



**OWL Phase A Review - Garching - 2<sup>nd</sup> to 4<sup>th</sup> Nov 2005**

# **Non-adaptive Wavefront Control**

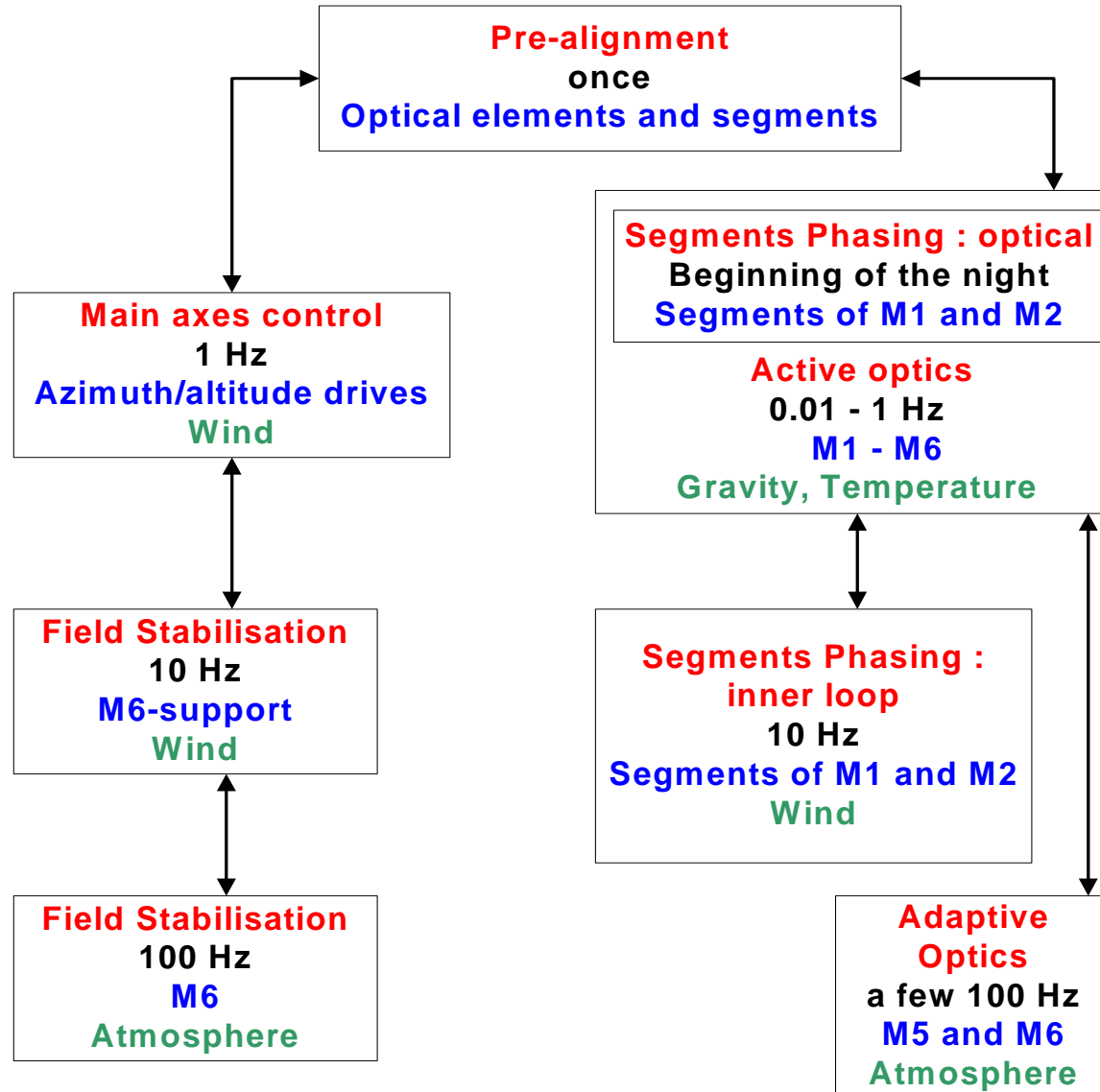
**(Presented by L. Noethe)**



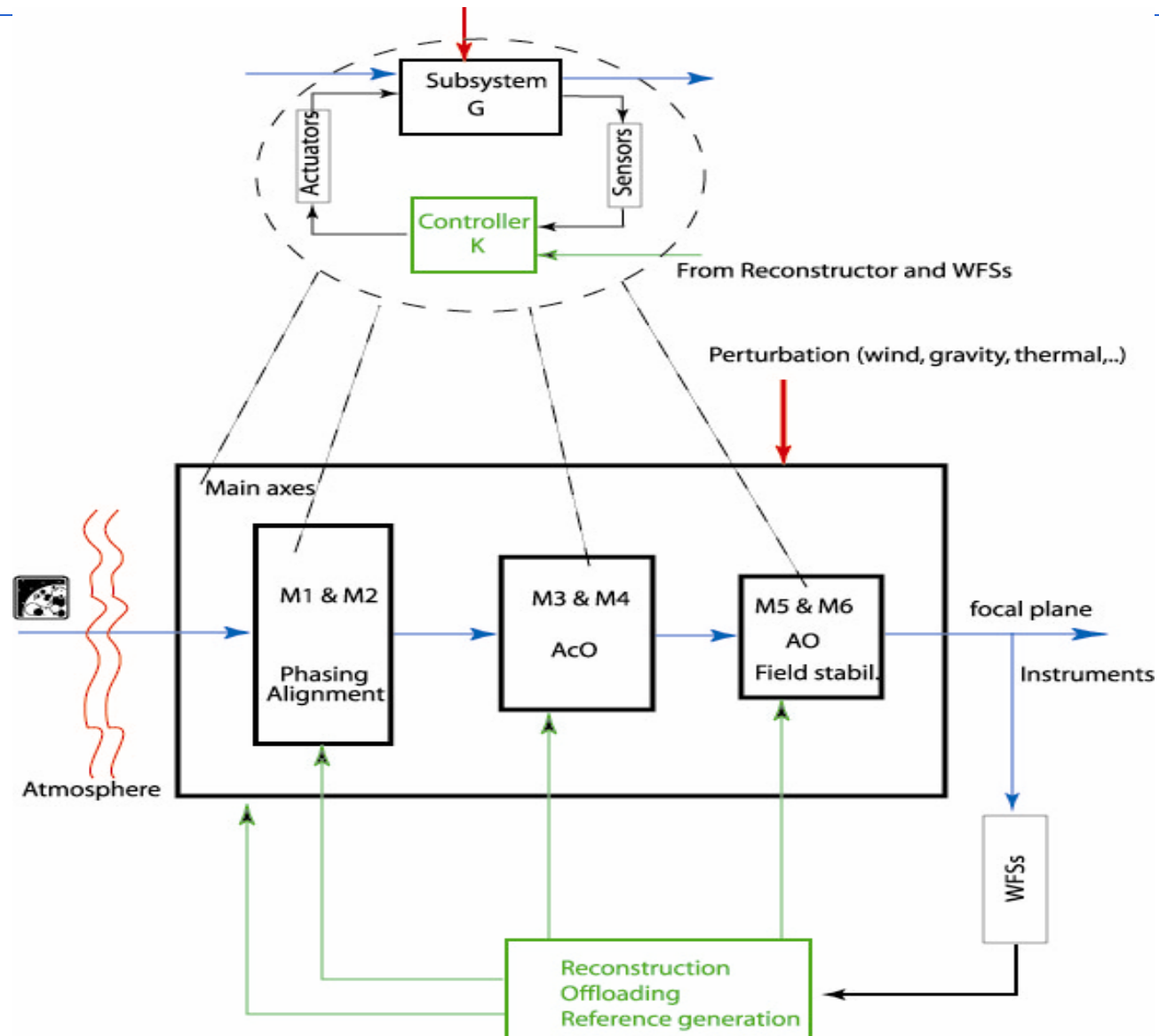
# Specific problems in ELTs and OWL

- Concentrate on problems which are specific for ELTs and, in particular, for OWL
- Alignment and shape correction of optical surfaces
  - Large number of segments in segmented mirrors
  - Six optical surfaces
  - Two segmented mirrors
  - Requires mainly further development of already existing active optics techniques
- Operation in open air
  - Advantages: thermal equilibrium and predictable wind loads
  - Disadvantage: larger wind loads
  - Feasible with extensive use of fast control loops

# Overview : Wavefront Control



# Control Architecture



# Pre-alignment

## ■ Alignment of segments

- Use of a spherometer to align a new segment relative to its neighbours
- Stacking of the images produced by individual segments

