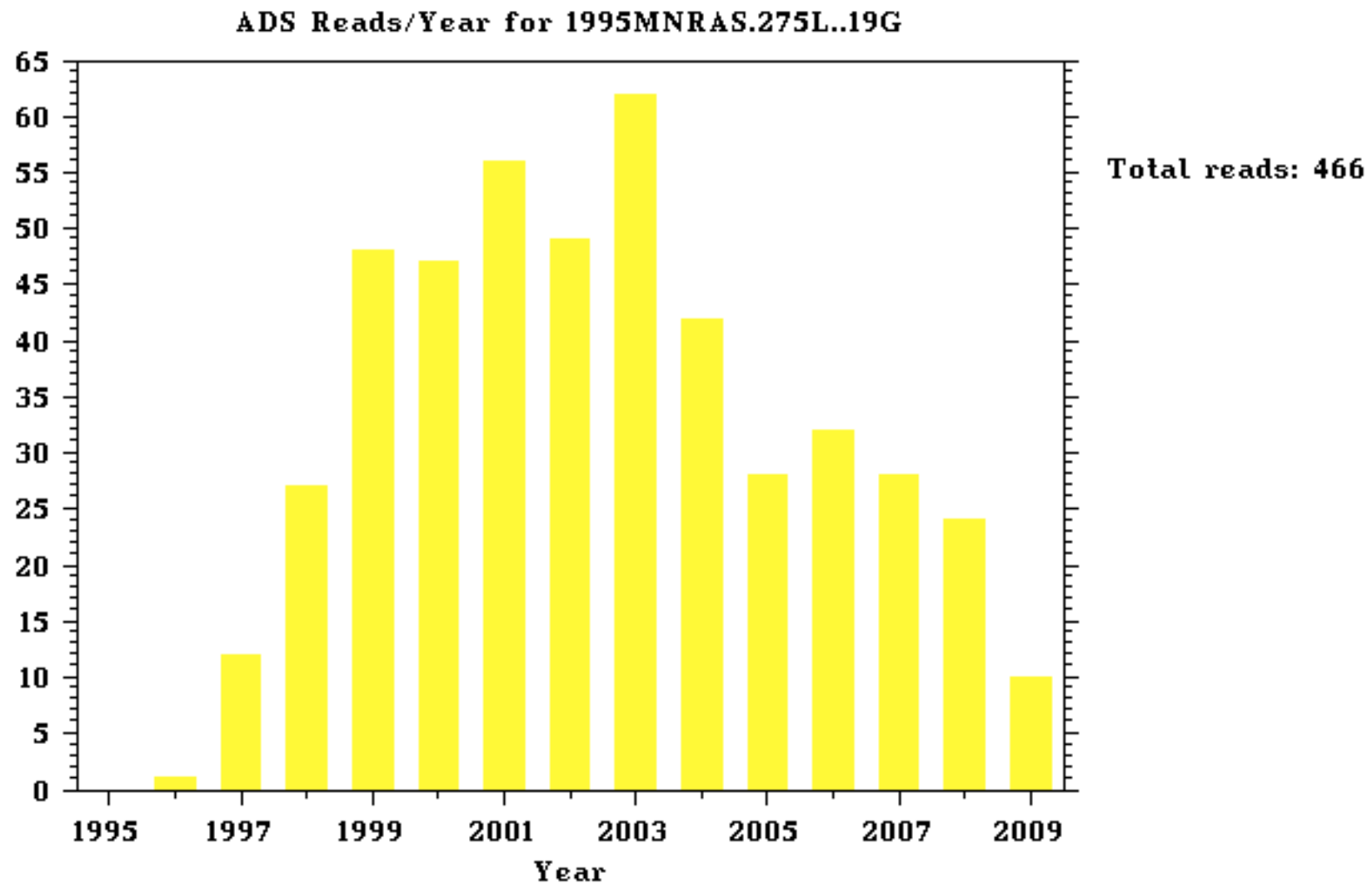
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http://arxiv.org/abs/astro-ph/0508407

