

DETECTING EXO-SOLAR SYSTEMS

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Imaging system

- ELT
- High Strehl Adaptive Optics
- Rotation shearing coronagraph
- Imaging & Differential imaging

Limiting noises

- Speckle noise
- Star photon noise
- Planet photon noise
- Readout noise

Speckle noise

- Without adaptative system: pure speckle noise ($\sigma^2 = I^2$, Dainty – 1974)
- With partial adaptative correction (Canales & Gagigal, 1999)
- Experimental approach (Racine et al.1999)
- Non stationnary phase structure fonction with spatial & spectral correlations (Chelli & Tatulli, in preparation)

Signal to noise ratio

$$SNR \approx \frac{f \bar{K} \Delta \lambda S Q_1}{\sqrt{4 \bar{K}^2 \Delta \lambda^2 \frac{\tau}{T} (1 - S)^2 g^2 \left(\frac{r_0}{D}\right)^4 \left(\frac{4}{\pi}\right)^2 L^2(\gamma, \bar{\lambda}) Q_2^2 + 2 \bar{K} (1 - S) g \Delta \lambda c^2 \left(\frac{r_0}{D}\right)^2 L(\gamma, \bar{\lambda}) + f \bar{K} \Delta \lambda S Q_1 + \sigma^2}}$$

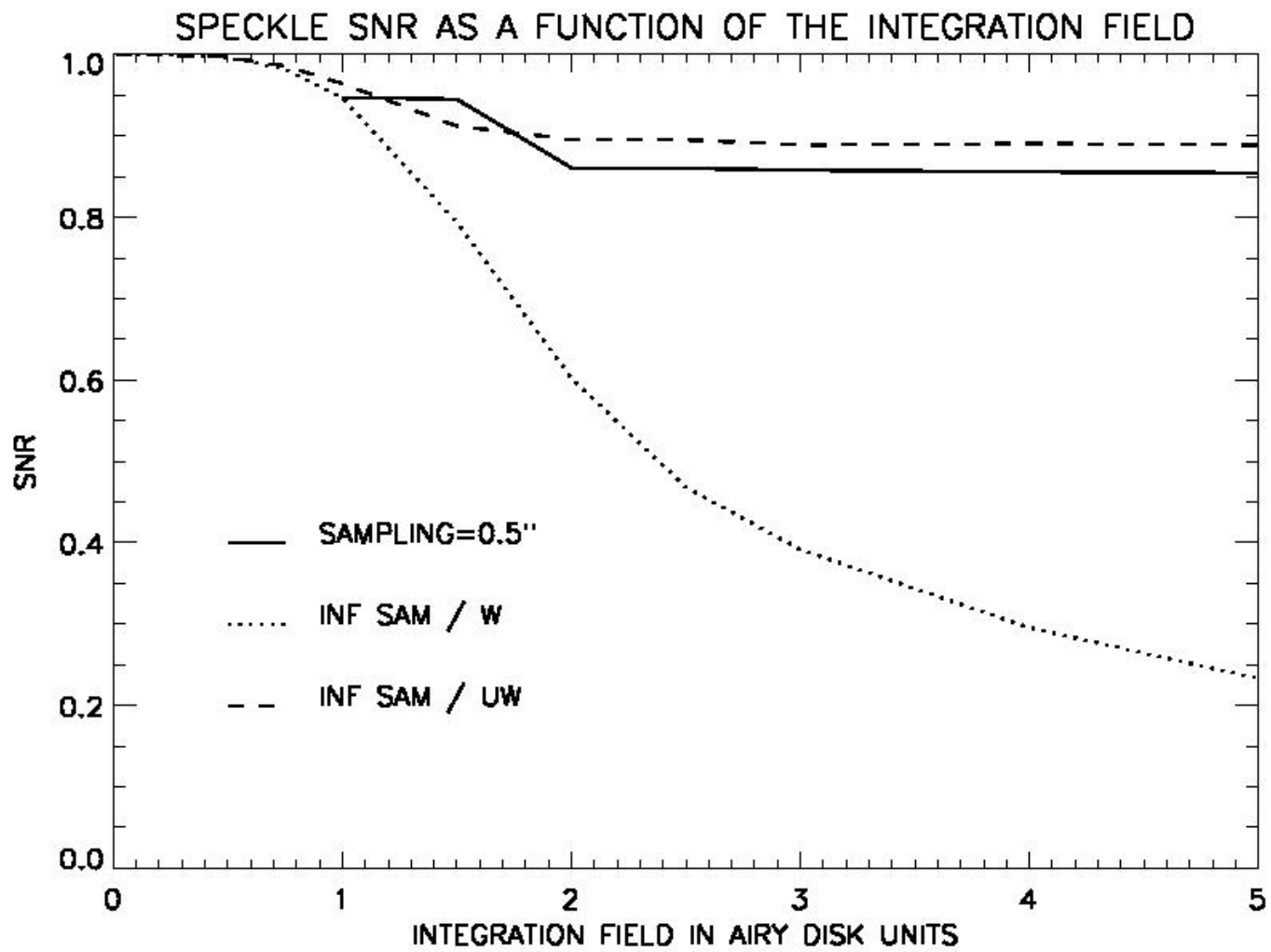
$$SNR \approx \frac{f}{2} \frac{S}{1 - S} \left(\frac{D}{r_0}\right)^2 \frac{1}{L(\gamma, \bar{\lambda})} \frac{Q_1}{Q_2} \sqrt{\frac{T}{\tau}}$$