## ■ Alignment of optical elements

- Use of a fibre extensometer to be developed within the FP6 ELT study

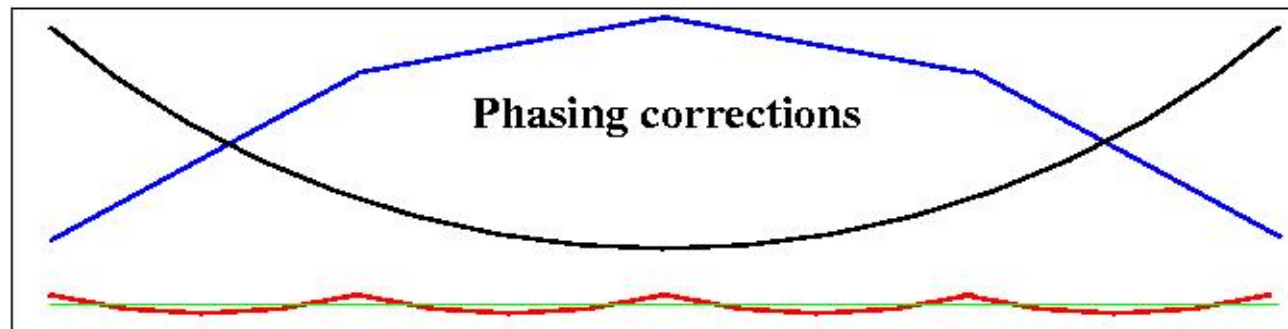
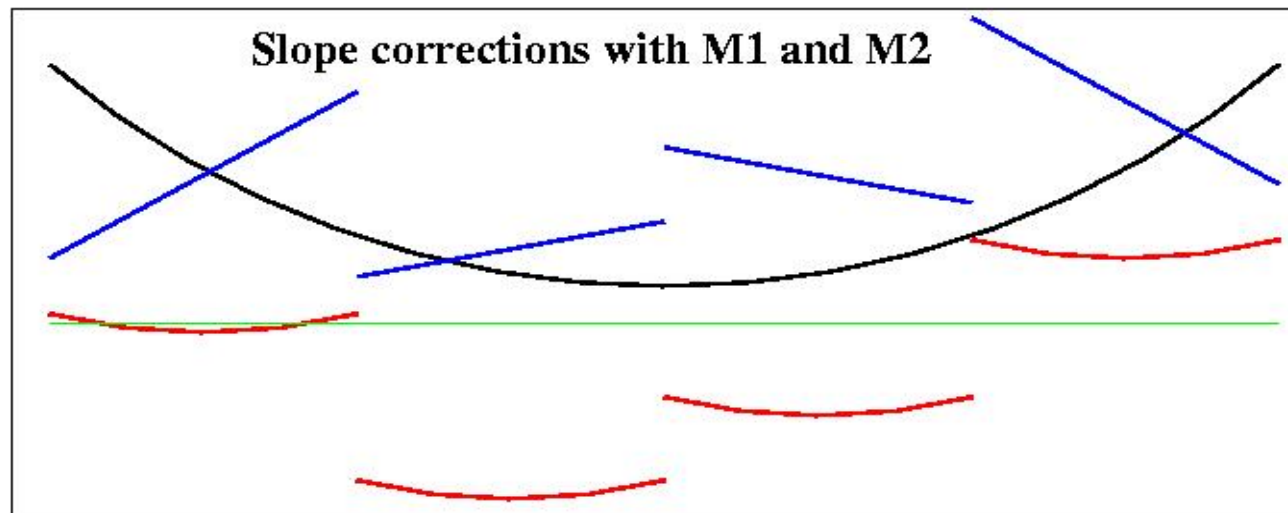
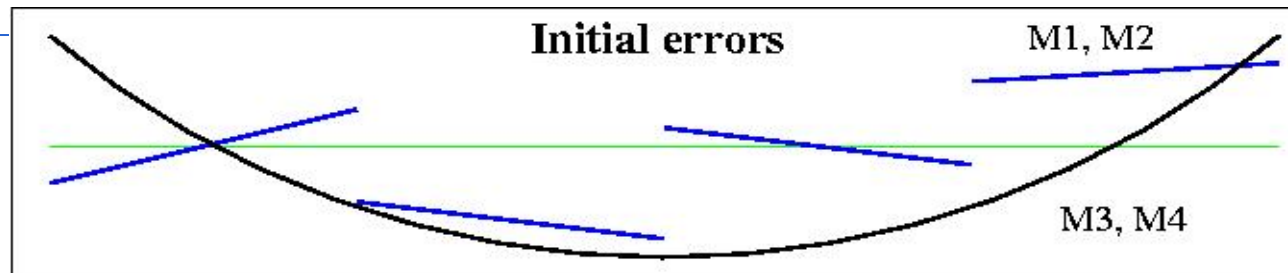
## ■ Residual errors after pre-alignment

- Positions M1 – M6 : ~1 mm
- Tilts M1 - M6 : ~ 1 arcsec
- Piston errors of segments : ~ 2  $\mu\text{m}$
- Deformations M3 and M4 : ~ 30  $\mu\text{m}$

# Correction strategy

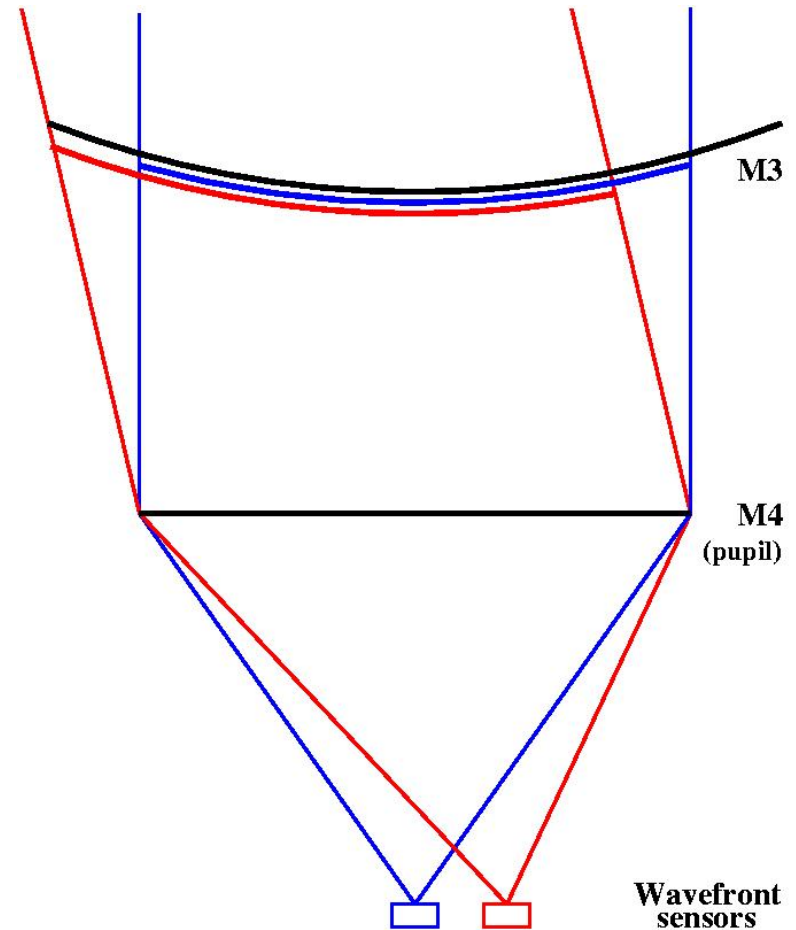
- Complete correction in one step
  - Measurements from several Shack-Hartmann sensors and one phasing sensor
  - Inversion of the matrix relating the actuator degrees of freedom to the measured parameters by singular value decomposition
  - Calculation of the actuator commands from the measured signals with the inverted matrix
  - Disadvantage : requires a very large matrix
- Alternative approach : split the correction into several steps
  - More than one possible strategy
- One sequence of correction steps
  - Correction of slope errors with the segments
  - Phasing of the segments
  - Correction of misalignments and deformations of M1 to M6

# Active optics corrections with one Shack-Hartmann and one phasing wavefront sensors



# Full correction with wavefront sensors in several field positions

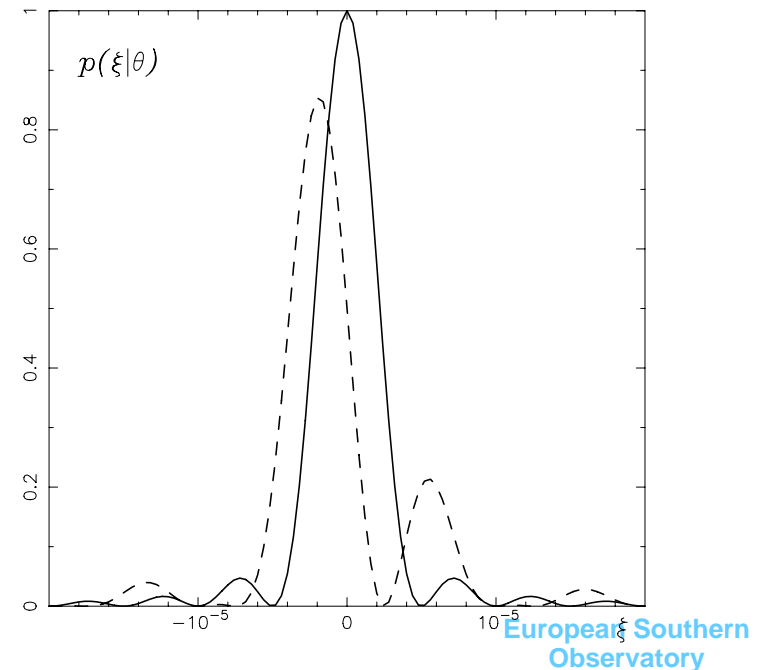
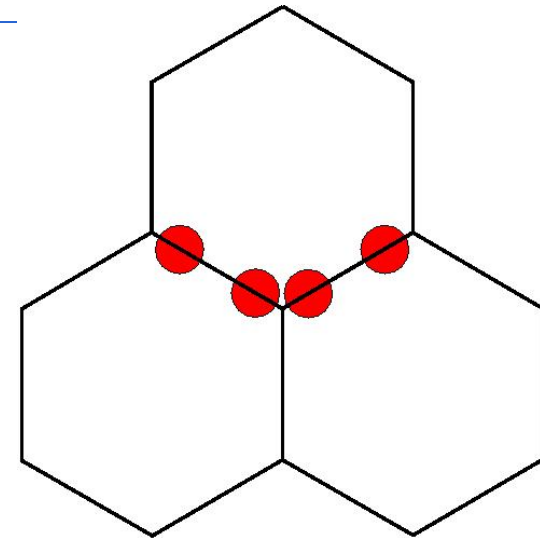
- Aberrations generated by
  - Misalignments of the mirrors
  - Deformations of the meniscus mirrors
  - Characteristic patterns of additional field aberrations
  - Correction with an in-pupil mirror only possible for one field angle
- Required wavefront sensors
  - 1 Shack-Hartmann sensor with 19 lenslets per M1 segment
  - At most 6 Shack-Hartmann sensors with 20 by 20 subapertures covering M1
  - 1 (baseline) or 2 (optional) phasing sensors