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Astrophysics

The Gemini Deep Deep Survey. VII. The Redshift Evolution of the Mass-Metallicity Relation

S. Savaglio (1), K. Glazebrook (1), D. Le Borgne (2), S. Juneau (3,4), R. Abraham (2), H.-W. Chen (5), D. Crampton (3), P. McCarthy (6), R. Carlberg (2), R. Marzke (7), K. Roth (8), I. Jorgensen (8), R. Murowinski (4) ((1) Johns Hopkins Univ., (2) Univ. of Toronto, (3) NRS Herzberg Institute, (4) Univ. of Montreal (5) MIT, (6) Carnegie Observatories, (7) San Francisco State Univ, (8) Gemini Observatory)

(Submitted on 18 Aug 2005)

We have investigated the mass-metallicity ($M-Z$) relation using galaxies at $0.4 < z < 1.0$ from the Gemini Deep Deep Survey and Canada-France Redshift Survey. Deep K and z' band photometry allowed us to measure stellar masses for 69 galaxies. From a subsample of 56 galaxies, for which metallicity of the interstellar medium is also measured, we identified a strong correlation between mass and metallicity, for the first time in the distant Universe. This was possible because of the larger base line spanned by the sample in terms of metallicity (a factor of 7) and mass (a factor of 400) than in previous works. This correlation is much stronger and tighter than the luminosity-metallicity, confirming that stellar mass is a more meaningful physical parameter than luminosity. We find clear evidence for temporal evolution in the $M-Z$ relation in the sense that at a given mass, a galaxy at $z=0.7$ tends to have lower metallicity than a local galaxy of similar mass. We use the $z=0.1$ SDSS $M-Z$ relation, and a small sample of $z=2.3$ Lyman break galaxies with known mass and metallicity, to propose an empirical redshift-dependent $M-Z$ relation, according to which the stellar mass and metallicity in small galaxies evolve for a longer time than in massive galaxies. This relation predicts that the generally metal poor damped Lyman-alpha galaxies have stellar masses of the order of $10^{8.8} M_{\odot}$ (with a dispersion of 0.7 dex) all the way from $z=0.2$ to $z=4$. The observed redshift evolution of the $M-Z$

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The Gemini Deep Deep Survey. VII. The Redshift Evolution of the Mass-Metallicity Relation

Authors: [Savaglio, S.](#); [Glazebrook, K.](#); [Borgne, D. Le](#); [Juneau, S.](#); [Abraham, R.](#); [Chen, H. -W.](#); [Crampton, D.](#); [McCarthy, P.](#); [Carlberg, R.](#); [Marzke, R.](#); [Roth, K.](#); [Jorgensen, I.](#); [Murowinski, R.](#)

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Comment: ApJ in press

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Astrophys.J.635:260-279,2005
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15 [Erb, Dawn K.](#); [Shapley, Alice E.](#); [Pettini, Max et al](#) (2006-02-21) oai:arXiv.org:astro-ph/0602473

We use a sample of 87 rest-frame UV-selected star-forming galaxies with mean spectroscopic redshift $z=2.26$ to study the correlation between metallicity and stellar mass at high redshift. Using stellar masses determined from SED fitting to 0.3-8 micron photometry, we divide the sample into six bins in ... Comment: 18 pages, 9 figures, 2 tables; accepted for publication in ApJ

Velocity-Metallicity Correlation for high- z DLA Galaxies: Evidence for a Mass-Metallicity Relation? [[Abstract](#), [12 Cites](#), [Pre-print PDF](#)]