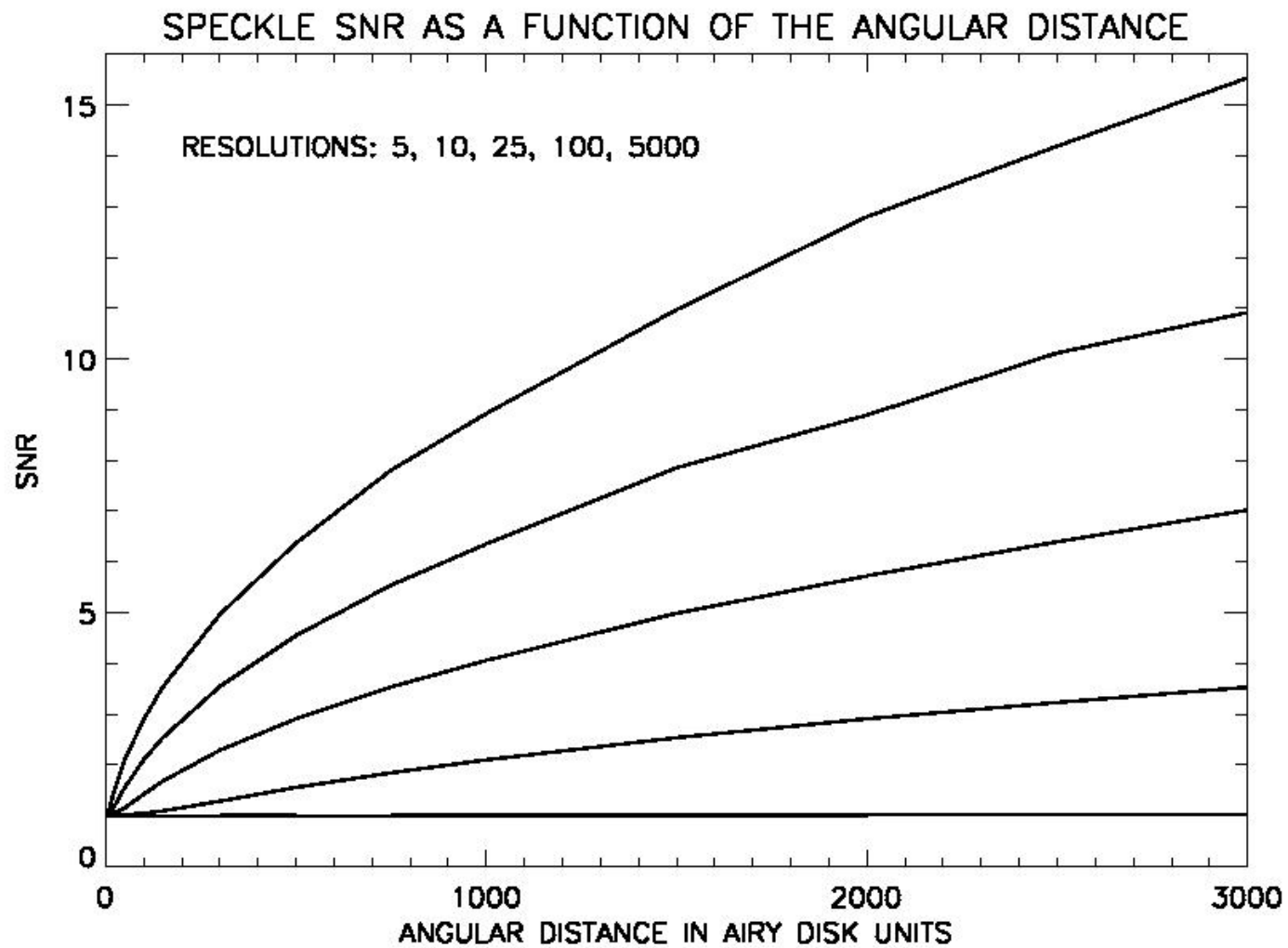
Differential imaging

- Scaling

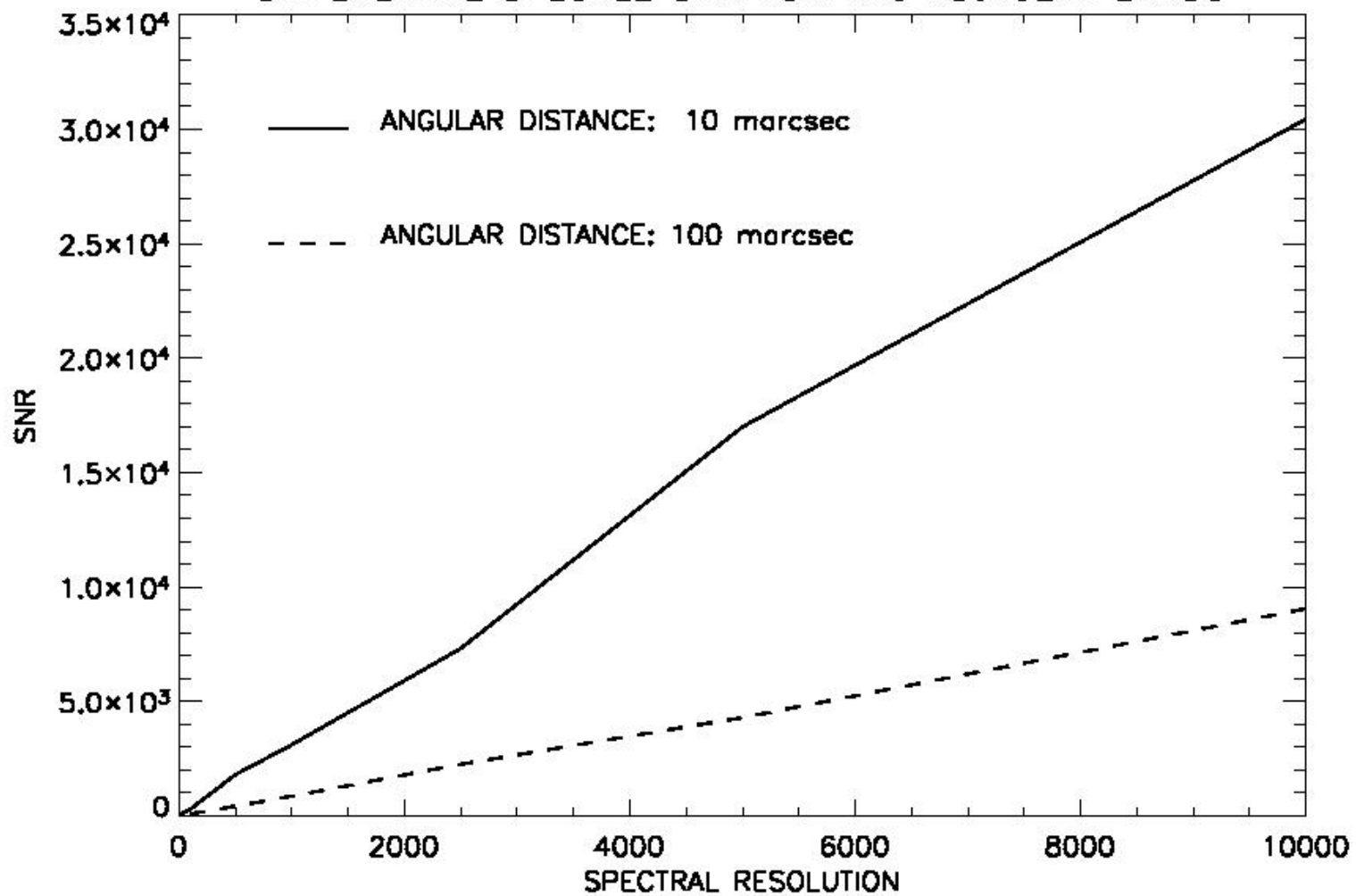
- If $\beta\Delta\lambda/\lambda > 1$ then the 2 contributions of the planet are fully separated