# Phasing wavefront sensors

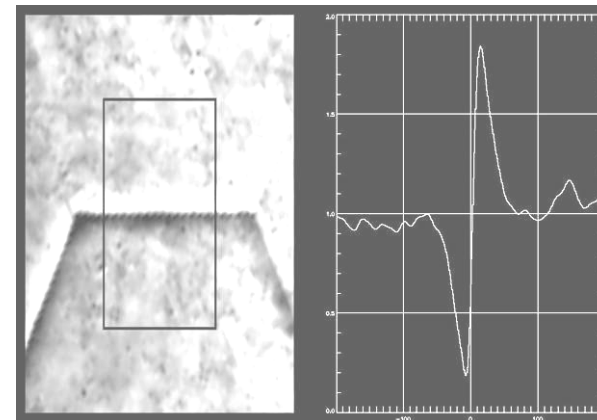
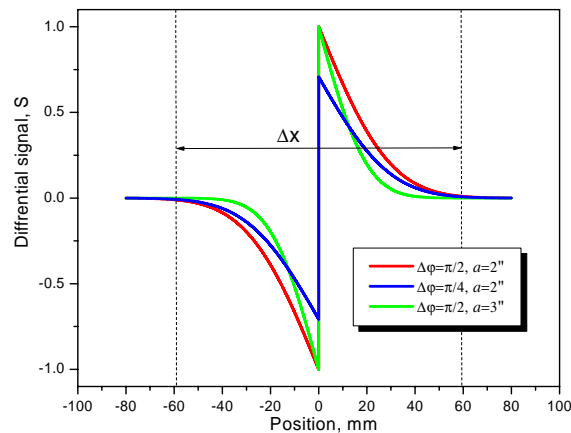
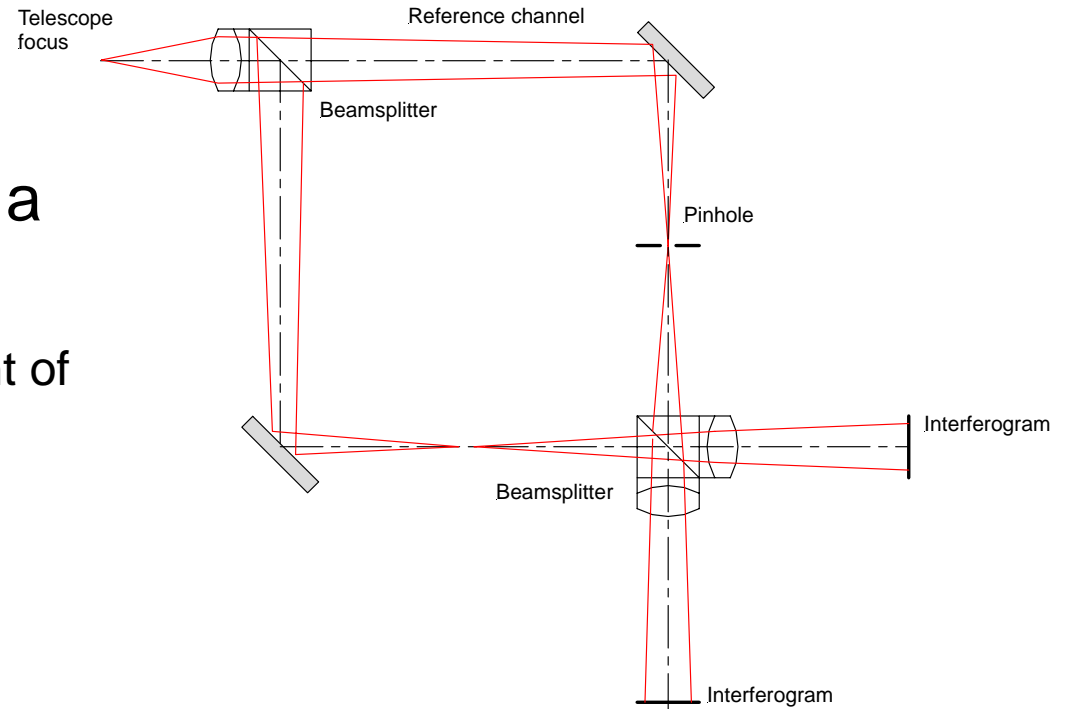
- Multi-wavelengths techniques
  - Reduce the wavefront piston steps of  $2\ \mu\text{m}$  to less than  $100\ \text{nm}$
- Narrowband techniques
  - Shack-Hartmann: lenslets covering subapertures centered on segment borders
  - Information contained in the position of the maximum and in the shape of the diffraction pattern
  - Applied in the Keck Telescope extracting the shape information
  - Problem : exact positioning of a large lenslet array in the re-imaged pupil



# Phasing wavefront sensors: Mach-Zehnder

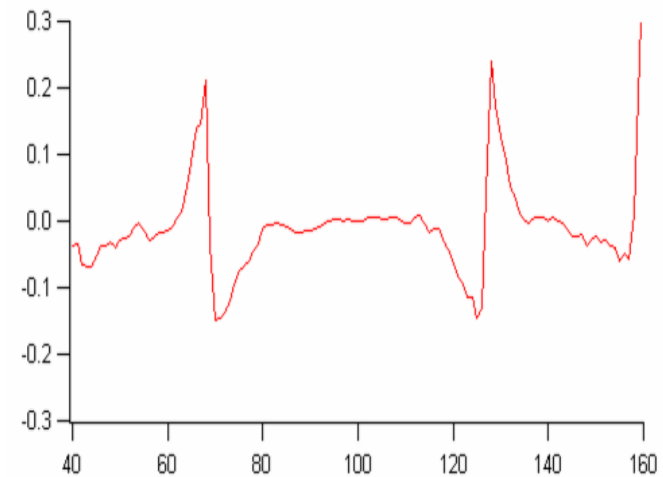
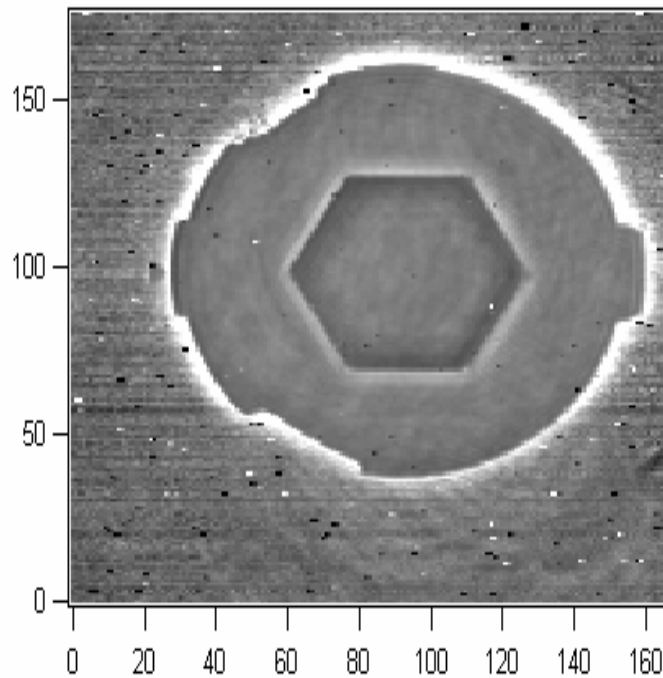
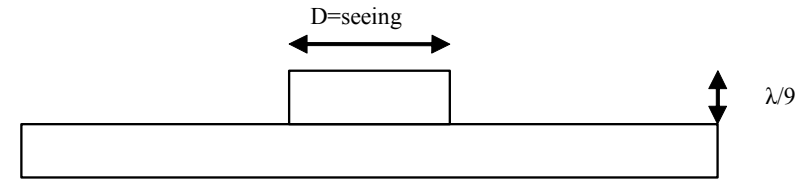
## ■ Spatial filtering in a focal plane