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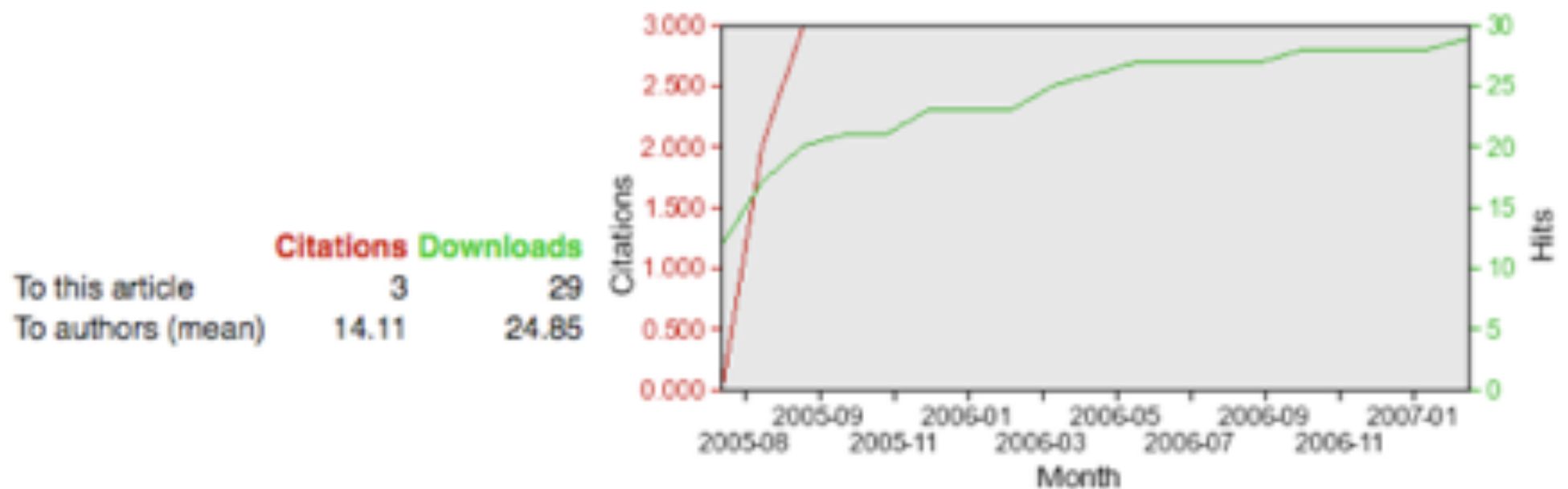
We used our database of VLT-UVES quasar spectra to build up a sample of 70 Damped Lyman-alpha (DLA) or strong sub-DLA systems with total neutral hydrogen column densities of $\log N(\text{H I}) > 20$ and redshifts in the range

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Part 2:

Tools for Organizing References

The Task

Scientists write papers → reference lists

- ▶ different references for each paper
- ▶ different citation styles

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Astrophysics

The Galaxy Population Hosting Gamma-Ray Bursts

S. Savaglio (MPE), K. Glazebrook (Swinburne University), D. Le Borgne (CEA/Saclay)

(Submitted on 18 Mar 2008 (v1), last revised 29 Dec 2008 (this version, v3))

We present the most extensive and complete study of the properties for the largest sample (46 objects) of gamma-ray burst (GRB) host galaxies. The redshift interval and the mean redshift of the sample are $0 < z < 6.3$ and $z = 0.96$ (look-back time: 7.2 Gyr), respectively. 89% of the hosts are at $z < 1.6$. Optical-near-infrared (NIR) photometry and spectroscopy are used to derive stellar masses, star formation rates (SFRs), dust extinctions and metallicities. The average stellar mass is $1.0 \pm 0.3 \times 10^{11} M_{\odot}$, with a 1 sigma dispersion of 0.8 dex. The average metallicity for a subsample of 17 hosts is about 1/6 solar and the dust extinction in the visual band (for a subsample of 10 hosts) is $A_V = 0.5$. We obtain new relations to derive SFR from [OII] or UV fluxes, when Balmer emission lines are not available. SFRs, corrected for dust extinction, aperture-slit loss and stellar Balmer absorption are in the range $0.01\text{--}36 M_{\odot} \text{ yr}^{-1}$. The median SFR per unit stellar mass (specific SFR) is 0.8 Gyr^{-1} . Equivalently the inverse quantity, the median formation timescale is 1.3 Gyr. Most GRBs are associated with the death of young massive stars, more common in star-forming galaxies. Therefore GRBs are an effective tool to detect star-forming galaxies in the universe. Star-forming galaxies at $z < 1.6$ are a faint and low-mass population, hard to detect by conventional optical-NIR surveys, unless a GRB event occurs. There is no compelling evidence that GRB hosts are peculiar galaxies. More data on the subclass of short GRB are necessary to establish the nature of their hosts.

