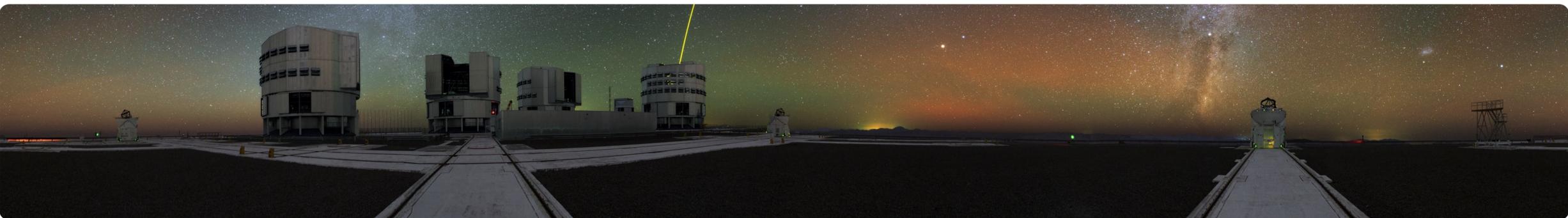


Metrics and quality control with adaptive optics instruments at the VLT



There are many **Adaptive Optics (AO)** fed instruments at the Paranal Observatory and many more to come. To monitor their performances and assess the quality of the scientific data, we are developing a scheme and a set of tools and metrics adapted to the various flavours of AO and variety of data products. In Service Mode (SM), our decisions to repeat observations or not depends heavily on this immediate quality control "zero" (QC0). Atmospheric parameters monitoring can also help predict performances. At the end of the chain, the user must be able to find the data that correspond to his/her needs translated into a set of requirements based on simulations done with an exposure time calculator (ETC). Predictions and real performances must match and the assessment must be intelligible to the community. We will emphasize on the difficulties encountered to perform quality control with SPHERE and the need for different metrics at various levels of wavefront: i.e Strehl ratio and FWHM for SCAO/LTAO/MCAO, contrast for an xAO, EE for GLAO, etc.

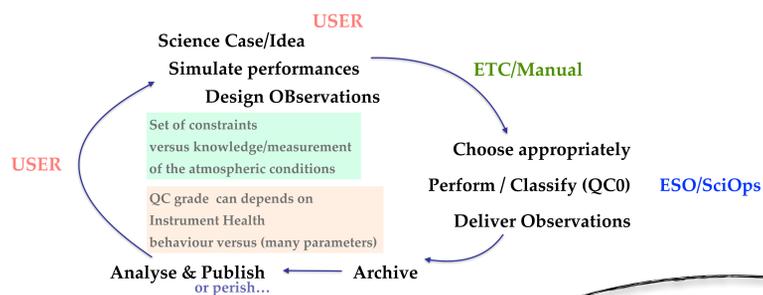
MANY AO FLAVOURS: IT'S NOW A ZOO!



SCAO, GLAO,
MCAO,
XAO, LTAO,
MOAO ...



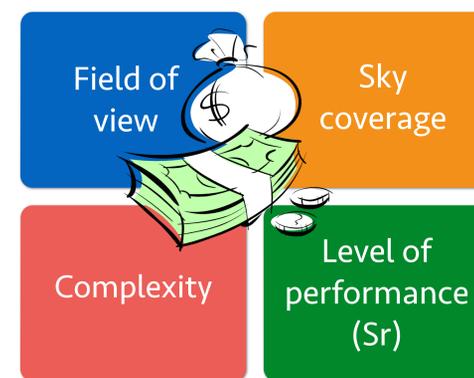
DATA PRODUCTION CHAIN: "SciOps" PERSPECTIVE



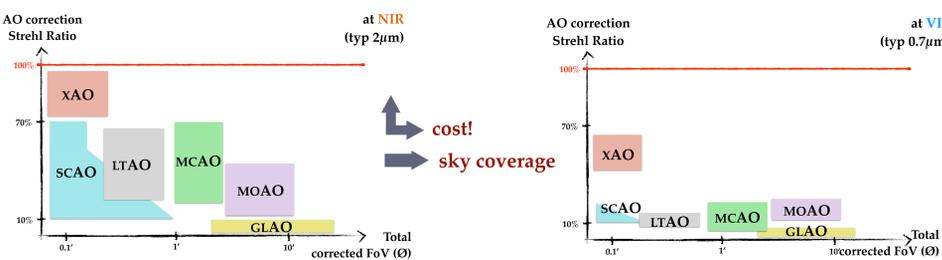
ADAPTIVE OPTICS EQUIPED INSTRUMENTS IN PARANAL

- NACO: 2001-2013 (UT4), 2015- (UT1)
- SINFONI/MACAO+LGS: 2003- VLT/UTs (MIDI, AMBER, PIONIER) + 4xMACAO: 2004 - MAD(emonstrator): 2008 (UT3, 1 month)
- CRILES(+): 2002-2014 (UT1) & UT3...
- SPHERE: 2014- (UT3)
- GRAVITY/CIAO: 2016- (VLT/UTs)
- MUSE/GALACSI: 2017- (UT4+DSM+4LGSF)
- HAWKI/GRAAL: 2017-(UT4+DSM+4LGSF)
- ERIS: ~2020
- MATISSE etc.