- Problem : Alignment of the optics

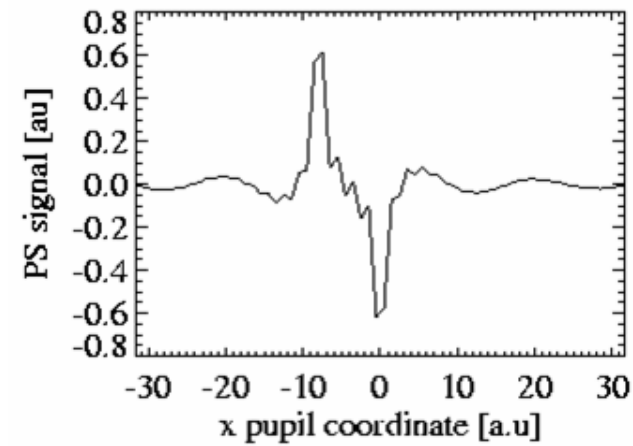
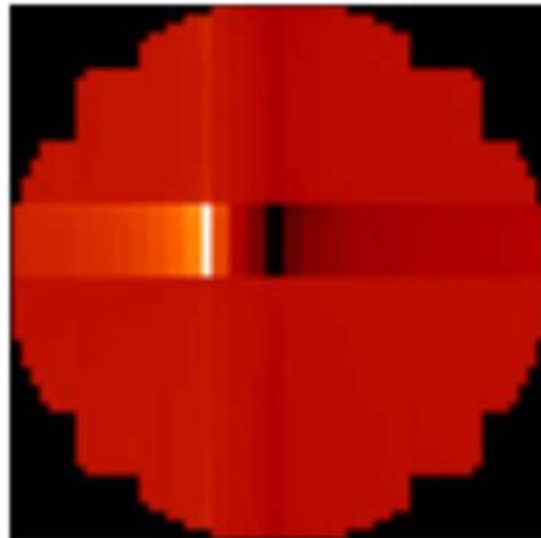
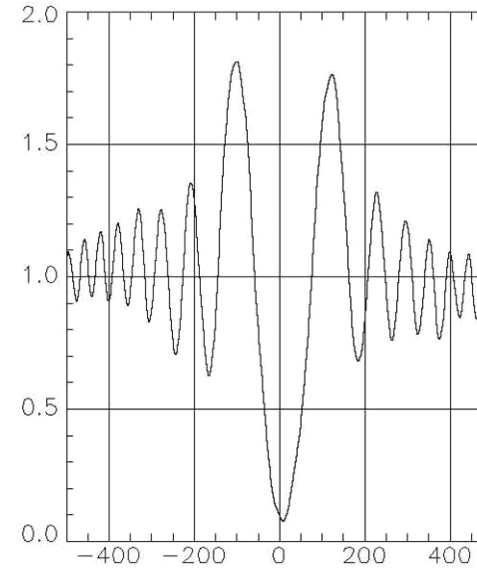
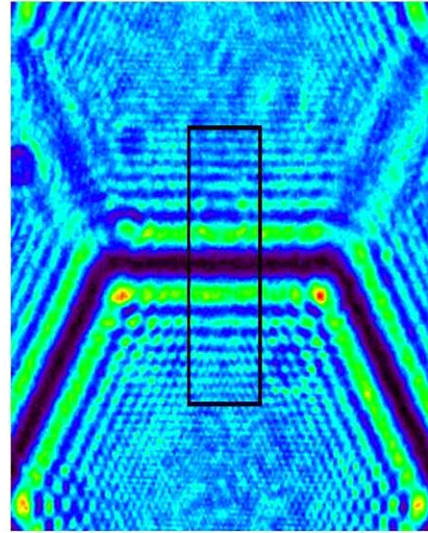


# Phasing wavefront sensors – phase filtering (LAM/IAC)

- Adding of a phase delay in the center of the image
  - Easier to align than the Mach-Zehnder sensor

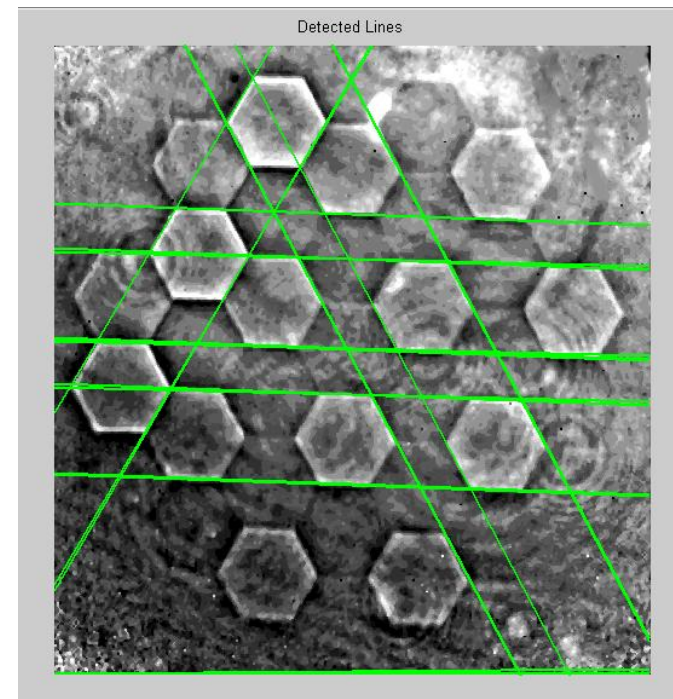
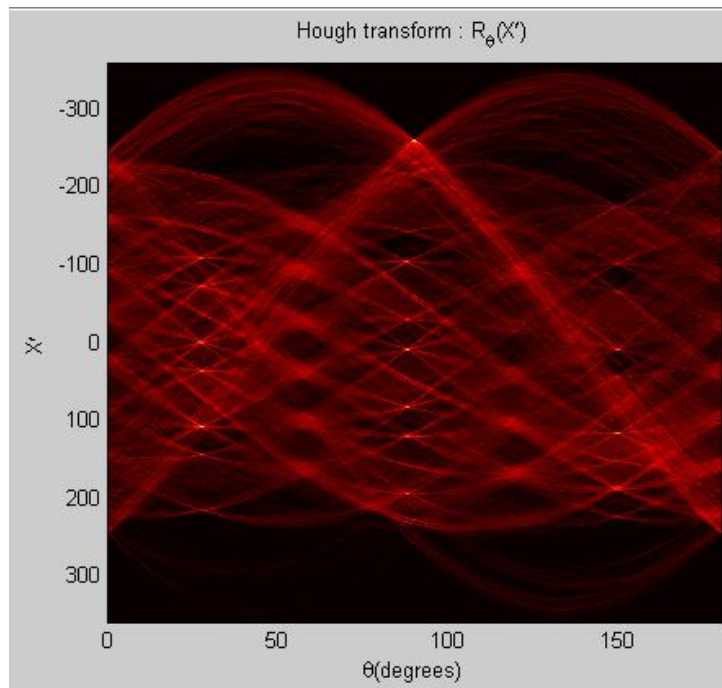


# Phasing wavefront sensors – defocusing (IAC) and pyramid sensor (Arcetri)



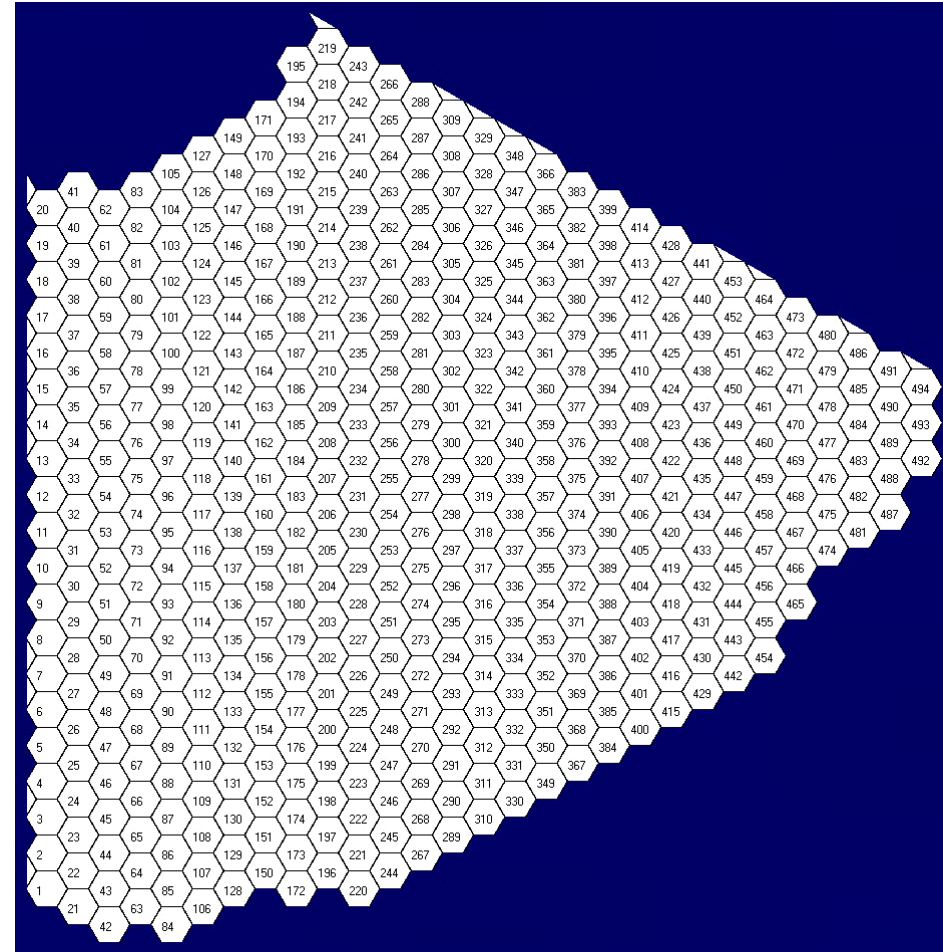
# Phasing wavefront sensors – identification of borders

- Possible algorithm : contrast enhancement and Hough transform
  - Promising for large piston steps
  - To be validated for small piston steps
- Imaging of the gaps
  - 10% reduction of the intensity for pixels covering gaps