Comments: 33 pages, 18 figures, 11 tables, final version to appear in ApJ, January 20 issue. Figure 18 corrected

Subjects: Astrophysics (astro-ph)

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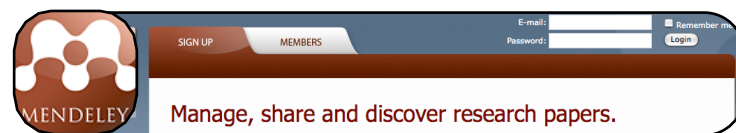
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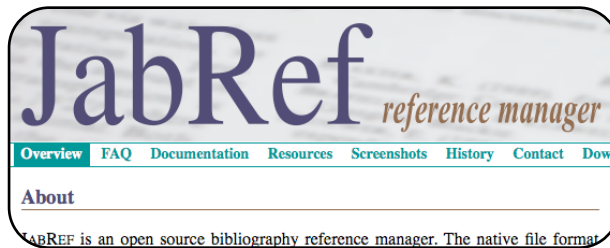
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2007, 377, 69

Abstract: With over 20 million records, the ADS citation database is regularly used by researchers and librarians to measure the scientific impact of individuals, groups, and institutions. In addition to the traditional sources of citations, the ADS has recently added references extracted from the arXiv e-prints on a nightly basis. We review the procedures used to harvest and identify the reference data used in the creation of citations, the policies and procedures that we follow to avoid double-counting and to eliminate contributions which may not be scholarly in nature. Finally, we describe how users and institutions can easily obtain quantitative citation data from the ADS, both interactively and via web-based programming tools.

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General Relativity and Gravitation

2008 vol. 40 pp. 221

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Title: The Galaxy Population Hosting Gamma-Ray Bursts

Authors: [Savaglio, S.](#); [Glazebrook, K.](#); [LeBorgne, D.](#)

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DOI: [10.1088/0004-637X/691/1/182](#)

Bibliographic Code: 2009ApJ...691..182S

Abstract

We present the most extensive and complete study of the properties for the largest sample (46 objects) of gamma-ray burst (GRB) host galaxies. The redshift interval and the mean redshift of the sample are $0 < z < 6.3$ and $z = 0.96$ (look-back time: 7.2 Gyr), respectively; 89% of the hosts are at $z \leq 1.6$. Optical-near-IR (NIR) photometry and spectroscopy are used to derive stellar masses, star formation rates (SFRs), dust extinctions, and metallicities. The average stellar mass is $10^{9.3} M_{\text{sun}}$, with a 1 σ dispersion of 0.8 dex. The average metallicity for a subsample of 17 hosts is about 1/6 solar and the dust