- Weighting

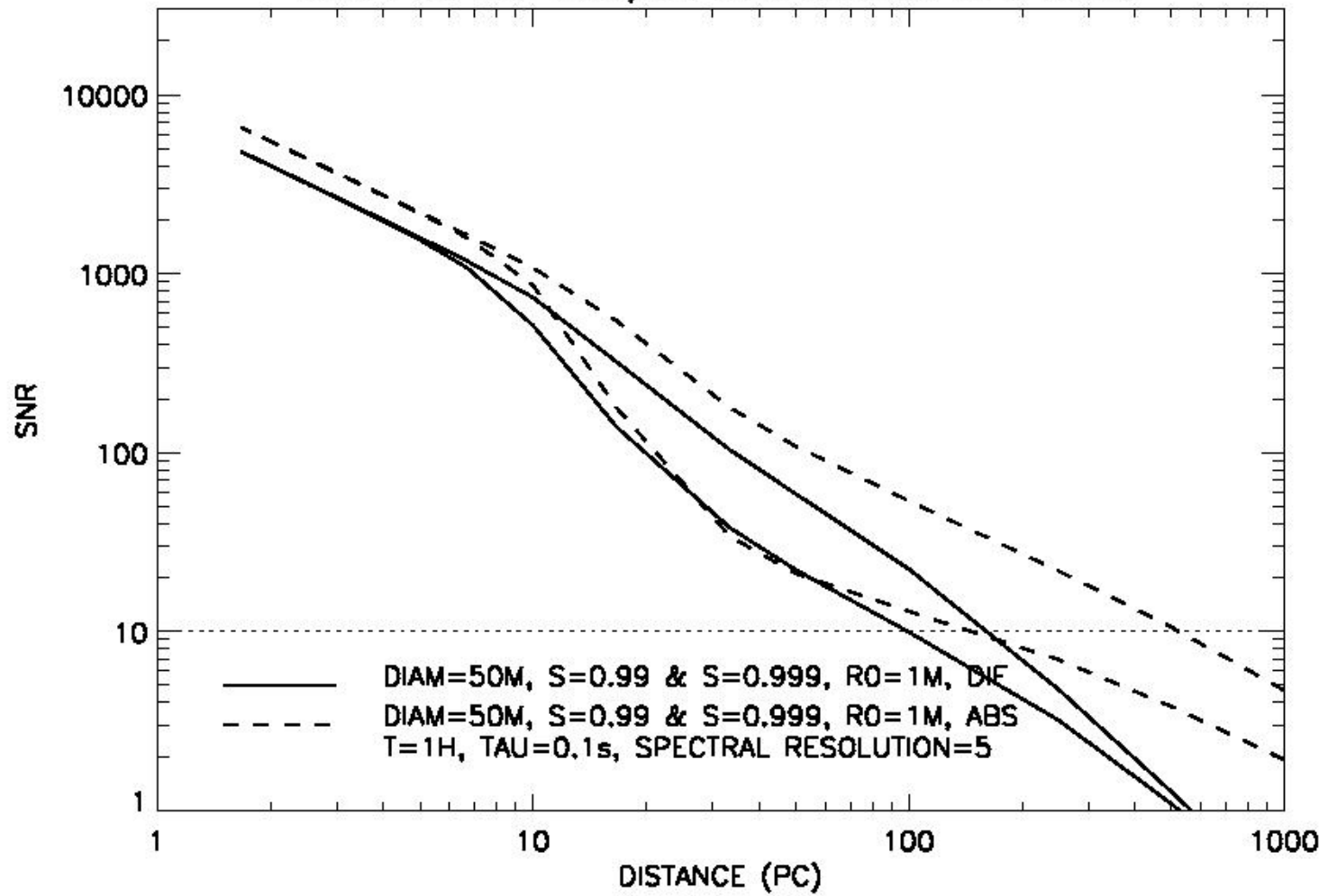