# Phasing of petals

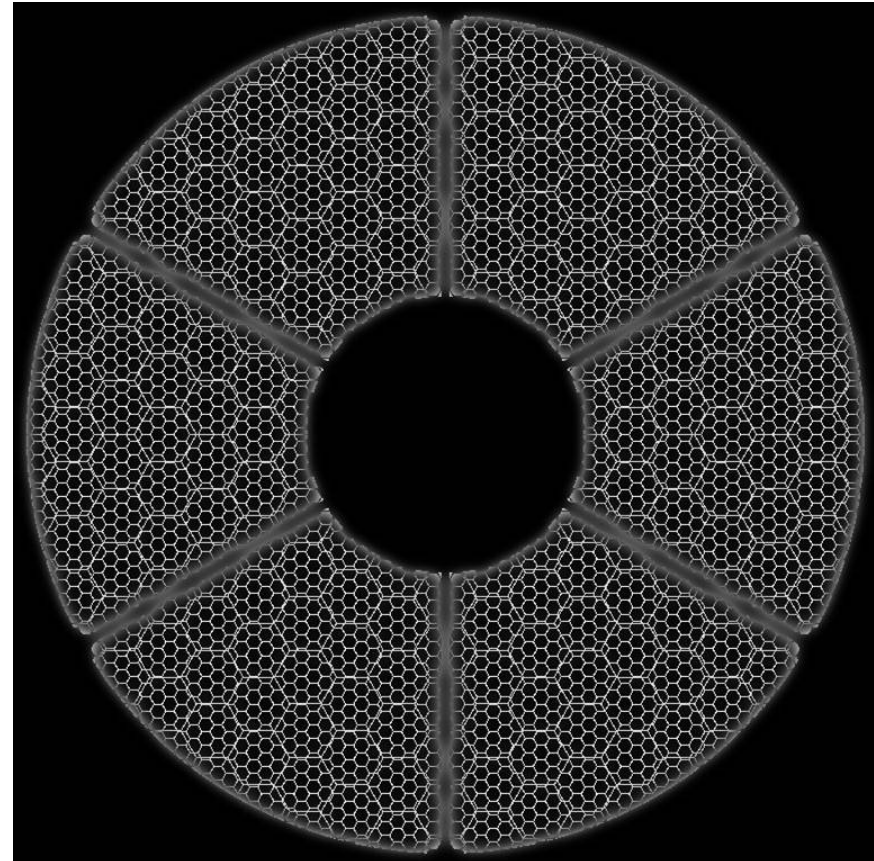
- The M2/corrector support structure may optically divide M1 into six petals.
- Struts thin enough to allow optical phasing of the full segmented mirror
- Backup solution : additional special wavefront sensor for the phasing of the petals relative to each other





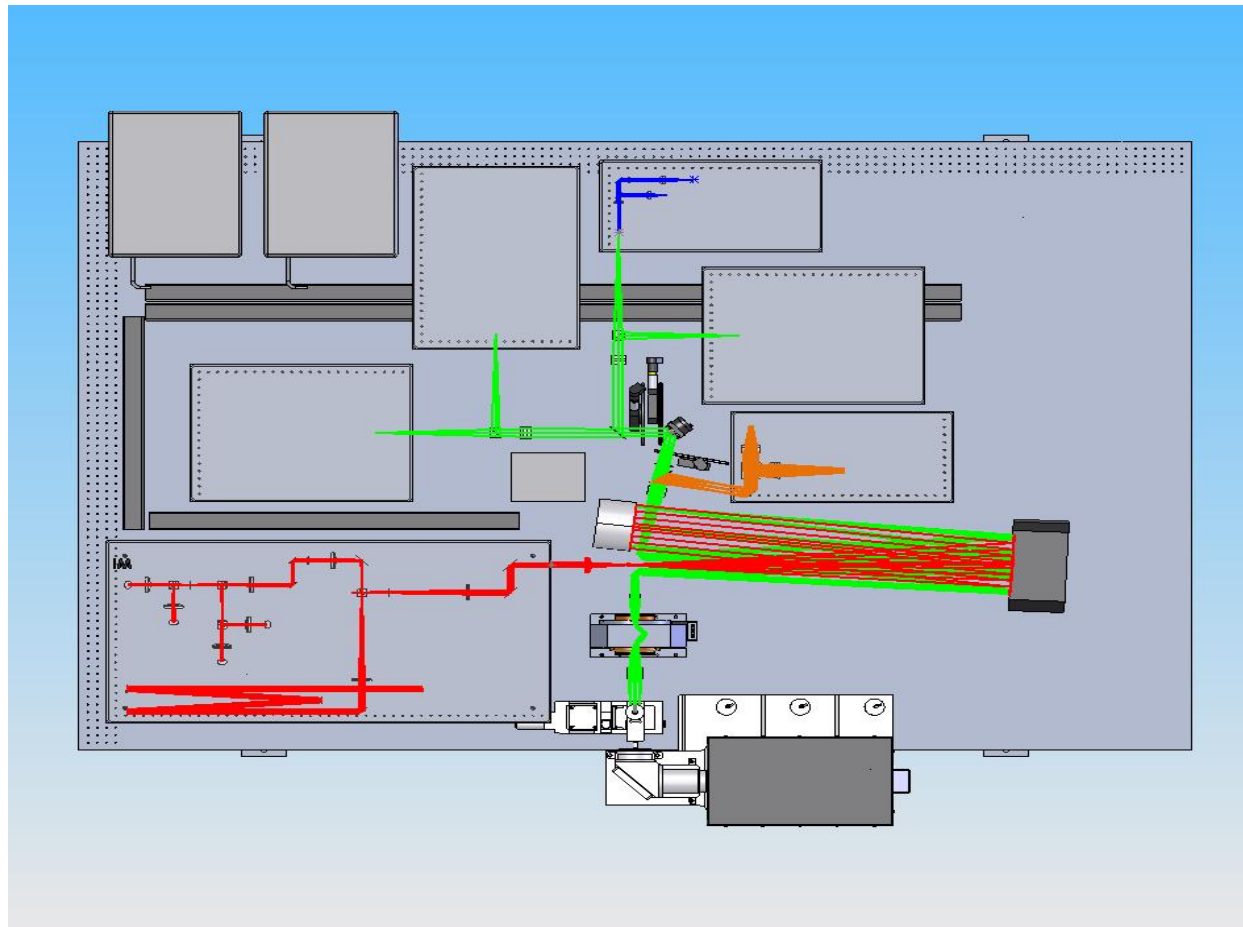
## Phasing – disentangling M1/M2 segmentation

- Segmentation patterns originating from M1 and M2 are superimposed in the detection planes of the wavefront sensors
- Disentangling could be done by
  - Spatial filtering in the Fourier space
  - Use of two or three phasing wavefront sensors in the field



# Active Phasing Experiment

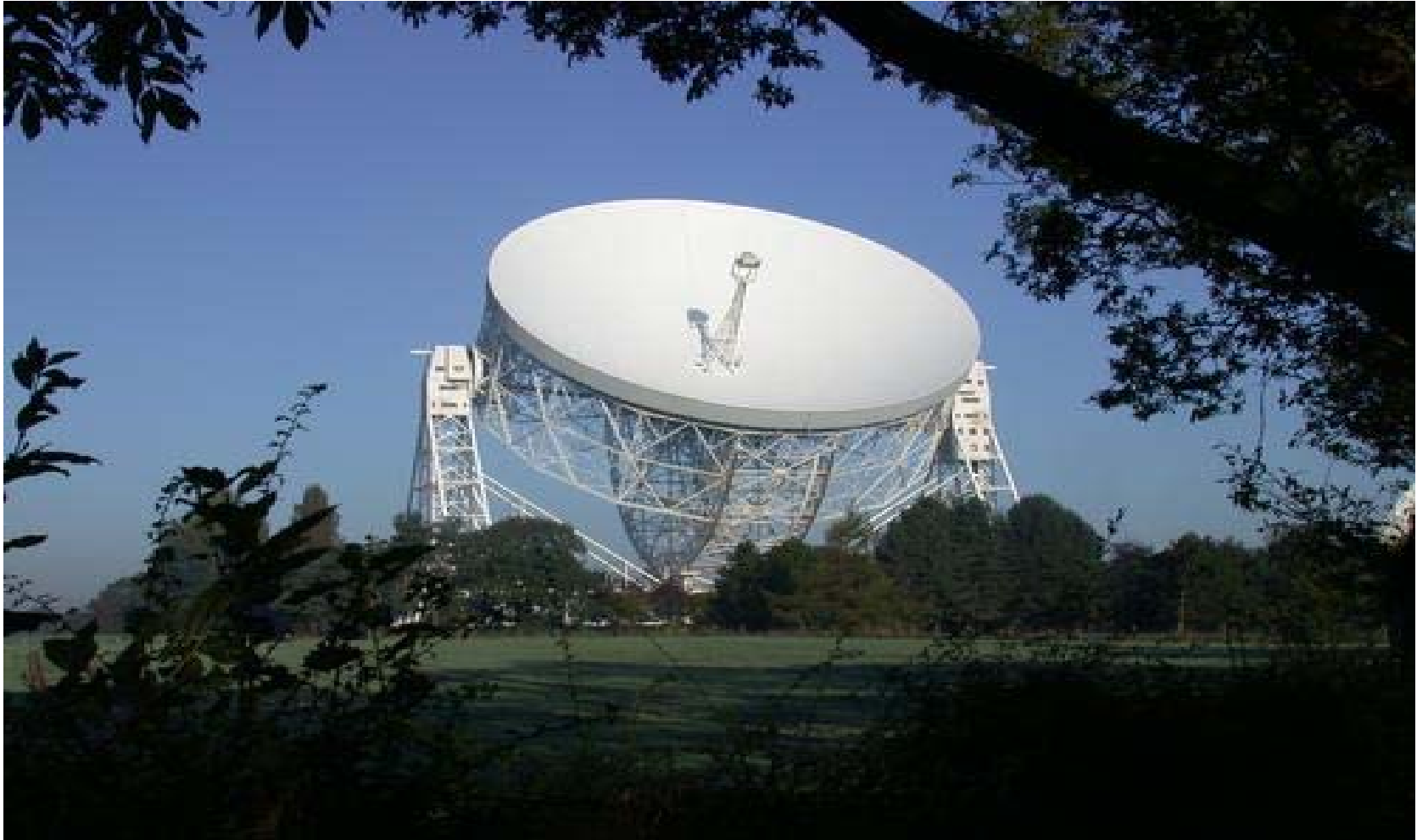
- Comparison of different phasing wavefront sensors
- Test of simultaneous correction of wavefront errors generated by segmented and flexible meniscus mirrors