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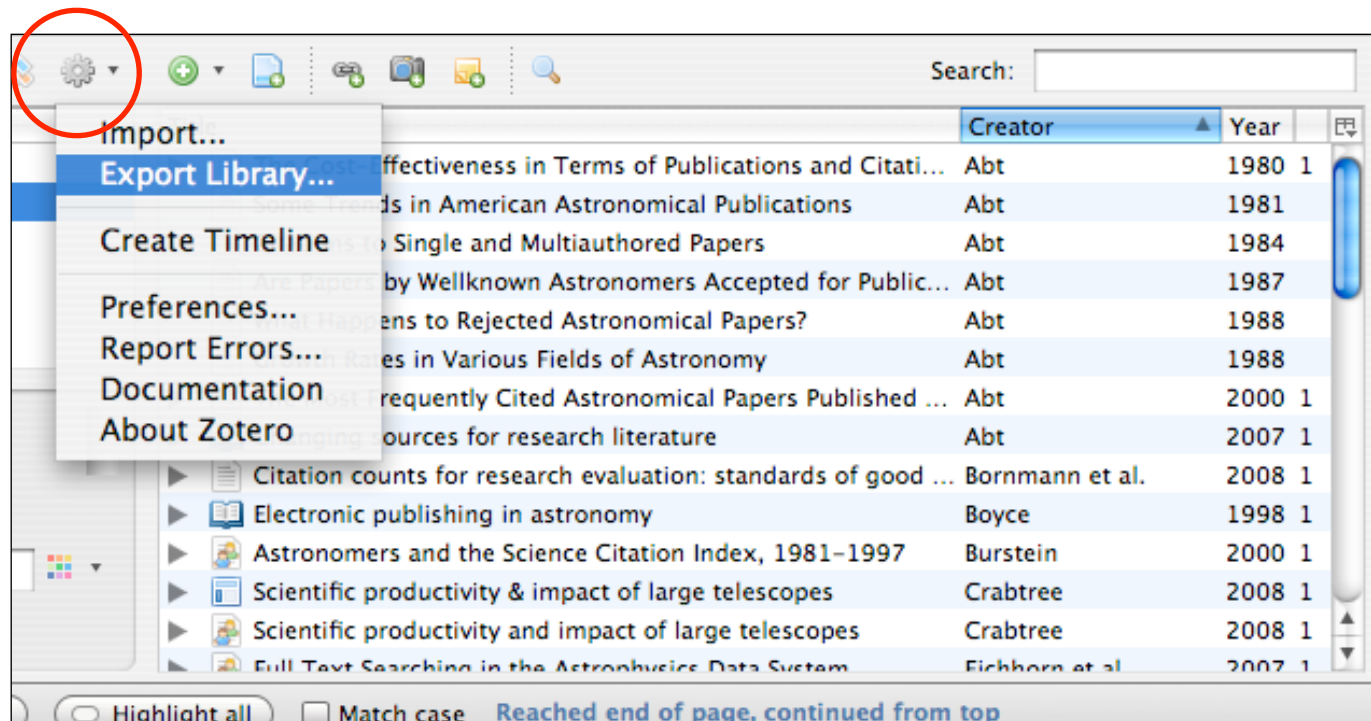
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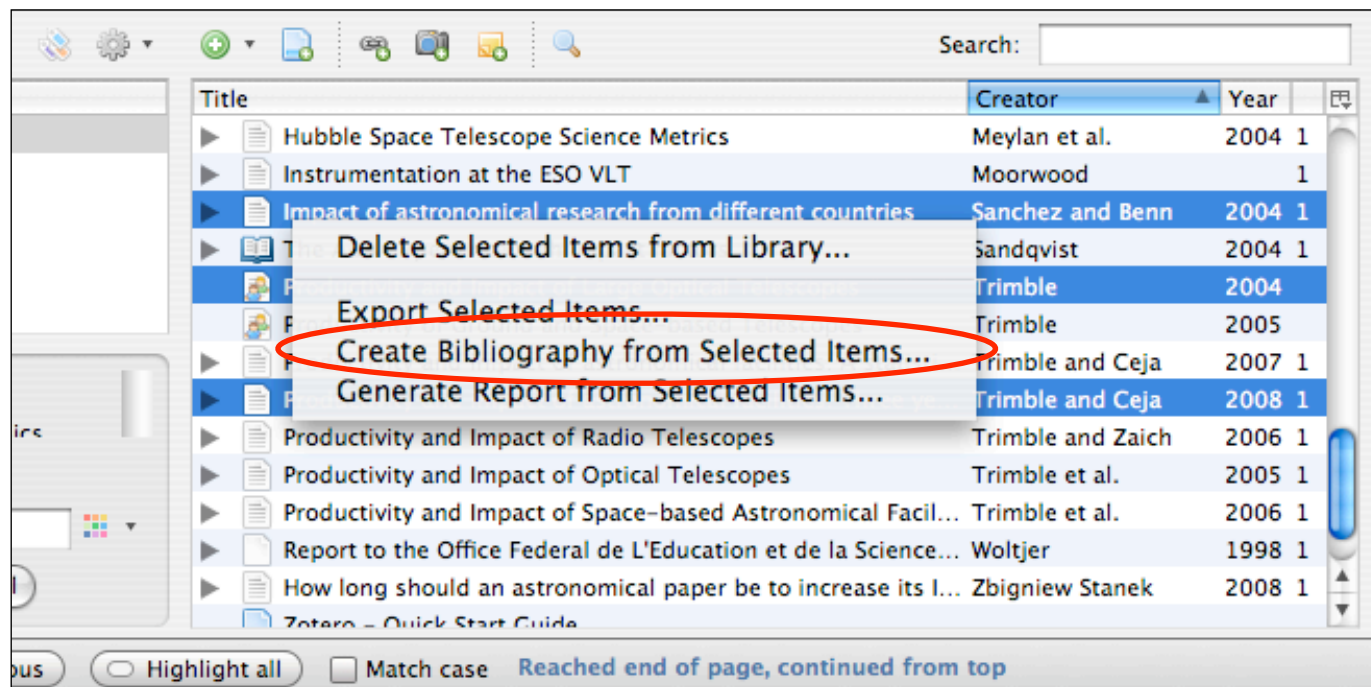
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For decades, distant galaxies that emitted their light six billion years ago were no more than small specks of light on the sky. With the launch of the Hubble Space Telescope in the early 1990s, astronomers were able to scrutinise the structure of distant galaxies in some detail for the first time. Under the superb skies of Paranal, the VLT's FLAMES/GIRAFFE spectrograph (ESO 13/02) — which obtains simultaneous spectra from small areas of extended objects — can now also resolve the motions of the gas in these distant galaxies (ESO 10/06).