AO SYSTEMS TRADE-OFFS



AO PERFORMANCE VERSUS CORRECTED FIELD OF VIEW



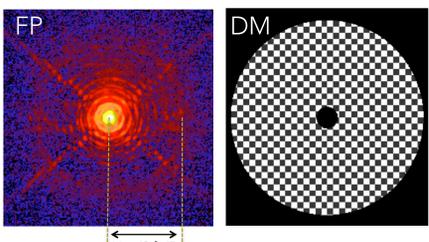
FOR EXTREME AO, THE CONTRAST IS THE BEST METRIC

High-density DM
High-density, faster WFS } 90% Strehl reached (H-band): almost the diffraction limit

$$\text{Contrast} \approx (1-Sr) / N_{act}^2$$

With $N_{act} \times N_{act}$ we can correct up to $N_{act}/2$ cycles / pup

Highest possible spatial frequency

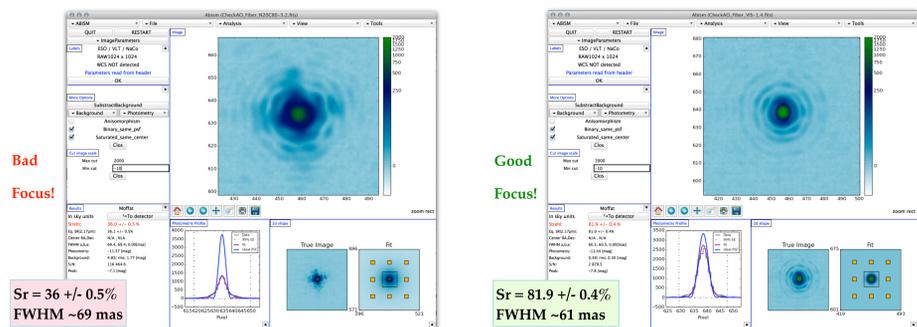


FOR GLAO, THE ENCIRCLED ENERGY or Energy concentration improvement in a given box

The Sr would be ~1%, insensitive



WHEN THE STREHL RATIO IS A GOOD METRIC AND THE FWHM ISN'T



Sr here is more sensitive than FWHM when $10\% < Sr < 90\%$

COMMON MODULAR STRATEGY

IMPOSSIBLE to use the same metric for all AO-fed instruments, modes !!!!

DIFFICULT to understand, measure => data mining, correlations

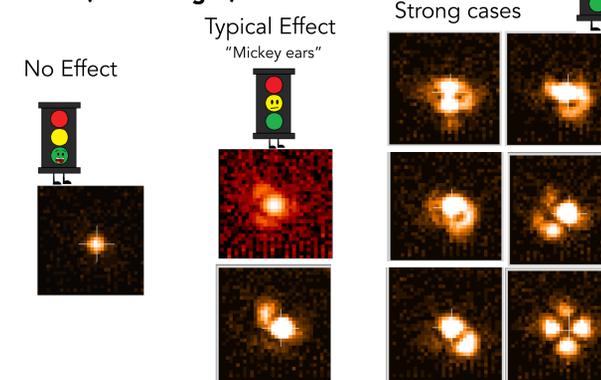
NECESSARY to develop a common but versatile QC / QC0 strategy / PHILOSOPHY

=> Analysis Tools
=> Trustworthy ASM 2.0
=> Training!



LOW WIND EFFECT (LWE) AS SEEN BY SPHERE

SPHERE PSF (DTS images)



The LWE (~15% of the time) makes the QC0 assessment based on external conditions impossible

RELATED CONTRIBUTIONS AT THIS CONFERENCE

- ABISM: an interactive image quality and Strehl ratio measurement tool for adaptive optics instruments (Girard et al. 2016, 9909-303, June 30th 2016 (poster))
- Versatile quality control scheme for the adaptive optics instruments at the VLT (Girard et al. 2016, 9910-108, June 29th 2016 (poster))
- SPHERE on-sky performance compared with budget predictions (Dohlen et al. 2016, 9910-108, June 29th 2016 (poster))
- SPHERE: on-sky results (Beuzit et al. 2016, 9910-108, June 29th 2016)
- SPHERE on-sky results: final performance, lesson learned, and possible upgrades (Fusco et al. 2016, 9910-108, June 29th 2016 (Invited talk))
- Data flow operations and quality control of SPHERE (Fusco et al. 2016, 9910-108, June 29th 2016 (Invited talk))
- Training telescope operators and support astronomers at Paranal (Boffin et al. 2016, 9910-108, June 29th 2016)