# Full scale pressure measurements

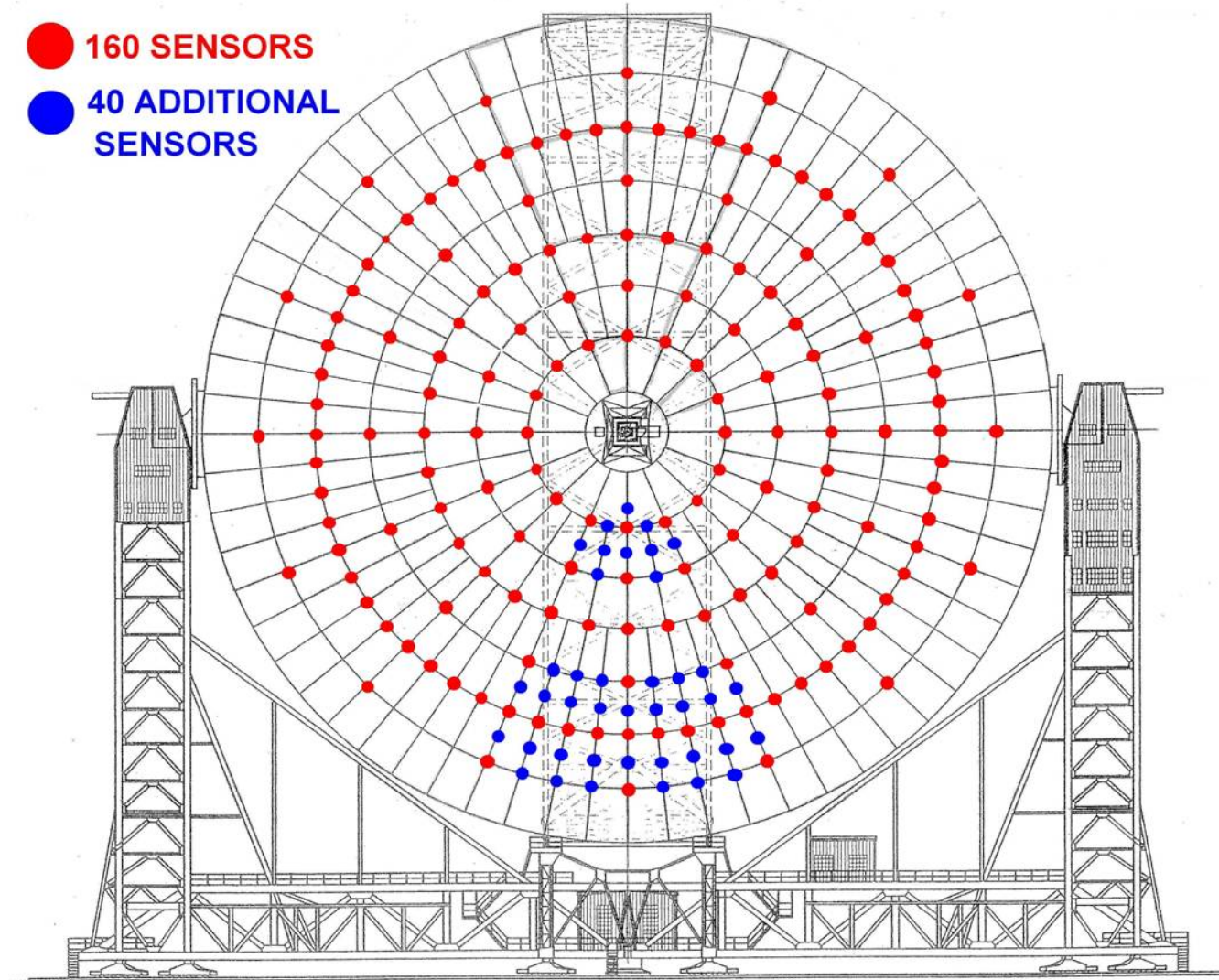
Jodrell Bank Radio Telescope  
Diameter : 76 m



# Jodrell Bank – Location of pressure sensors

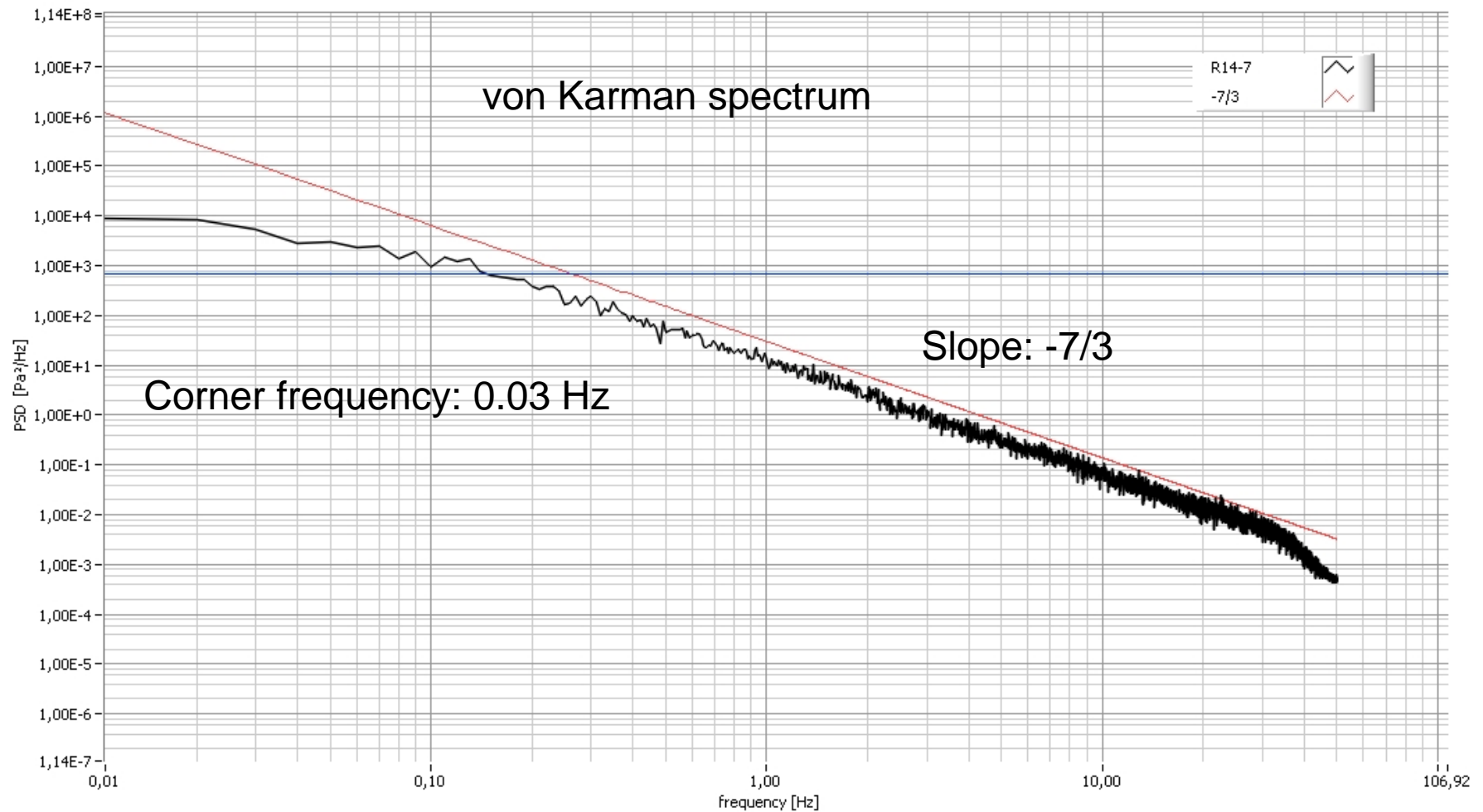
● 160 SENSORS

● 40 ADDITIONAL SENSORS



# Jodrell Bank – Power spectrum

Wind speed: 10 m/sec, Sampling rate : 8 kHz, Integration time : 78 min

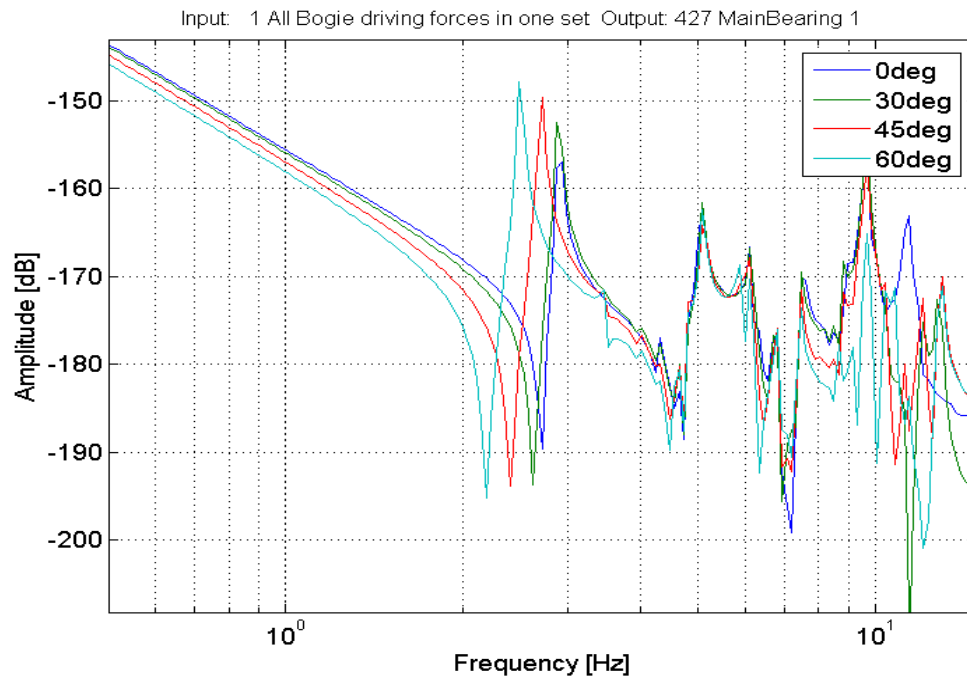
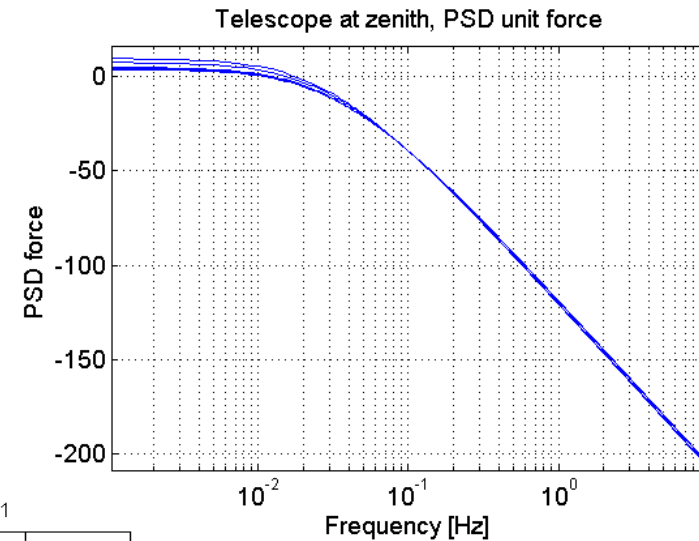