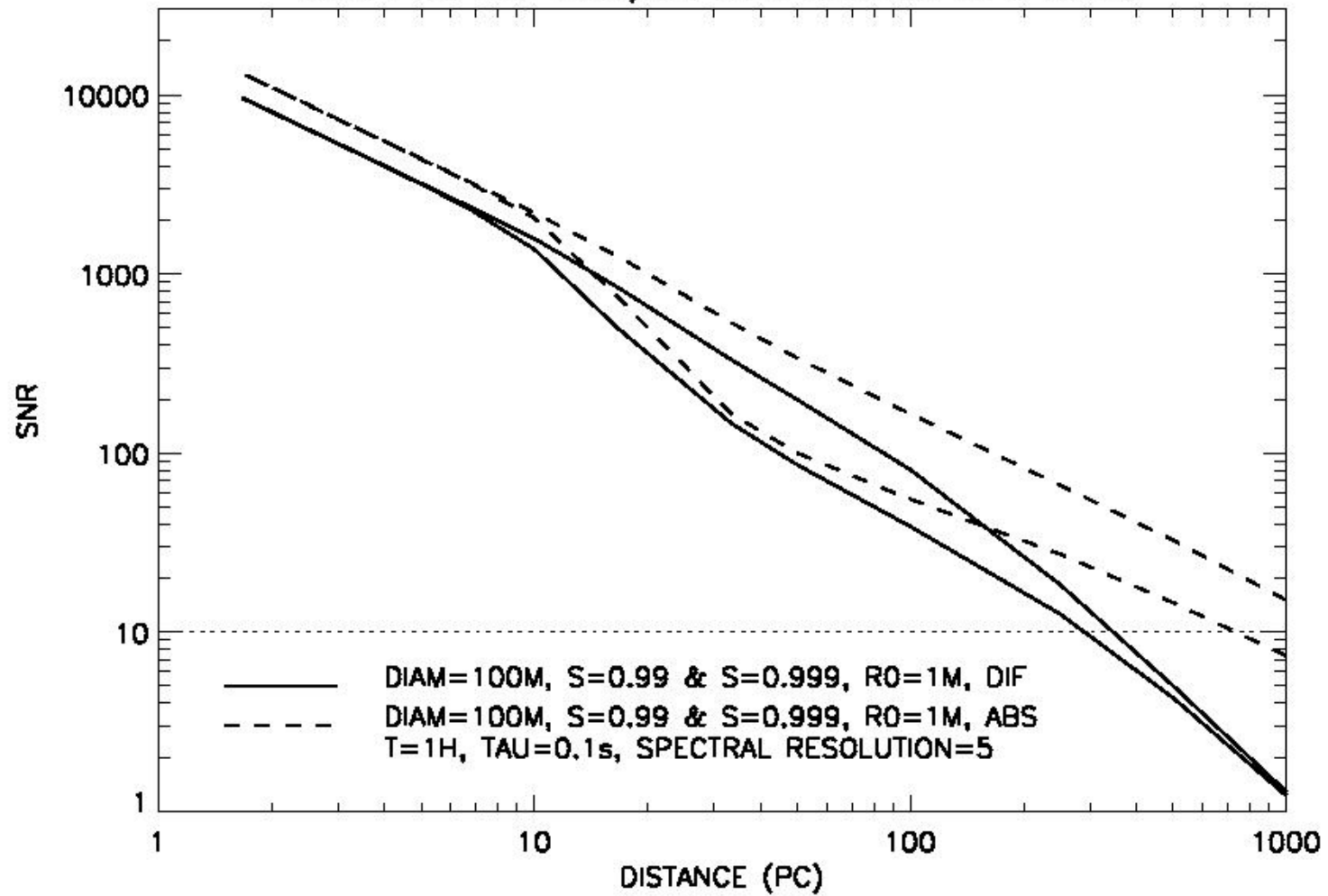
DIFFERENTIAL SPECKLE SNR FOR TWO ADJACENT BANDS