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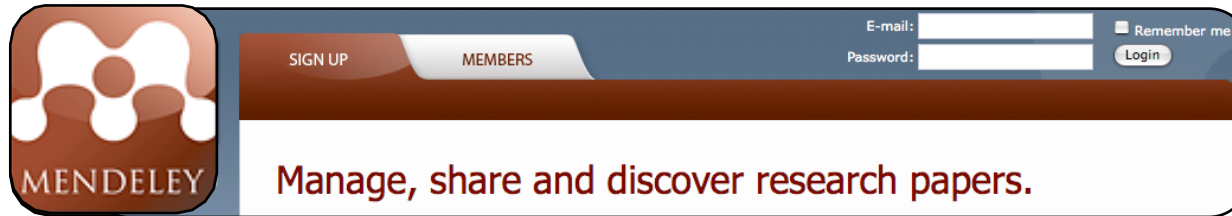
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Emotional Gestalts: Appraisal, Change, and the Dynamics of Affect

Paul Thagard
Philosophy Department
University of Waterloo

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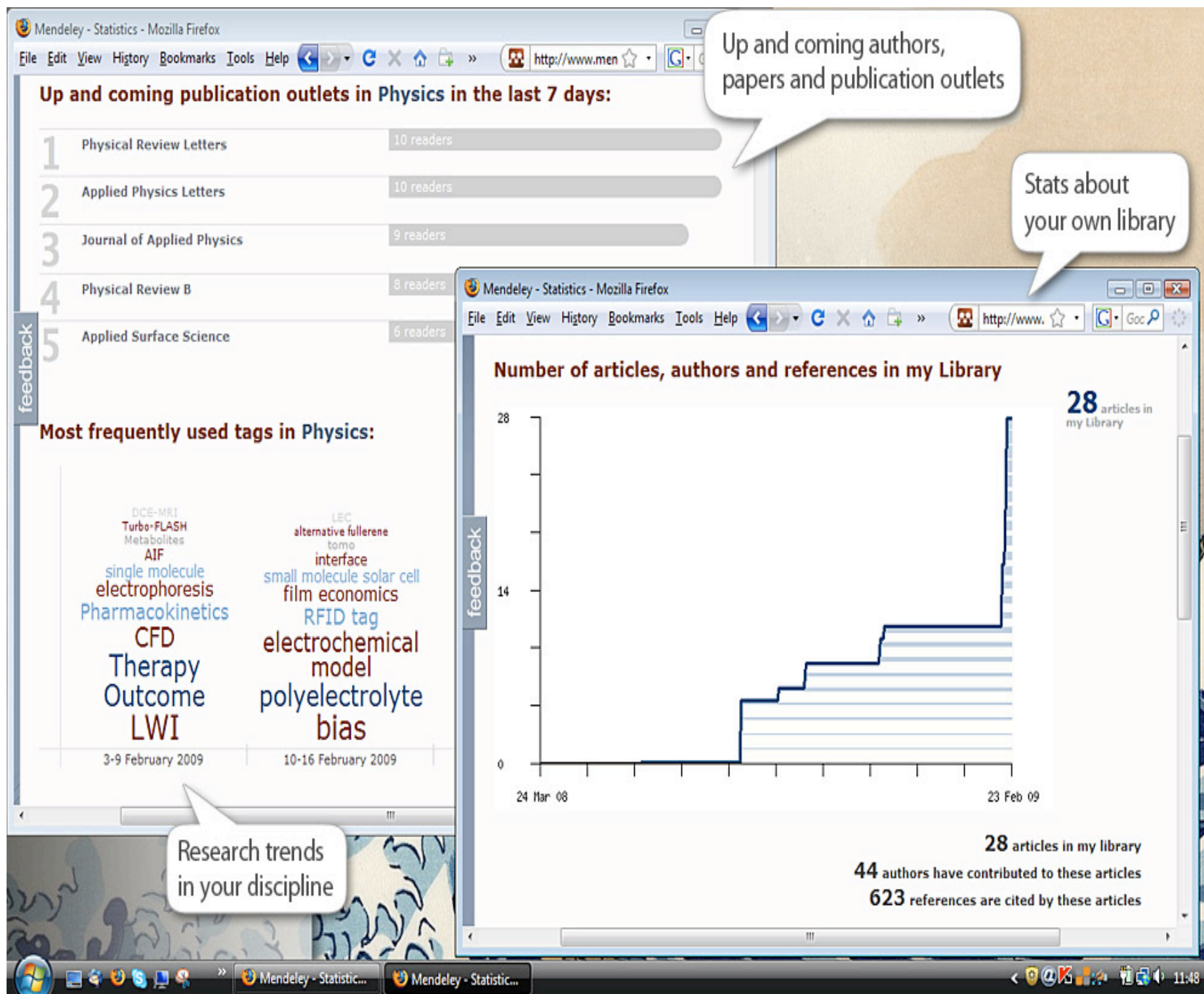
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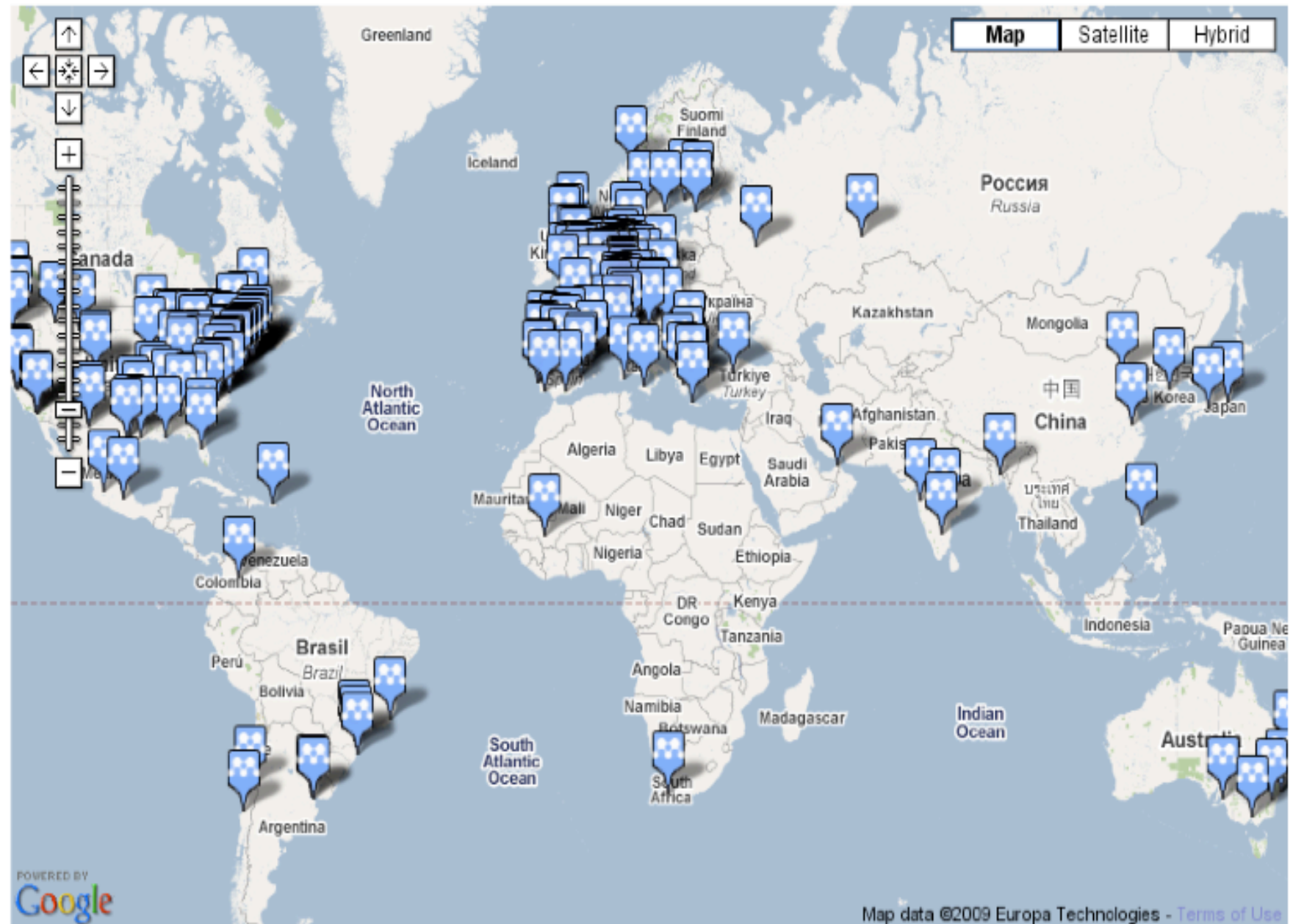
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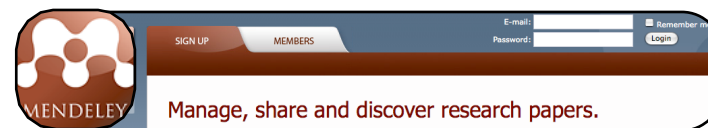
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