# Main axes control - wind spectra and altitude transfer function

## ■ Von Karman wind spectra

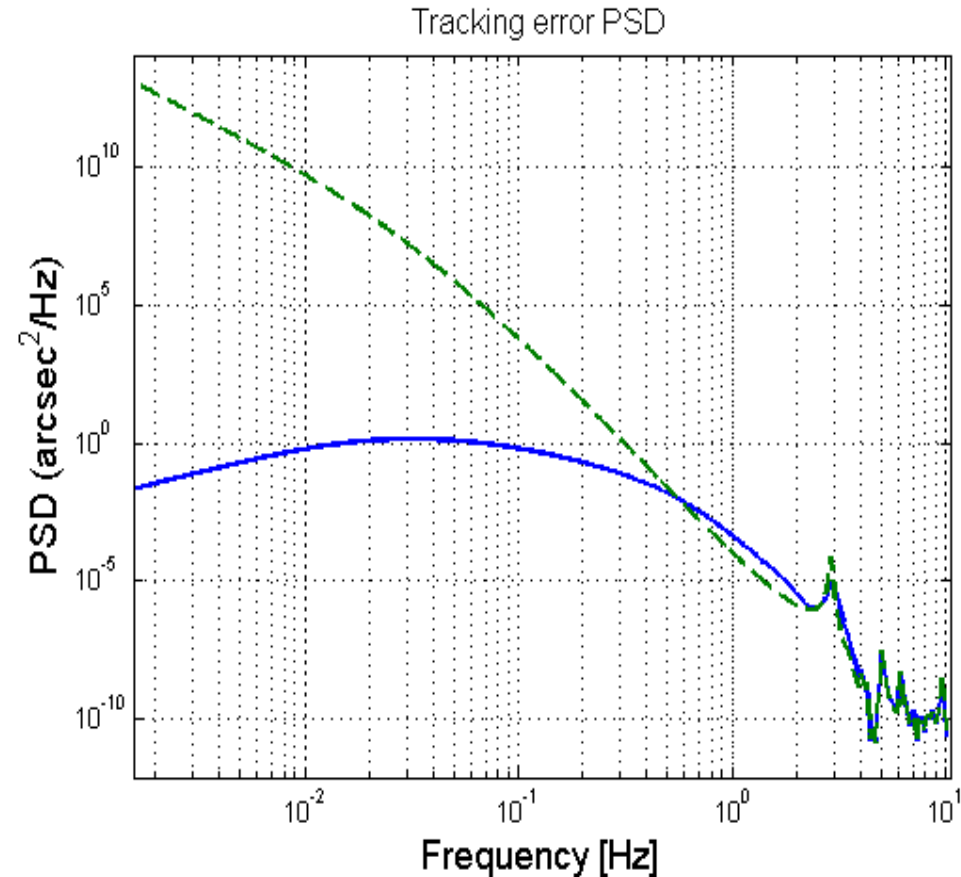
- Wind speed 10 – 14 m/sec
- Turbulence intensity 0.15
- Corner frequency 0.02 Hz



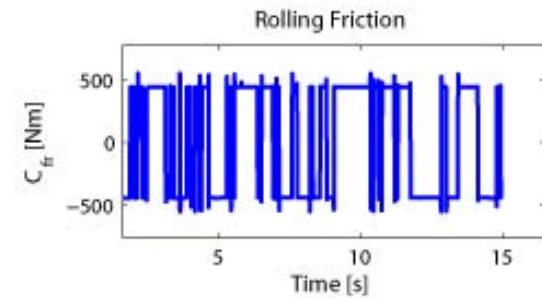
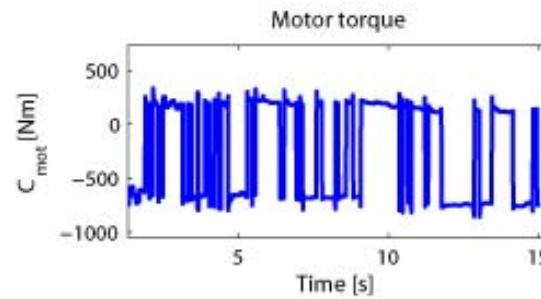
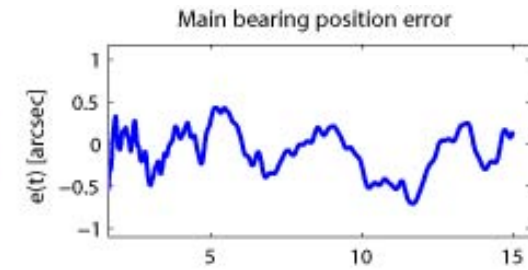
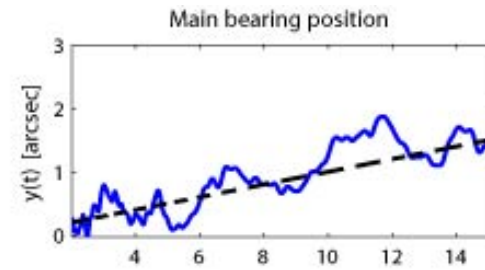
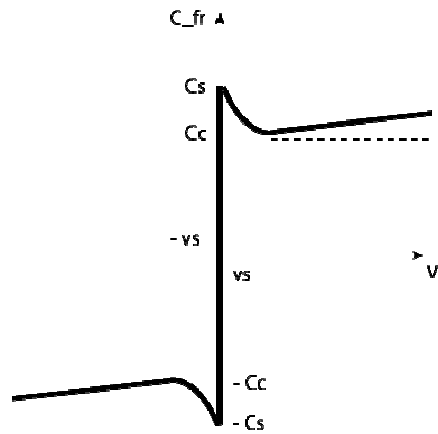
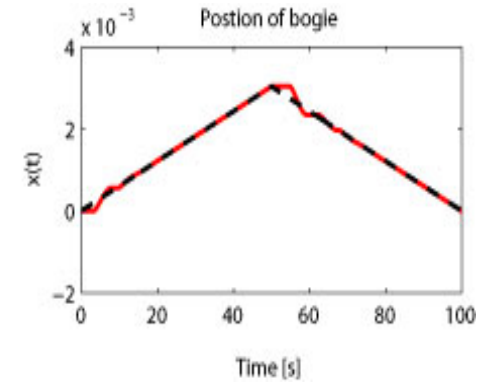
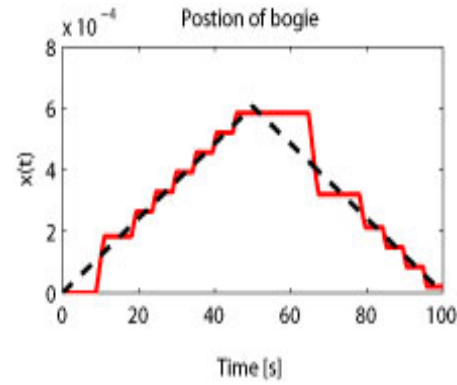
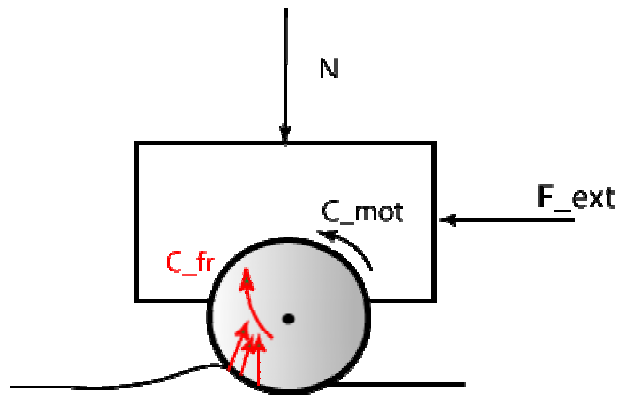
## ■ Altitude axis transfer functions

# Main axes control - Residual tracking errors

- Goal: Residual tracking errors be within the correction range of the fast tip-tilt corrections with the M6-support
- Closed-loop bandwidth: 1 Hz
- Robust design with sufficient modulus margins
- Residual RMS errors:  
0.19 arcsec for 0 deg  
0.32 arcsec for 60 deg