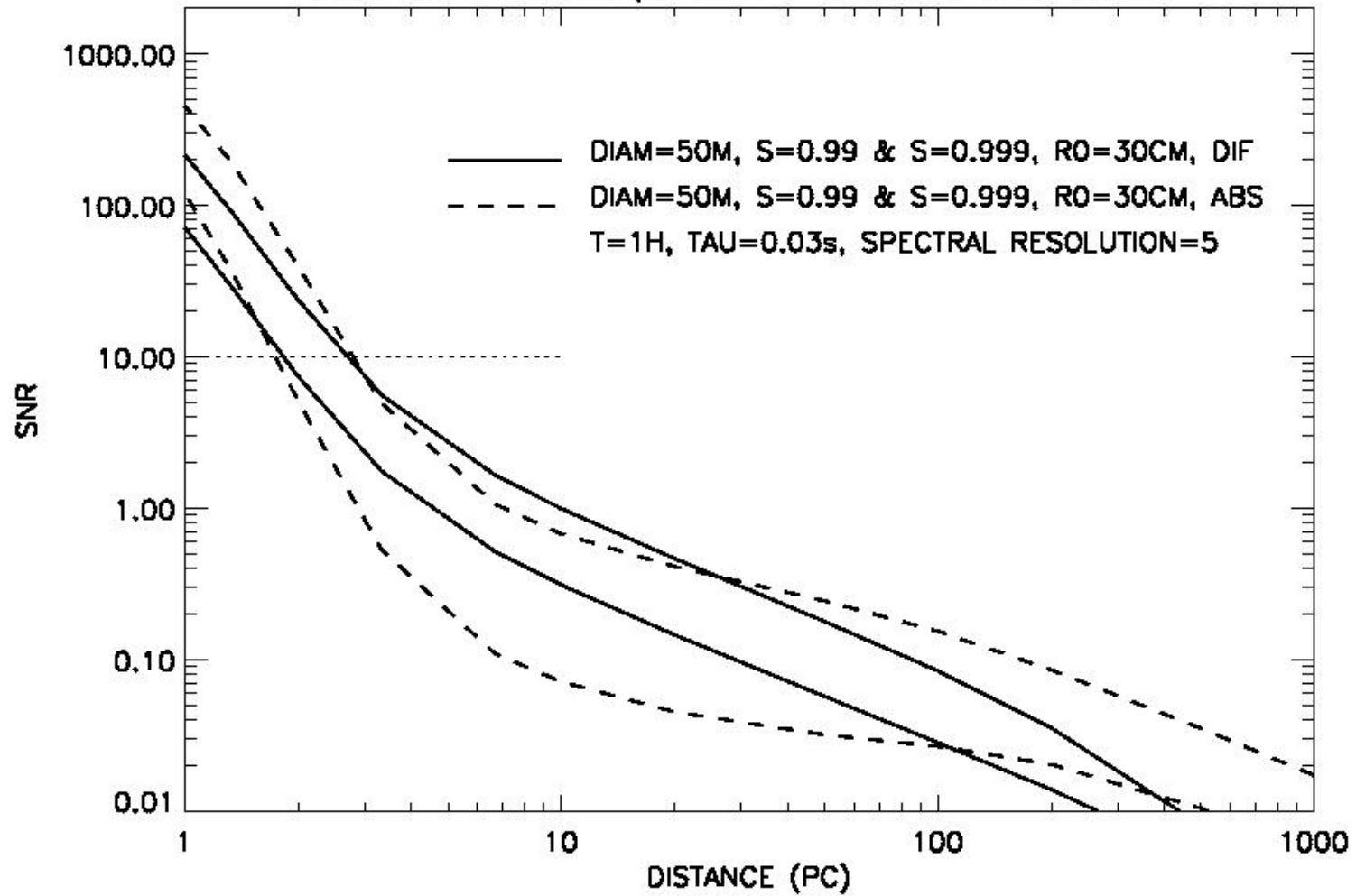
SNR FOR THE SUN/JUPITER SYSTEM IN H BAND



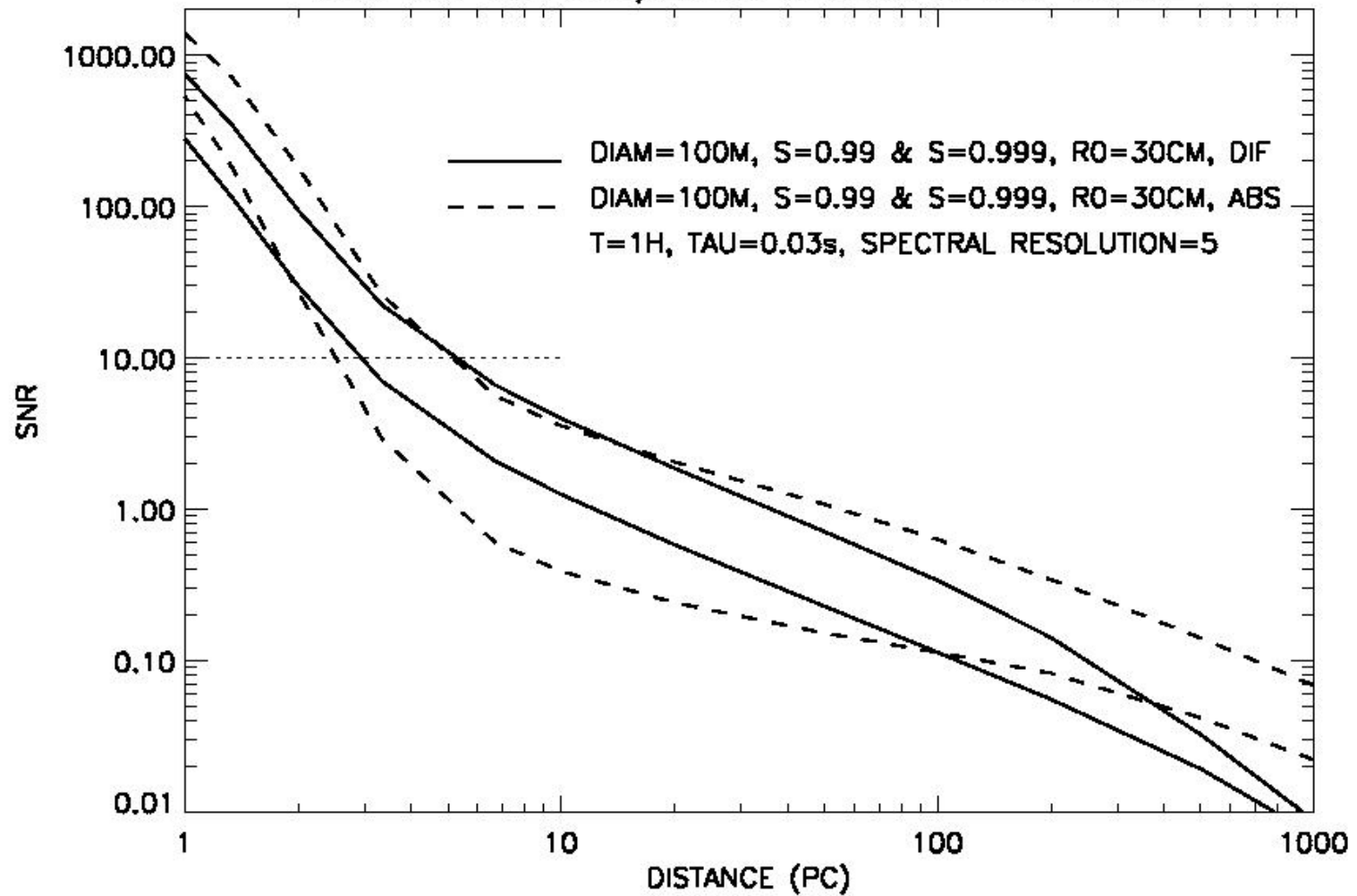
SNR FOR THE SUN/JUPITER SYSTEM IN H BAND