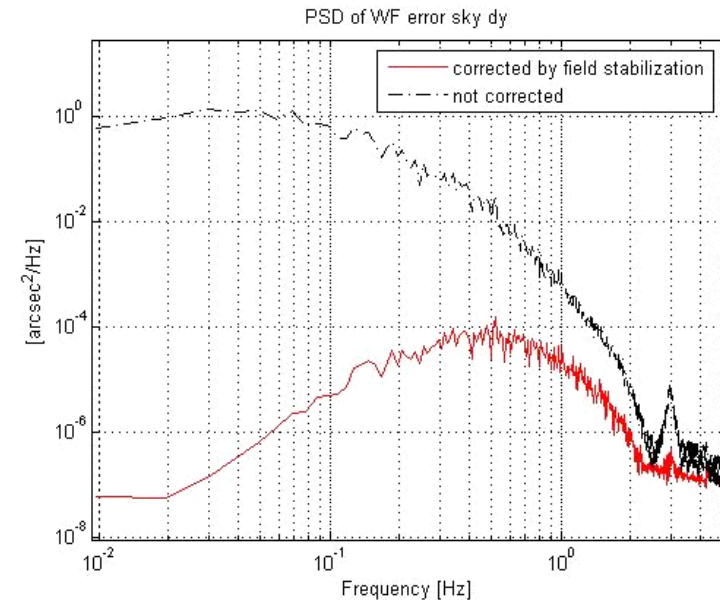
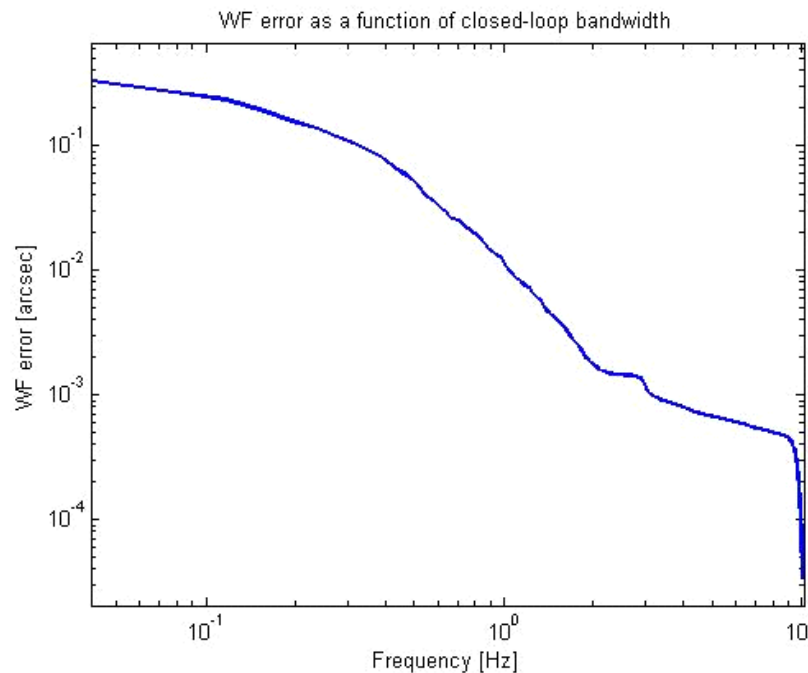
# Rolling friction effects of Bogies





# Correction of tracking errors by M6-support

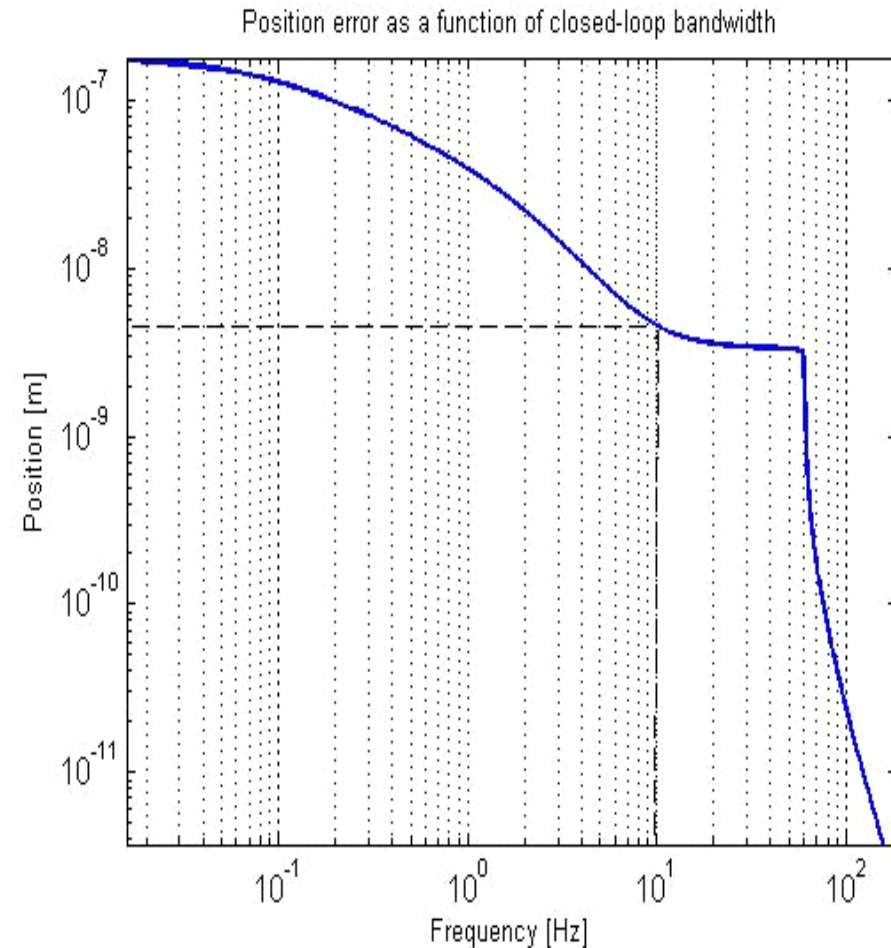
- Power spectral density of the tilt error with and without closed-loop corrections by the M6-support (closed-loop bandwidth: 3.5 Hz)



- Residual tilt error as a function of the closed-loop bandwidth
- Will be corrected by adaptive M6

# Control of segment position

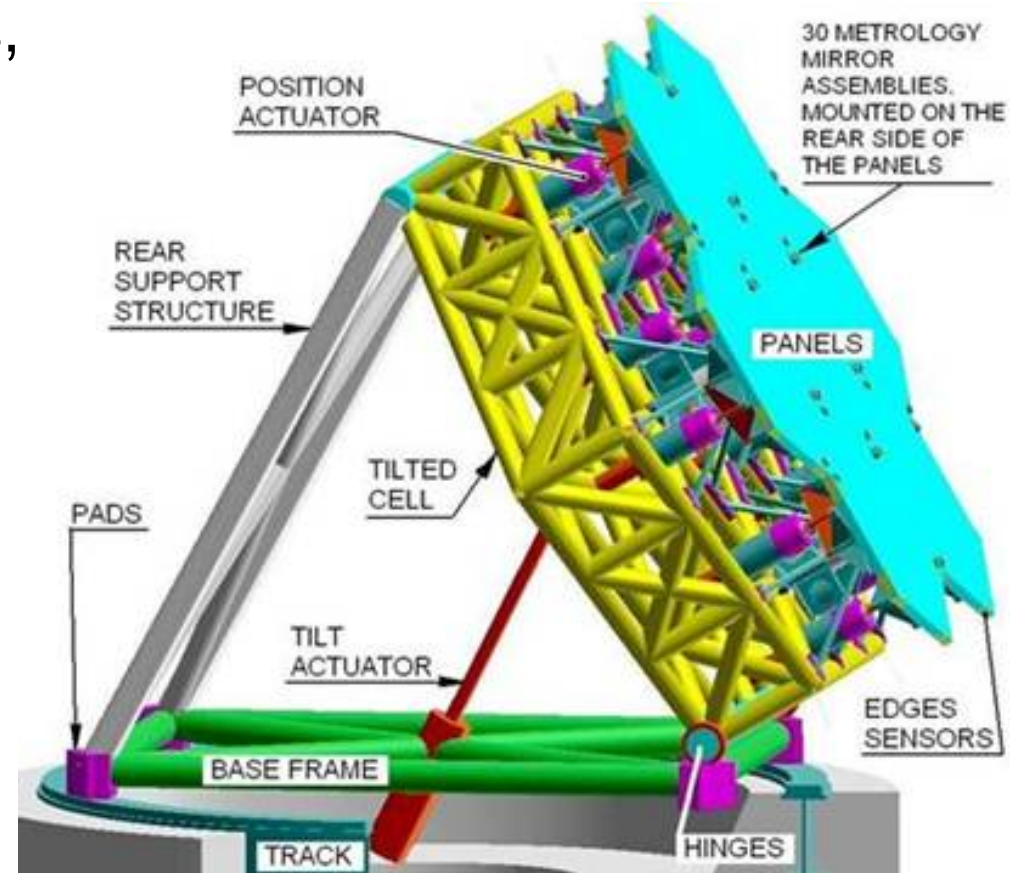
- Control of the segment position based on signals from the edge sensors
- Wind speed : 10 m/sec, corner frequency: 0.07 Hz, turbulence intensity: 0.15
- Lowest segment piston frequency : 60 Hz
- Residual closed-loop error with a 10 Hz bandwidth : ~ 7 nm
- Further reduction may be possible with acceleration feedback control





# Wind Evaluation Breadboard

- Test edge sensors, position actuators and control algorithms under observatory conditions



# Conclusions

- Further development required for phasing wavefront sensors
  - Disentangling of overlapping segmentation patterns
  - Detection of the segmentation pattern
  - Optionally continuous wavefront sensing
  - Questions are addressed by the APE experiment
- Operation of the telescope seems feasible in open air
  - Segmented mirrors controlled by fast feedback loops
  - Large meniscus mirrors shielded from the wind
- No fundamental problems with the non-adaptive wavefront control