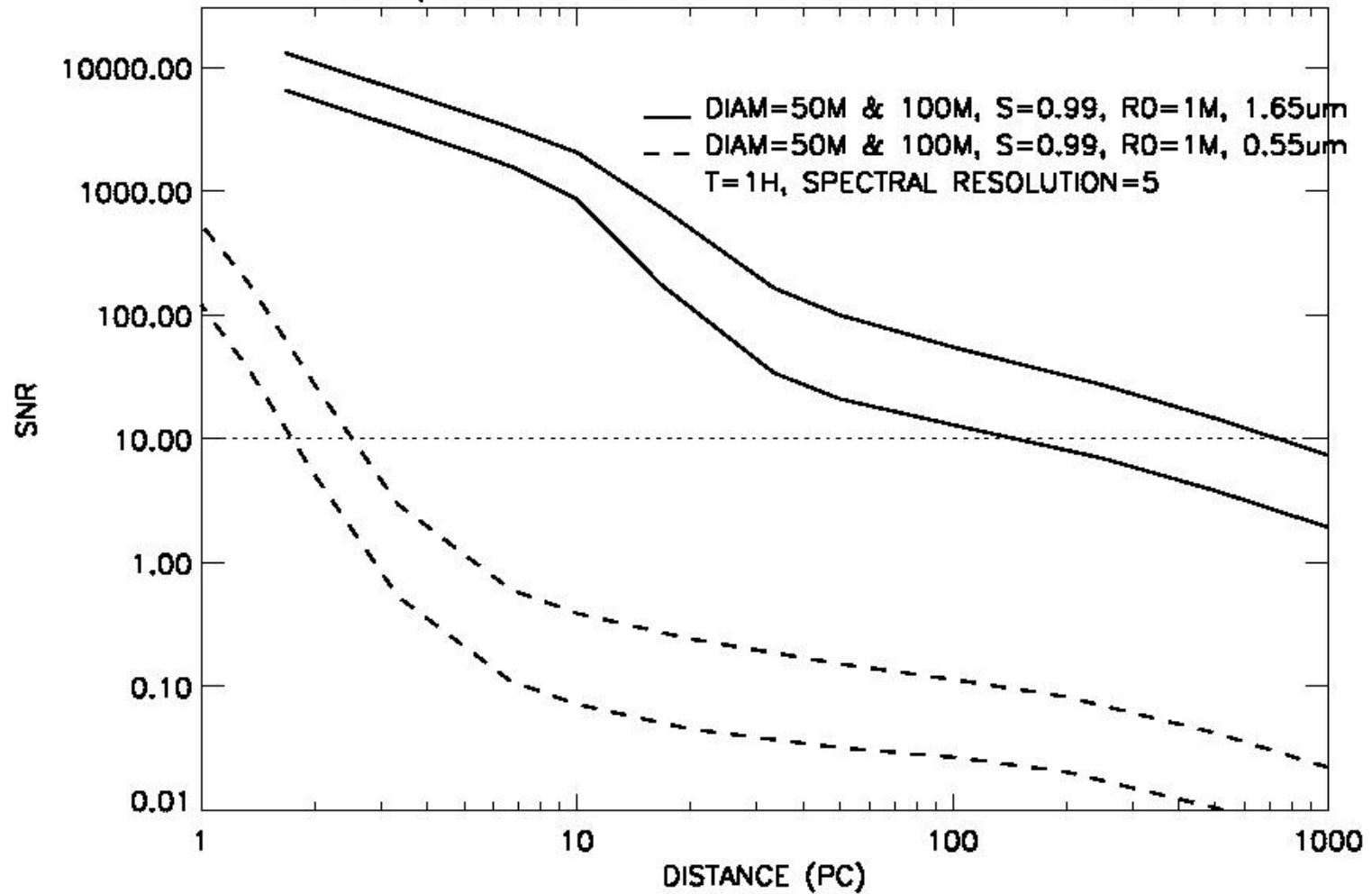
SNR FOR THE SUN/EARTH SYSTEM IN THE VISIBLE



SNR FOR THE SUN/EARTH SYSTEM IN THE VISIBLE



SUN/JUPITER & EARTH SYSTEM IN H & V BAND



Conclusions

- Rigorous & complete treatment of speckle noise
- Science cases:
 - Absolute magnitude
 - Flux ratio
 - Separation
- SNR on absolute & differential detection